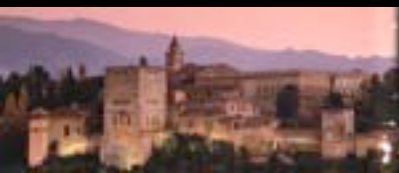


# What have we learned from observational cosmology ? (observer's point of view)

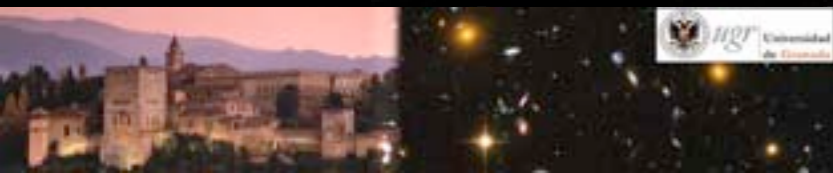


Jean-Christophe Hamilton  
CNRS - IN2P3 - APC - Paris



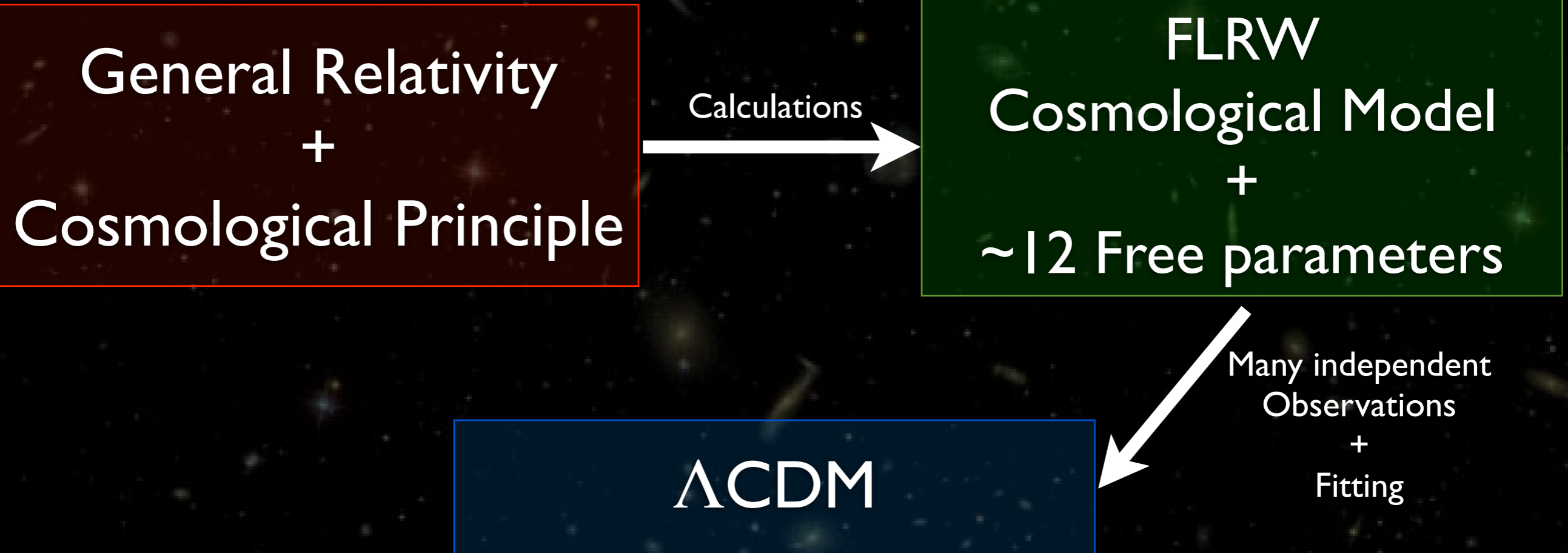
# Outline

- $\Lambda$ CDM: the concordance model
- Observational evidences for  $\Lambda$ CDM
  - ★ The Cosmic Microwave Background and the early Universe
  - ★ Dark Matter
  - ★ Dark Energy
  - ★ Tests of the cosmological principle
  - ★ Tests of the Big-Bang paradigm
- Summary

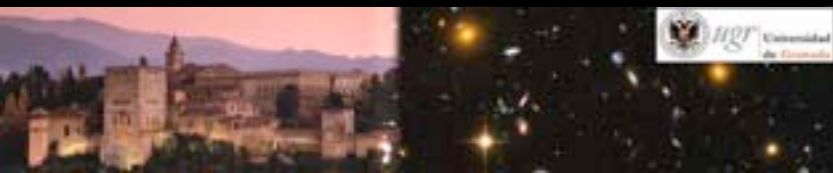


# $\Lambda$ CDM foundations

Assumptions

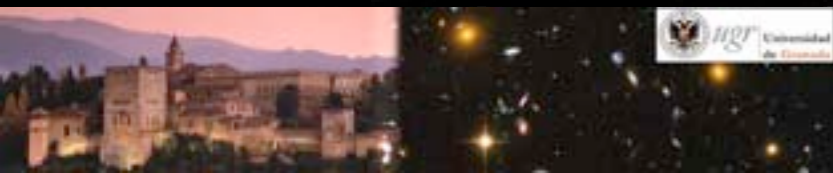


- There is not much room out of this scheme
  - ★ Recusing the cosmological principle (obs. constraints)
  - ★ Extensions/modifications to G.R. (obs. constraints)
  - ★ Modelling of the energy content of the Universe (obs. constraints + particle physics)



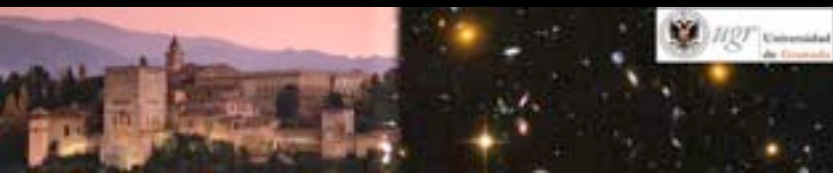
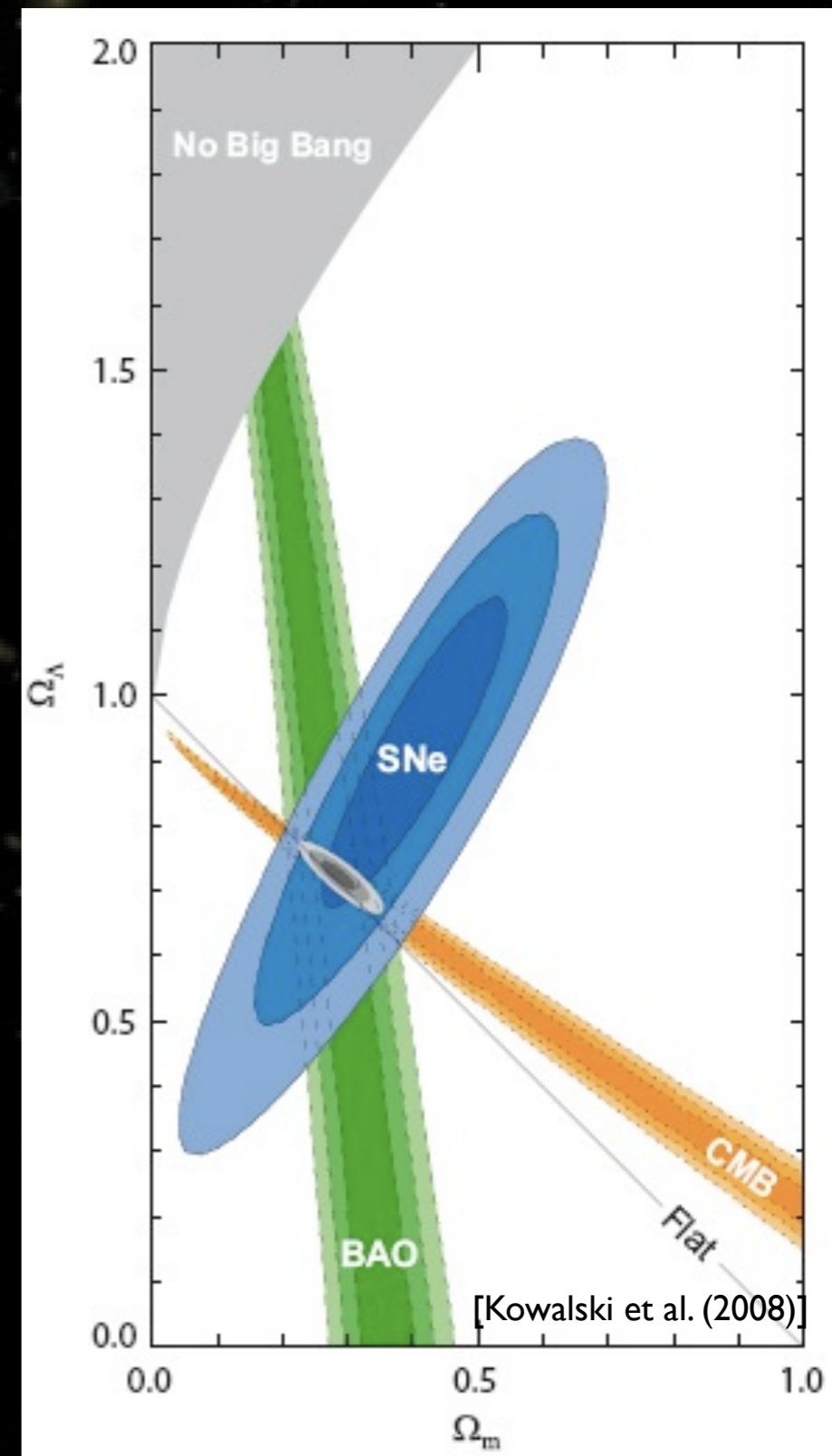
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- ★ FLRW expanding Universe (Hot Big-Bang)
  - General Relativity + Cosmological Principle
  - Hubble constant  $\sim 70 \text{ km.s}^{-1}.\text{Mpc}^{-1}$
  - CMB blackbody, BigBang Nucleosynthesis
- ★ The Universe is  $\sim$  flat :  $\Omega_{\text{tot}} \approx 1$ 
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- ★ It contains  $\sim 22\%$  of Dark Matter (unknown)
  - Galaxies rotation curves, Clusters X, weak-lensing, Structure formation, CMB
- ★ It contains  $\sim 74\%$  of Dark Energy (unknown)
  - SNIa fainter than expected  $\Rightarrow$  further  $\Rightarrow$  acceleration
  - CMB+H, direct measurements of  $\Omega_m$ , ISW effect



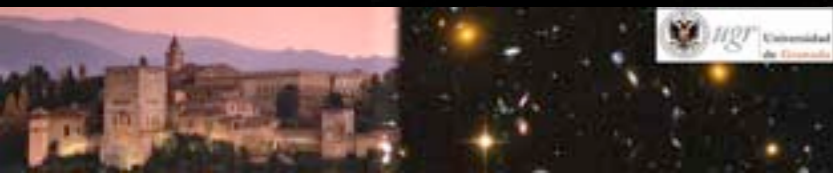
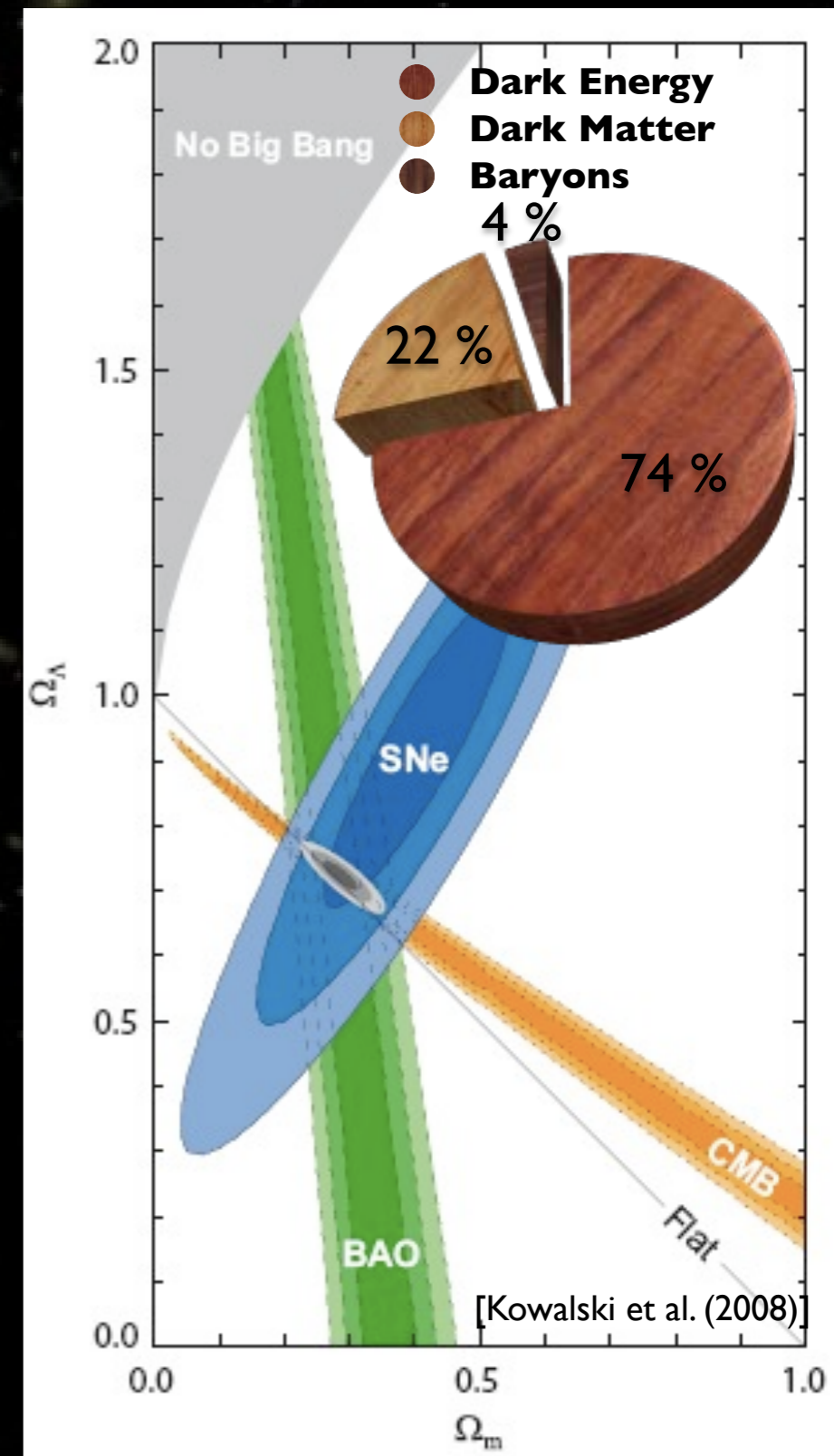
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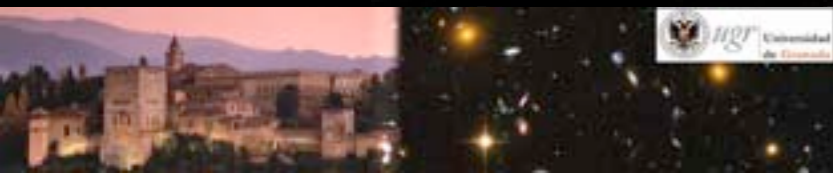
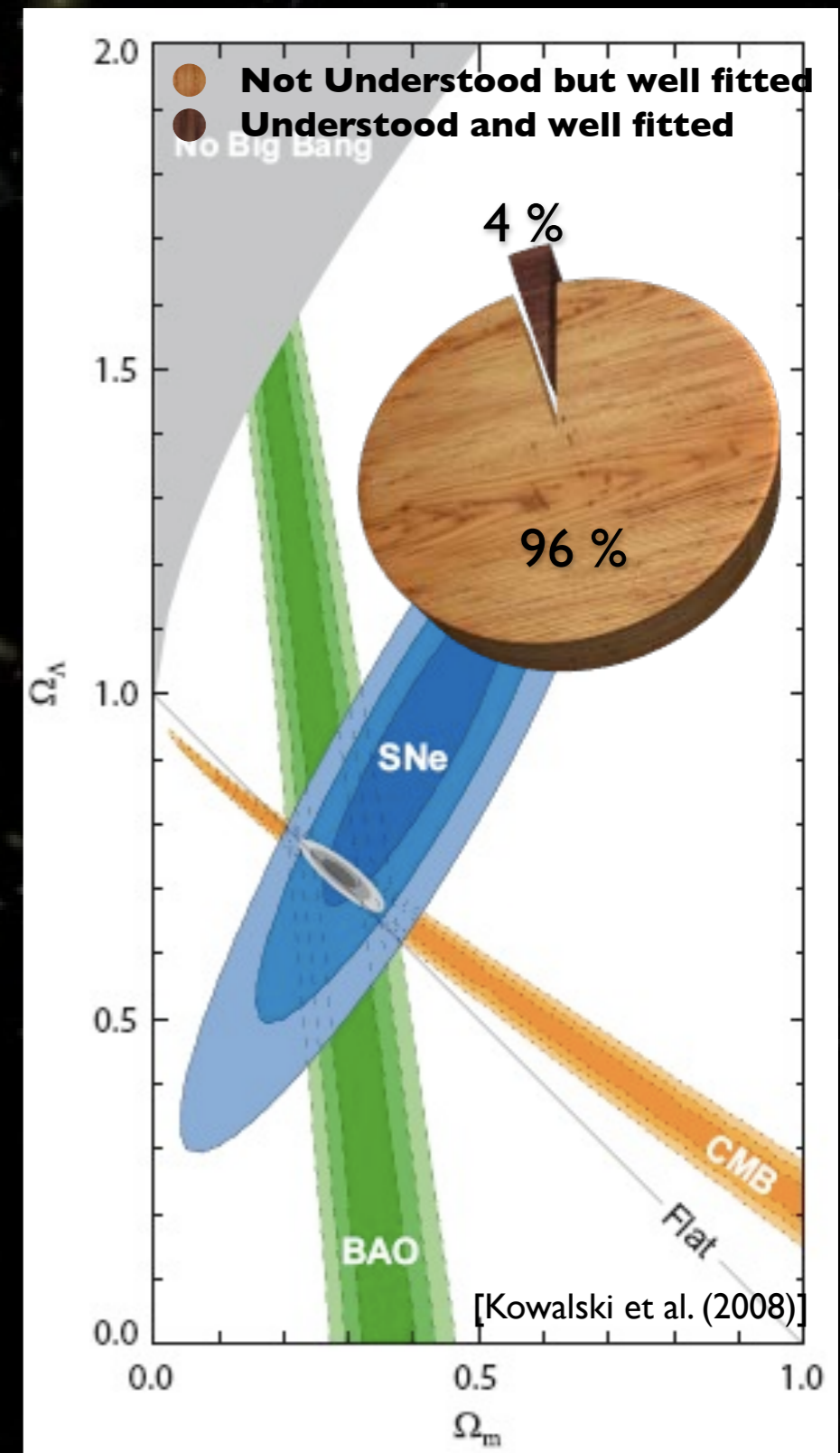
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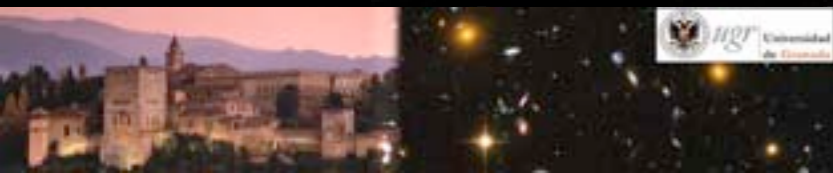
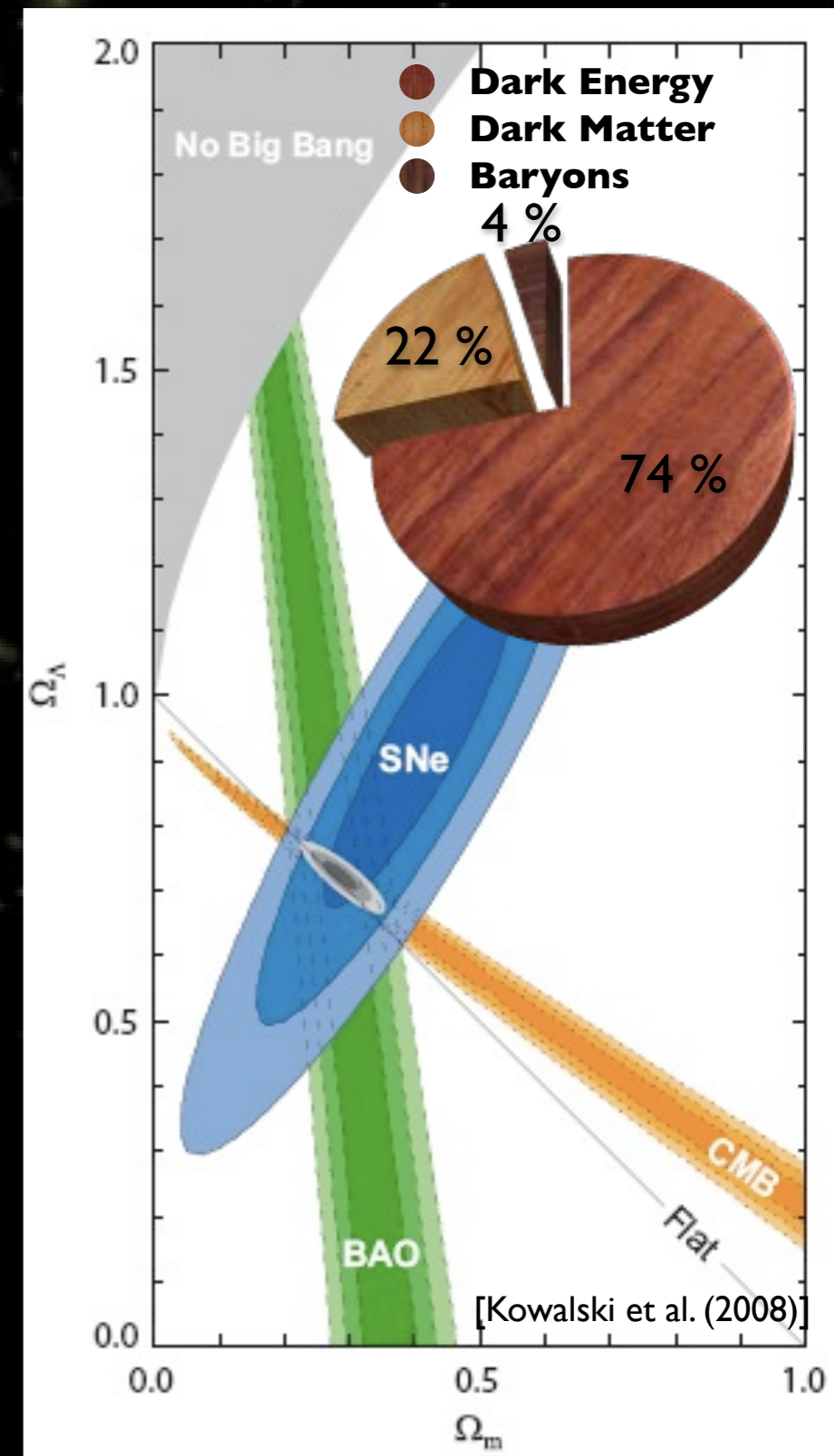
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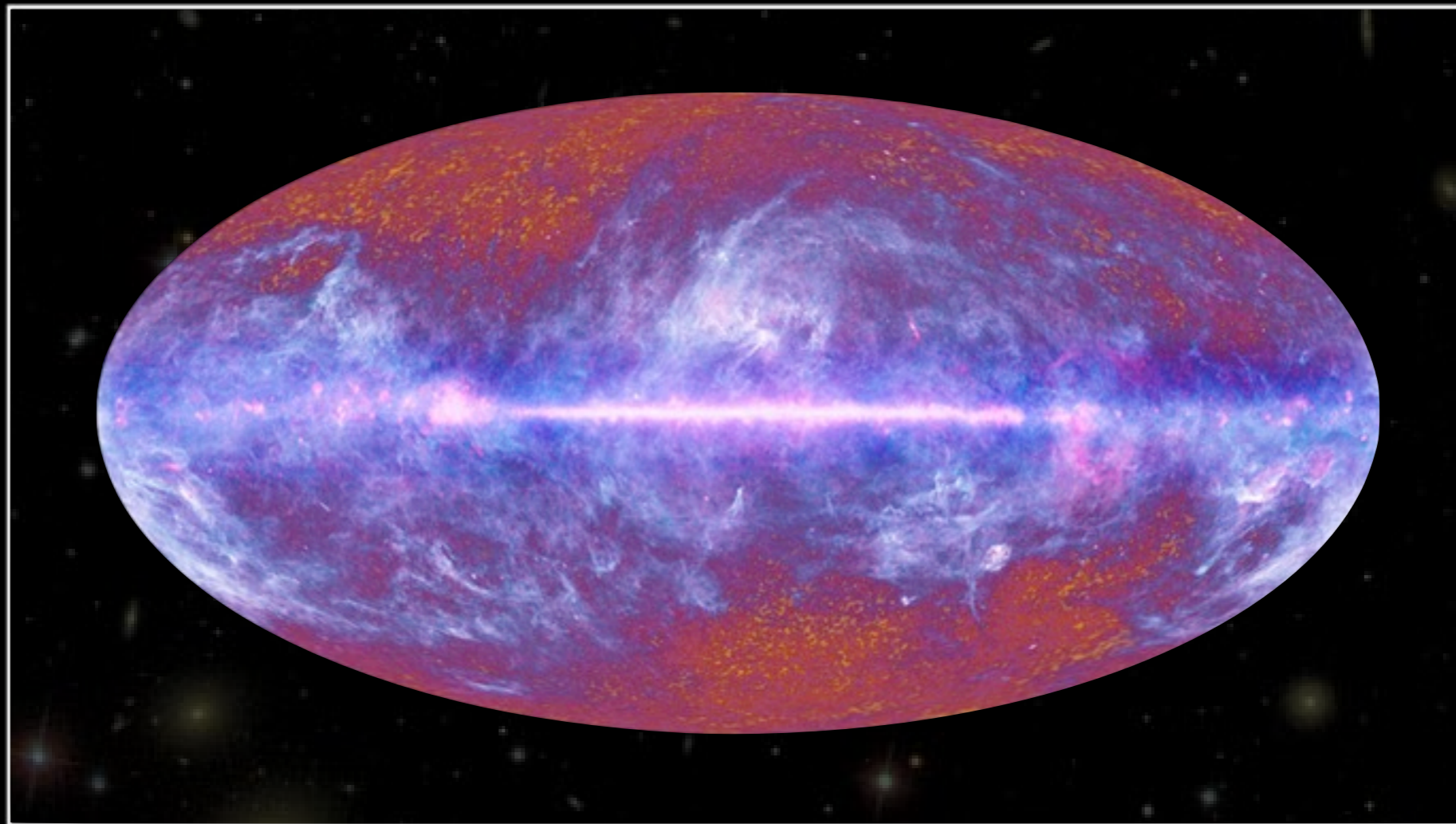
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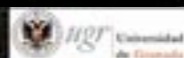
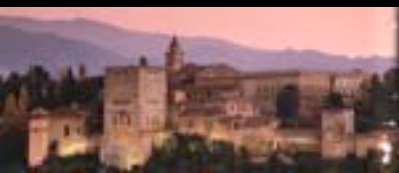
# Cosmic Microwave background



The Planck one-year all-sky survey



(c) ESA, HFI and LFI consortia, July 2010



# CMB Physics

- Origin

- ★ Early Universe

- Ionized  $\Rightarrow$  opaque to photons
- thermal equilibrium

- ★  $T \ll 13.6 \text{ eV}$

- Neutral  $\Rightarrow$  matter/radiation decoupling
- CMB emitted. Blackbody at 3000K ( $z=1000$ )
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- Shape

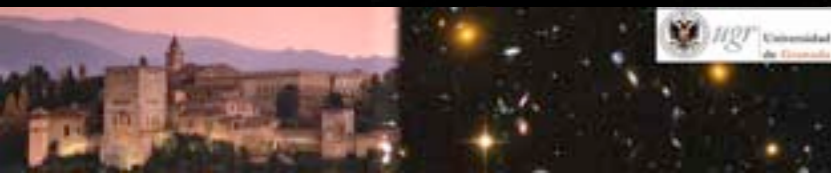
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# CMB Physics

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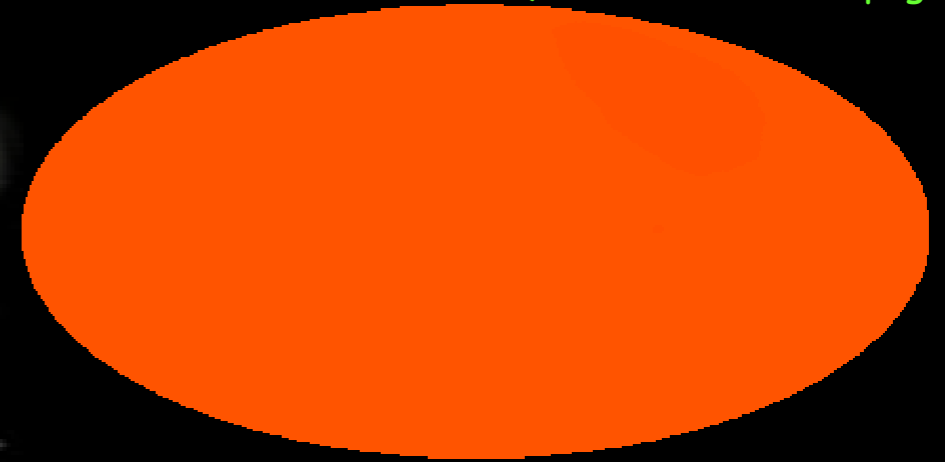
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(COBE/DMR homepage)



