# Predictability Crisis in Early Universe Cosmology

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|              |                      |                                  |             |            |
| Motivatin    | g Questions          |                                  |             |            |

- Is cosmology a science? What kind of science is cosmology?
  - Fundamental / law-seeking
  - Historical / descriptive
- What are the appropriate aims of cosmology?
  - Detailed reconstruction of the physical properties, evolution within our Hubble volume
  - Assessment of probability of that account with respect to fundamental theory

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- Discovery and justification of new laws

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Predictability Crisis

In an eternally inflating universe, anything that can happen will happen; in fact, it will happen an infinite number of times. Thus, the question of what is possible becomes trivial — anything is possible, unless it violates some absolute conservation law. To extract predictions from the theory, we must therefore learn to distinguish the probable from the improbable. (Guth 2007)

#### Questions

- Does eternal inflation make any predictions, and in what sense? (More generally, predictions from the multiverse?)
- Output Bound We define probabilities in cosmology?
- I how should we characterize empirical success of a cosmological theory?

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| Outline      |                      |                                  |             |            |

# Background

- Testability of Inflation
- From Inflation to Eternal Inflation
- Predictability Crisis
  - Recipe for Predictions in Eternal Inflation
  - Measure Problem
  - Status of the Measure
- 8 Responses
  - Response 1: Probabilities in Cosmology
  - Response 2: Reconsidering Empirical Success

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| Testability of Inflat | ion                 |                                  |             |            |
| Successes             | for Inflation       |                                  |             |            |



WMAP angular power spectrum (Spergel et al. 2006)

- Uniformity
- Flatness
- Spectrum of density perturbations (nearly scale invariant, Gaussian, adiabatic)

Consequences of dynamical evolution of "inflaton," scalar field  $\phi$ 

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| Testability of Inflat | ion                  |                                  |             |            |
| Indifferenc           | ce Principle         |                                  |             |            |



Ernan McMullin (1924-2011)

This first version [Diogenes Laertius] of the cosmogonic indifference principle contains two elements: no special setting of the initial state is required (a 'chaos' will do), and no subsequent intervention of a purposive agency of any sort is required for order to appear out of the original disorder. The normal operation of what a later generation would call mechanical law suffices (McMullin 1993)

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| Testability of Inflatio | n                    |                                  |             |            |
| Indifference            | e and Inflation      |                                  |             |            |

• Conventional Wisdom: results of inflation "independent" of initial state, fixed by dynamical evolution of  $\phi$ , replaces "finely-tuned" initial state

Issues

- How to make sense of claim that initial state in standard cosmology is "improbable" or "unnatural"?
- How probable is it that inflation occurred? (Penrose 1986; Hollands and Wald, Turok et al.)
- Exchange fine-tuning of ICs for specific properties of  $V(\phi),$  initial state of  $\phi$

- Chaotic / Eternal Inflation
  - Linde: response to "fine-tuning" of  $V(\phi), \phi$

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| Inflation to Eterna | Inflation            |                                  |             |            |
| Eternal In          | flation              |                                  |             |            |

# "Inflation is generically eternal"

- Heuristic arguments: volume expansion rate >> rate of false vacuum decay
- Leads to universe with:
  - Regions of false vacua
  - "Pocket universes"



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| Inflation to Eternal | Inflation            |                                  |                           |            |
| Eternal In           | flation              |                                  |                           |            |

- Pocket universes with different low energy physics
- Variation based on "meta-law" governing generation of pocket universes



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Image: Andrei Linde

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| Inflation to Eternal Infl | ation               |                                  |                    |            |
| Consequenc                | es of Eterna        | al Inflation                     |                    |            |

### "... anything is possible"

Scope of variation depends upon "meta-law" governing generation.

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#### Predictions and Probabilities

- In what sense can we make predictions?
- What do we need to introduce to do so?