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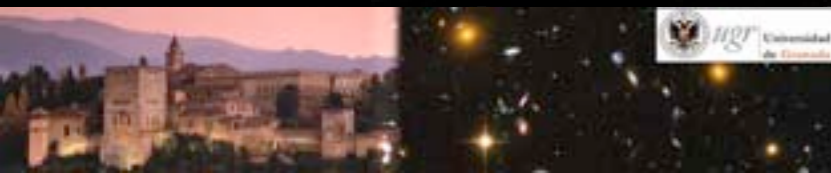
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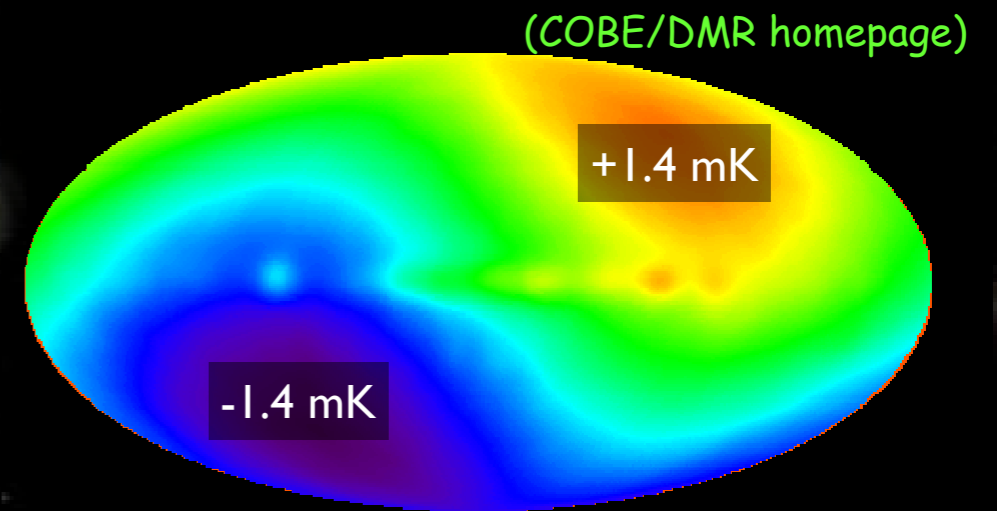
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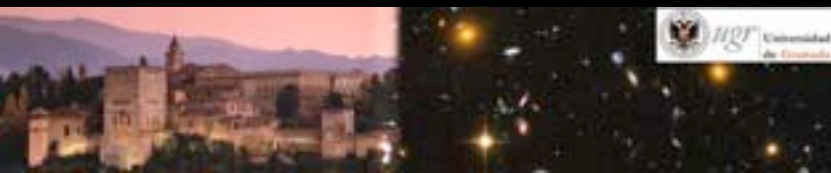
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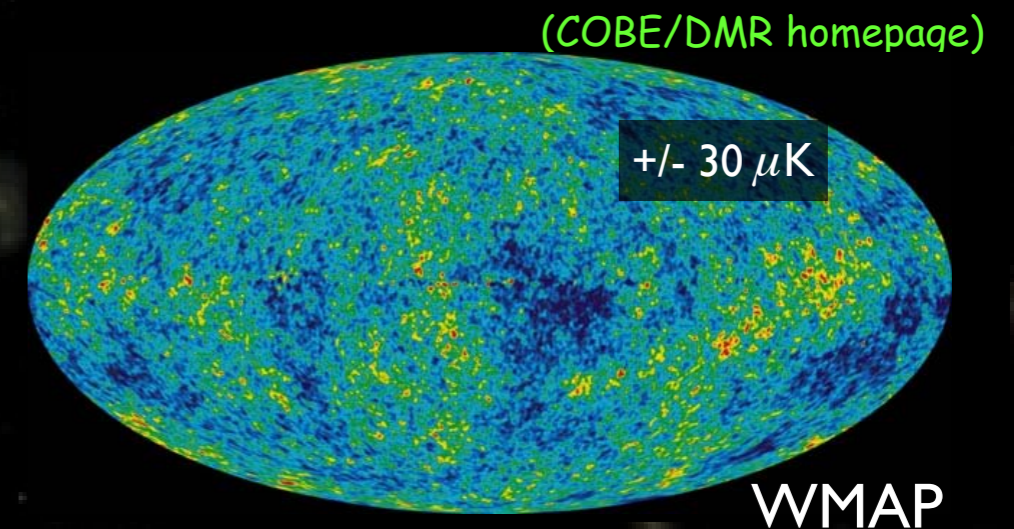
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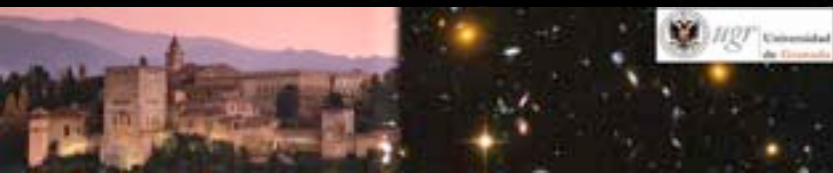
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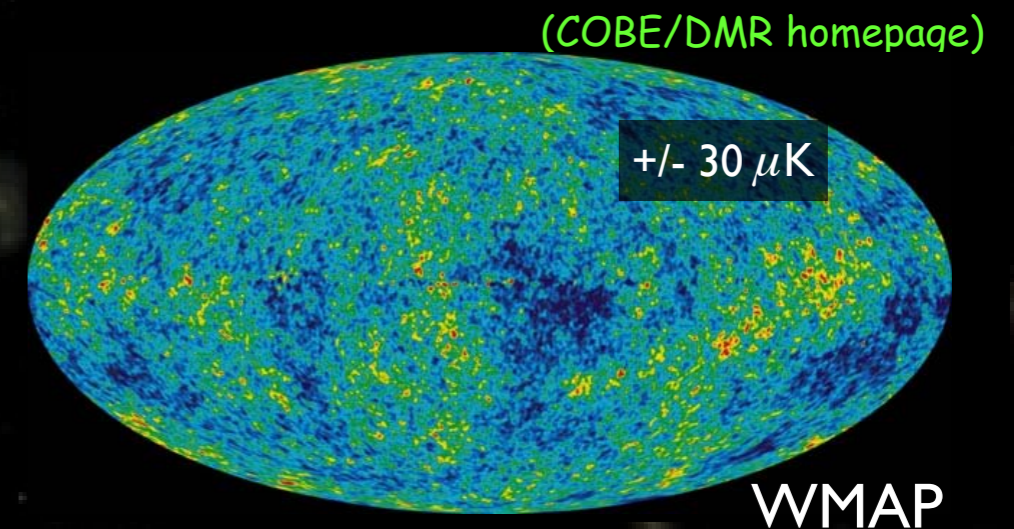
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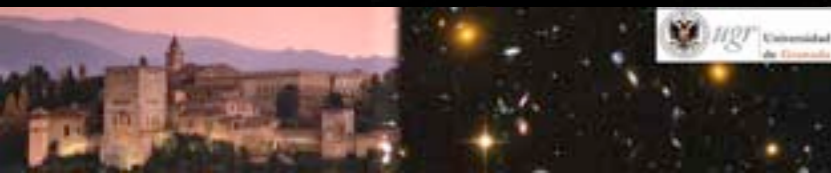
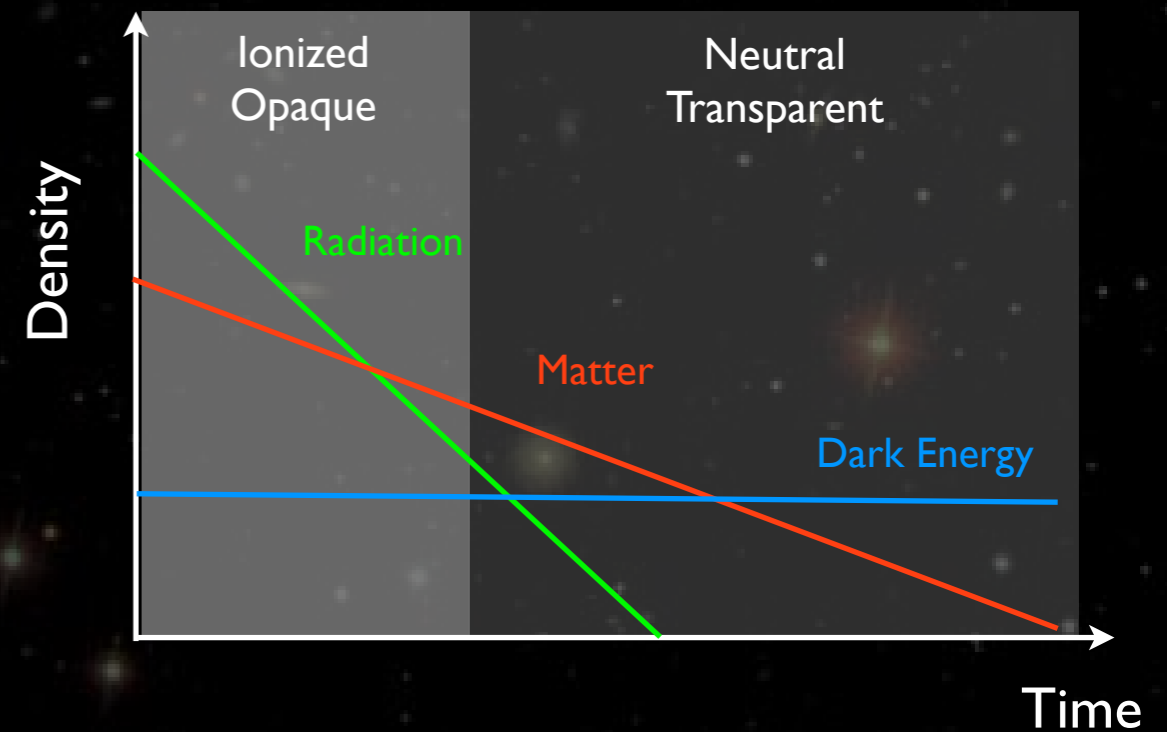
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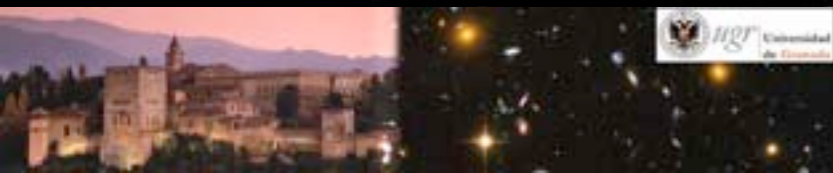
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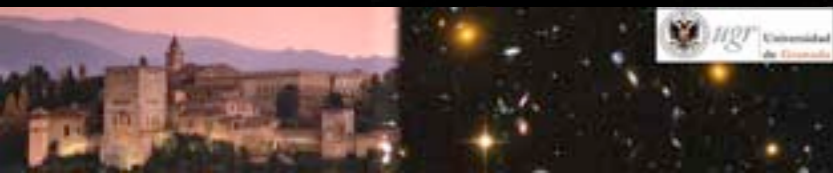
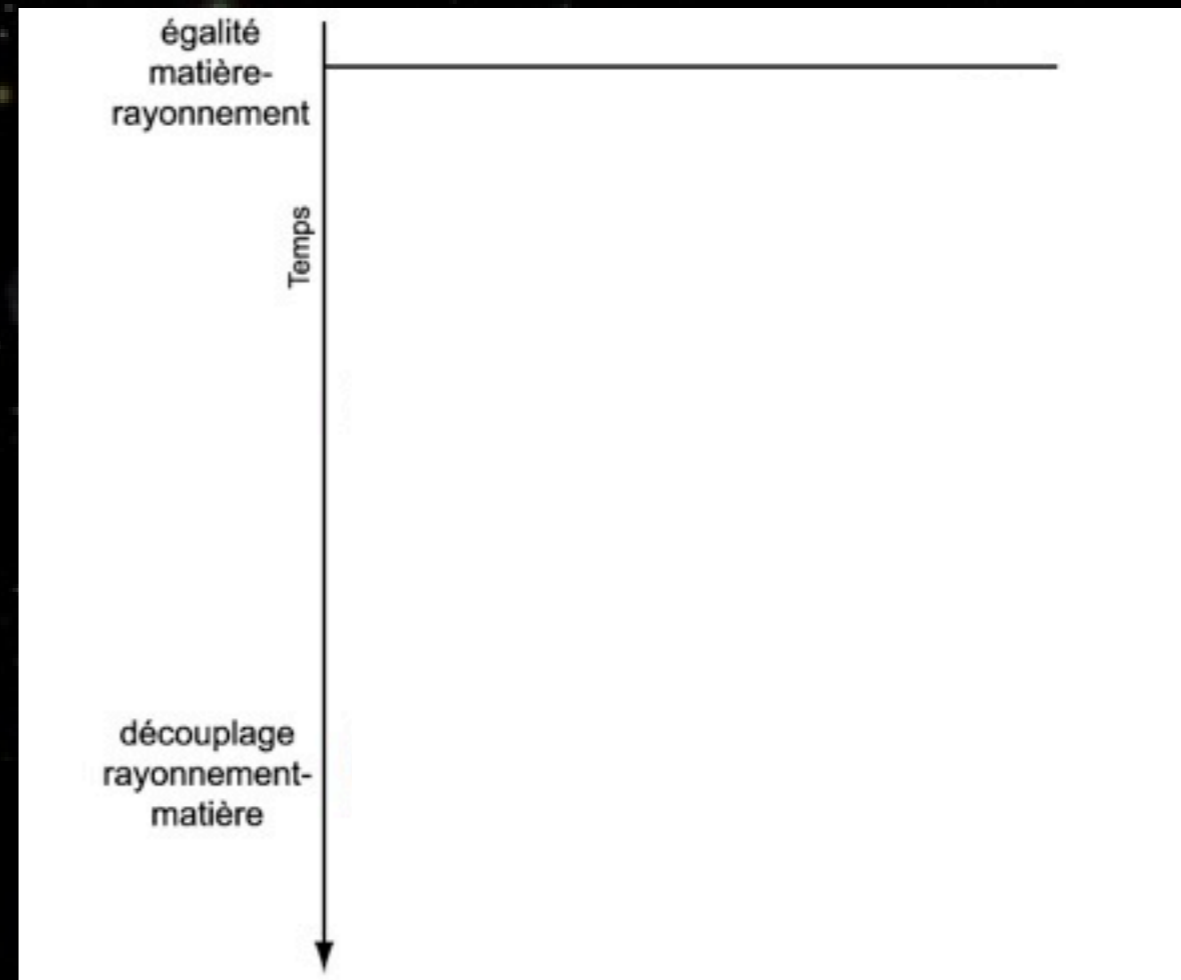
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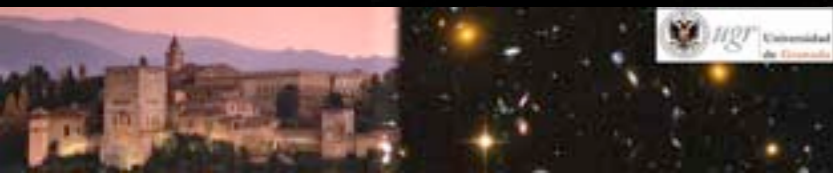
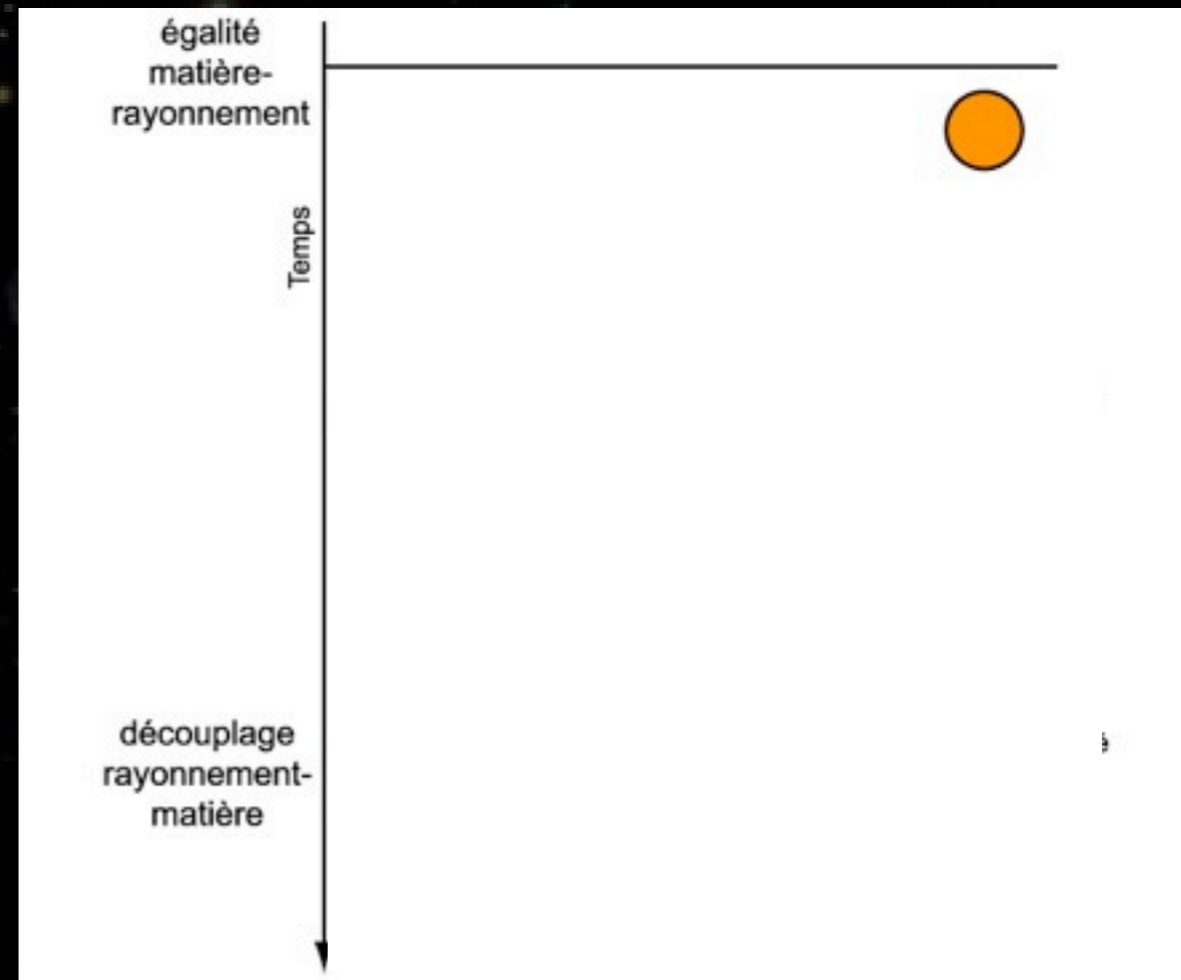
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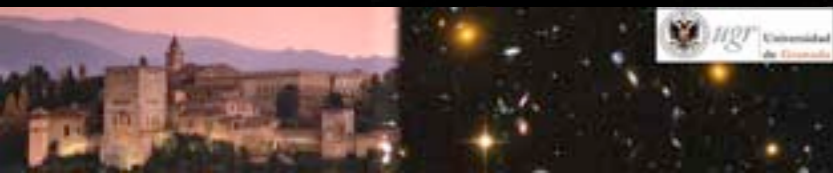
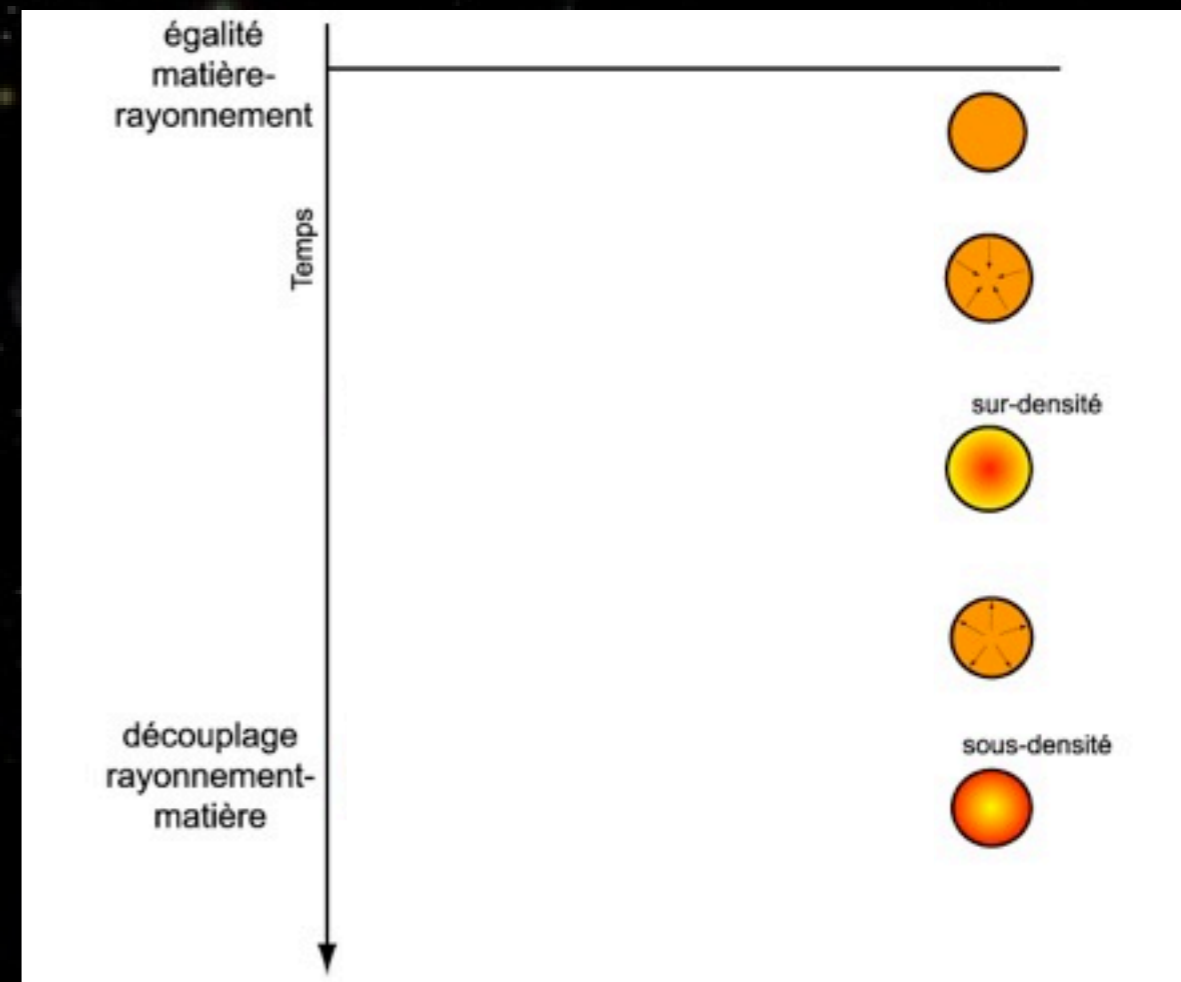
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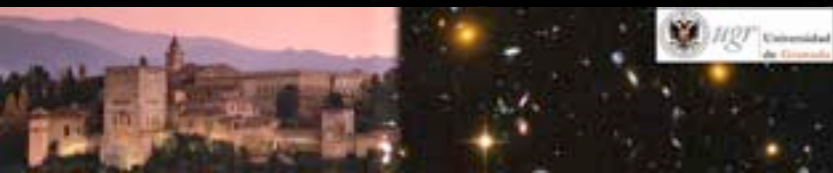
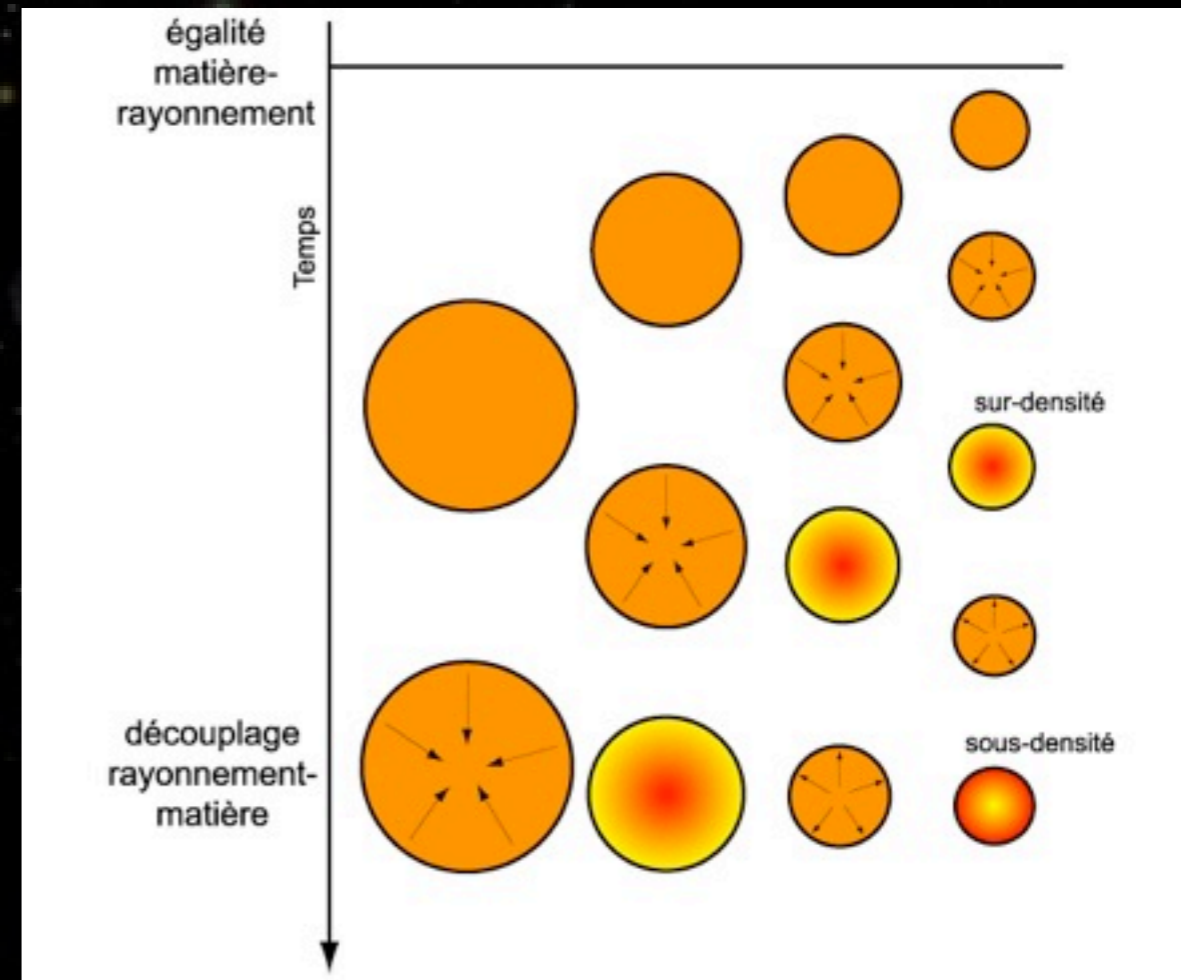
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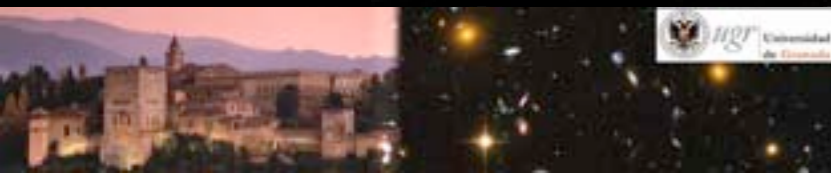
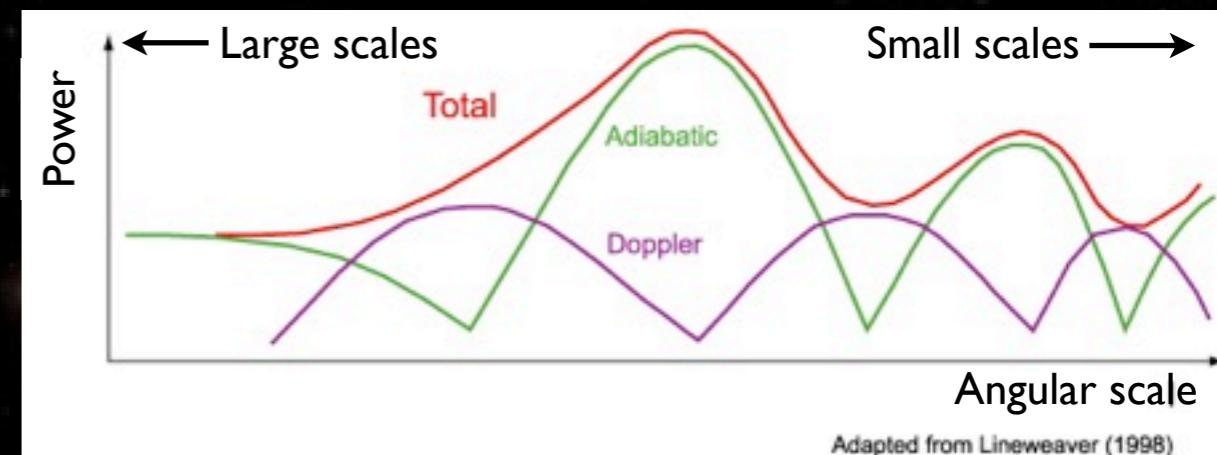
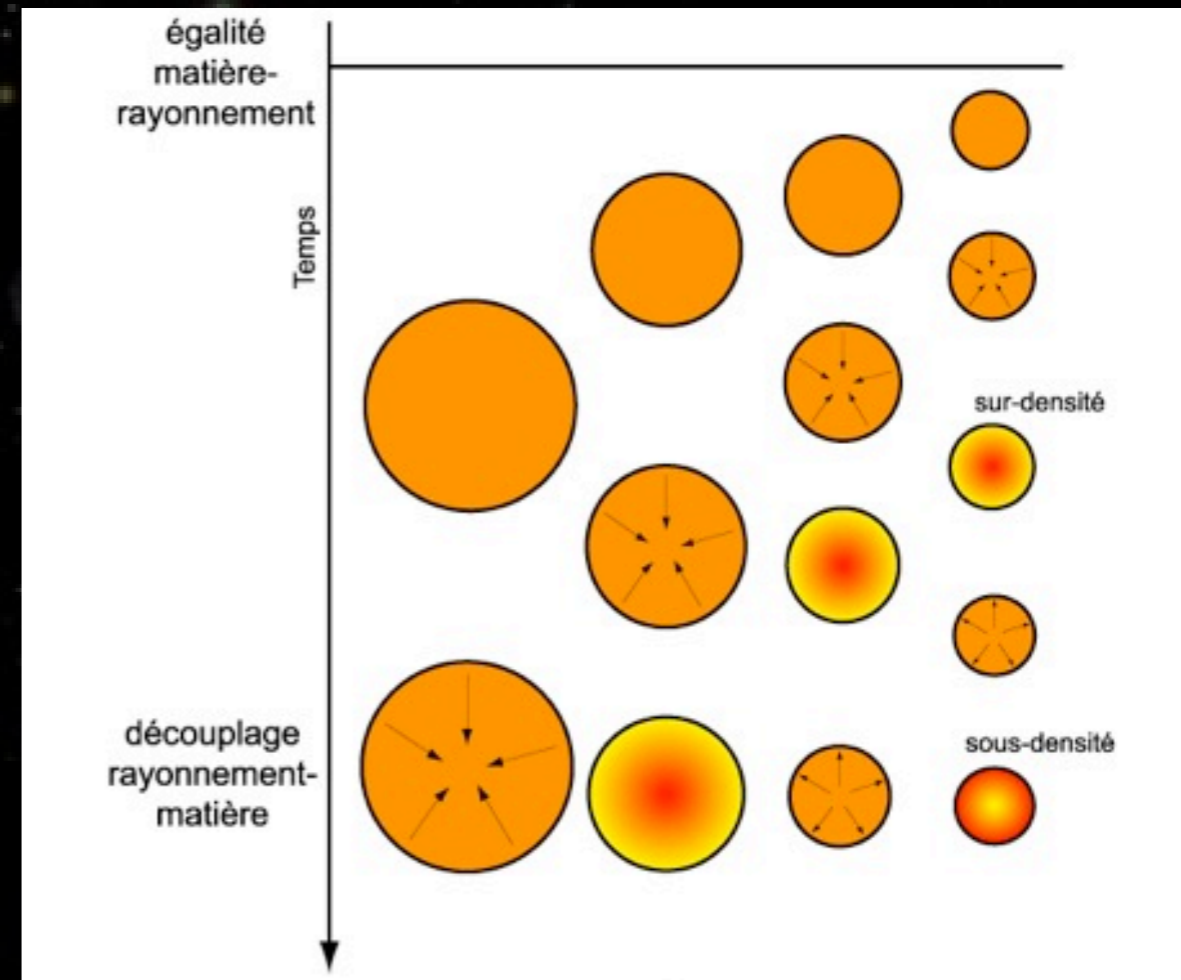
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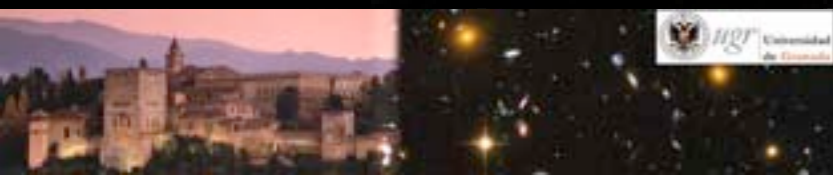
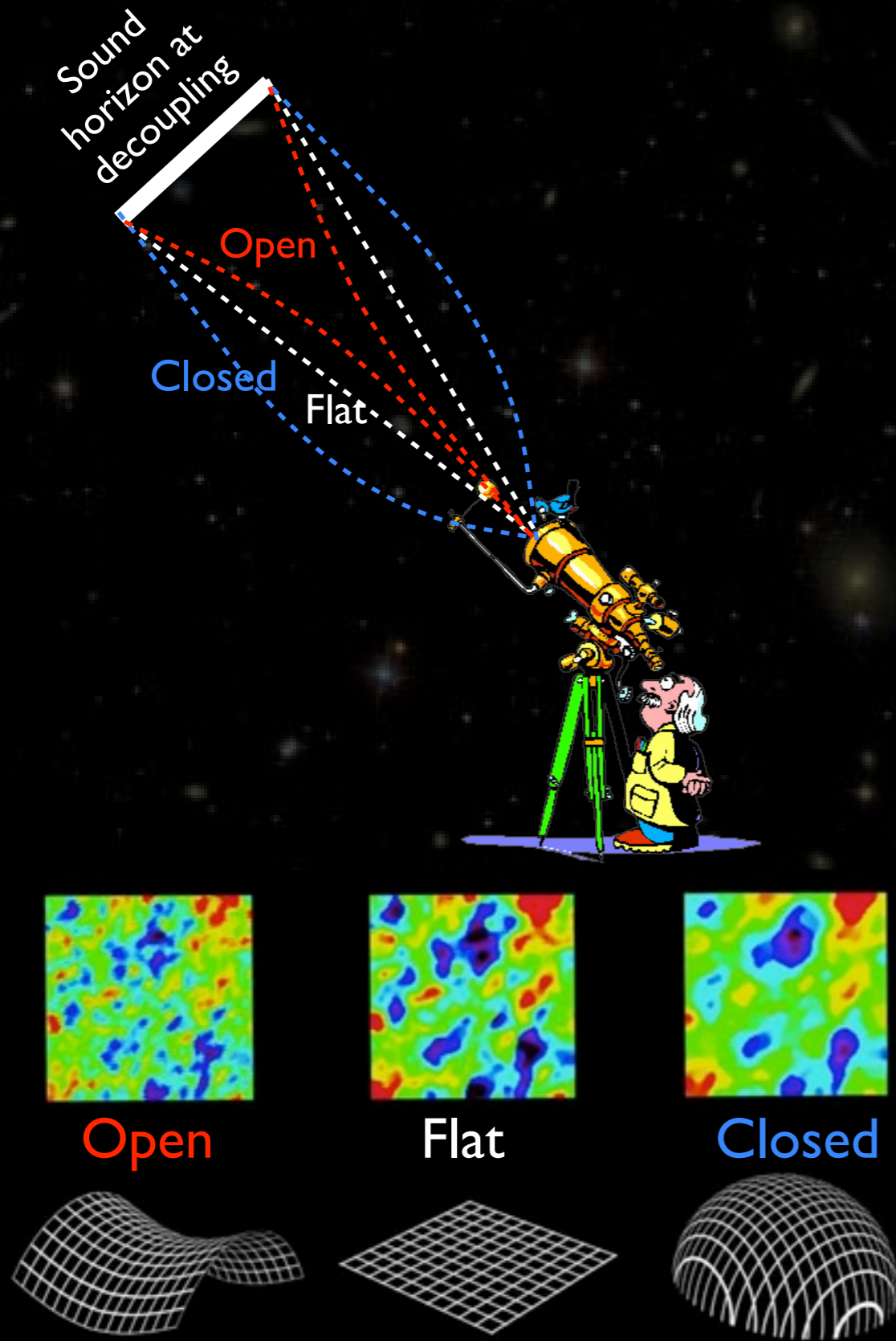
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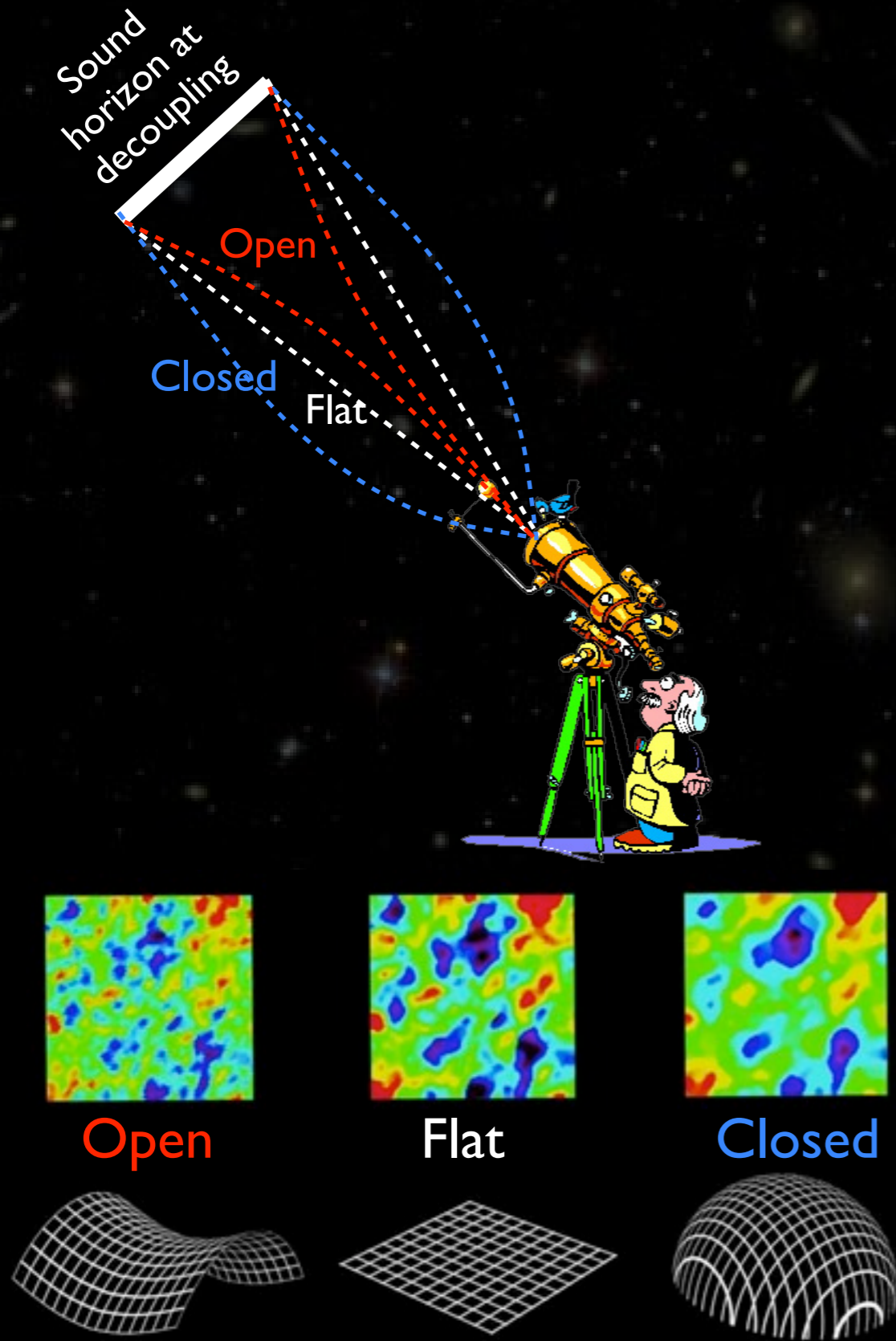
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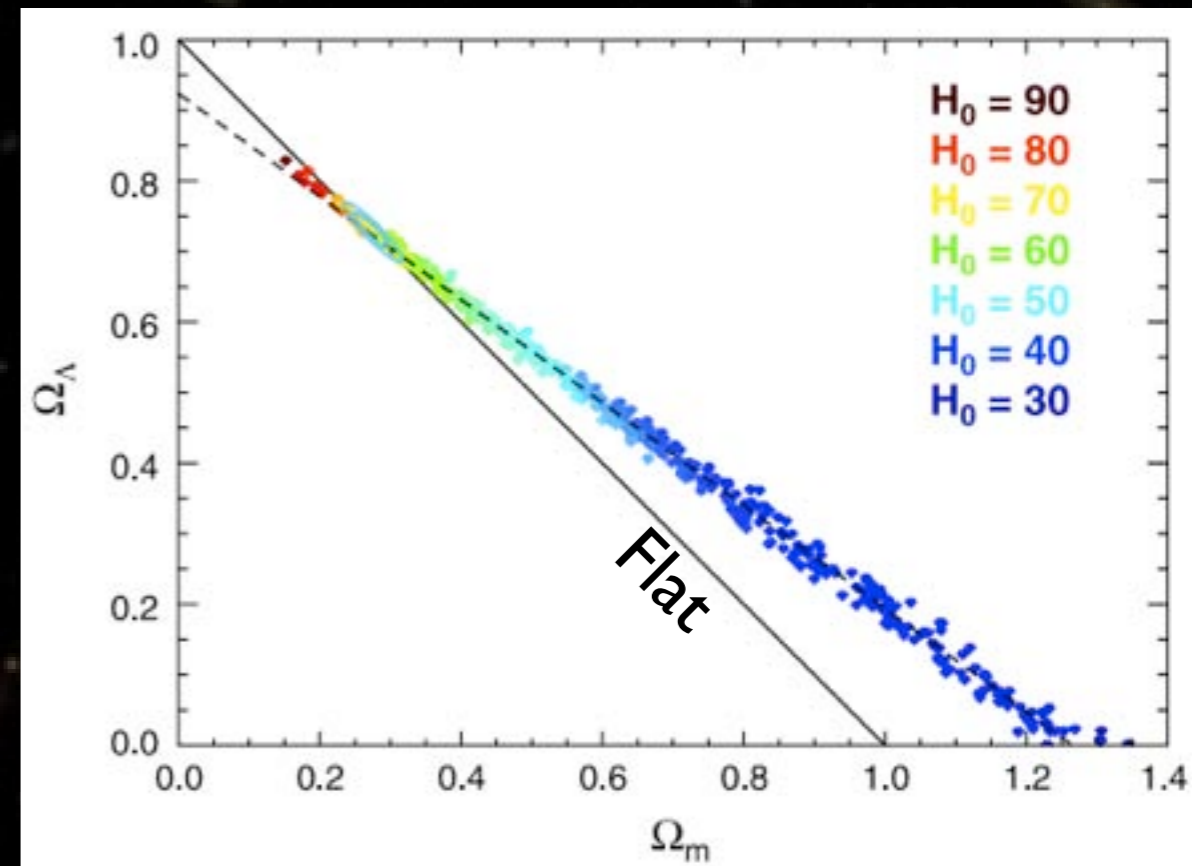
# Influence of the cosmological parameters



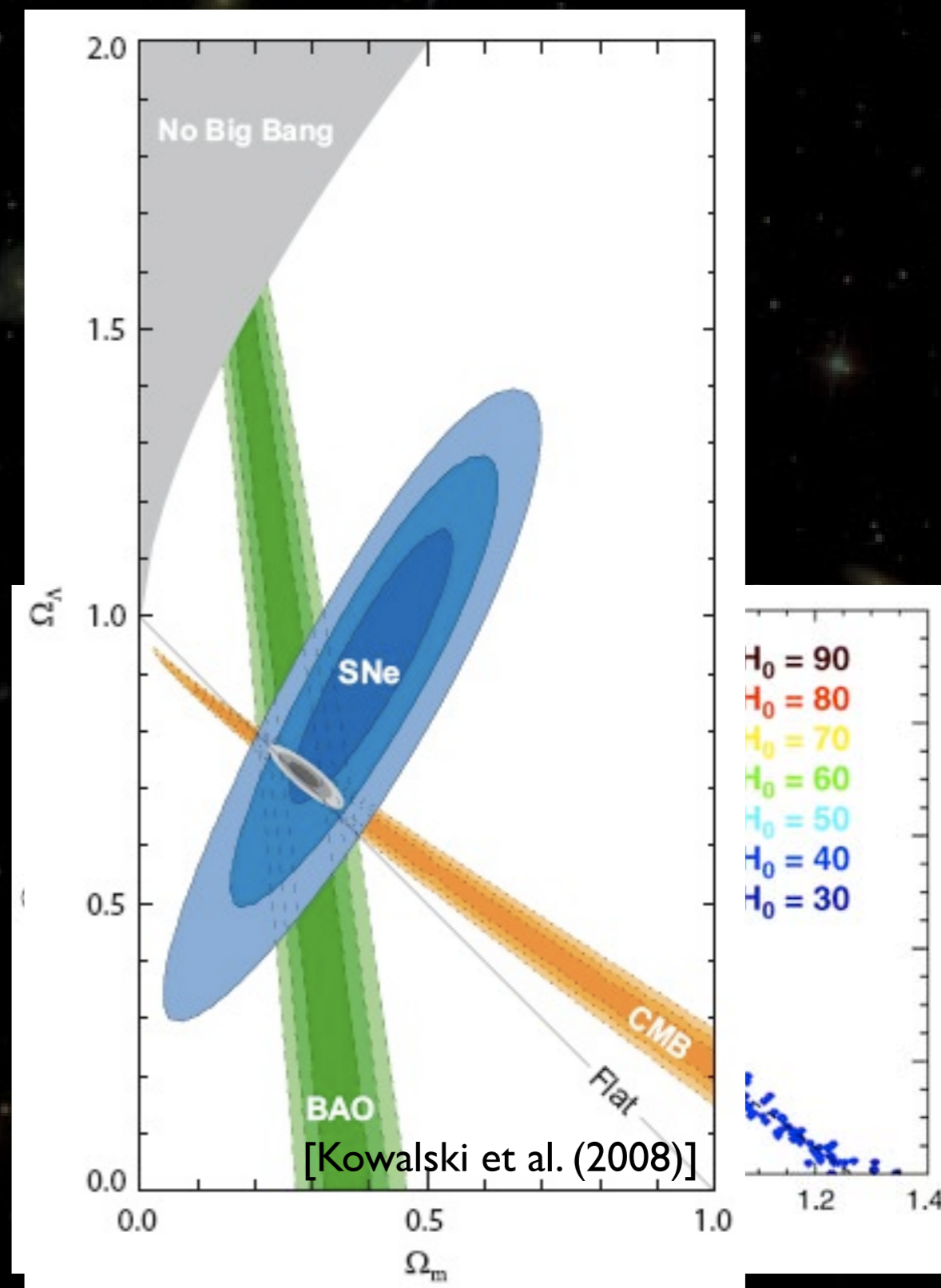
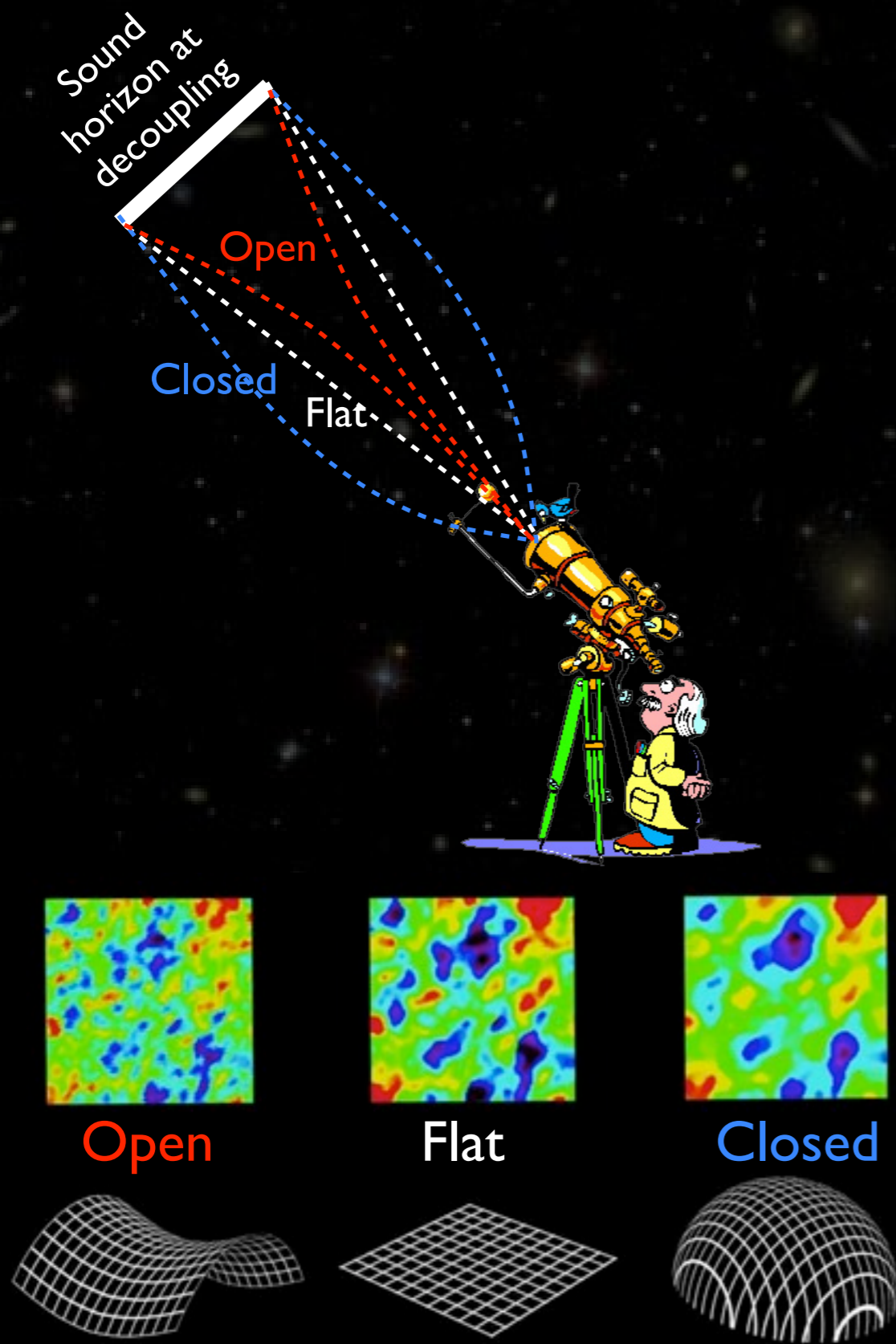
# Influence of the cosmological parameters



Actually slightly more complex due to the degeneracy with  $H$  et  $\Lambda$



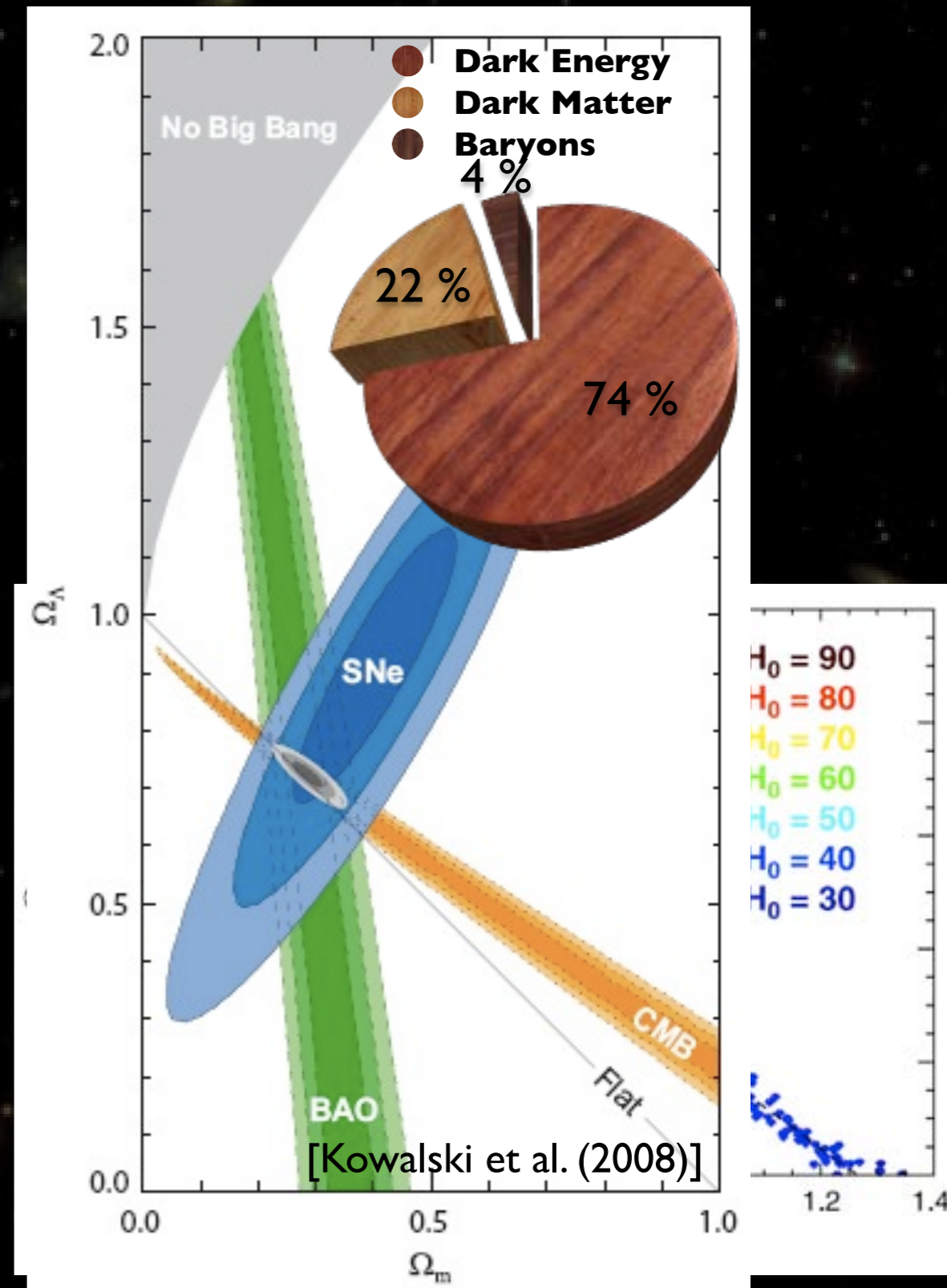
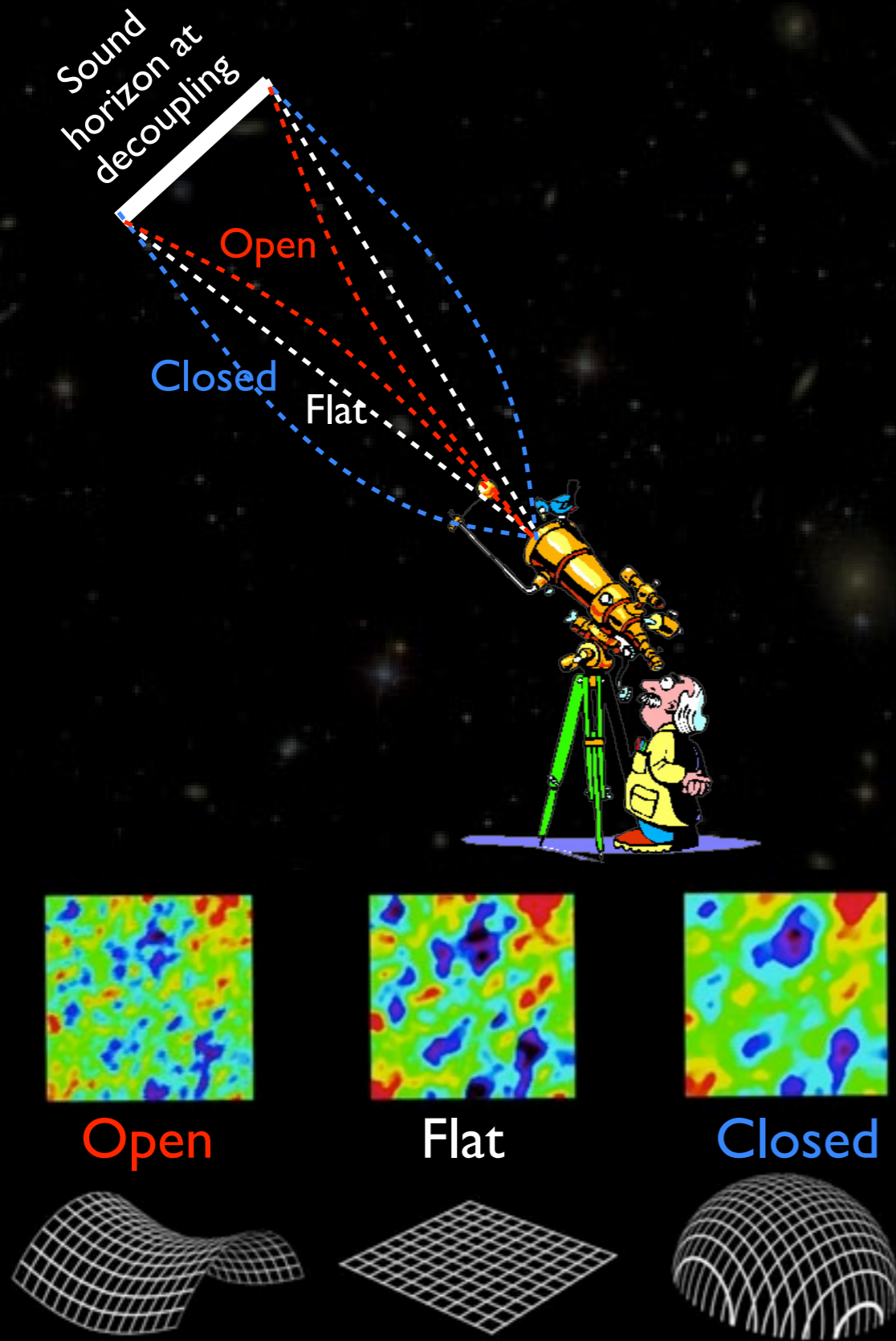
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WMAP 7y



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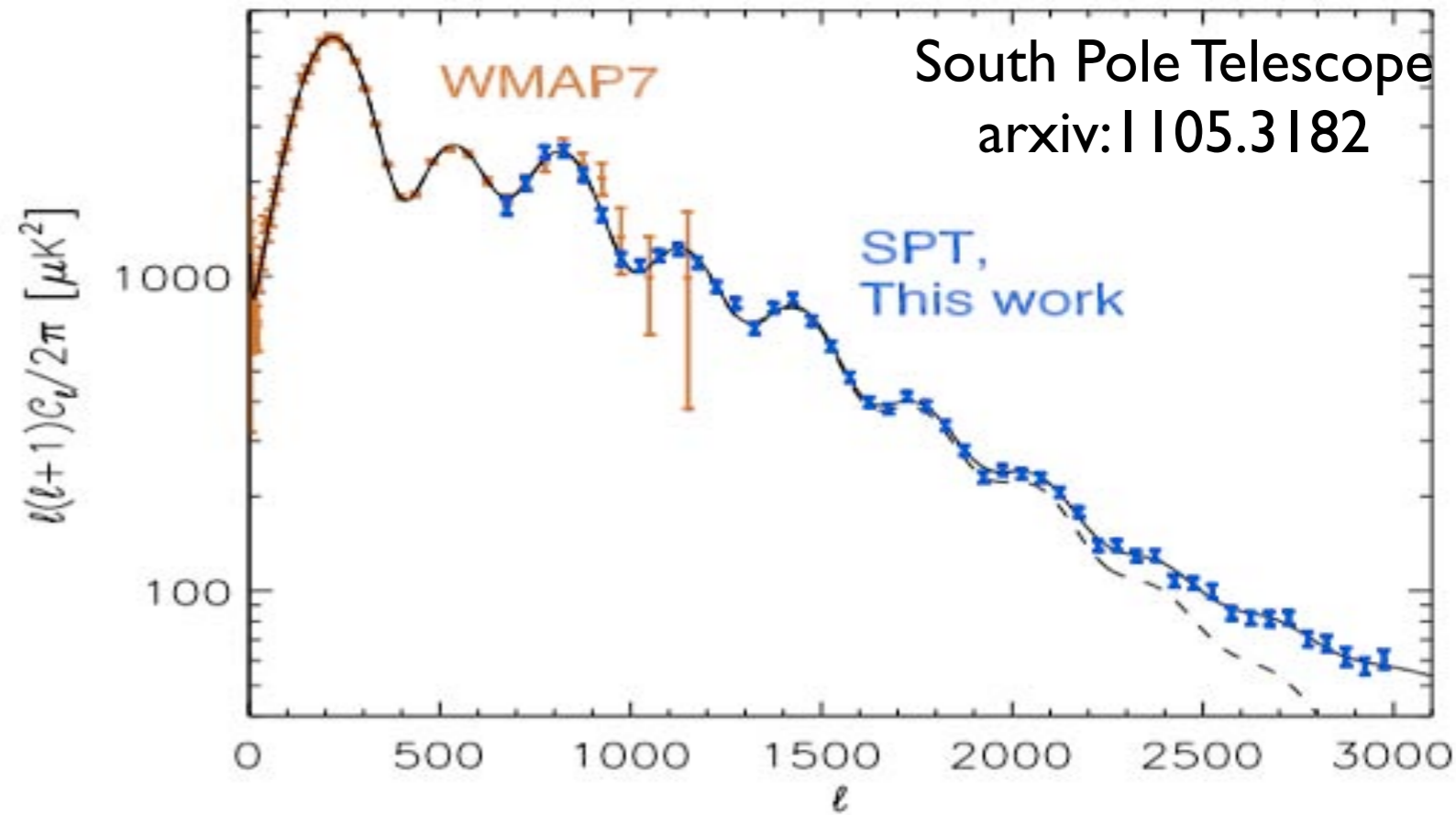
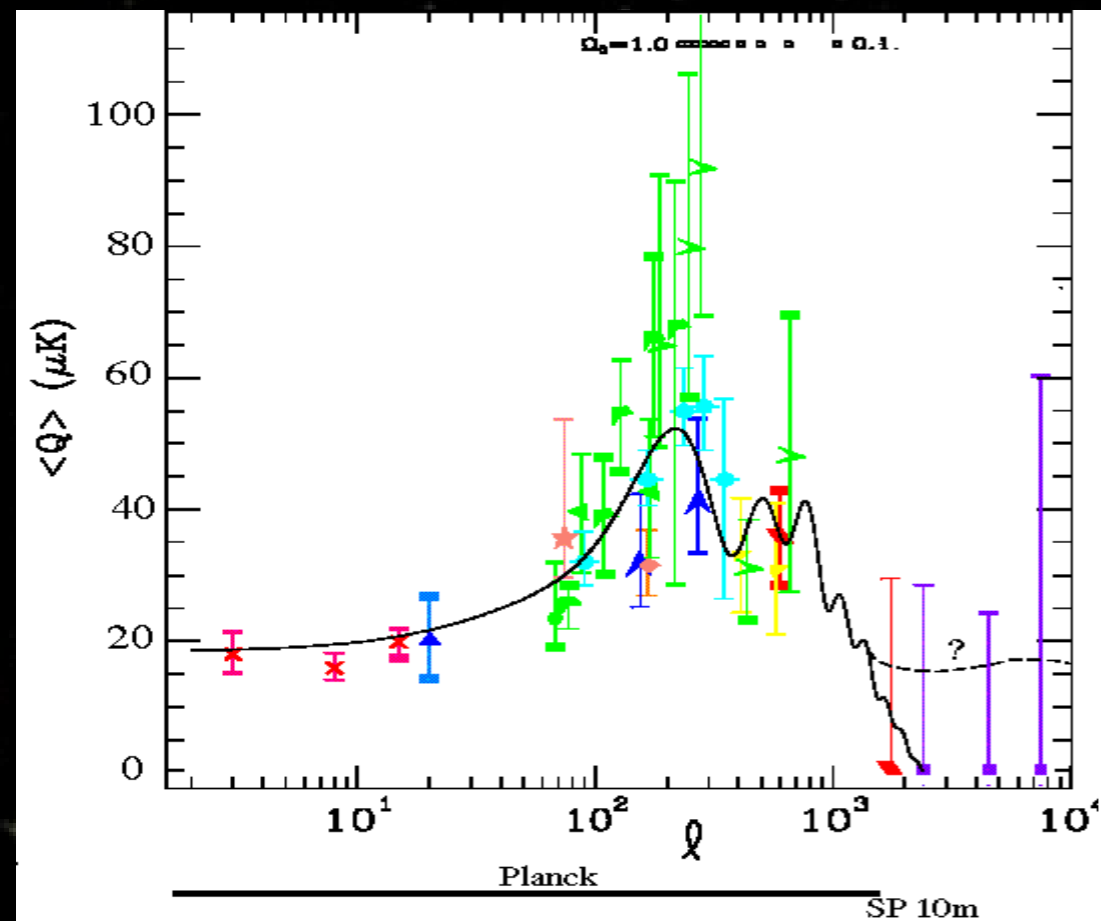


WMAP 7y





# Tremendous progress over the last decade

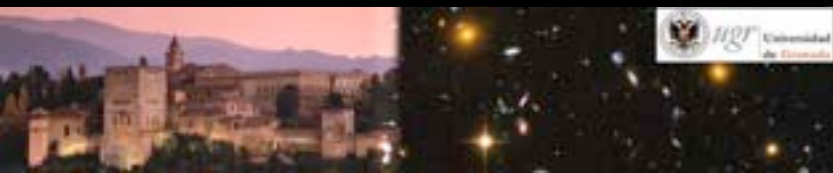


1999

2011

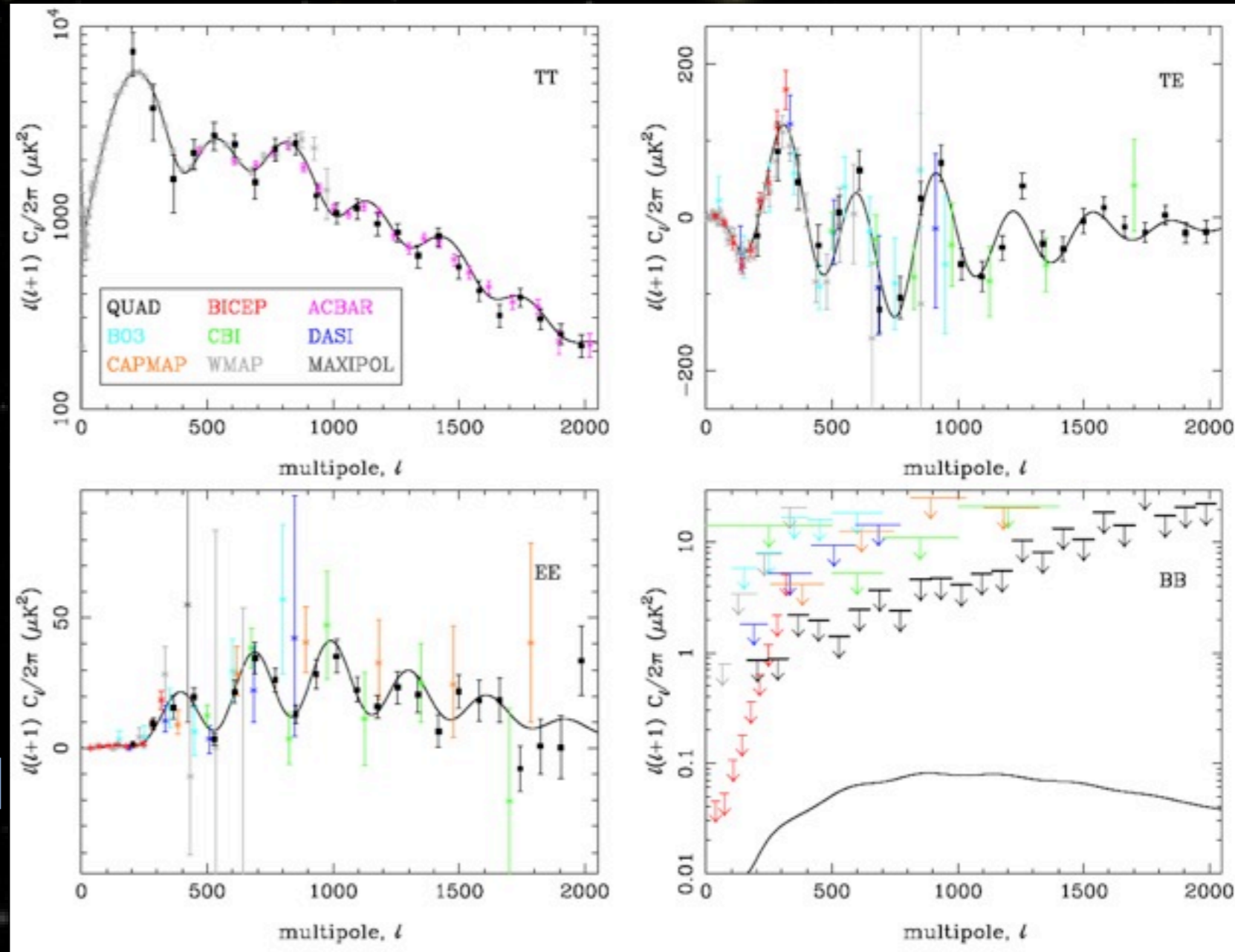
Huge success : thousands of independant points fitted with less than 10 parameters and a  $\chi^2/\text{ndf}$  about 1

Theoretical curve predicted in 1987 [Bond & Efstathiou] without any data ...

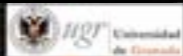
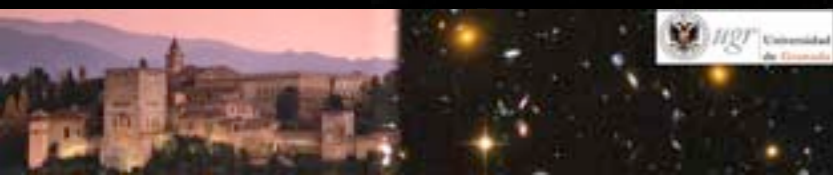


# CMB Polarization

- Predicted long ago
  - ★ electrons/photons scattering before decoupling
- Detection 2001
  - ★ DASI et CBI (interferometers)
- Later measurements:
  - ★ WMAP, QUAD, BICEP ...
  - ★ Perfect agreement with temperature measurements
- Correspondance between TT peaks and EE troughs
  - ★ Typical of adiabatic primordial fluctuations (generated by inflation for instance ...)

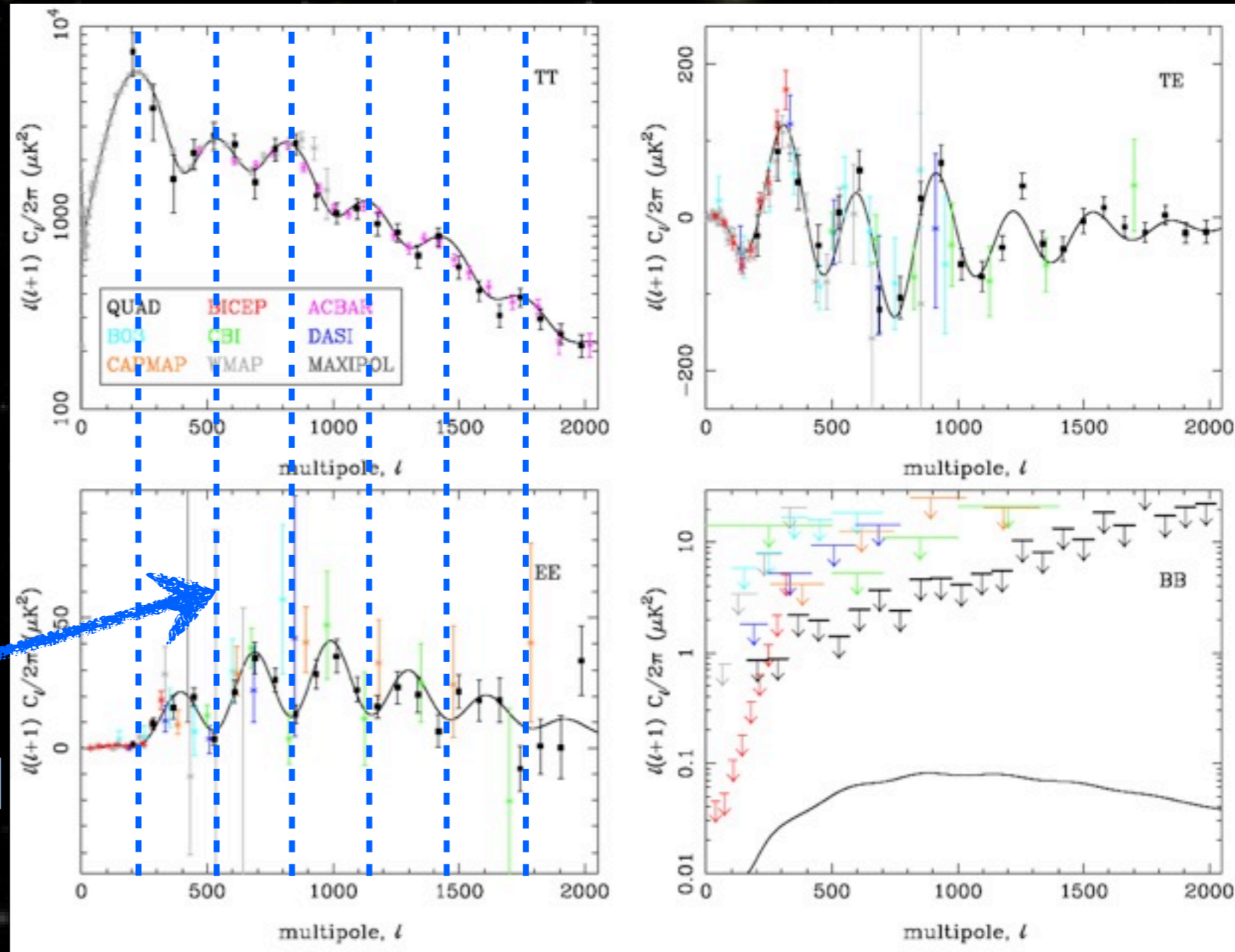


[QUAD Collaboration: Arxiv:0906.1003]

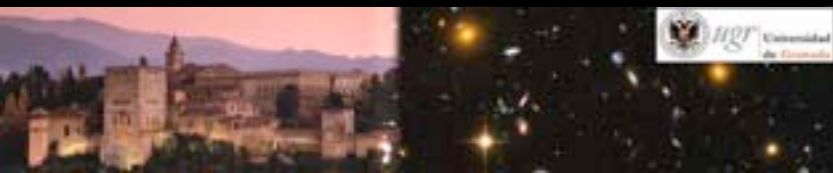


# CMB Polarization

- Predicted long ago
  - ★ electrons/photons scattering before decoupling
- Detection 2001
  - ★ DASI et CBI (interferometers)
- Later measurements:
  - ★ WMAP, QUAD, BICEP ...
  - ★ Perfect agreement with temperature measurements
- Correspondance between TT peaks and EE troughs
  - ★ Typical of adiabatic primordial fluctuations (generated by inflation for instance ...)