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| Predictions  |                      |                                  |              |            |
| Recipe for   | r Multiverse I       | Predictions (e.g.                | Aguirre 2006 | 5)         |

- O: reference class for conditionalization, e.g. "observers" (or some proxy)
- $lpha_i$ : parameters taken to vary in different regions of multiverse
- $N_O(\alpha_i)$ : number of "observers"
- $P(\alpha_i)$ : prior probability

"What a typical member of reference class will observe" (Principle of Mediocrity):

$$P_O(\alpha_i) = N_O(\alpha_i)P(\alpha_i) \tag{1}$$

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| Predictions  |                      |                                  |                    |            |
| Exemplar:    | Weinberg or          | η <b>Λ</b>                       |                    |            |

- O: large gravitationally bound systems (as proxy for observers)
- $\alpha$ : consider varying  $\Lambda$ , other parameters all fixed
- $N_O(\alpha) =:$  only non-zero in small window, due to  $\Lambda$ 's effect on structure formation
- $P(\alpha) =:$  expect this to be uniform, because anthropically allowed region small compared to particle physics energy scales

Result: expect to something close to "median value" of  $\Lambda$  (calculation in 2005)

$$\rho_{\rm v} = 13.3\rho_m \tag{2}$$

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Observed value:  $\rho_v = 2.3 \rho_m$ . Probability = .156.

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| Predictions  |                |                       |           |            |
| Questiona    | able Ingredien | ts                    |           |            |

- $N_O(\alpha_i)$ : number of "observers"
- $P(\alpha_i)$ : prior probability
- Reasonable estimates of  $N_O(\alpha_i), P(\alpha_i)$ ?
- Principle of Mediocrity?
- "Measure Problem": implicit choice of measure. What is the appropriate measure over the multiverse?

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| Measure      |                      |                                  |             |            |
| Measure Pr   | roblem               |                                  |             |            |

- What should be given "equal weight" by the measure?
  - Distinct pocket universe
  - Spacetime volume
  - Each distinct pocket universe generated from a given starting region
  - Length of a given world-line in each distinct pocket universe

• Dealing with infinity: require some way of "regulating" infinities

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| Measure      |                      |                                  |             |            |
| Example      |                      |                                  |             |            |



Figure: Vilenkin 2006

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| Measure      |                      |                                  |             |            |
| Debates r    | egarding Mea         | asures                           |             |            |

Desiderata for Measure

- Independent of Initial Conditions
- "Calculable"  $(N_O(\alpha_i))$
- Foliation-independent

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# Paradoxes

- Youngness paradox
- Q catastrophe
- Boltzmann brains / freak observers

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# State of the Debate (?)

- "Testing" different proposed measures by considering paradoxes
- Disagreement regarding desiderata

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| Measure      |                      |                                  |             |            |
| Getting Pre  | dictions ou          | t of Nothing?                    |             |            |

- "Phenomenological" approach
  - Measures proposed without underlying dynamics
- Indifference principle revisited
  - Independent of initial conditions (some take as a desiderata)

"[W]e require that [the probabilities] should be independent of the initial conditions at the onset of inflation. The dynamics of eternal inflation is an attractor; its asymptotic behavior has no memory of the initial state." (Vilenkin 2006)

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| Response 0    |                      |                       |                    |            |
| Eternal Infla | ation is Self-       | Undermining?          |                    |            |

- Does El predict anything?
  - Original predictions of inflation undermined
  - Response (?): don't abandon successful theory due to open problems

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- Reconsider claim that inflation  $\rightarrow$  El
- "Fair treatment" of competing theories

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| Response 1   |                      |                                  |                    |            |
| Probabilit   | ies in Cosmol        | logy                             |                    |            |

- Contrast with statistical mechanics
  - Active debates regarding nature of probability in SM
  - Measure plays a central role in predictive success of the theory
- Change inductive methodology in El?
  - Bayesian methodology to calculate conditional probabilities (subjective credences)
  - Should this depend on infinite extent of the universe? (cf. Neal 2006)

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| Response 2   |                      |                                  |                    |            |
| "Predictions | "?                   |                                  |                    |            |

- Given a suitable measure, accept Principle of Mediocrity, ...
- Possible response to incorrect "prediction" of parameter value  $\alpha_i$ ?
  - *Some* parameters will have unusual values
  - Reject measure
  - Reject estimate of  $N_O(\alpha_i)$ , consider different reference class
  - Reconsider calculation varying more parameters
- Contrast with more informative cases
  - Set of clearly motivated sufficient conditions to derive a particular claim
  - Systematic discrepancies revealing; possibility of refinement and further empirical testing

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| Concludin    |                      |                                  |                           |            |

- Pessimism about Measure Problem
  - Not clear how the debates about appropriate measure can be resolved
  - Extracting predictions from El requires accepting dubious principles such as the Principle of Mediocrity, specification of a reference class

- Connection with broader questions
  - Justifying new fundamental physics based on its role in reconstruction in cosmic history
  - Modalities in cosmology: laws and initial conditions, probabilities