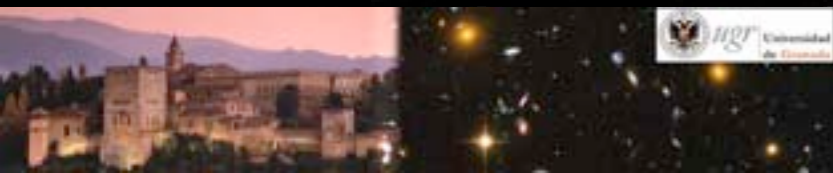


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# Major steps for CMB

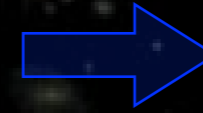
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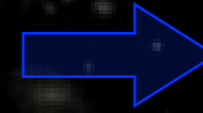
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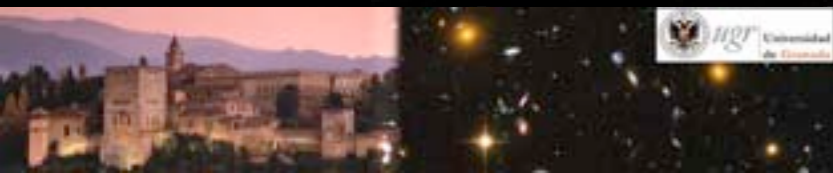
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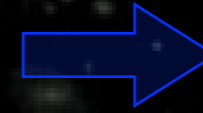
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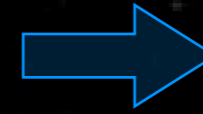
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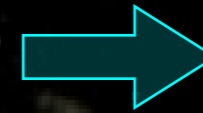
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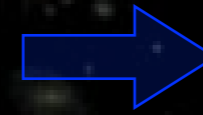
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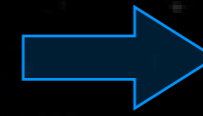
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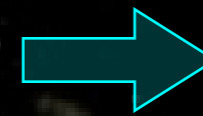
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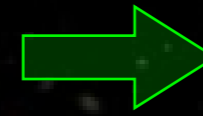
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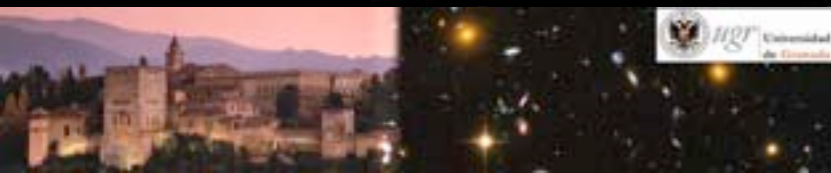
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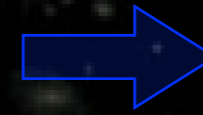




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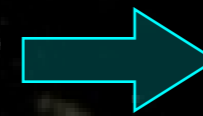
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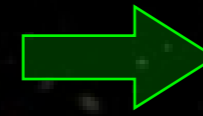
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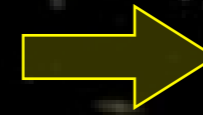
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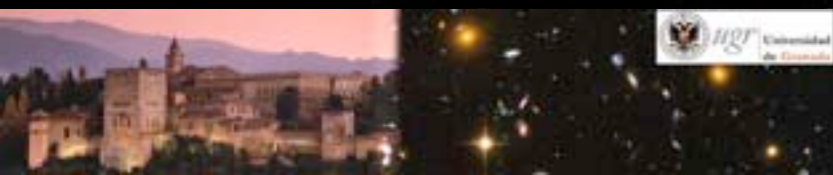
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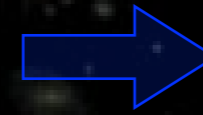
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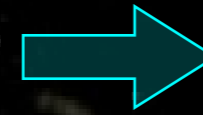
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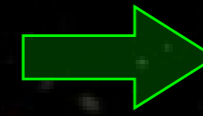
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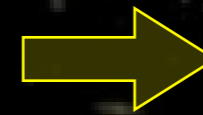
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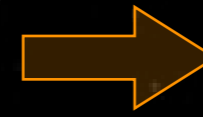
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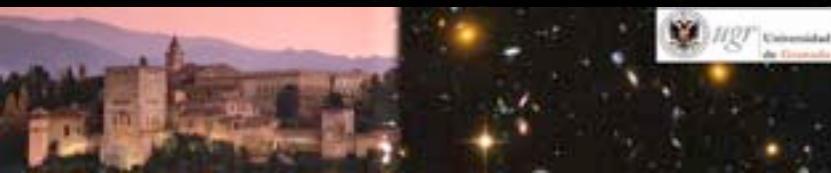
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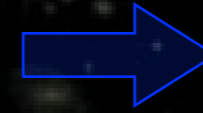
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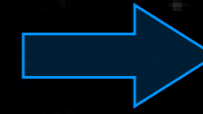
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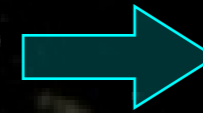
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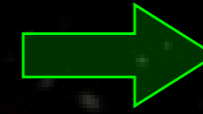
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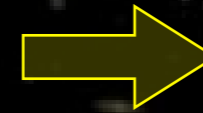
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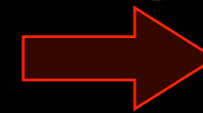
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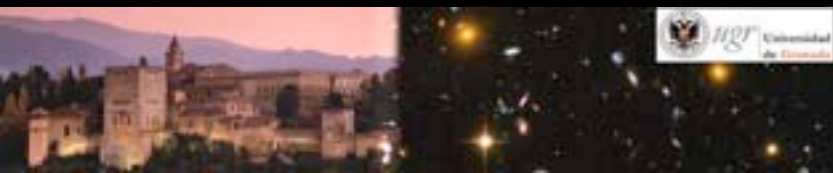
- Exploration of inflationary physics



# Primordial fluctuations: where are we standing ?


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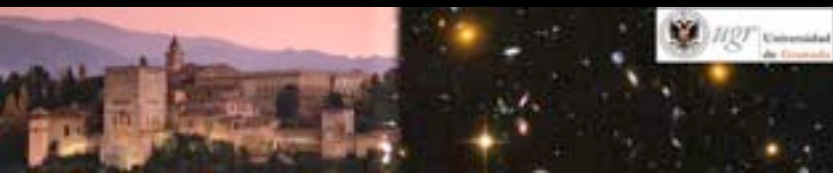
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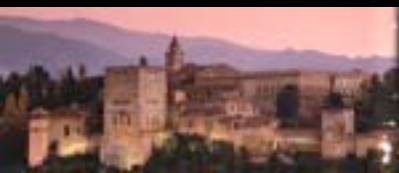
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# Dark Matter



Bullet Cluster



# Dark Matter

- Clusters dynamics:

- ★ Zwicky 1933

- velocity dispersion gives kinetic energy
- Distances between galaxies gives potential energy
- Virial theorem allows to infer mass
- 100-500 times larger than stellar mass

- ★ X emission from hot gaz in clusters

- gaz far more extended than galaxies in clusters
  - BUT:
    - gaz can account for a factor 2 in mass
    - gaz needs to be heated to  $10^7$ - $10^8$ K
- ➔ ~85% dark matter needed

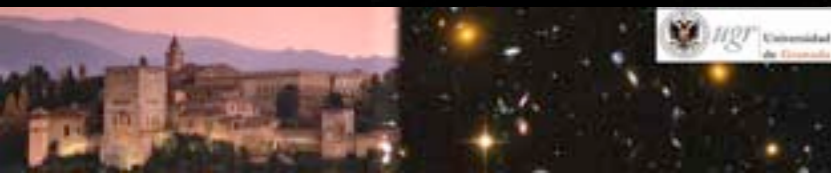
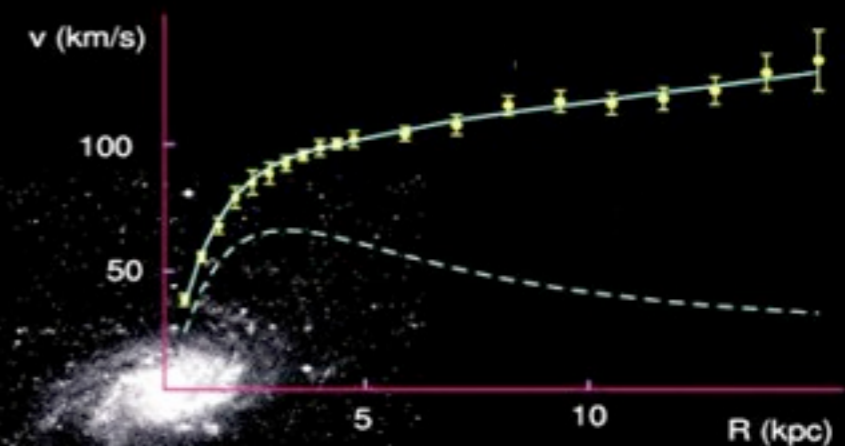


Coma cluster (image X: ROSAT)



- Galaxies rotation curves

- ★ Stars velocities too high in outer parts
- ★ well explained by a dark matter halo with ~ 200 times the stellar mass



# Dark Matter

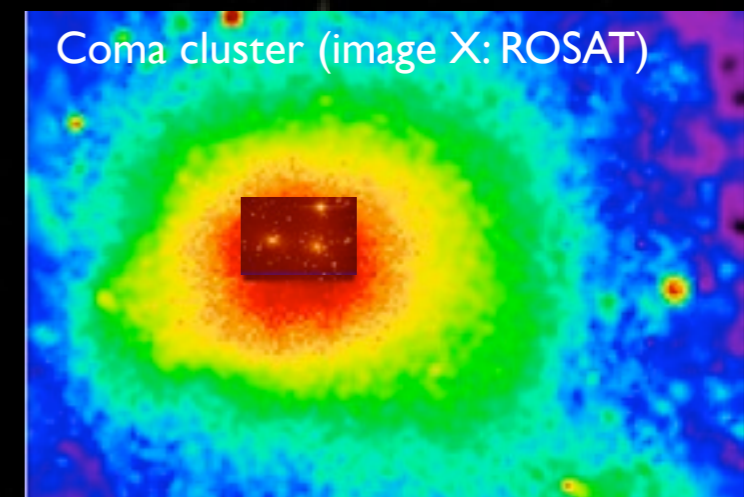
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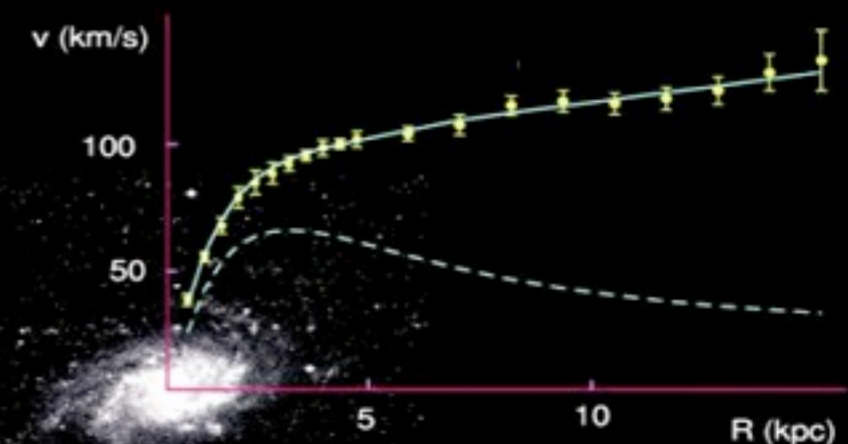
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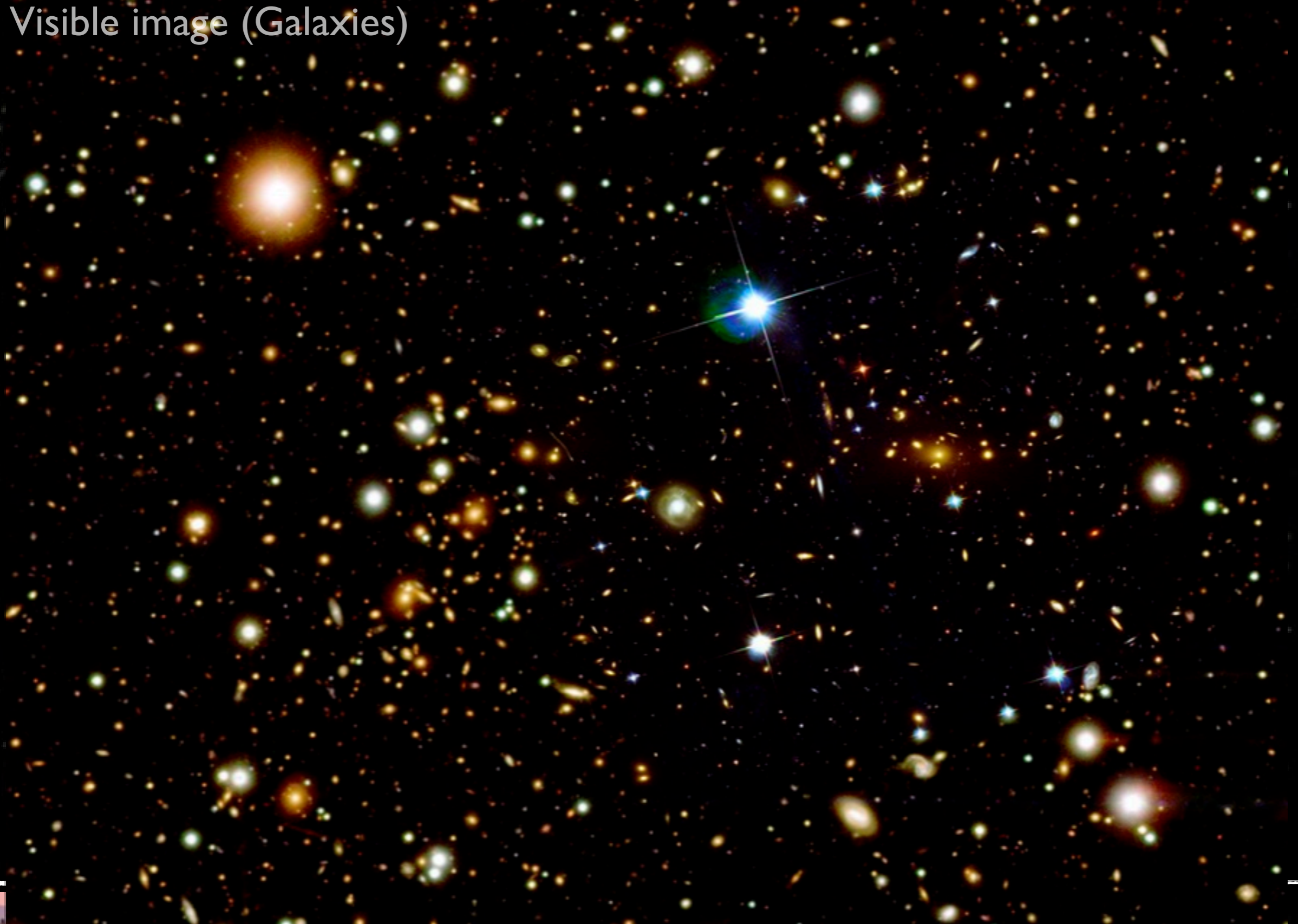
M33





# The Bullet Cluster

Visible image (Galaxies)



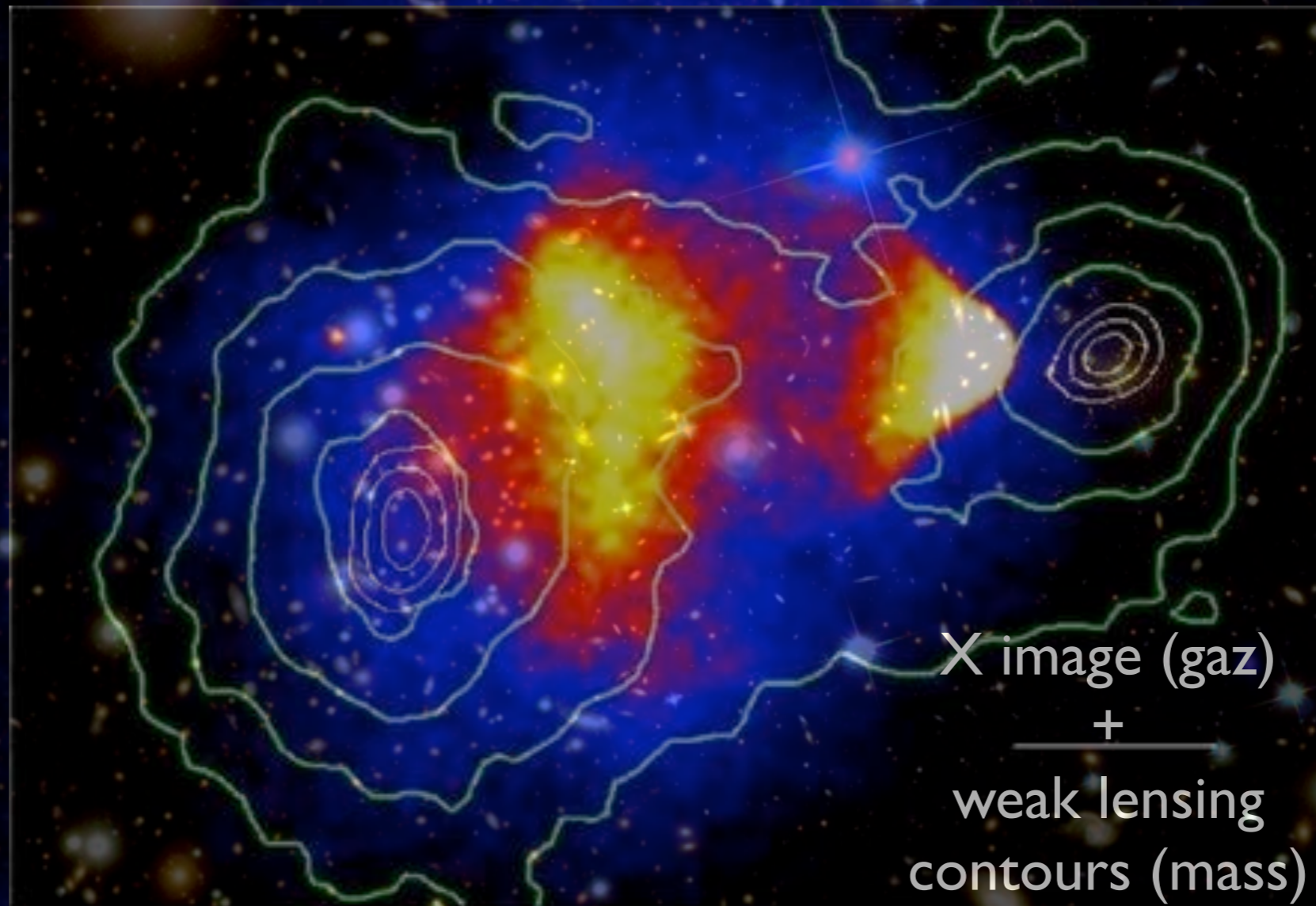
# The Bullet Cluster



X image (gaz) - Chandra



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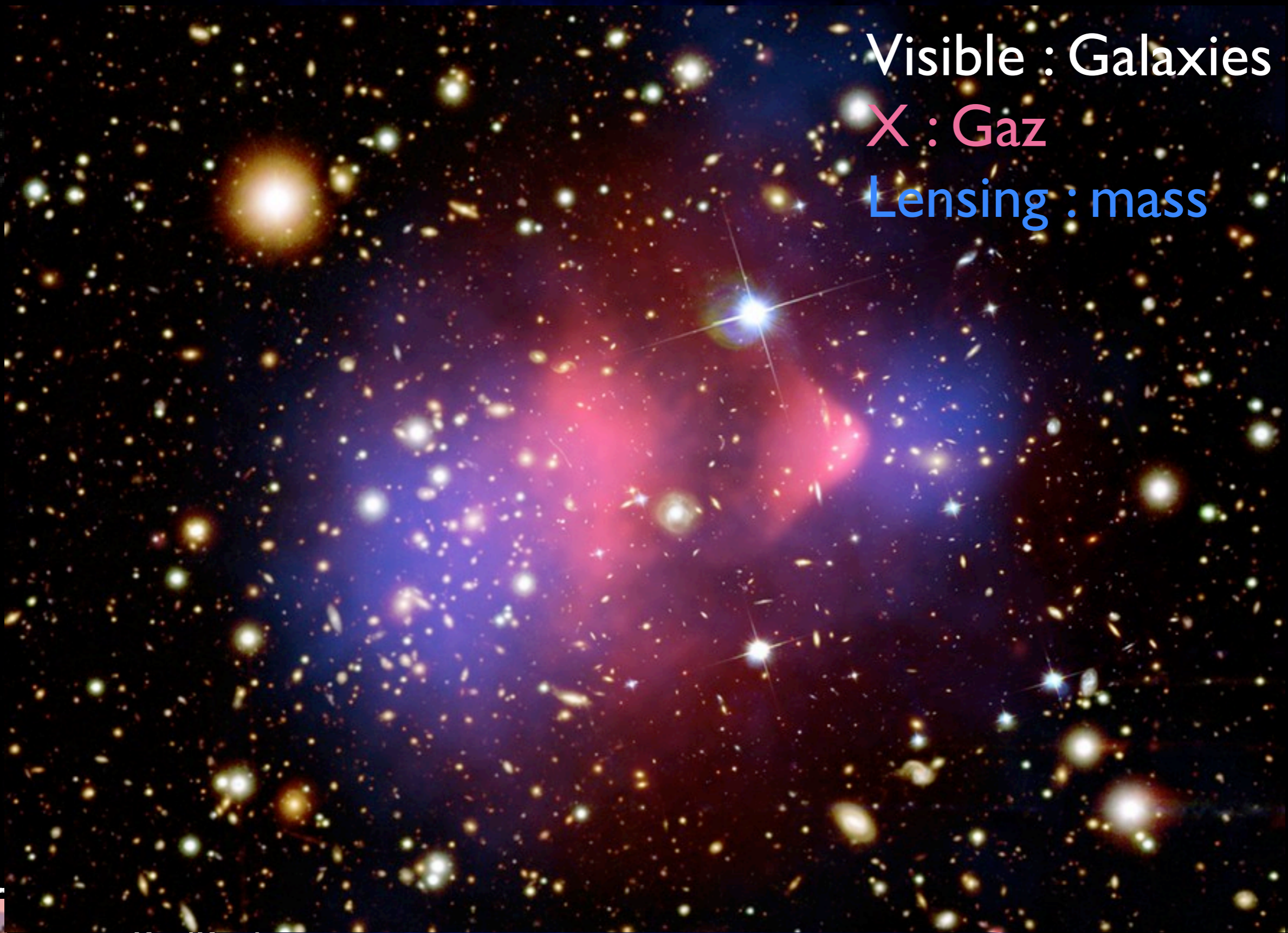


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Visible : Galaxies

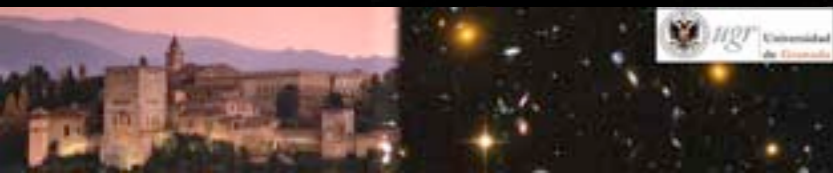
X : Gaz

Lensing : mass

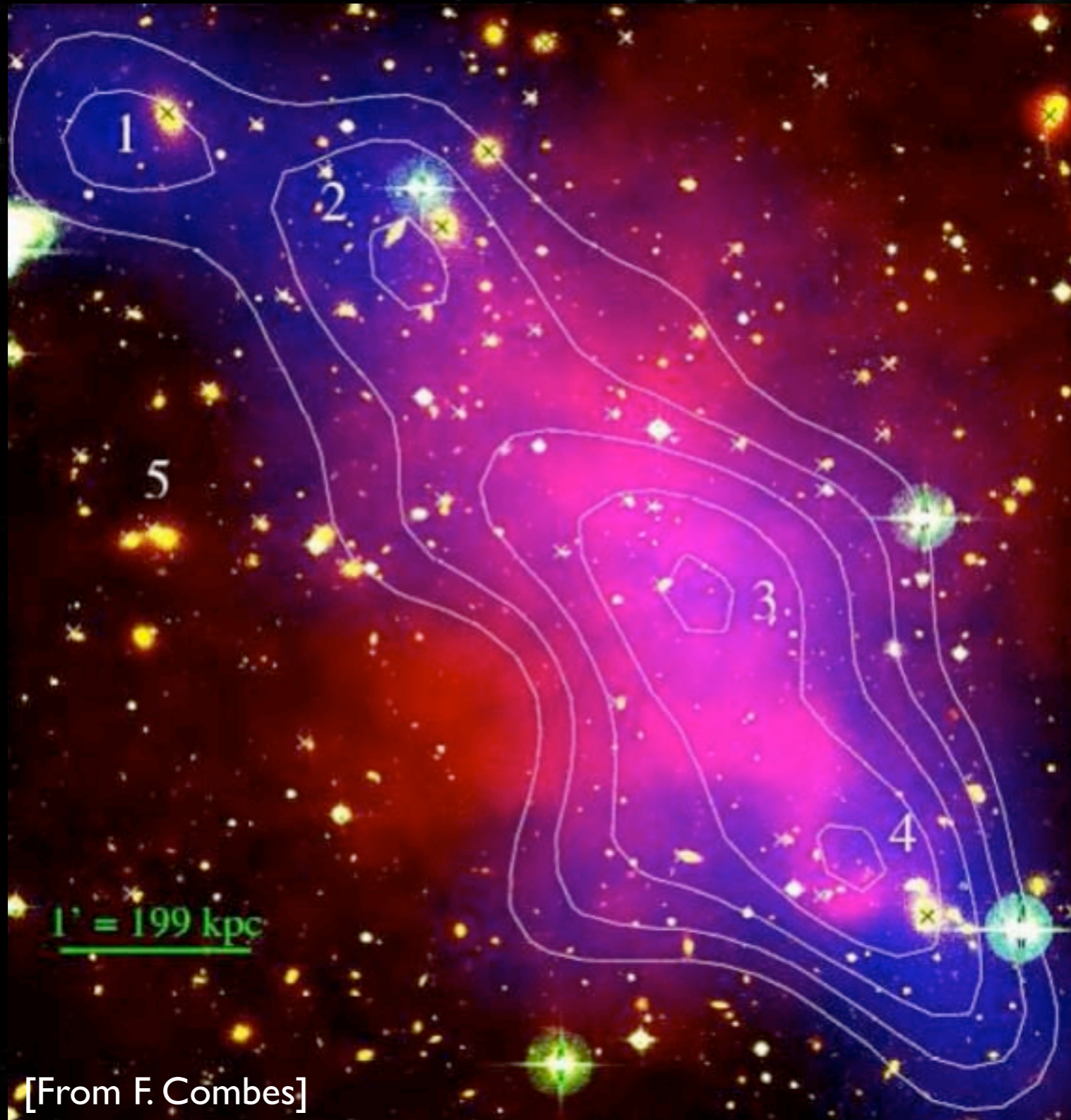


# Bullet Cluster

- Collision of two clusters
- Galaxies and dark matter are non collisional
  - ★ They pass without seeing each other
- Gaz is collisional
  - ★ it stays at the center and is heated
  - ★ Shock waves appear
- Strongest argument for Dark Matter ?



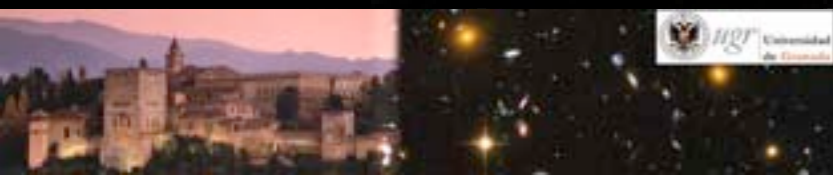
# but ... there exists a counter example...



Abell 520  
 $z=0.02$   
[Mahdavi et al. 2007]

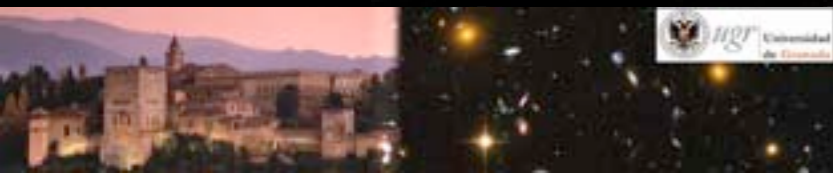
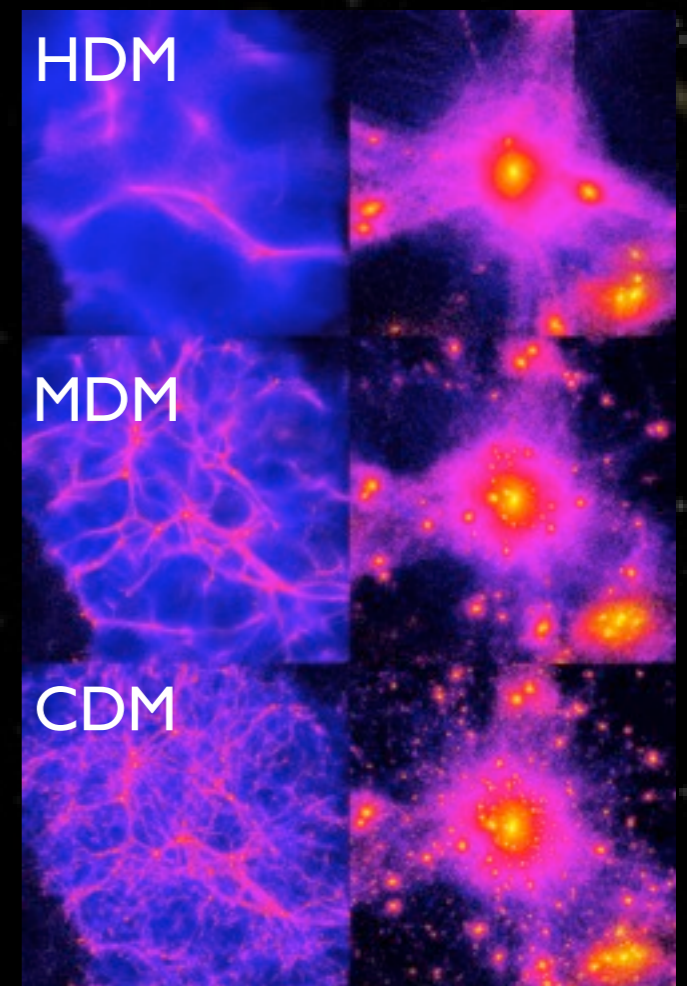
Dark Matter core  
coincides with gaz, not  
with galaxies...

[From F. Combes]



# Structure formation requires Dark Matter

- Primordial anisotropies
  - ★ very weak from CMB: 1/100 000
- Gravitational collapse
  - ★ starts at matter-radiation equality for ordinary matter
  - ★ Then expansion slows down the contraction
  - ★ This is not enough to explain observations
  - ★ More matter is needed
  - ★ Some matter needs to start collapsing earlier
    - Needs to be decoupled earlier than baryons...
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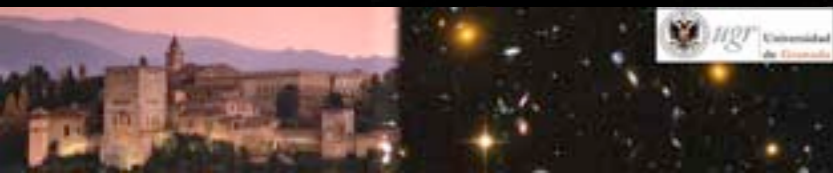
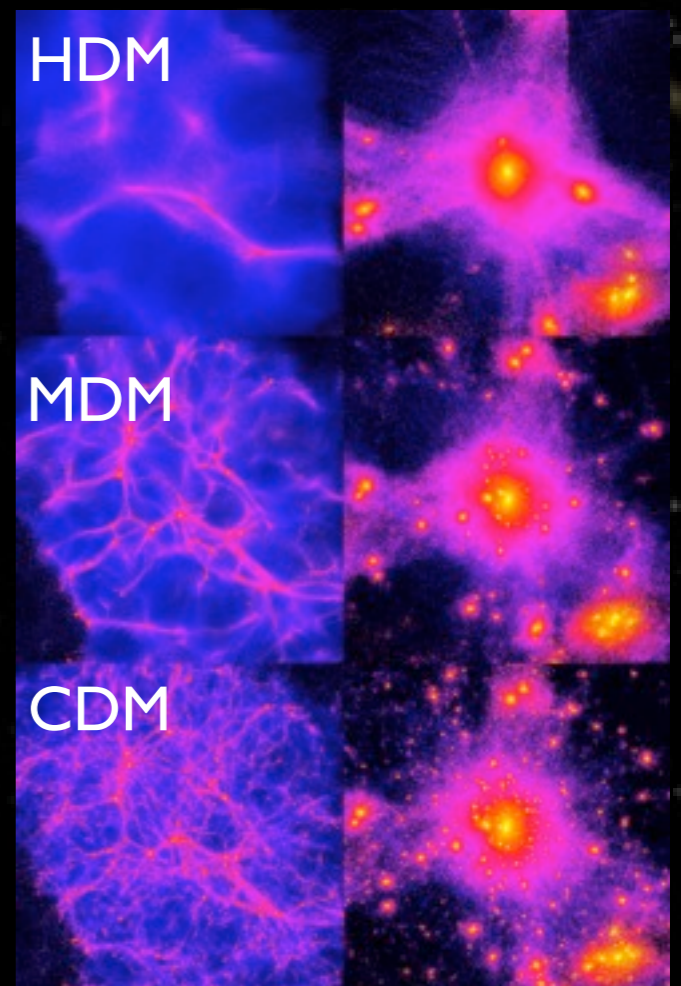
R = 6.0 Mpc

z = 10.155



a = 0.090

dLemard 2003





# Baryonic Acoustic Oscillations

## ● BAO:

### ★ Acoustic Oscillations:

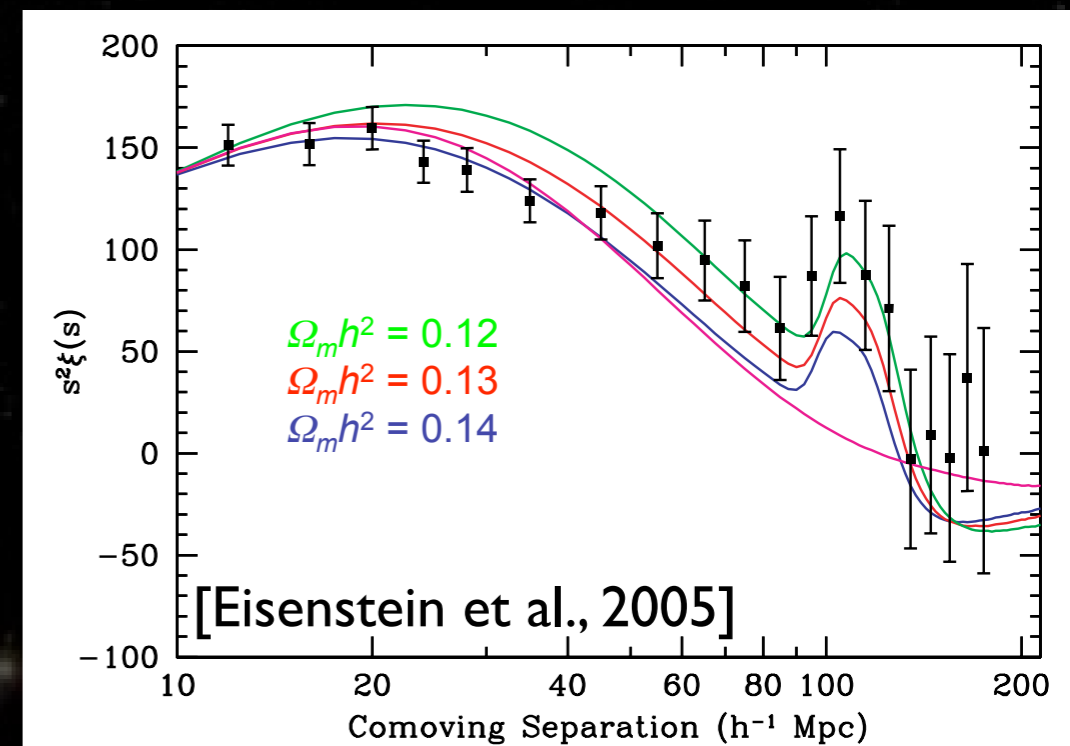
- between matter/radiation equality and decoupling
- Sound wave propagation
- Stops at decoupling
- ➔ Observed in the CMB

### ★ Now:

- Each D.M. peak surrounded by a spherical excess
- ➔ Galaxies preferably form there
- peak in the 2-pts correlation function of matter
- Standard ruler for angular distance test

### ★ Meaning:

- Prediction confirmed ! at the right place !
- This is among the strongest evidences for  $\Omega_m=0.3$  and the presence of Dark Matter



# Baryonic Acoustic Oscillations

## ● BAO:

### ★ Acoustic Oscillations:

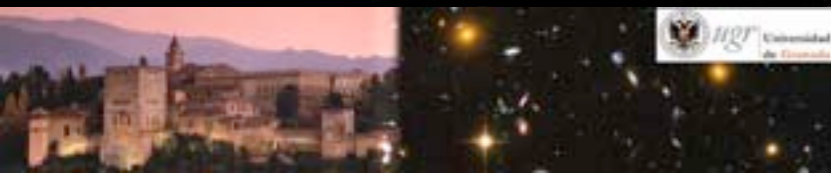
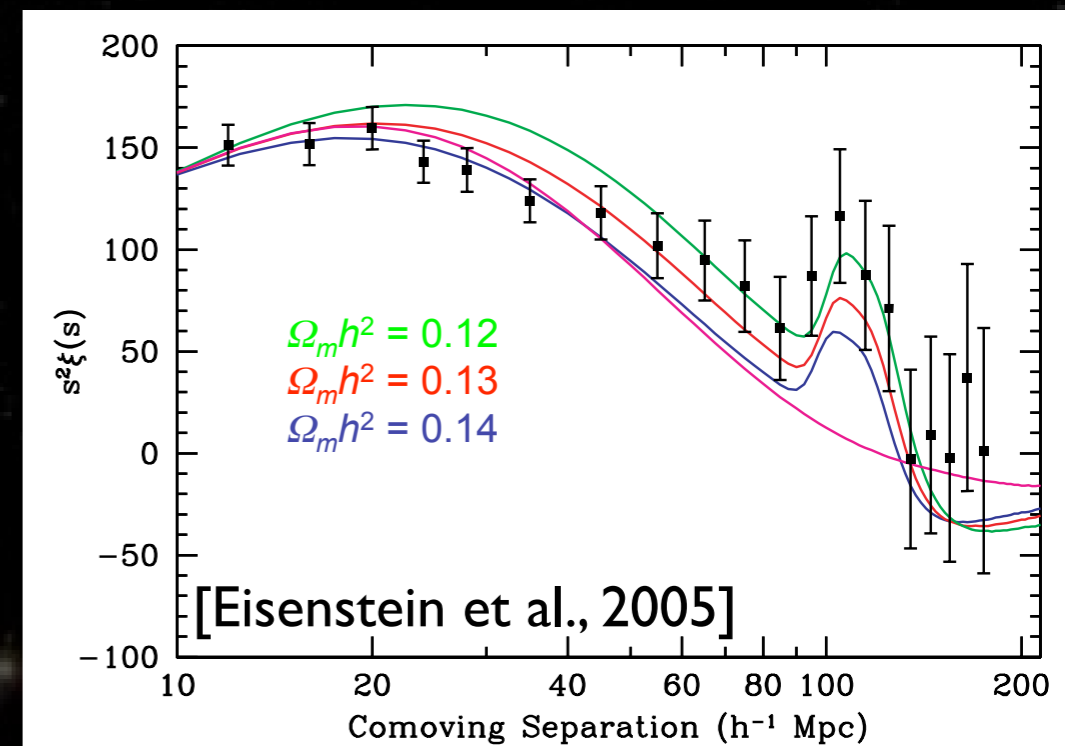
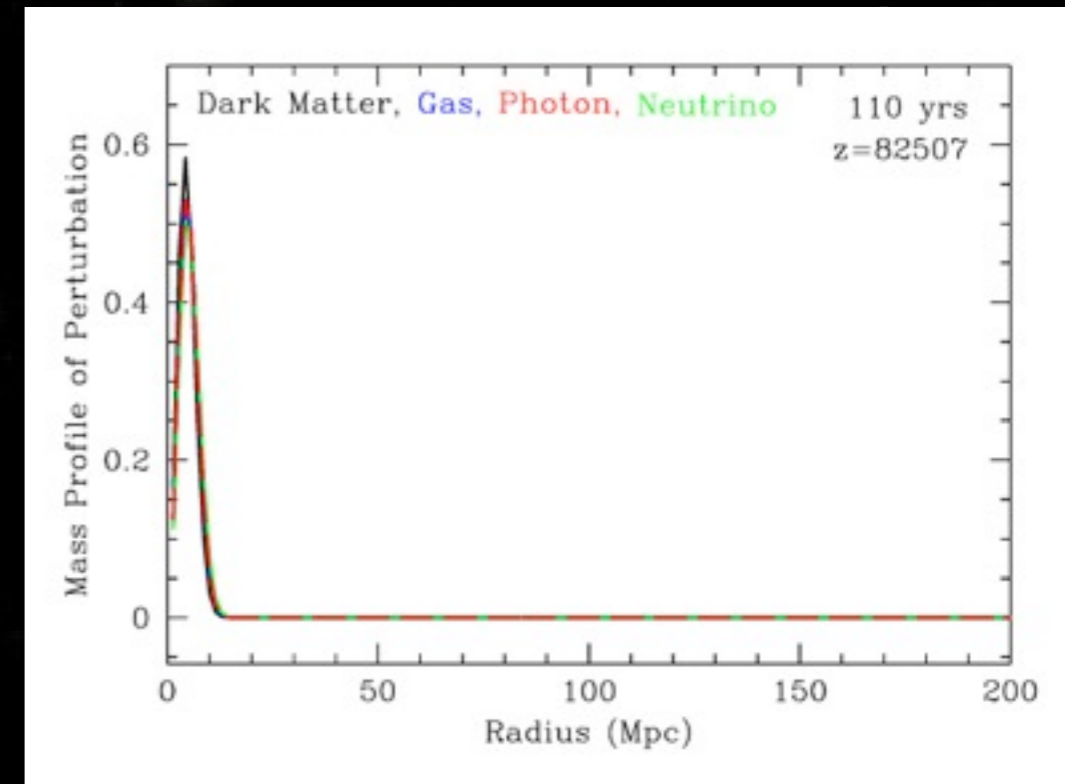
- between matter/radiation equality and decoupling
- Sound wave propagation
- Stops at decoupling
- ➔ Observed in the CMB

### ★ Now:

- Each D.M. peak surrounded by a spherical excess
- ➔ Galaxies preferably form there
- peak in the 2-pts correlation function of matter
- Standard ruler for angular distance test

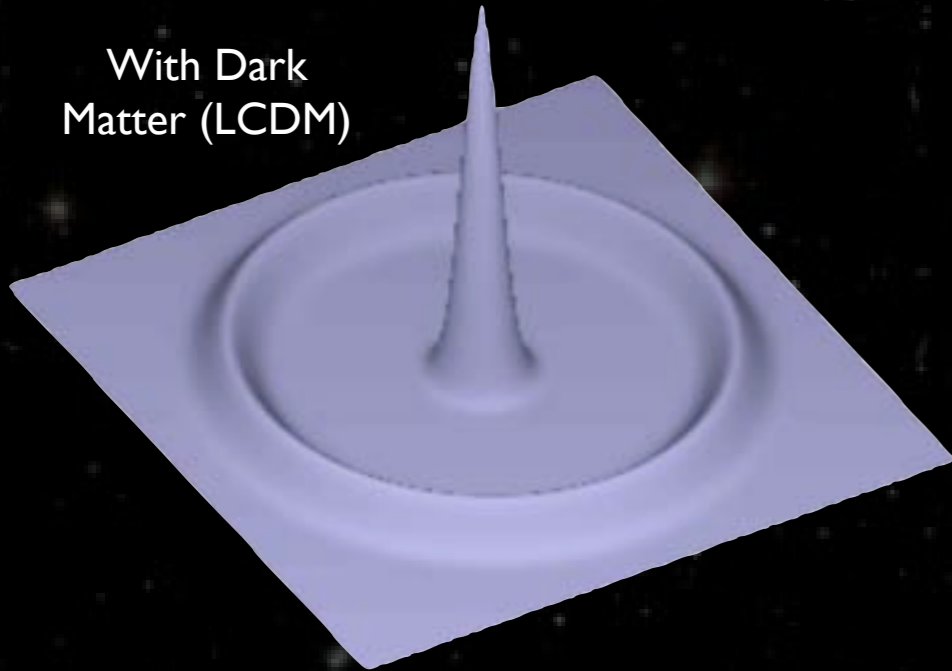
### ★ Meaning:

- Prediction confirmed ! at the right place !
- This is among the strongest evidences for  $\Omega_m=0.3$  and the presence of Dark Matter

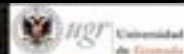
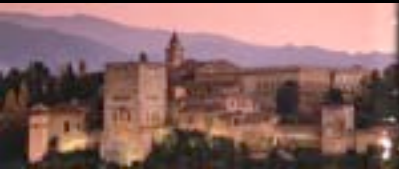
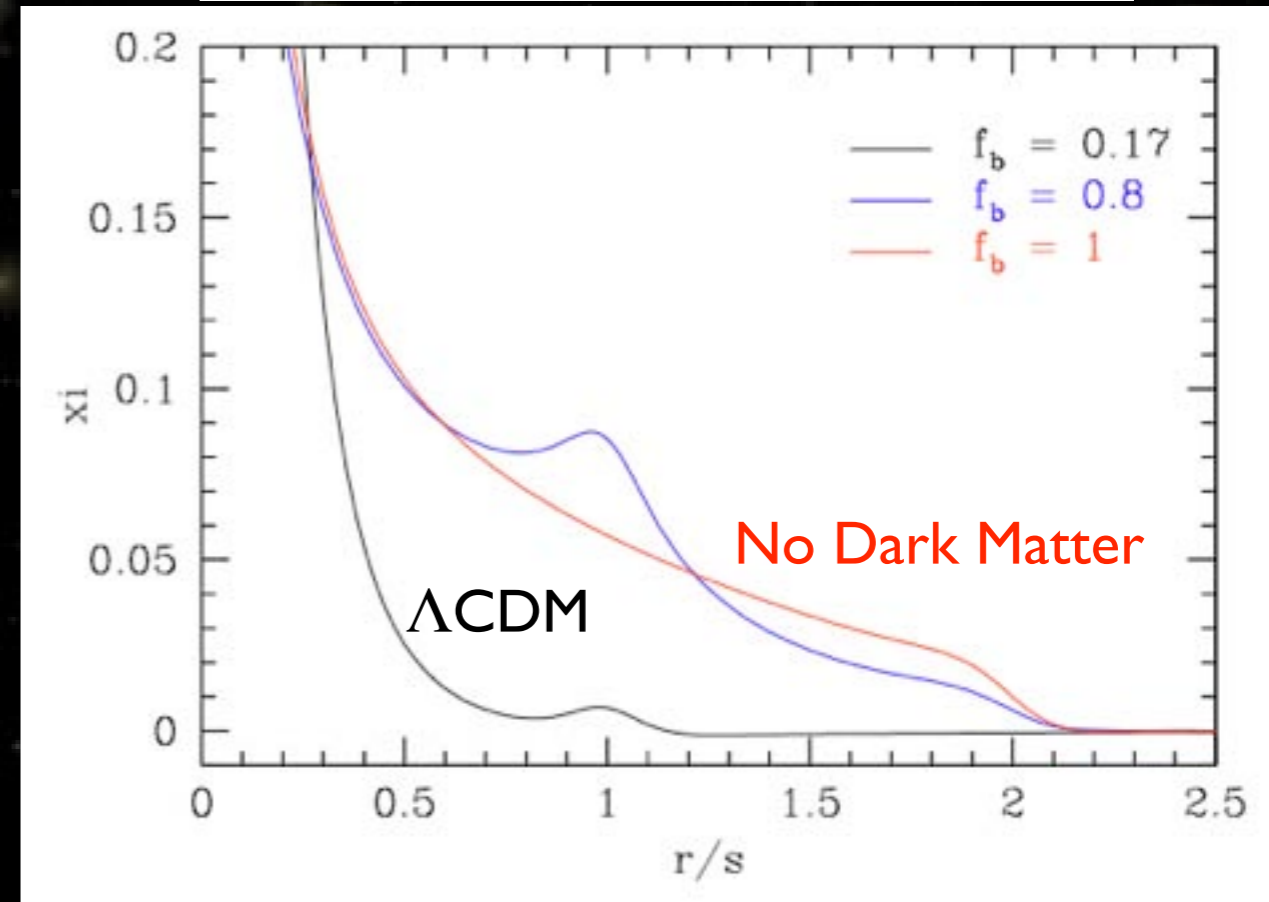
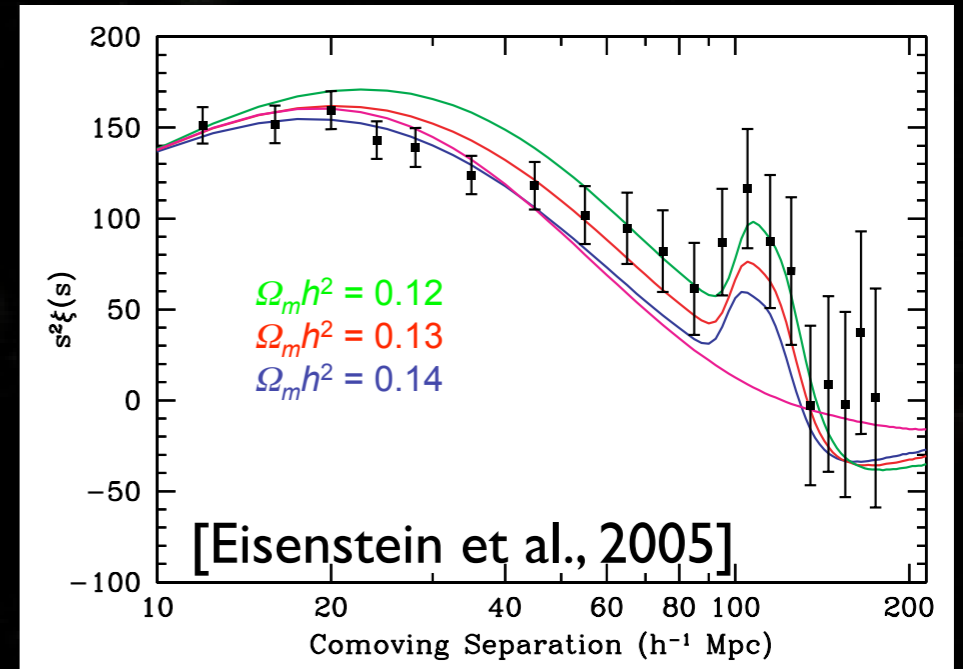
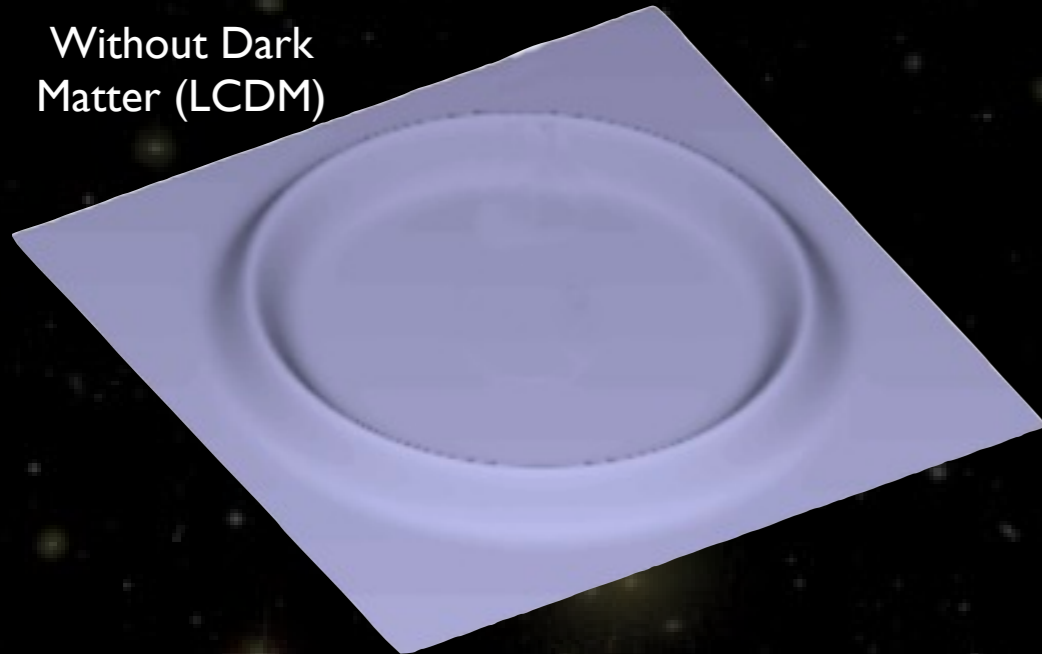


# BAO support Dark Matter

With Dark Matter (LCDM)

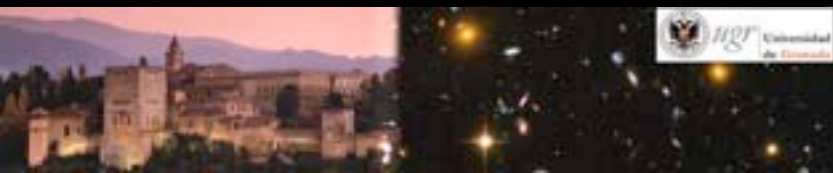
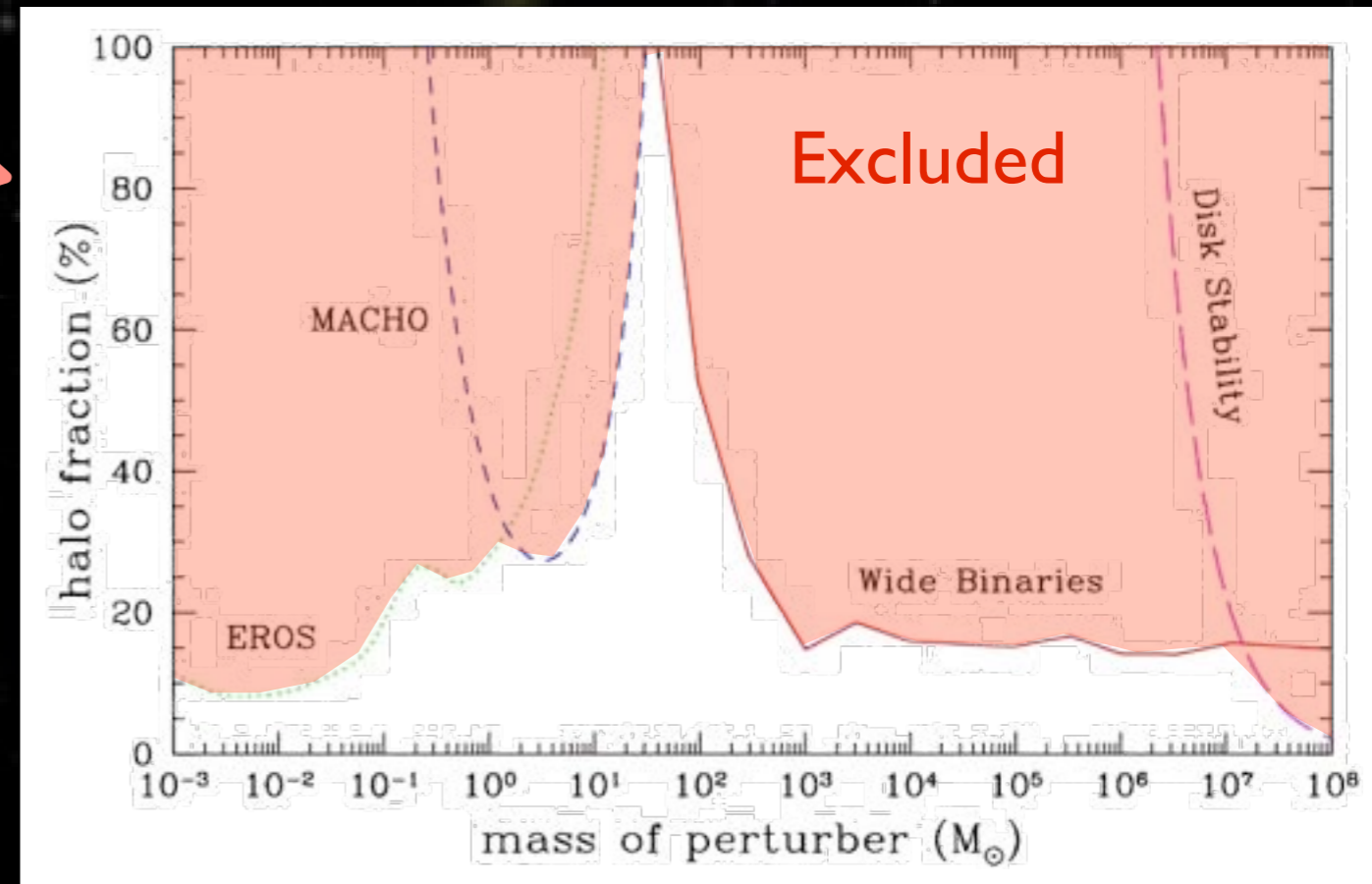


Without Dark Matter (LCDM)



# Candidates for Dark Matter

- Compact objects
  - Black holes, brown dwarfs
  - Essentially excluded in the late 90s
- Particle Physics
- Modifications of Gravity



# Candidates for Dark Matter

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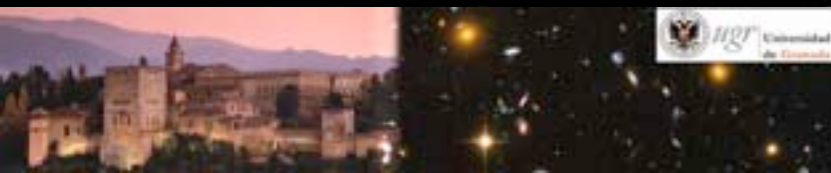
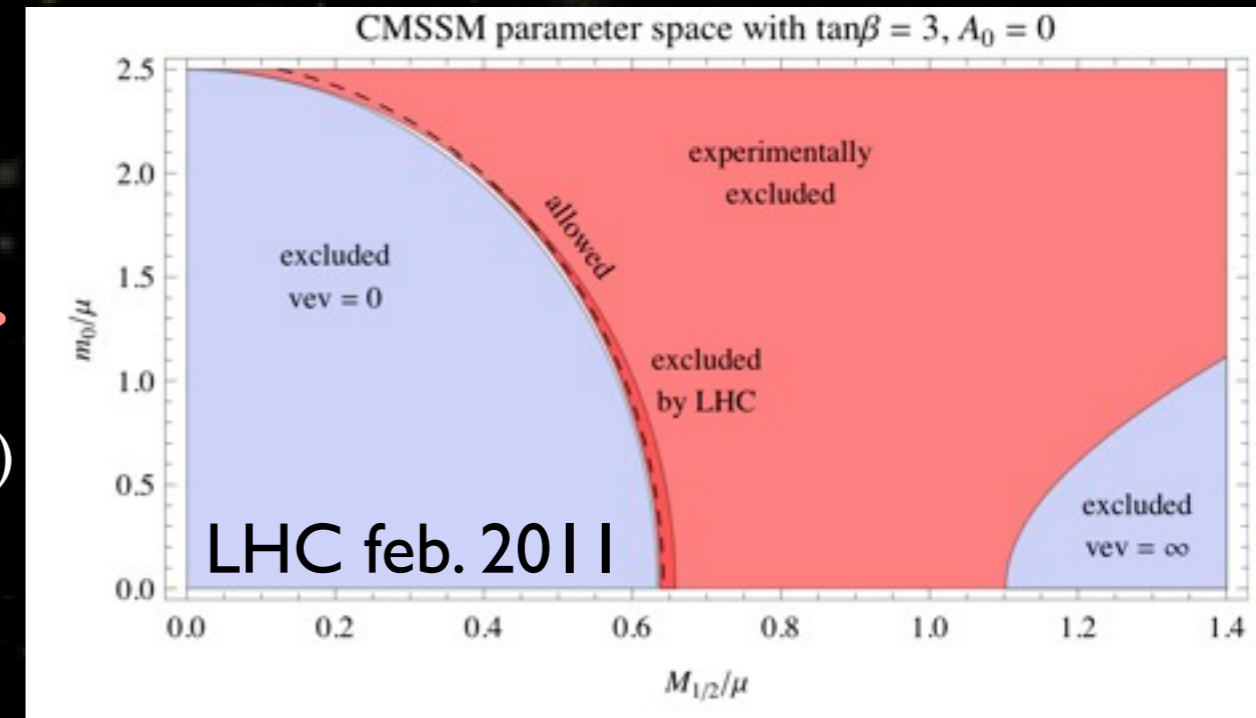
- Particle Physics

- ★ Supersymmetry

- minimal models seem disfavored by LHC
- Resists direct search (Edelweiss, CDMS, Xenon)
  - BUT: DAMA, COGENT, CRESST claim for detection

- ★ Extra dimensions, Axions

- Modifications of Gravity



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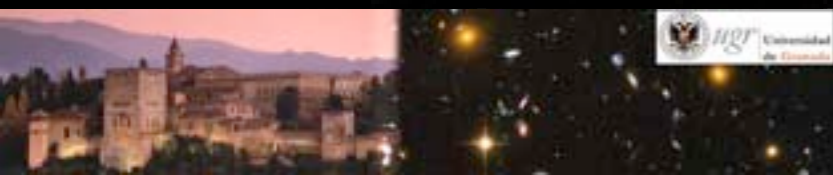
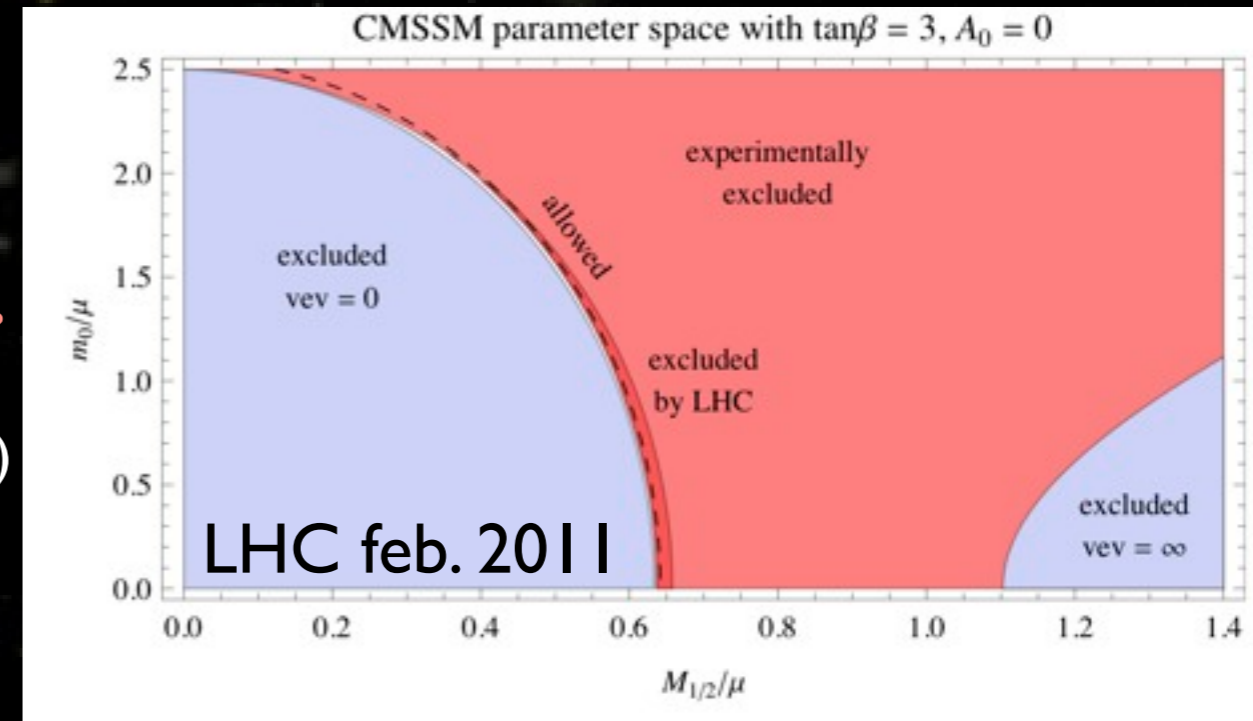
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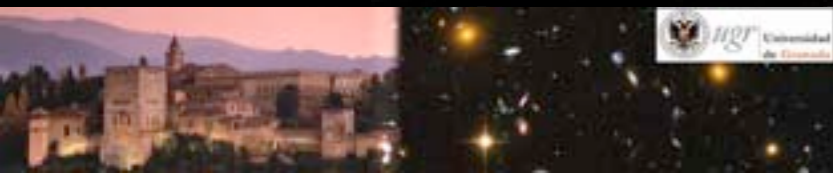
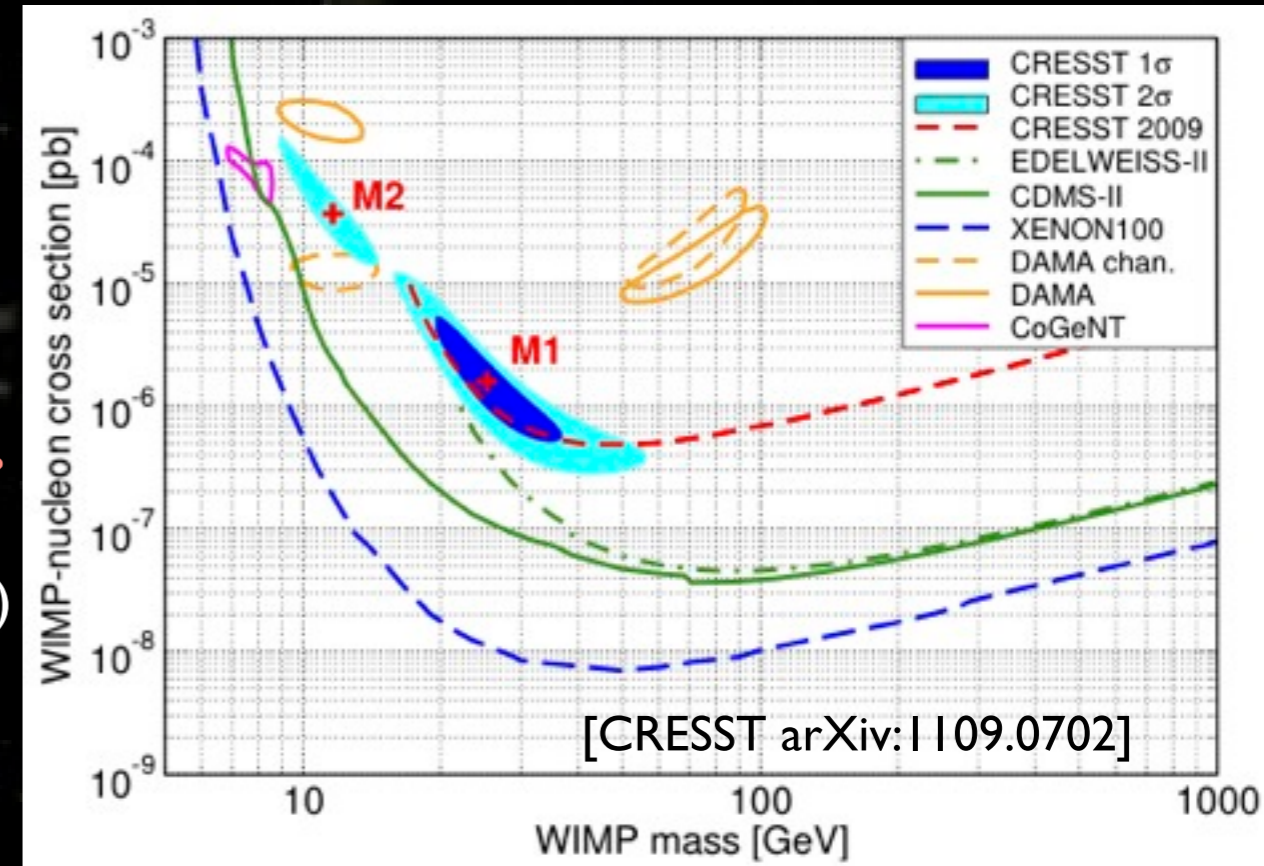
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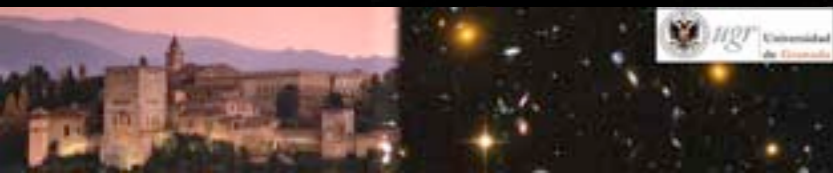
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  - ★ Extra dimensions, Axions
- Modifications of Gravity
  - ★ D.M. is only seen through gravitational effects
  - ★ MOND/TeVS





# MOND

- new fundamental principle of dynamics:  $F = \begin{cases} m \frac{a^2}{a_0} & \text{si } a < a_0 \\ ma & \text{si } a > a_0 \end{cases}$
- Newton force:  $F = \frac{GMm}{r^2}$
- Circular orbit:  $a = \frac{v^2}{r}$

## MOND

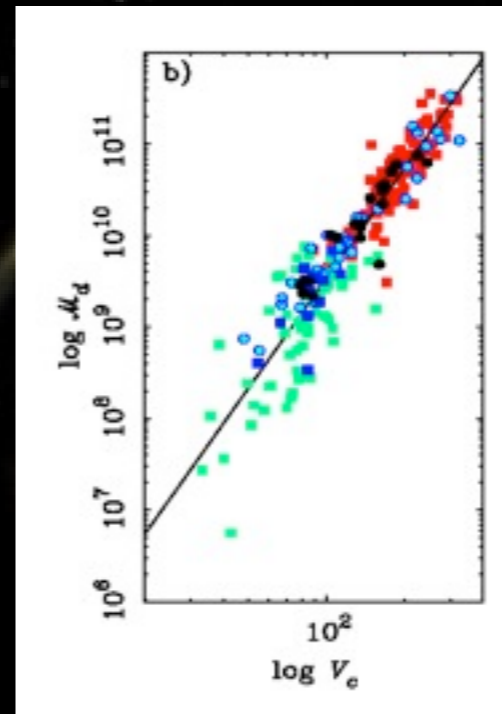
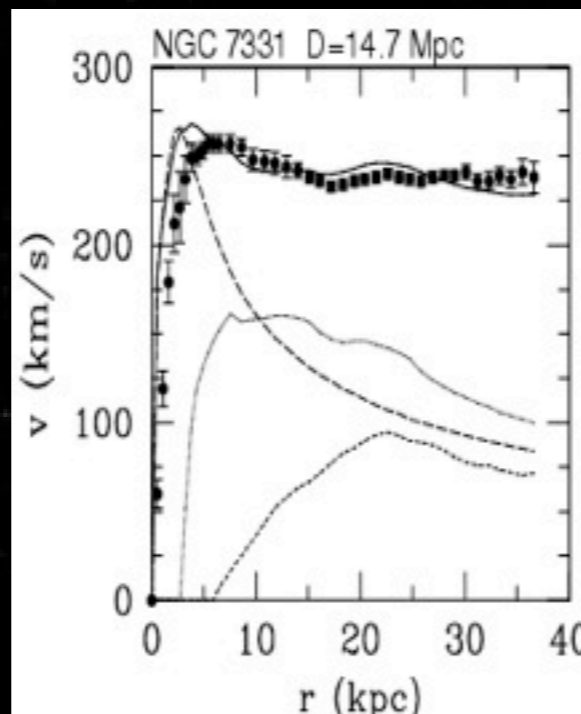
$$\Rightarrow \frac{GM}{r^2} = \frac{a^2}{a_0}$$

$$\Rightarrow a = \sqrt{GMa_0r}$$

$$\Rightarrow \frac{v^2}{r} = \sqrt{GMa_0r}$$

$$\Rightarrow v = (GMa_0)^{1/4}$$

$$\Rightarrow L \propto M \propto v^4$$



→ Flat rotation curves ← ~~X~~

→ Tully-Fisher relation ← ~~X~~

## Newton Dynamics:

$$\Rightarrow \frac{GM}{r^2} = a$$

$$\Rightarrow \frac{v^2}{r} = \sqrt{GMr^2}$$

$$\Rightarrow v = \sqrt{\frac{GM}{r}}$$

$$\Rightarrow L \propto M \propto rv^2$$

Impressive but MOND fails on everything else...



# Candidates for Dark Matter

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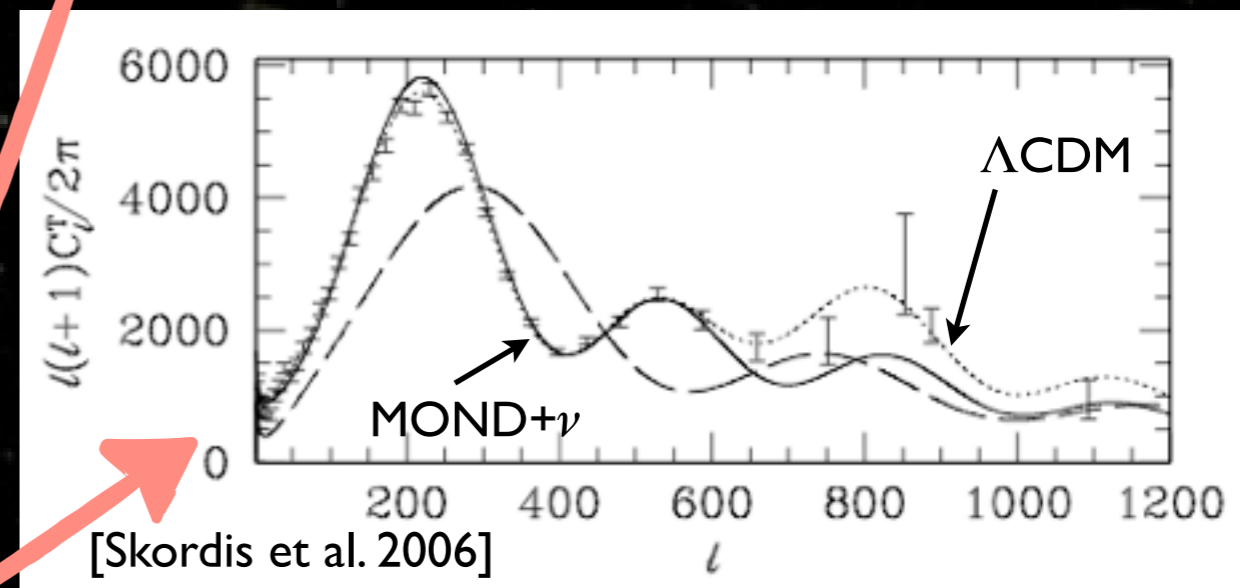
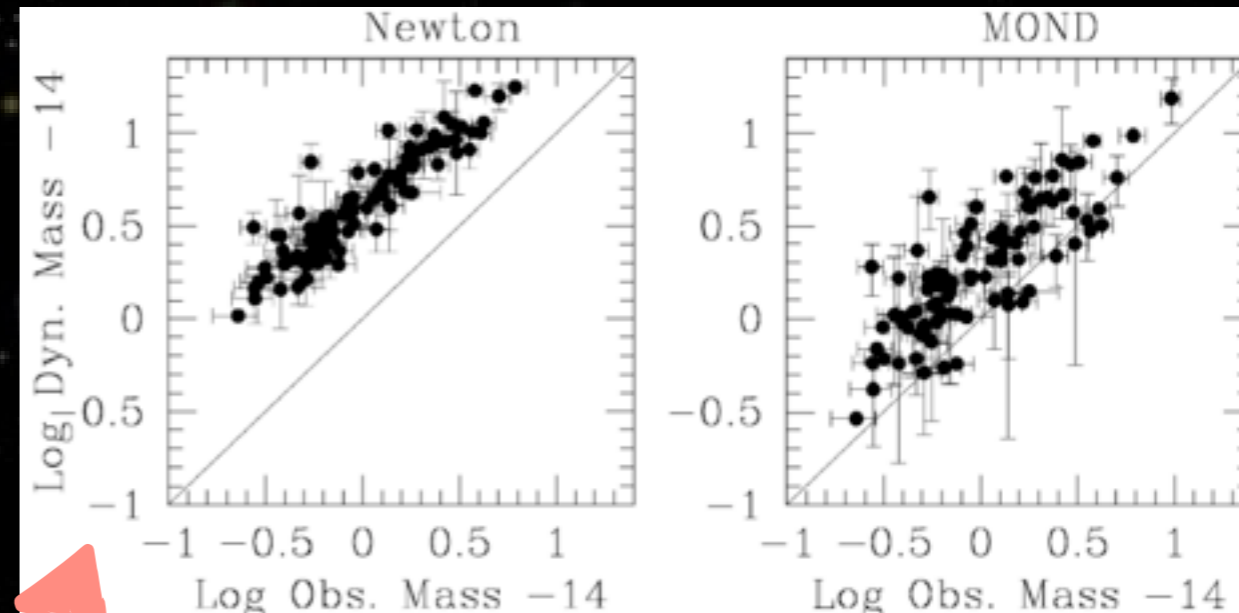
- ★ Extra dimensions, Axions

- Modifications of Gravity

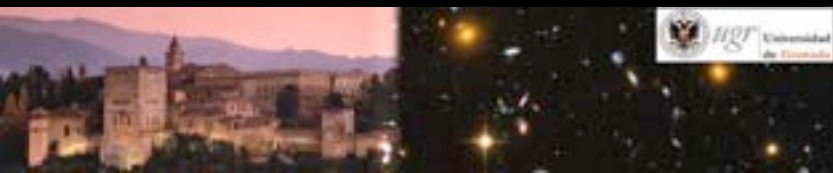
- ★ D.M. is only seen through gravitational effects

- ★ MOND/TeVS: appealing but ...

- still requires a lot of neutrinos to explain galaxy rotation curves and clusters and CMB
- CMB hard to fit now, obvious disagreement with BAO observed in distribution of galaxies
- Does it really seem more sensible than D.M.?

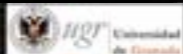
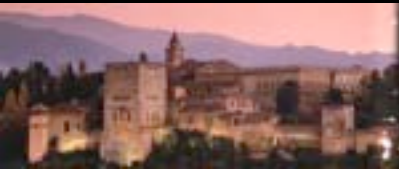


$$F = \begin{cases} m \frac{a^2}{a_0} & \text{si } a < a_0 \\ ma & \text{si } a > a_0 \end{cases}$$



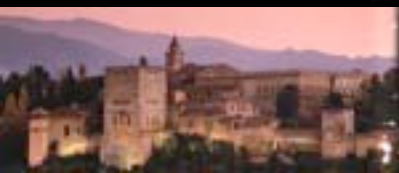
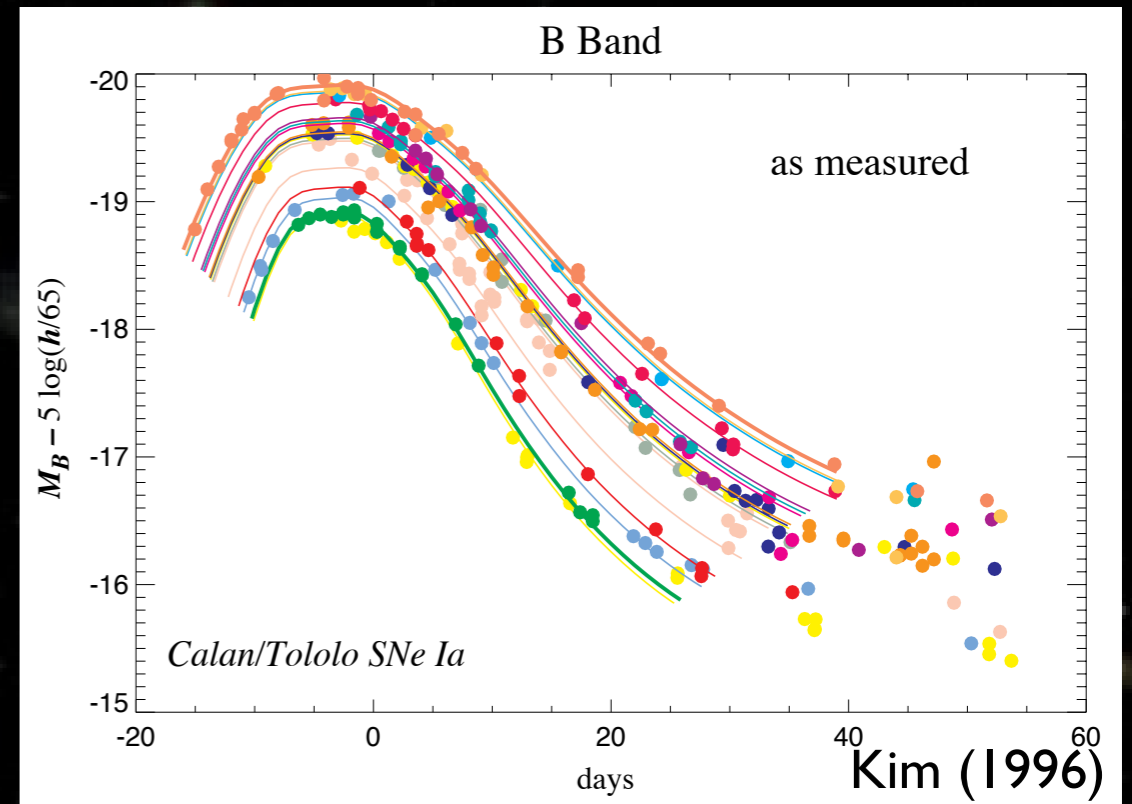


# Dark Energy



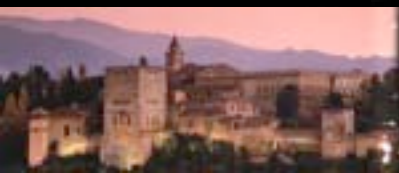
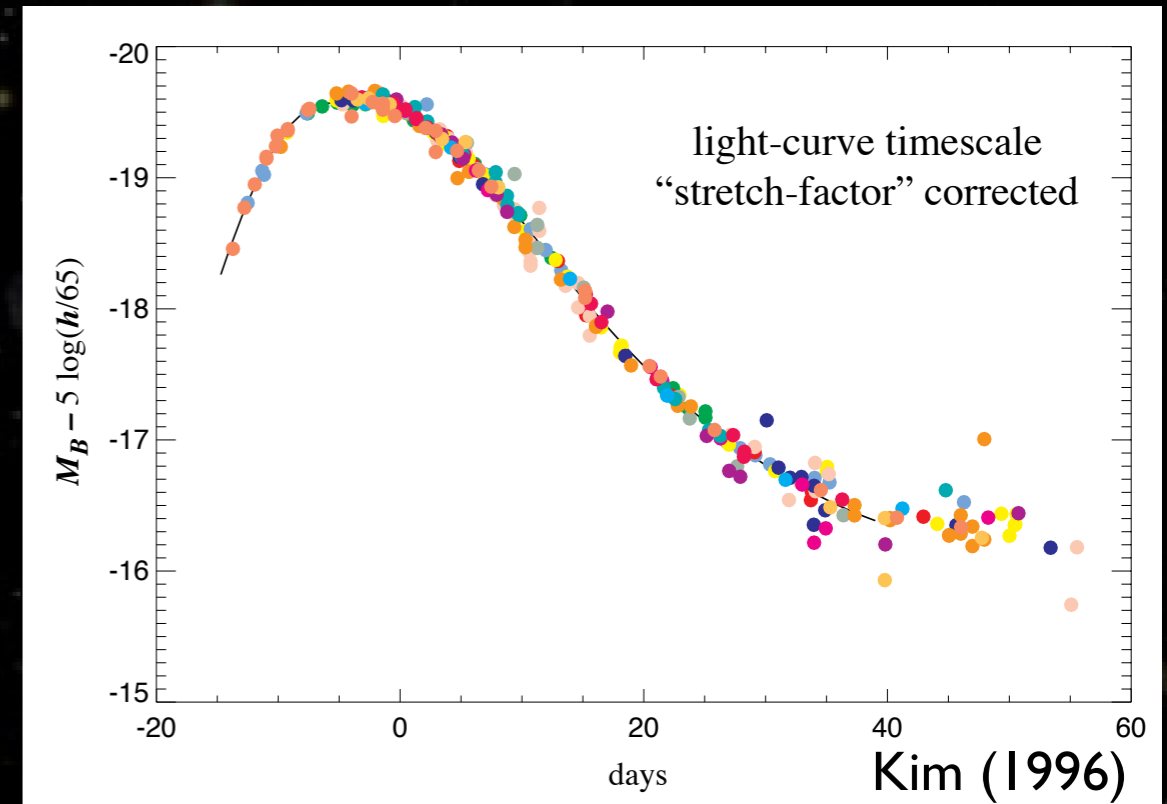
# SN Ia: accelerated expansion (1998)

- Standard candles
  - ★ bright (more than host galaxy)
  - ★ standardizable (within  $\sim 0.15$  mag)
- Luminosity distance Vs.  $z$ 
  - ★ SN Ia fainter than expected
  - ★ Possible systematics well controlled



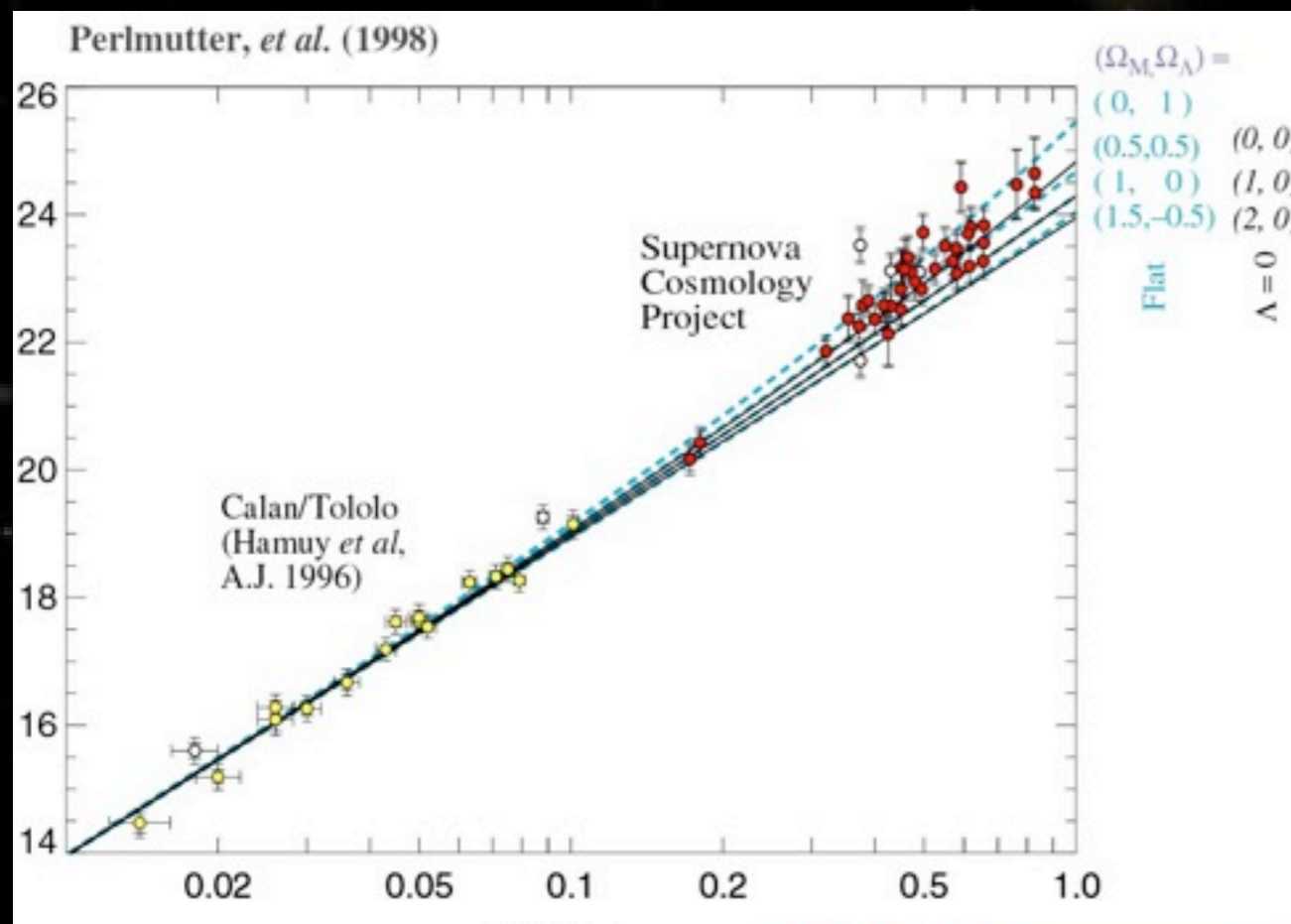
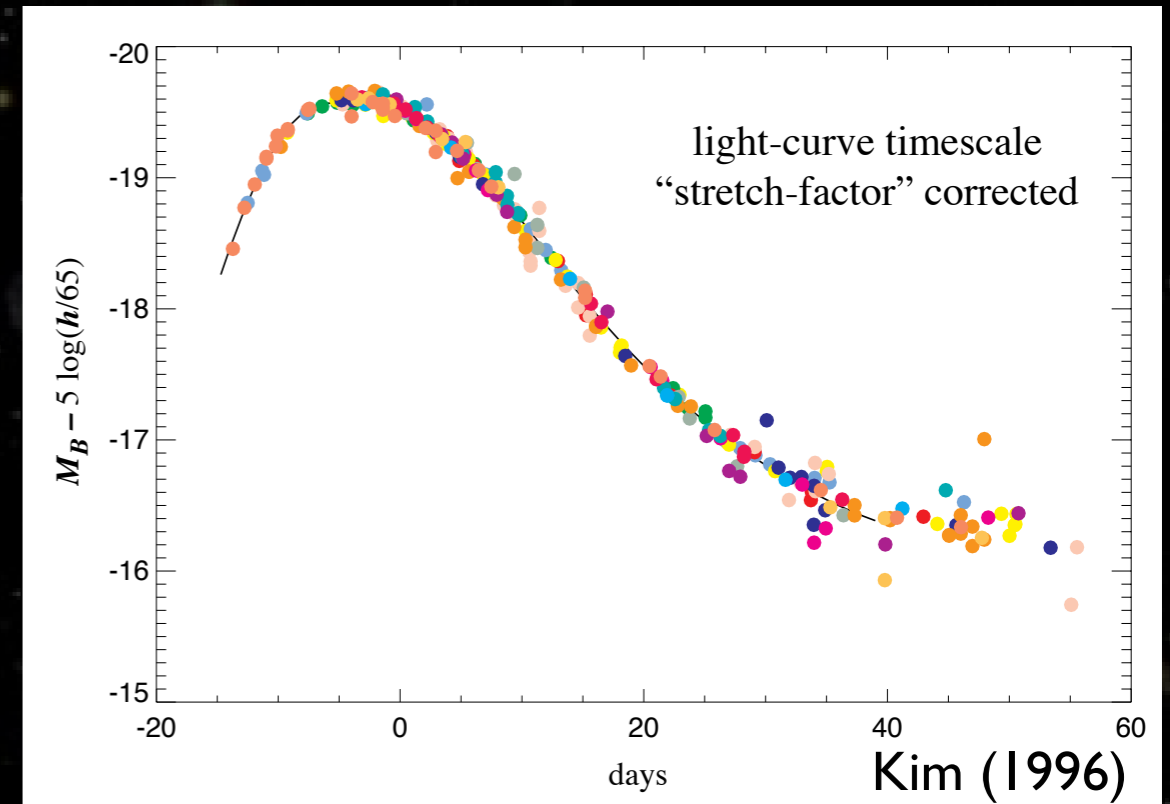
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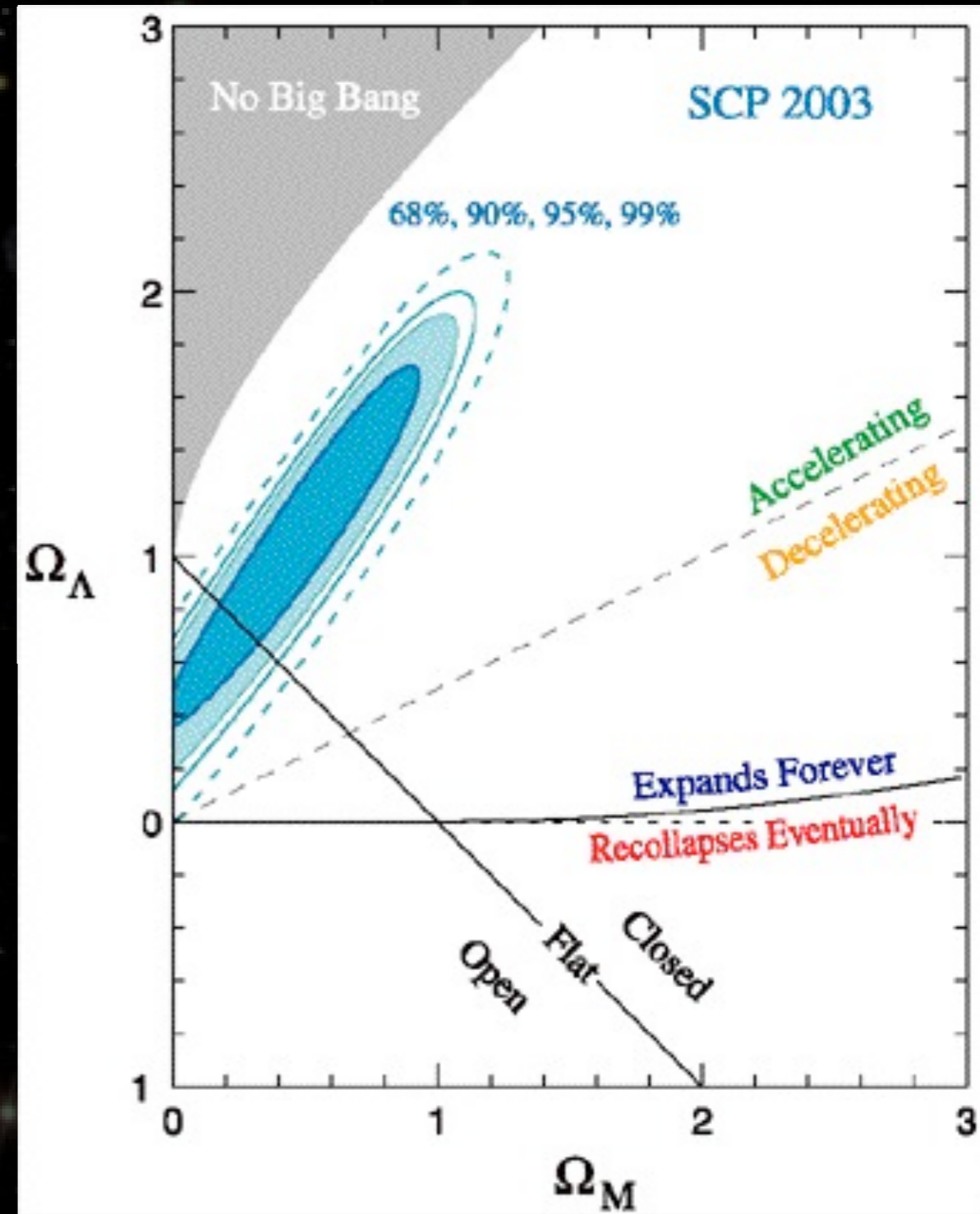
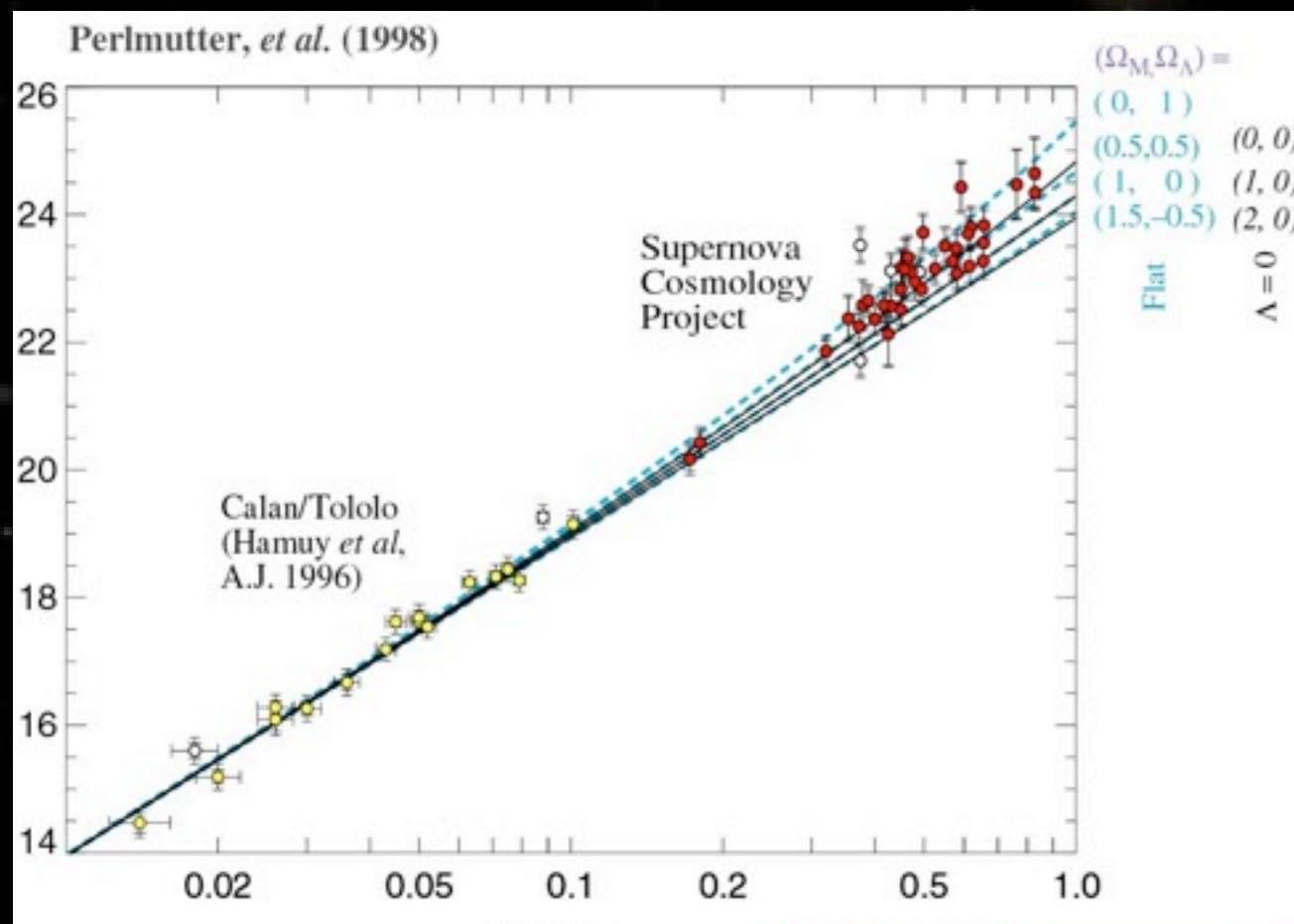
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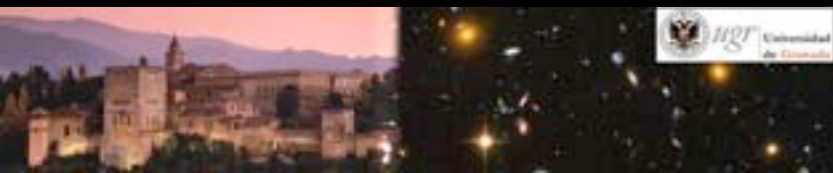
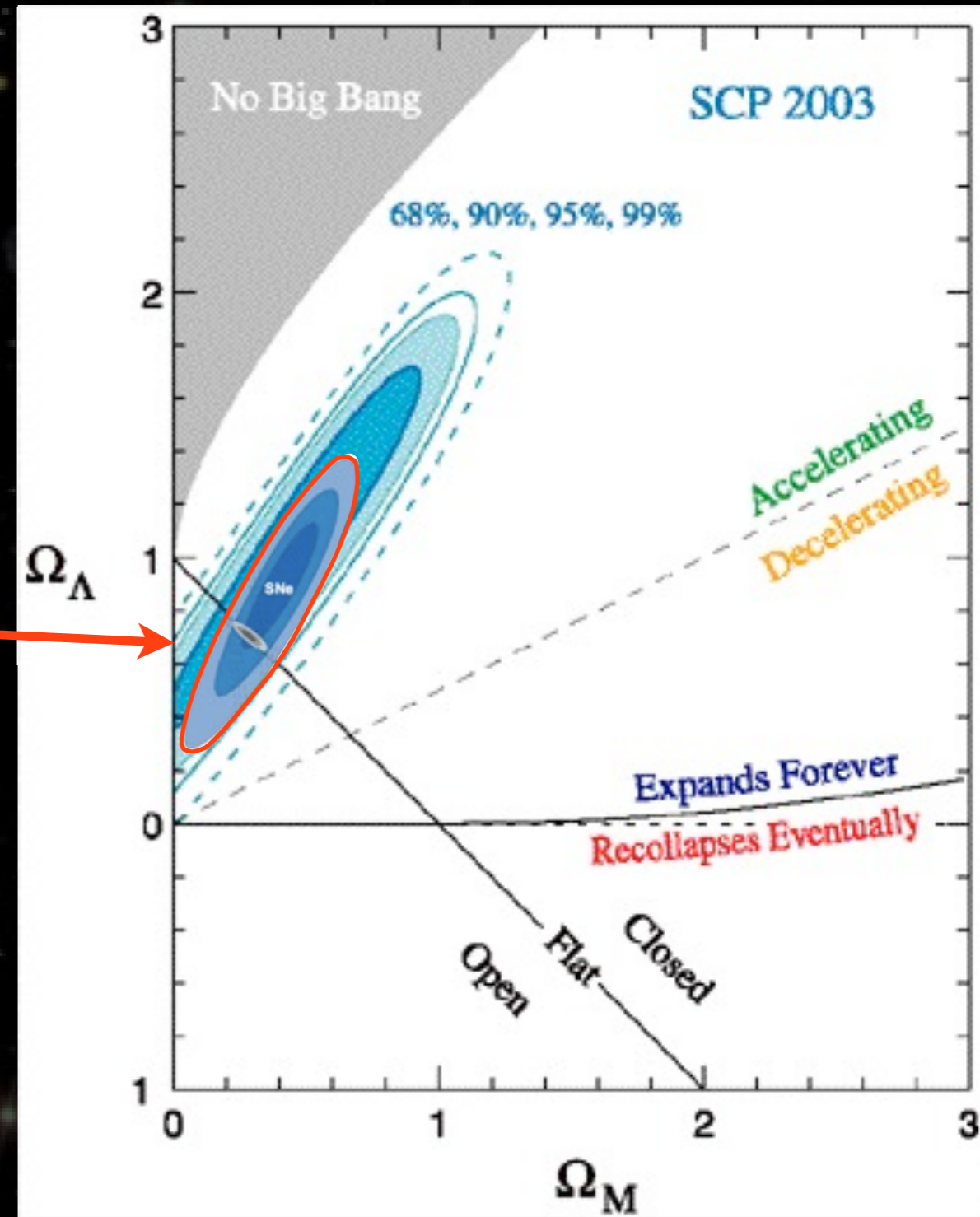
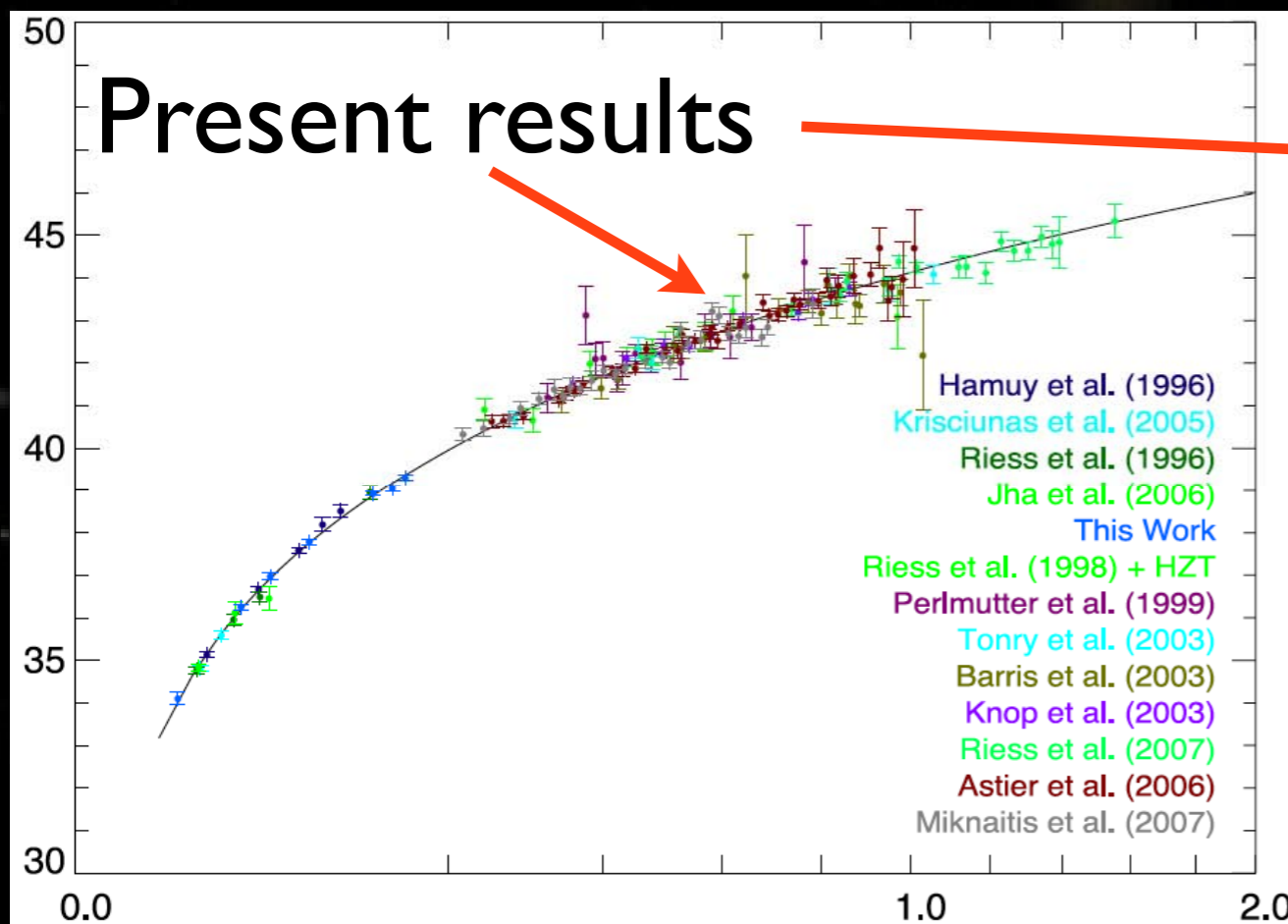
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# Actually, the idea was around...

## Letters to Nature

*Nature* **348**, 705-707 (27 December 1990) | doi:10.1038/348705a0; Accepted 31 October 1990

### The cosmological constant and cold dark matter

G. Efstathiou, W. J. Sutherland & S. J. Maddox

1. Department of Physics, University of Oxford, Oxford OX1 3RH, UK

**THE cold dark matter (CDM) model<sup>1-4</sup> for the formation and distribution of galaxies in a universe with exactly the critical density is theoretically appealing and has proved to be durable, but recent work<sup>5-8</sup> suggests that there is more cosmological structure on very large scales ( $l > 10 h^{-1} \text{ Mpc}$ , where  $h$  is the Hubble constant  $H_0$  in units of  $100 \text{ km s}^{-1} \text{ Mpc}^{-1}$ ) than simple versions of the CDM theory predict. We argue here that the successes of the CDM theory can be retained and the new observations accommodated in a spatially flat cosmology in which as much as 80% of the critical density is provided by a positive cosmological constant, which is dynamically equivalent to endowing the vacuum with a non-zero energy density. In such a universe, expansion was dominated by CDM until a recent epoch, but is now governed by the cosmological constant. As well as explaining large-scale structure, a cosmological constant can account for the lack of fluctuations in the microwave background and the large number of certain kinds of object found at high redshift.**

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#### SEARCH PUBMED FOR

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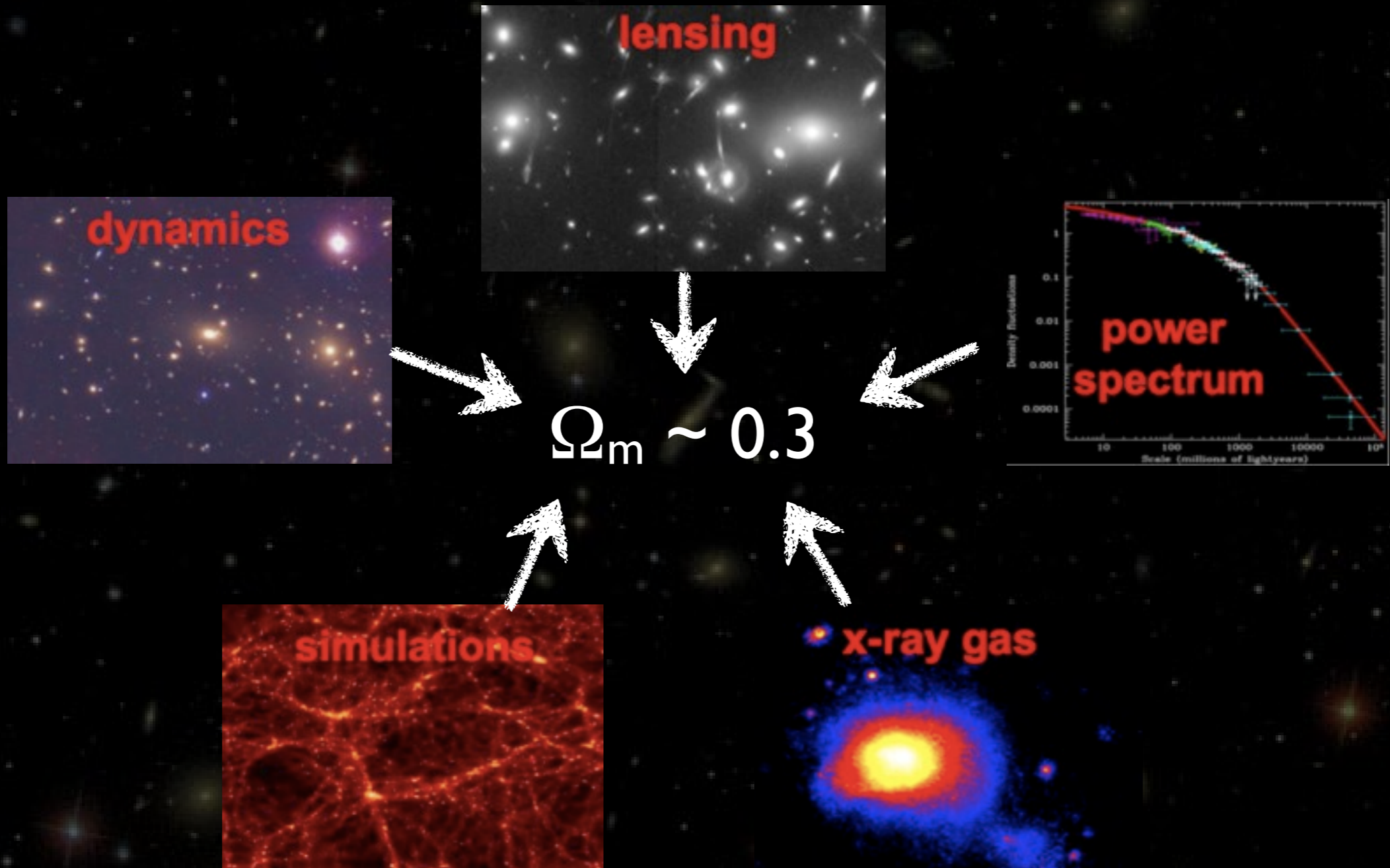
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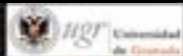
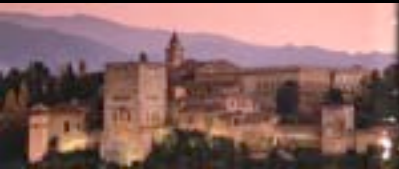
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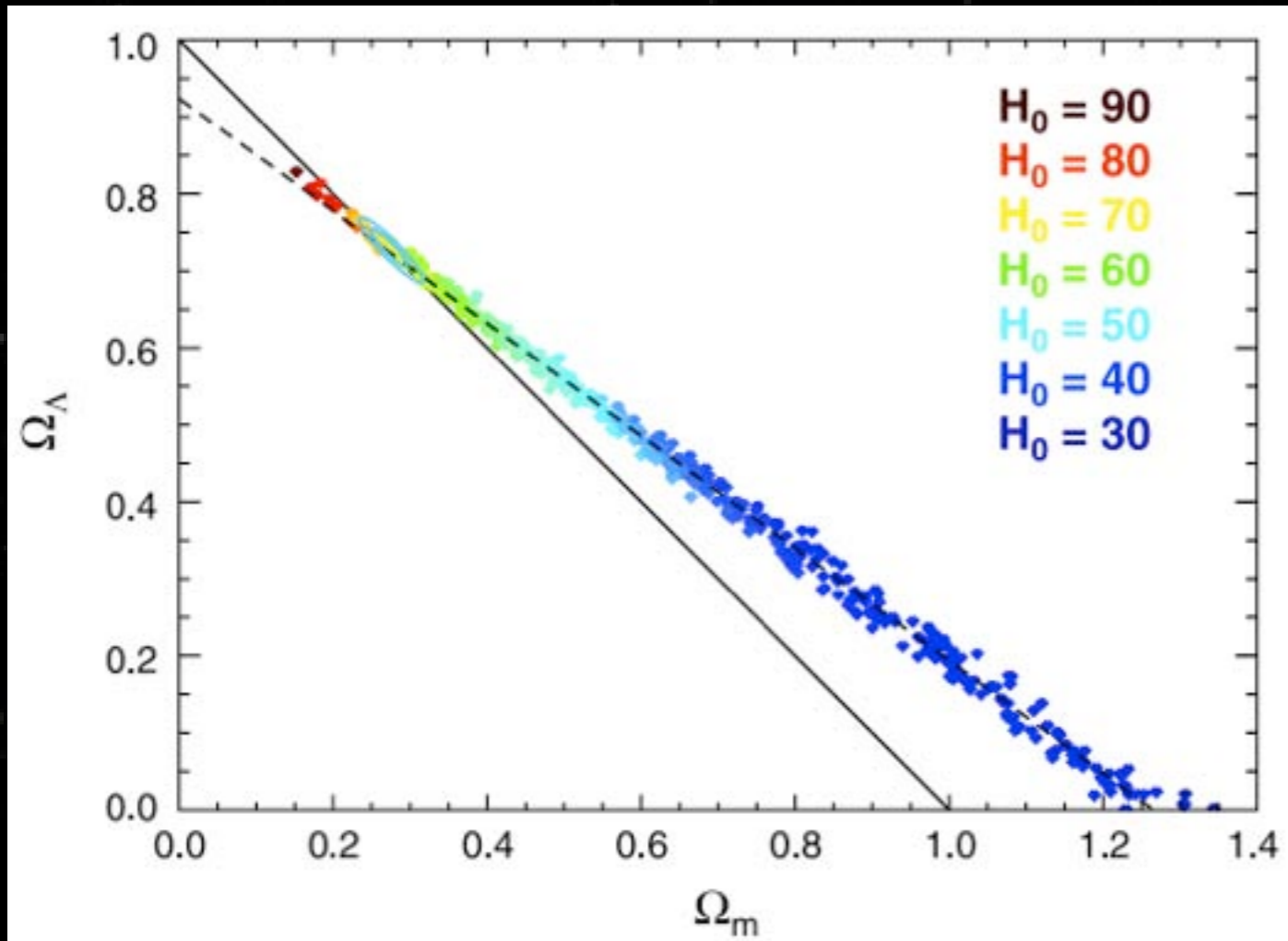
# Cosmic subtraction: $1 - 0.3 = 0.7$



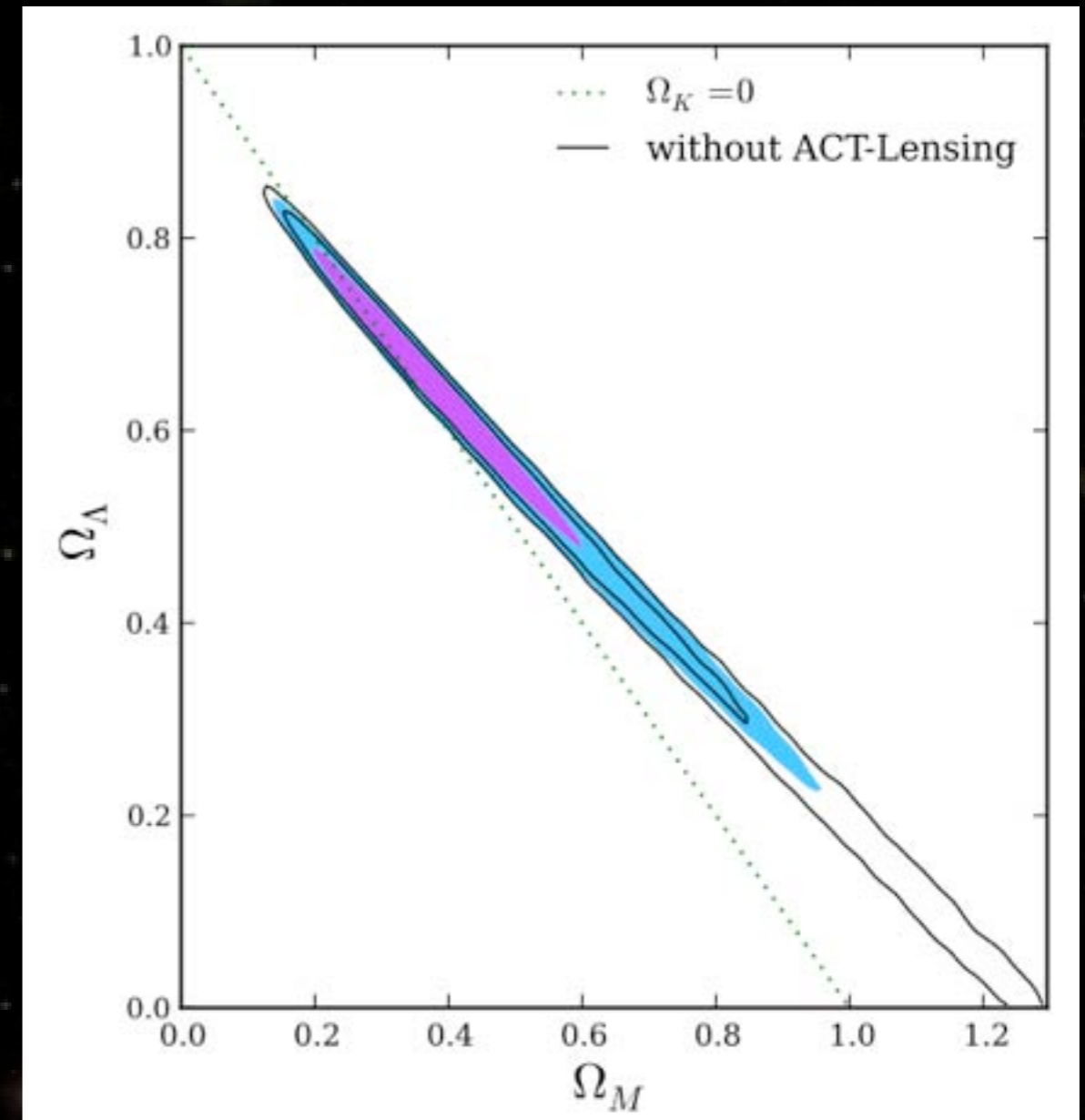
From R. Kolb



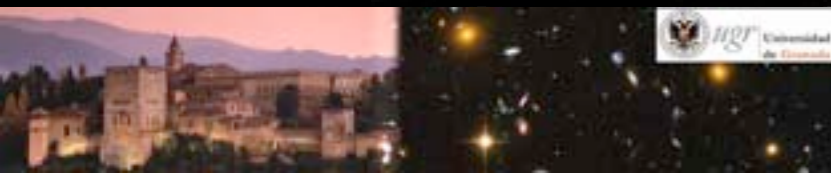
# evidence for D.E. from CMB alone



WMAP 7y



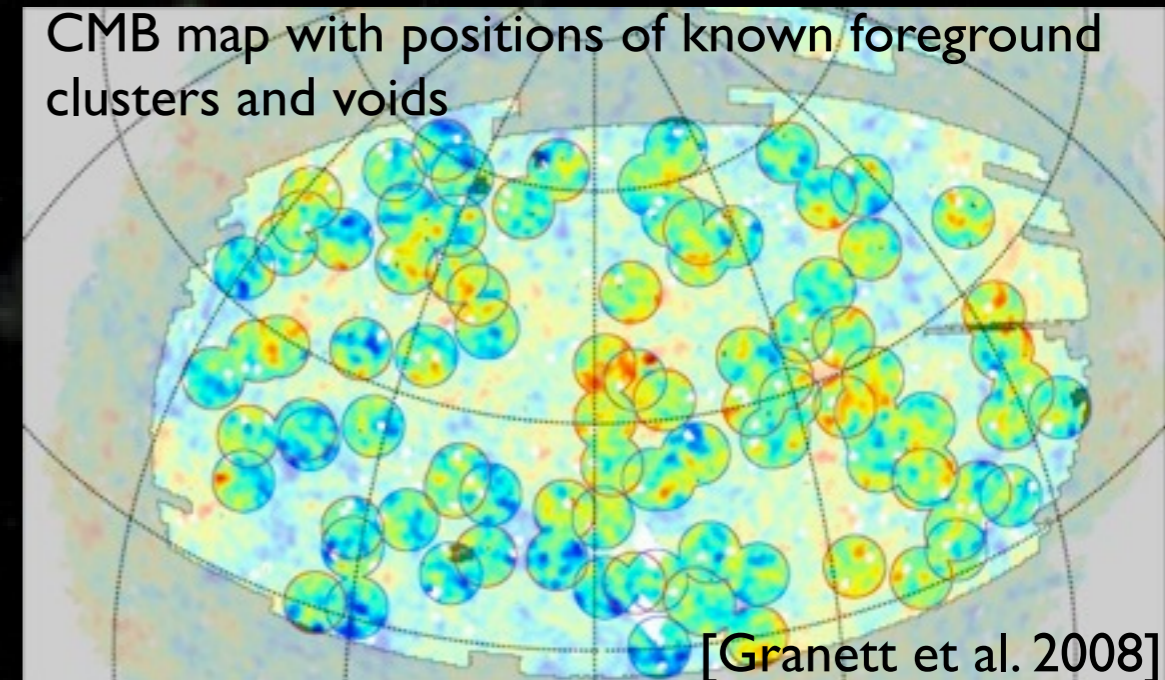
WMAP 7y + ACT (no  $H_0$ )  
[Sherwin et al. 2011]



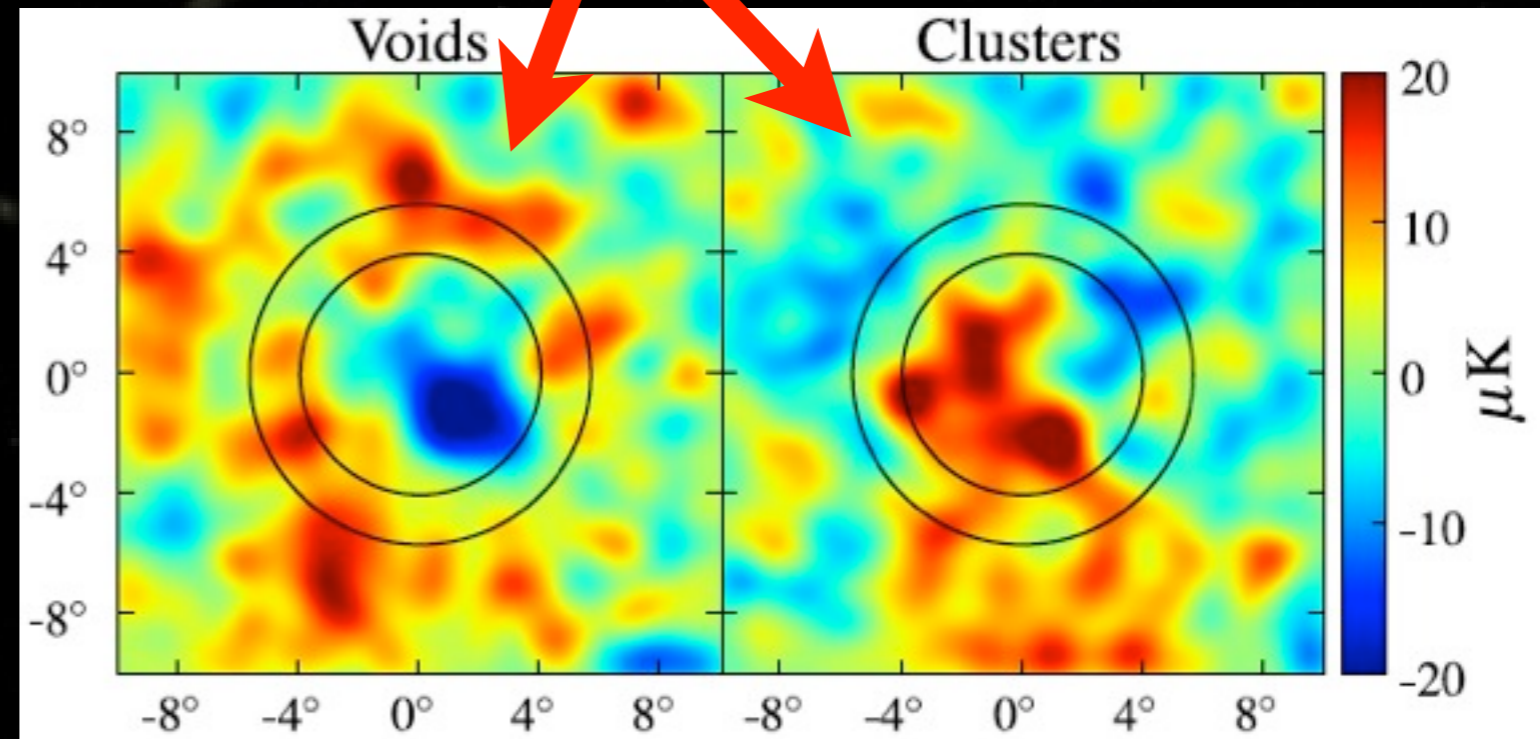
# ISW effect: independent probe

- Integrated Sachs-Wolfe effect

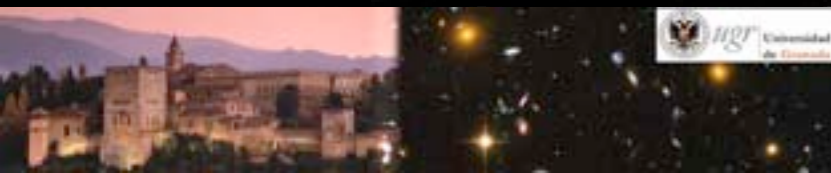
- ★ Photons gain energy falling into clusters potential wells
- ★ They lose energy escaping
- ★ In the presence of Dark Energy the net result is not zero
- ➔ Correlation between CMB hot/cold spot expected with foreground clusters/voids



- ➔  $4\sigma$  level detection by various teams
- ➔ N.B. expected effect is zero if no D.E. (or more generally if no acceleration)

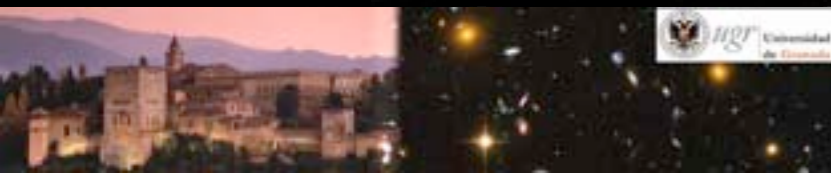
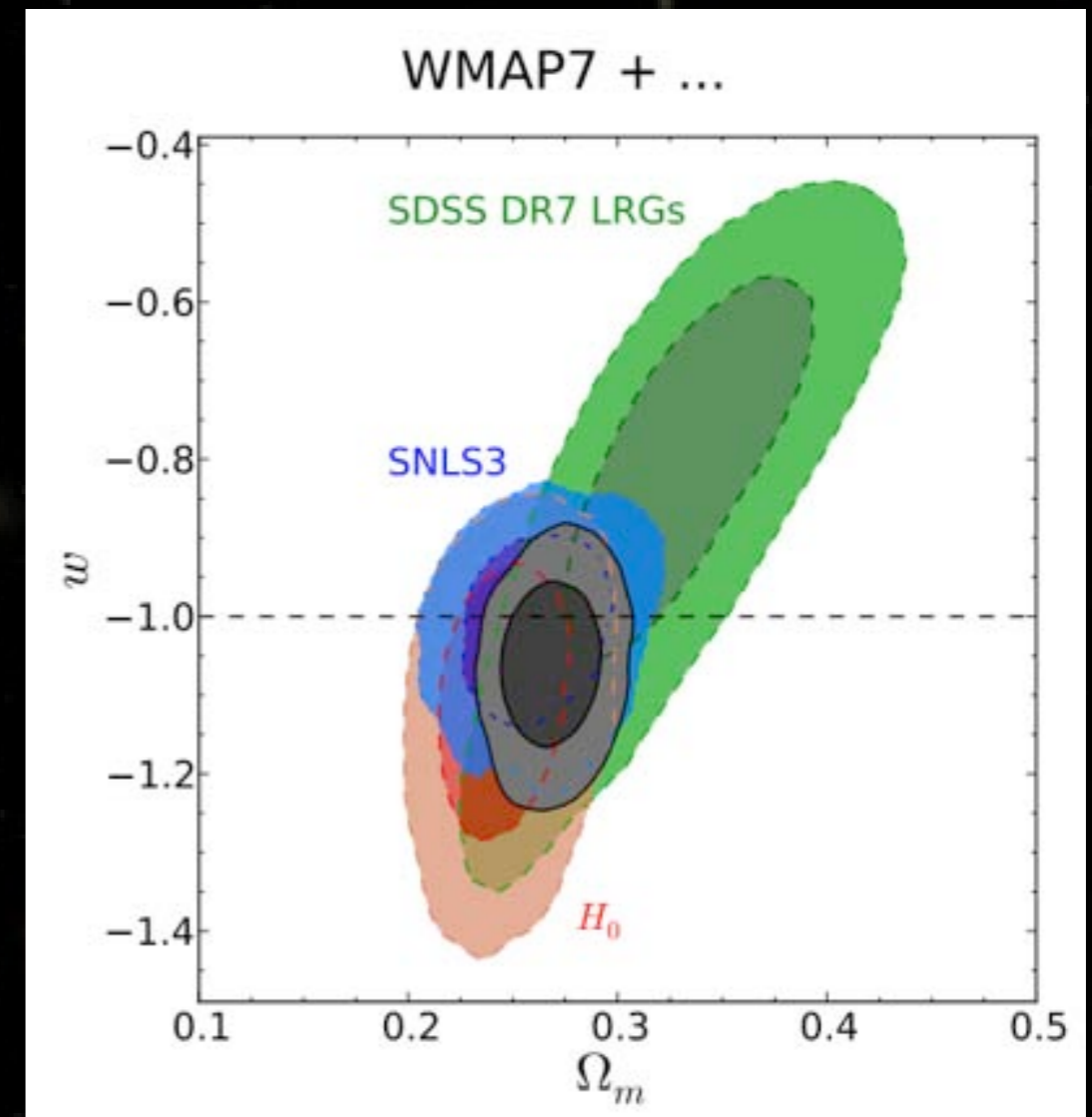
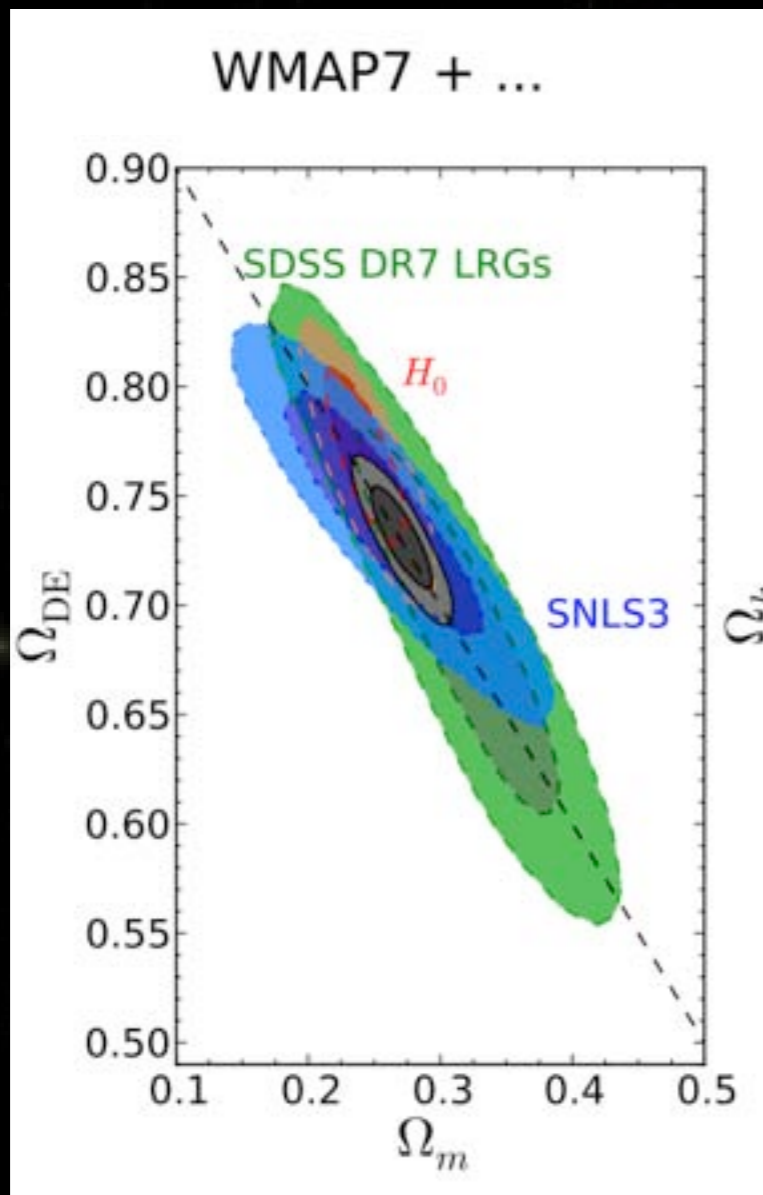


[Granett et al. 2008]



# Finally, with latest data

- Dark Energy Equation of state:  $p = w\rho$ 
  - ★  $w = -1 \Leftrightarrow$  pure cosmological constant (or scalar field)



# Bug or reality ?

- So the Universe really seems to be accelerating...
  - ★ «we are in a void» explanation excluded by lack of kSZ effect [Zhang and Stebbins 2011]

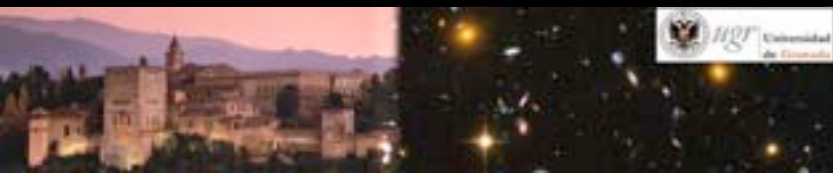
- $\Omega_{D.E.} \sim 0.7$  with an equation of state  $w \sim -1$

- ★ It really looks like a cosmological constant
- ★ or maybe a scalar field depending on which side in Einstein's Equation

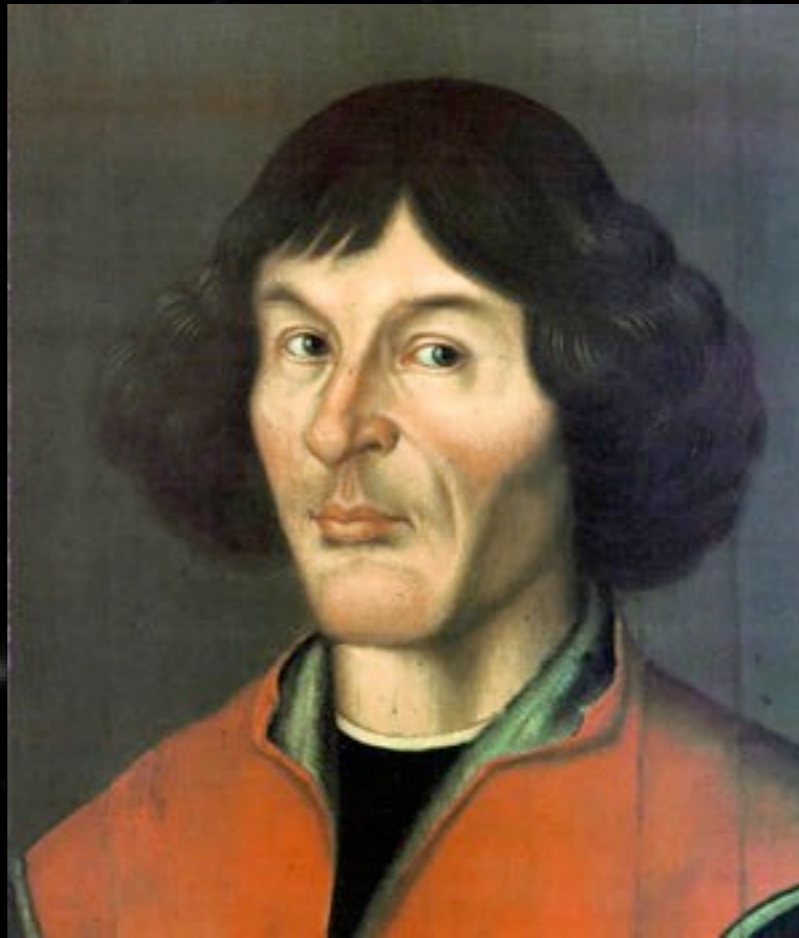
$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R - \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4}T_{\mu\nu} \quad \text{or} \quad R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = \frac{8\pi G}{c^4}T_{\mu\nu} + \Lambda g_{\mu\nu}$$

- Now ... is it really serious ?

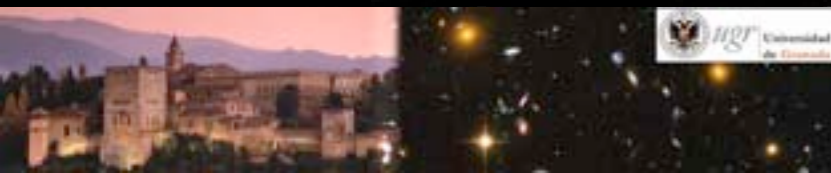
- ★ No scalar field has ever been observed
- ★ Strong disagreement with particle physics vacuum
  - From  $H_0$  and  $\Omega_\Lambda$ :  $\Lambda \sim 10^{-52} \text{ m}^{-2}$
  - From Planck scale physics:  $\Lambda \sim 10^{70} \text{ m}^{-2}$
- ★ Something deep is obviously not understood...



# Tests of the Cosmological principle



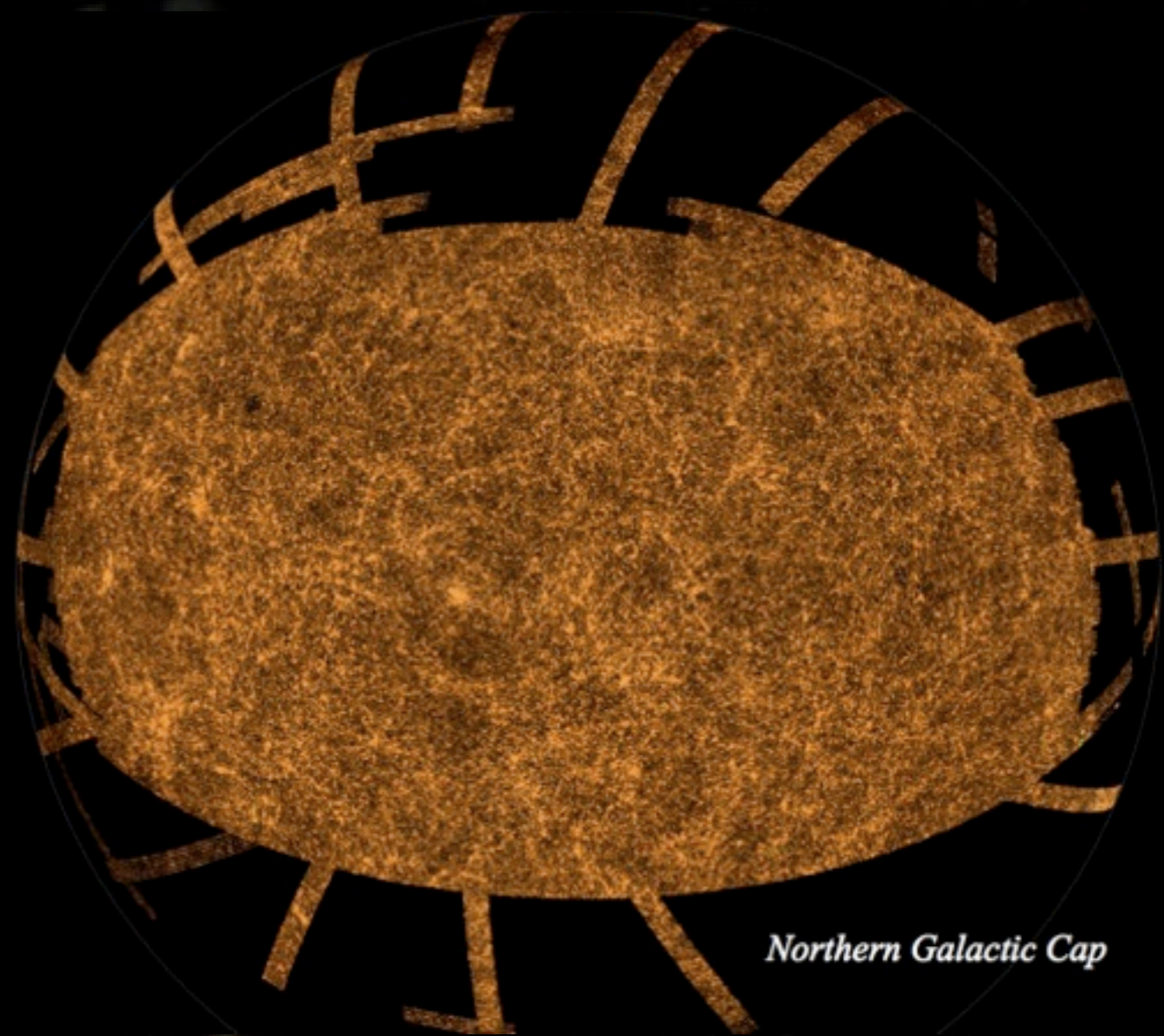
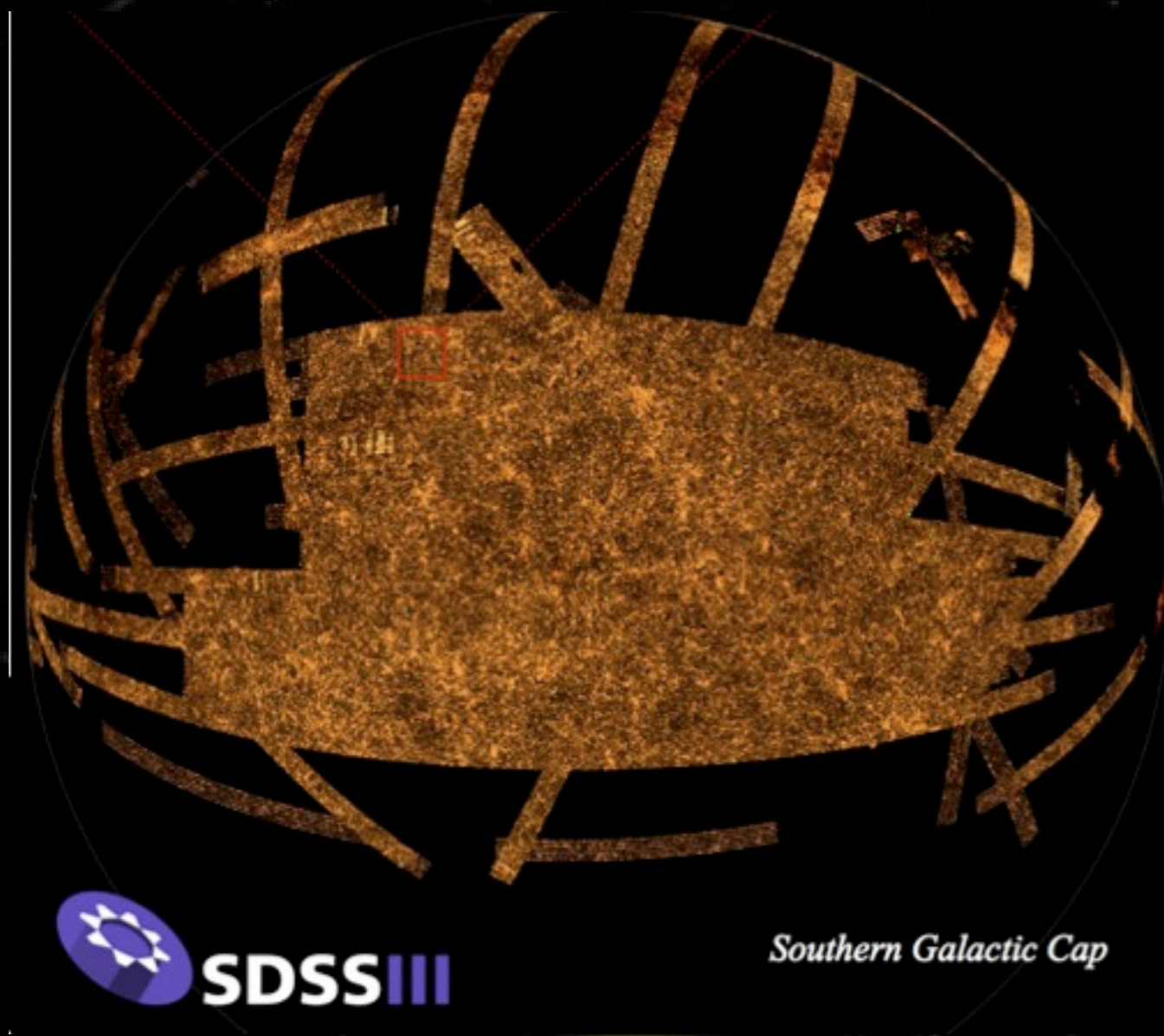
Mikolaj Kopernik



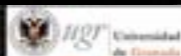
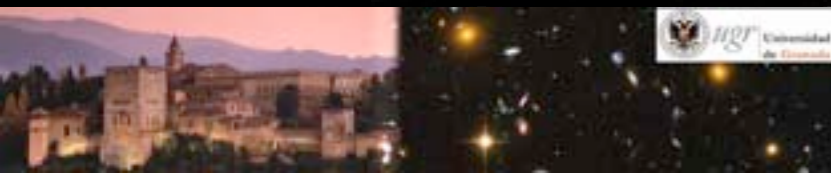


# Cosmological Principle tests

- At first a simplifying hypothesis
- Isotropy tested with increasing success (here  $z \sim 1$ )



470 Millions objects, 1.5 million on the way of being spectred

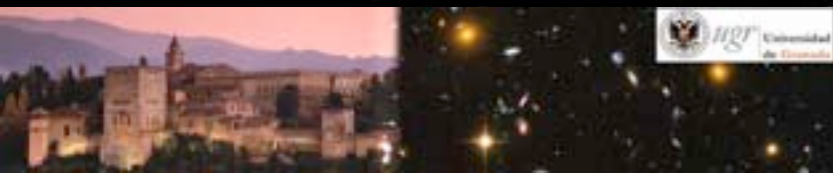


# Cosmological Principle tests

- At first a simplifying hypothesis
- Isotropy tested with increasing success (here  $z \sim 1000$ )

(COBE/DMR homepage)

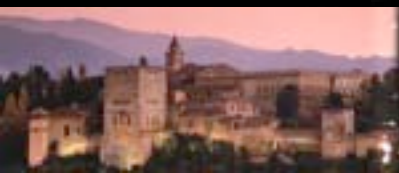
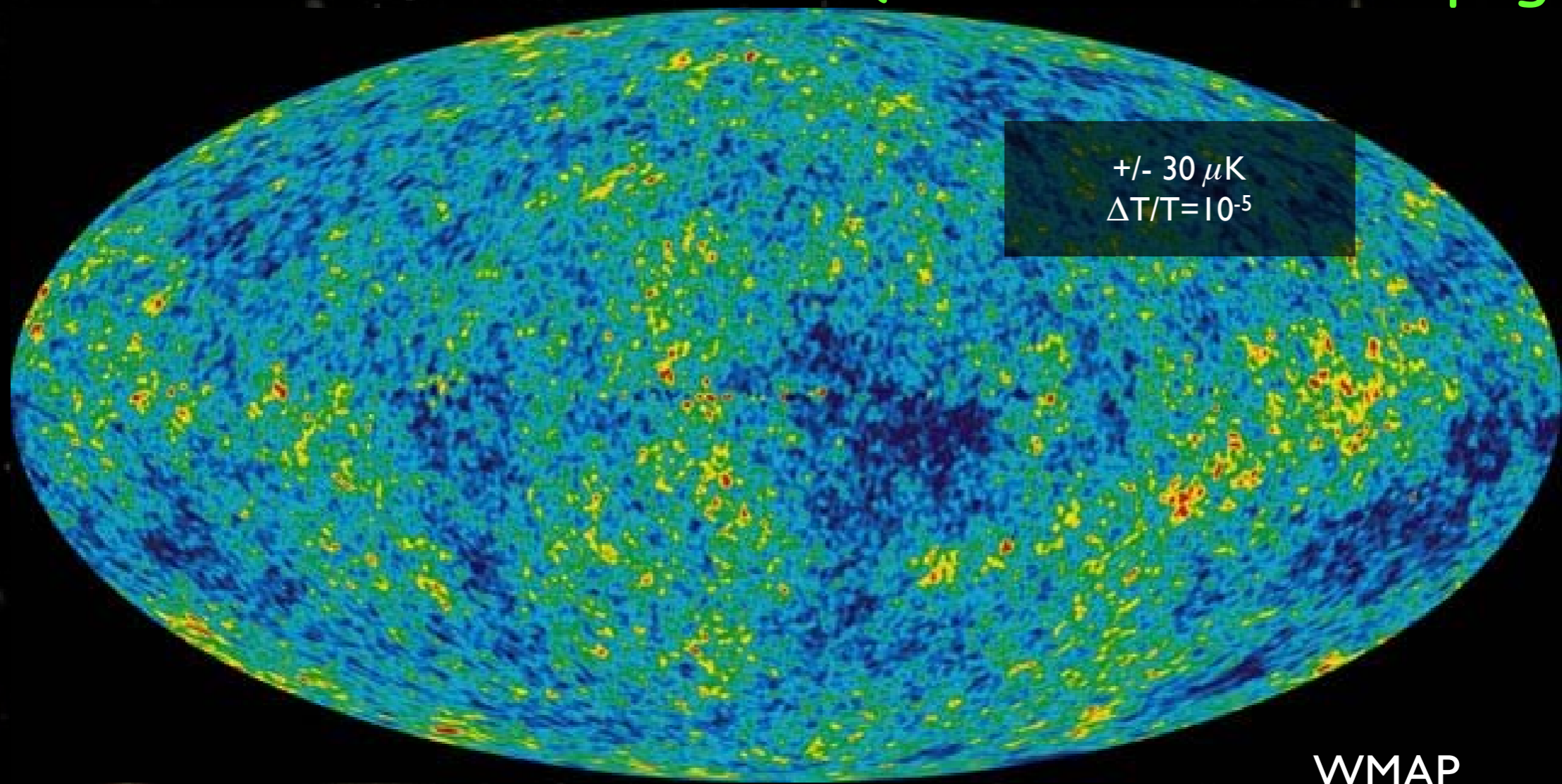
corps noir à 3 K



# Cosmological Principle tests

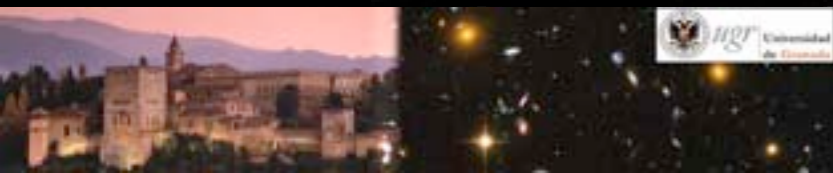
- At first a simplifying hypothesis
- Isotropy tested with increasing success (here  $z \sim 1000$ )

(COBE/DMR homepage)

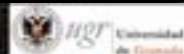
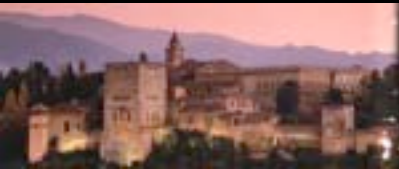
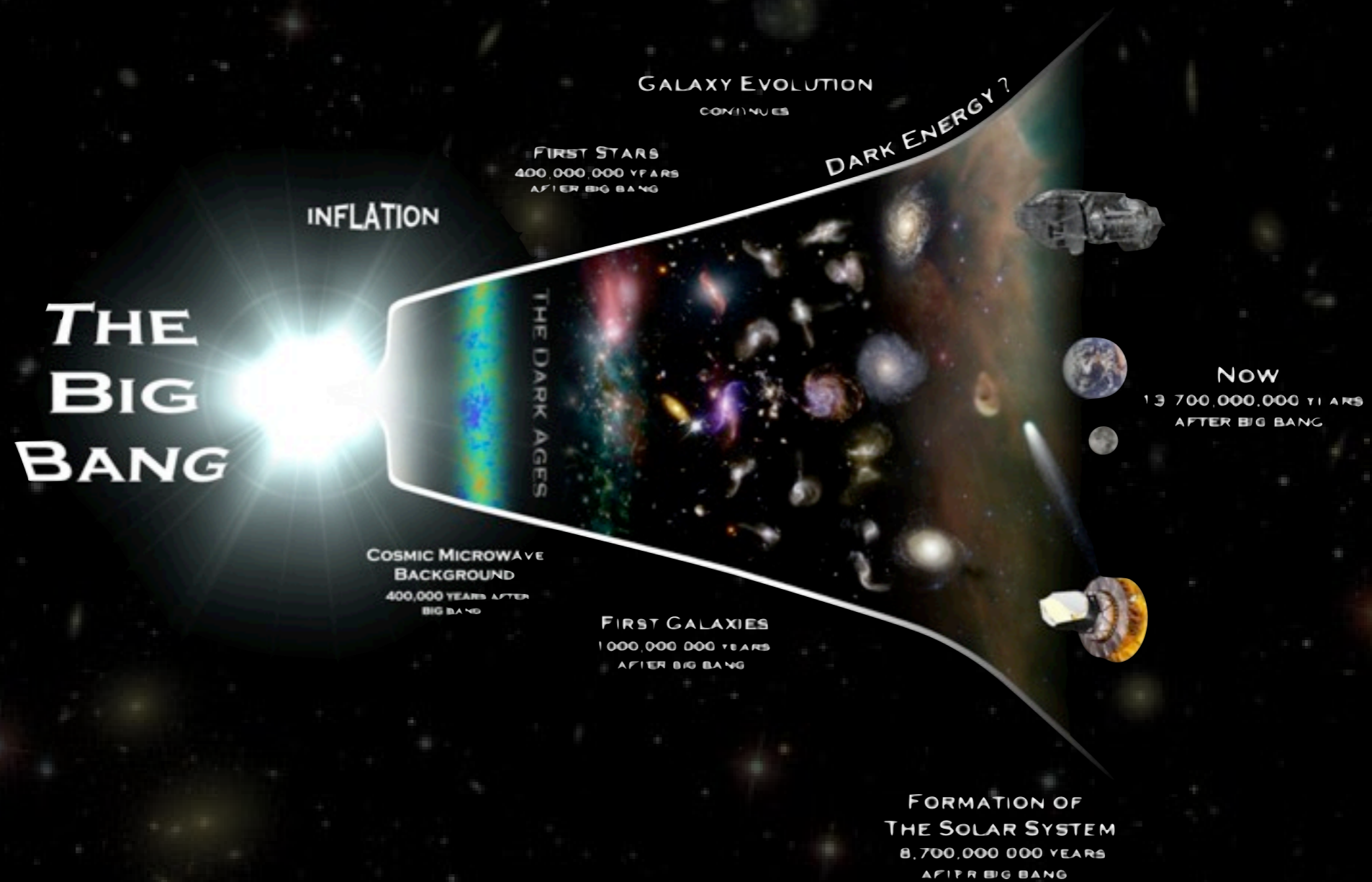


# Cosmological Principle tests

- At first a simplifying hypothesis
- Isotropy tested with increasing success
- If Copernican principle is assumed
  - ★ Every observer would see similar matter/CMB distribution
  - ➔ homogeneity follows [Maartens, 2011]
  - ★ But no strong observational basis for Copernican principle...
- Testing the Copernican Principle
  - ★ Spectral distortions on the CMB
    - Black-Body nature [Caldwell and Stebbins, 2008]
    - lack of kinetic-SZ effect on small scales [Zhang and Stebbins, 2010]
  - ★ future: time drifts of cosmological redshifts [Uzan et al, 2008]
- Basically it seems fine...

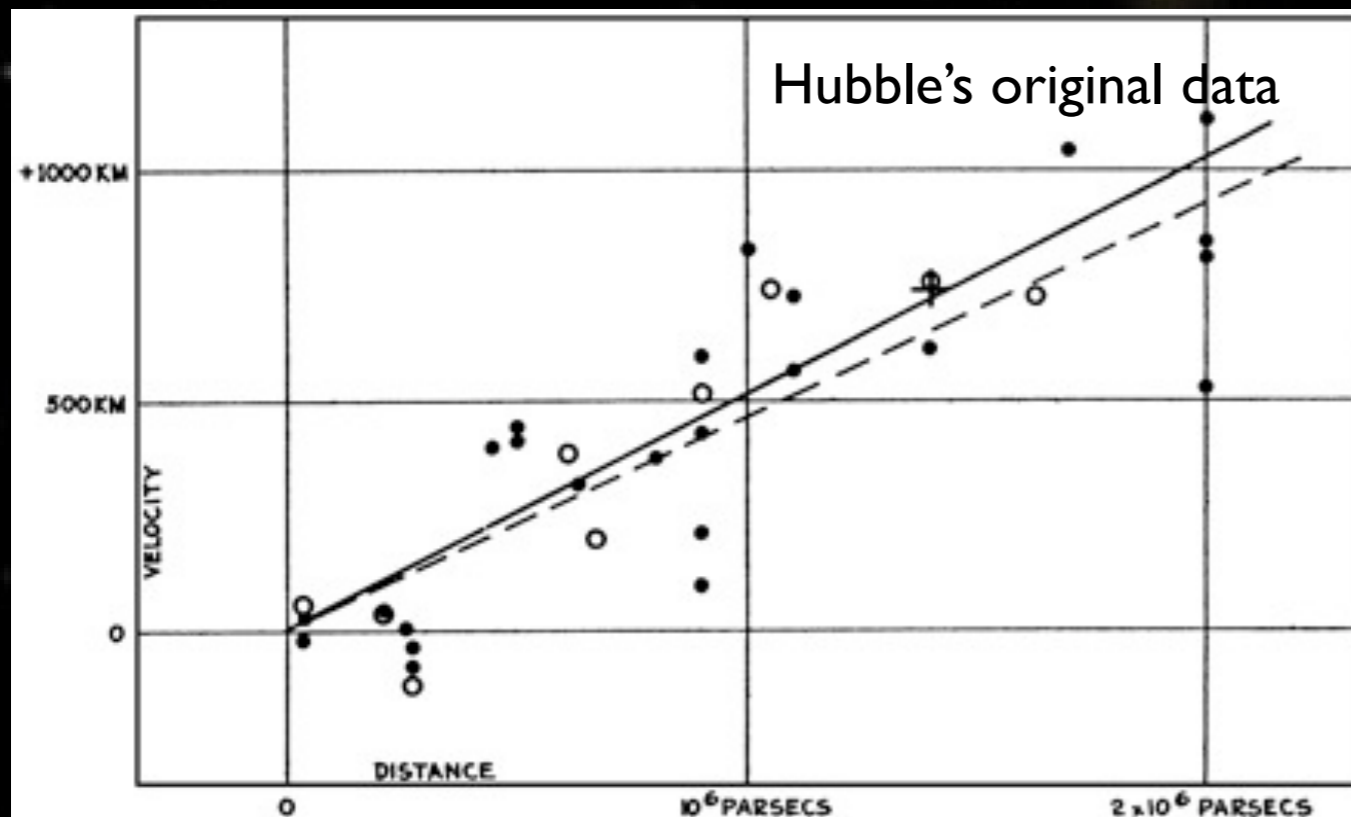


# Tests of the Hot Big-Bang paradigm

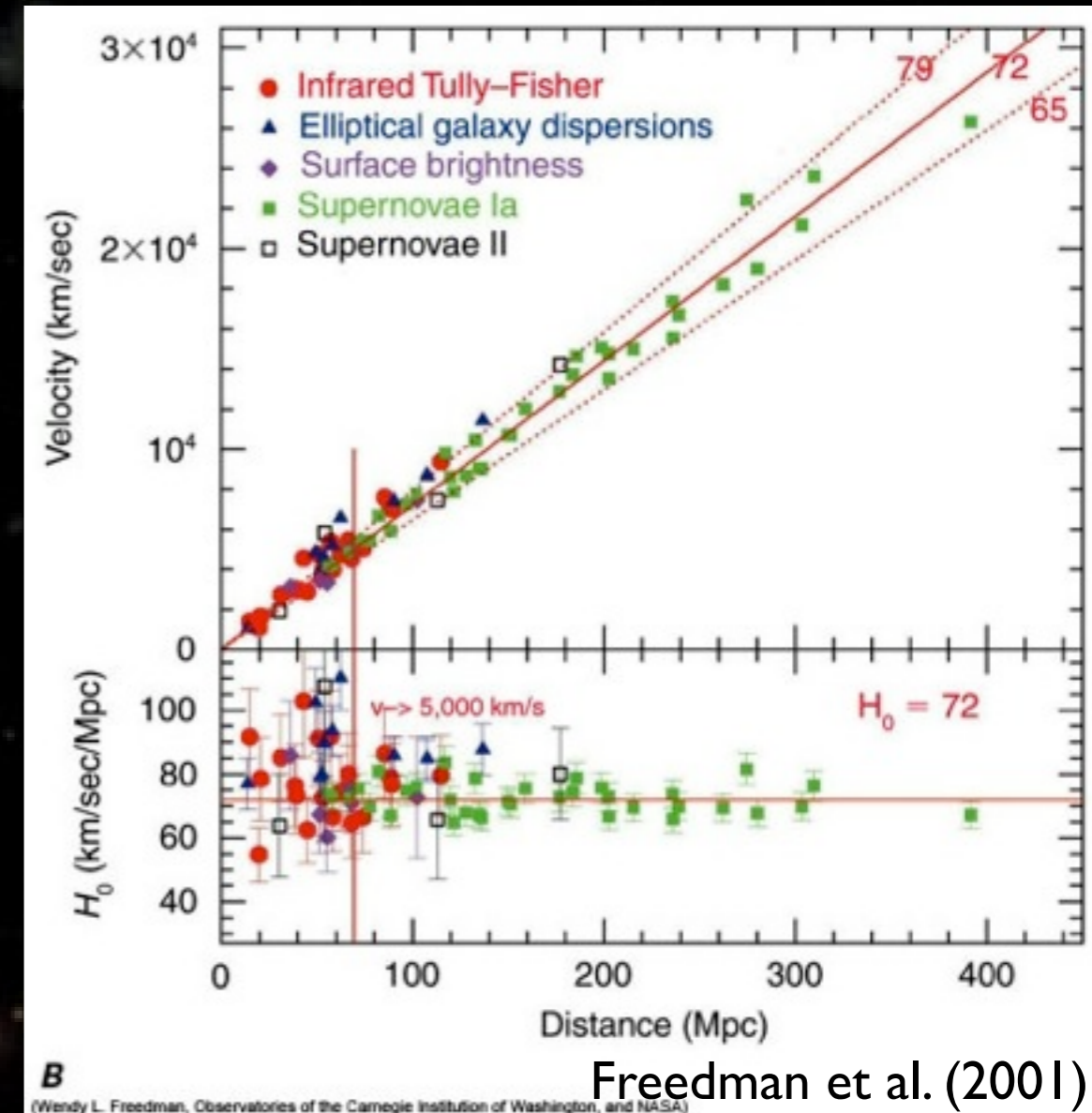


# Hubble's law

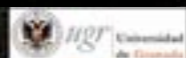
- Long debate on  $H_0$  value
  - Now stabilized:
  - $H_0 = 72 \pm 3(\text{stat}) \pm 7(\text{syst})$
- ★ main systematics:
- Cepheids metallicity
  - Distance to LMC



1927 :  $H_0 \sim 500 \text{ km/s/Mpc}$



2001 :  $H_0 = 72 \text{ km/s/Mpc}$



# Age of stars supports $\Lambda$ CDM

- Oldest stars in Milky Way globular clusters:

- ★ age  $> 11.2$  Gyr (95% C.L.) [Krauss & Chaboyer, 2003]

- Time since Big-Bang depends on content in FLRW

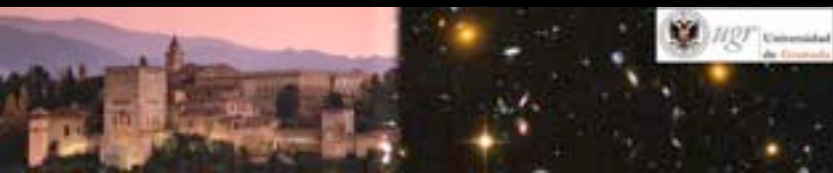
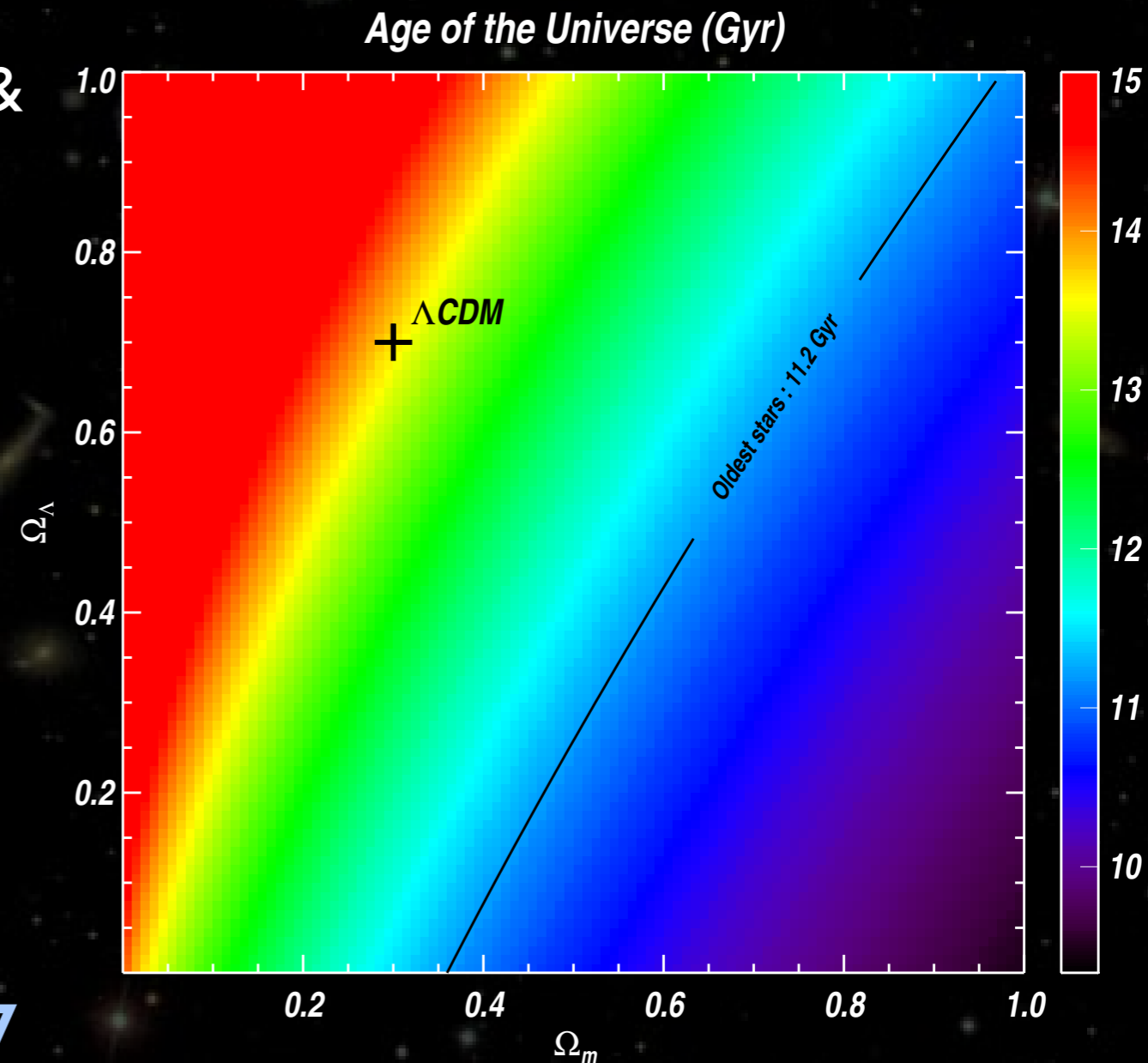
- ★ in open  $\Omega_m=0.3$  : 11.3 Gyr

- ★ in flat  $\Omega_m=1$  : 9.2 Gyr

- ★ in  $\Lambda$ CDM : 13.7 Gyr

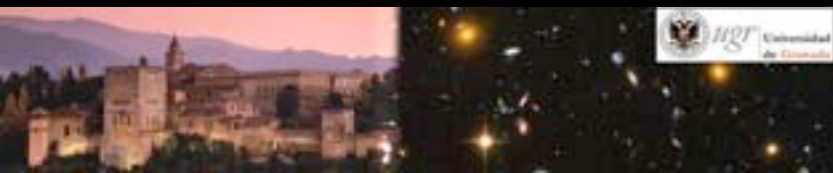
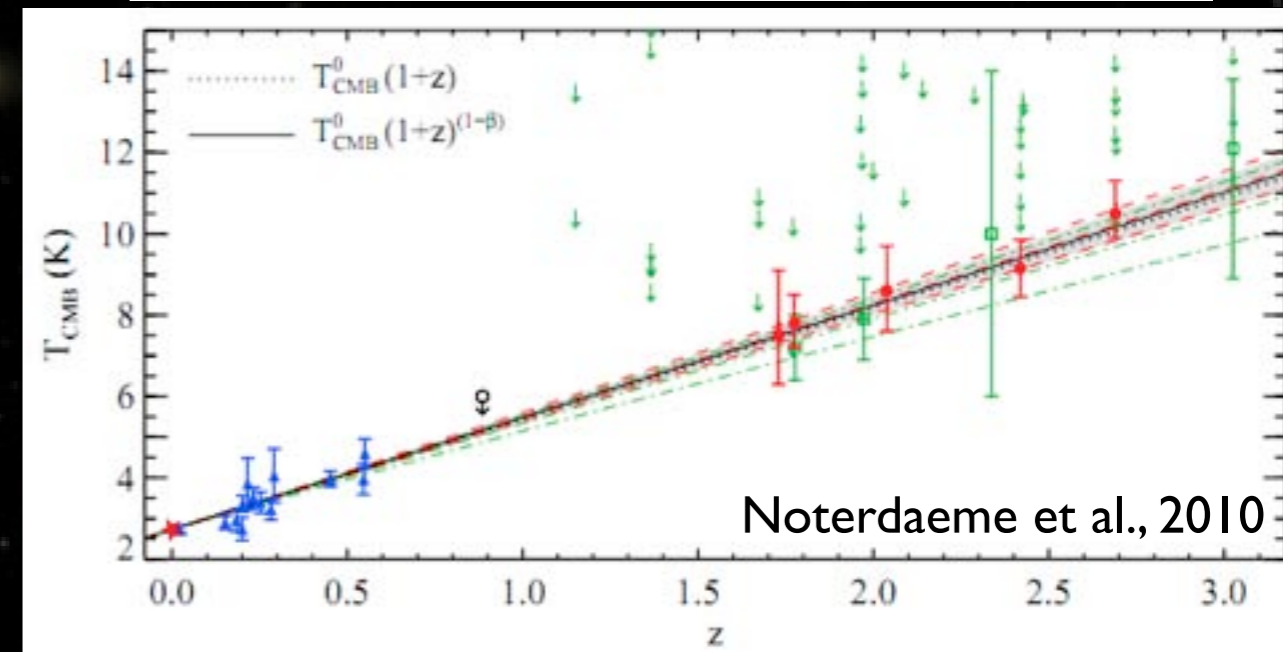
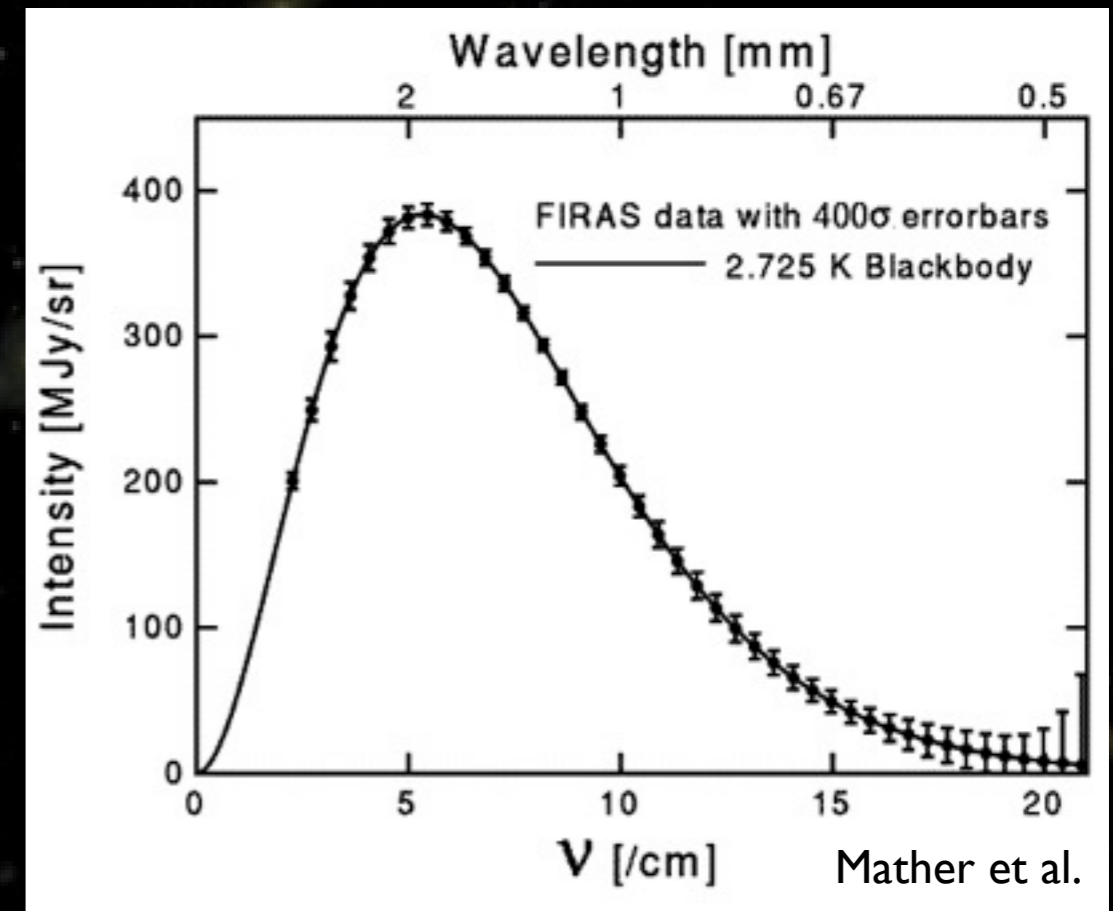
- But: SDSS J102915+172927

- ★ Ridiculously low metals, no Li, age  $\sim 13$  Gyr?



# Black Body spectrum of the CMB

- Thermal equilibrium in the past
  - ★ photons: small mean free path in the hot plasma
  - ★ decoupled when Universe becomes neutral
  - ★ Black Body observed today
- Tests of  $T_{\text{CMB}}$  as a function of  $z$ 
  - ★ Low- $z$ : S-Z effect in clusters (scattering of CMB photons on the hot gas)
  - ★ High- $z$ : rotational excitation of the CO molecules in QSOs absorption systems
- In perfect agreement with  $\Lambda\text{CDM}$





# Primordial nucleosynthesis

- **Gamow 1948:**

- ★ Earlier (high T) nuclei were broken
  - in early times:  $\gamma, p, n$
  - form nuclei when T is low enough
  - Nuclear reactions stop when T too low
- ★ Abundances calculated from theory
- ★ free parameter: fraction of baryons  $\Omega_b$

**The Origin of Elements and the Separation  
of Galaxies**

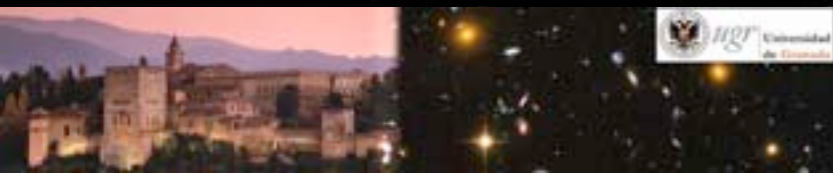
G. GAMOW  
*George Washington University, Washington, D. C.*  
June 21, 1948

- **Difficult measurements:**

- ★ Fusion in stars modify the picture
  - Metal poor stars, Ly- $\alpha$  forest of QSOs

- **Excellent agreement**

- ★ **Matches CMB  $\Omega_b$**
- ★ possible problem:  ${}^7\text{Li}$ 
  - Maybe... but measurement is extremely difficult (could be burnt in stars)
- ★ N.B.:  $\Omega_b h^2 \sim 0.02$  is small...  
➡ open Universe or dark stuff ?



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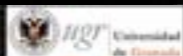
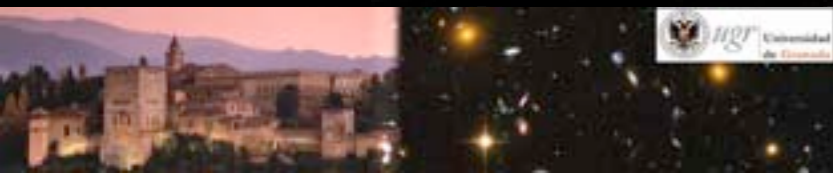
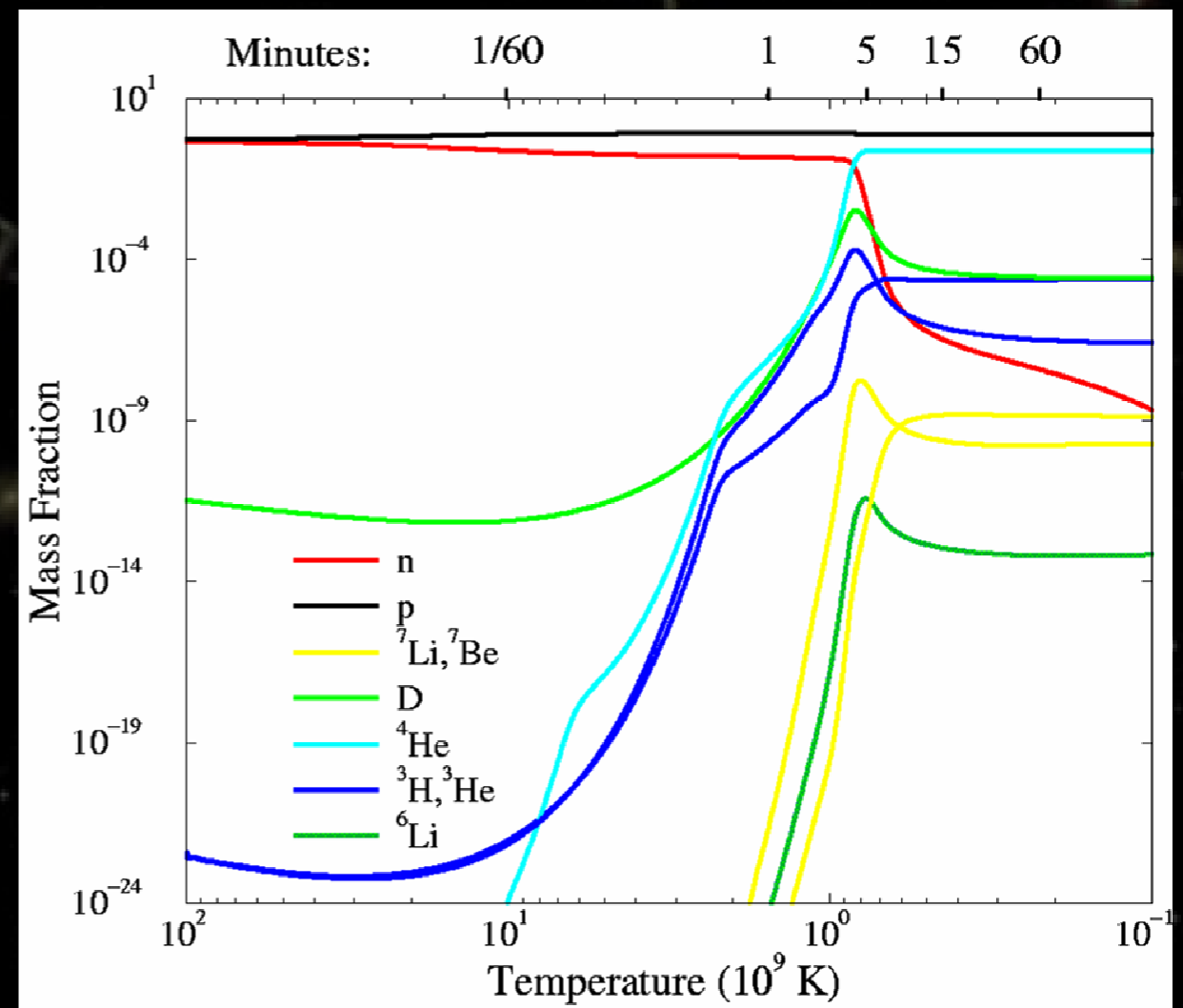
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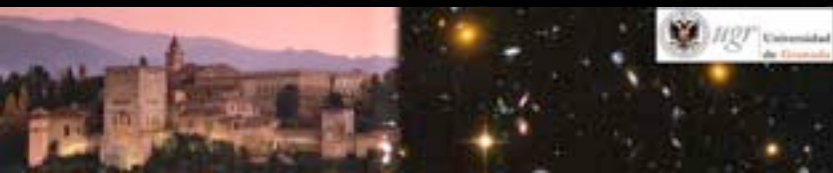
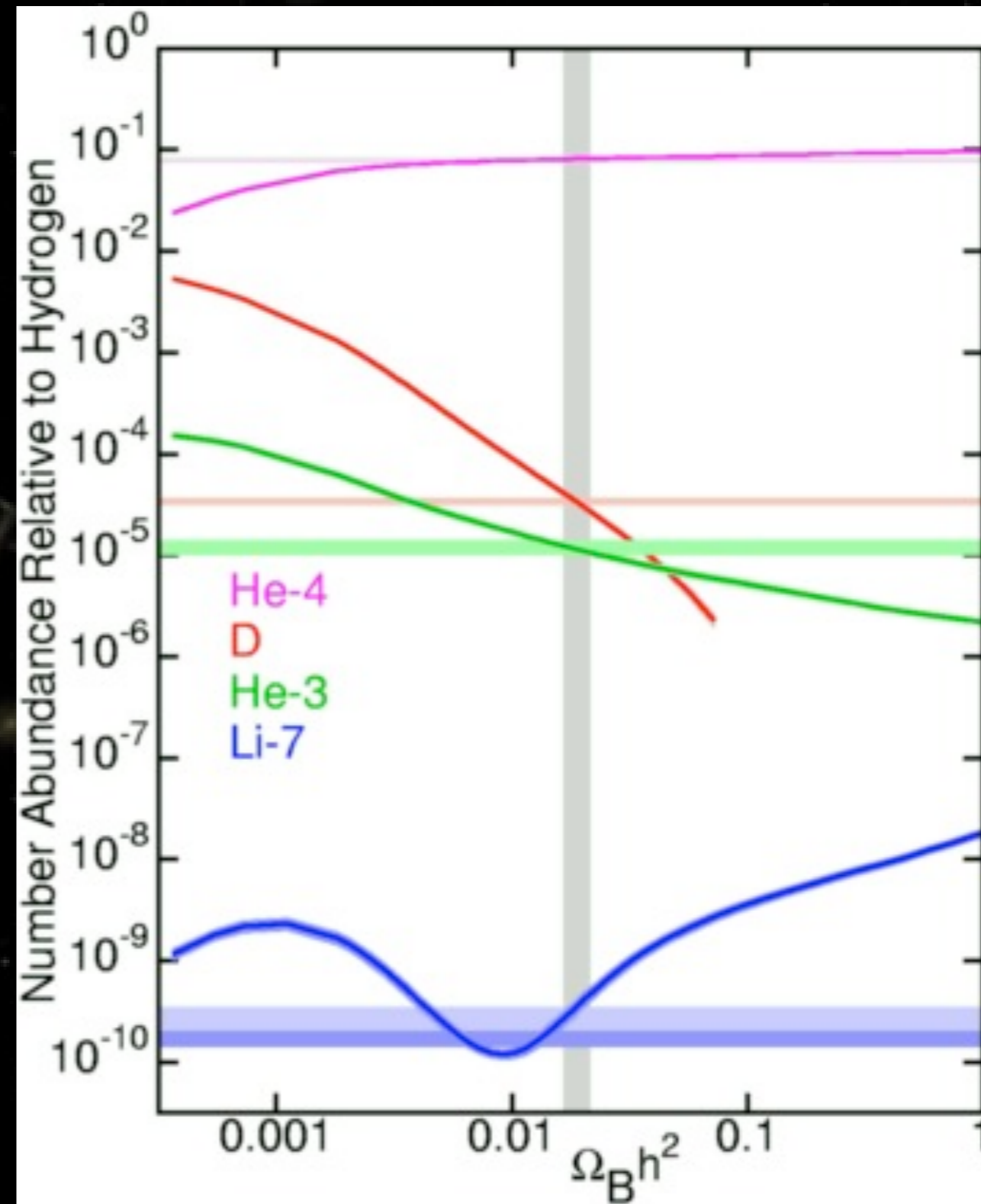
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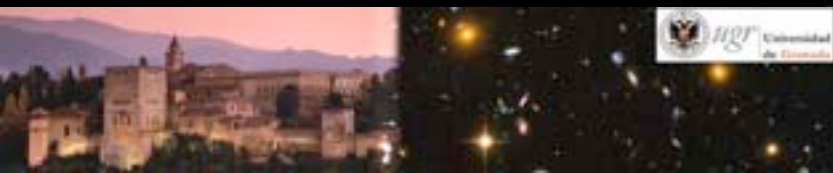
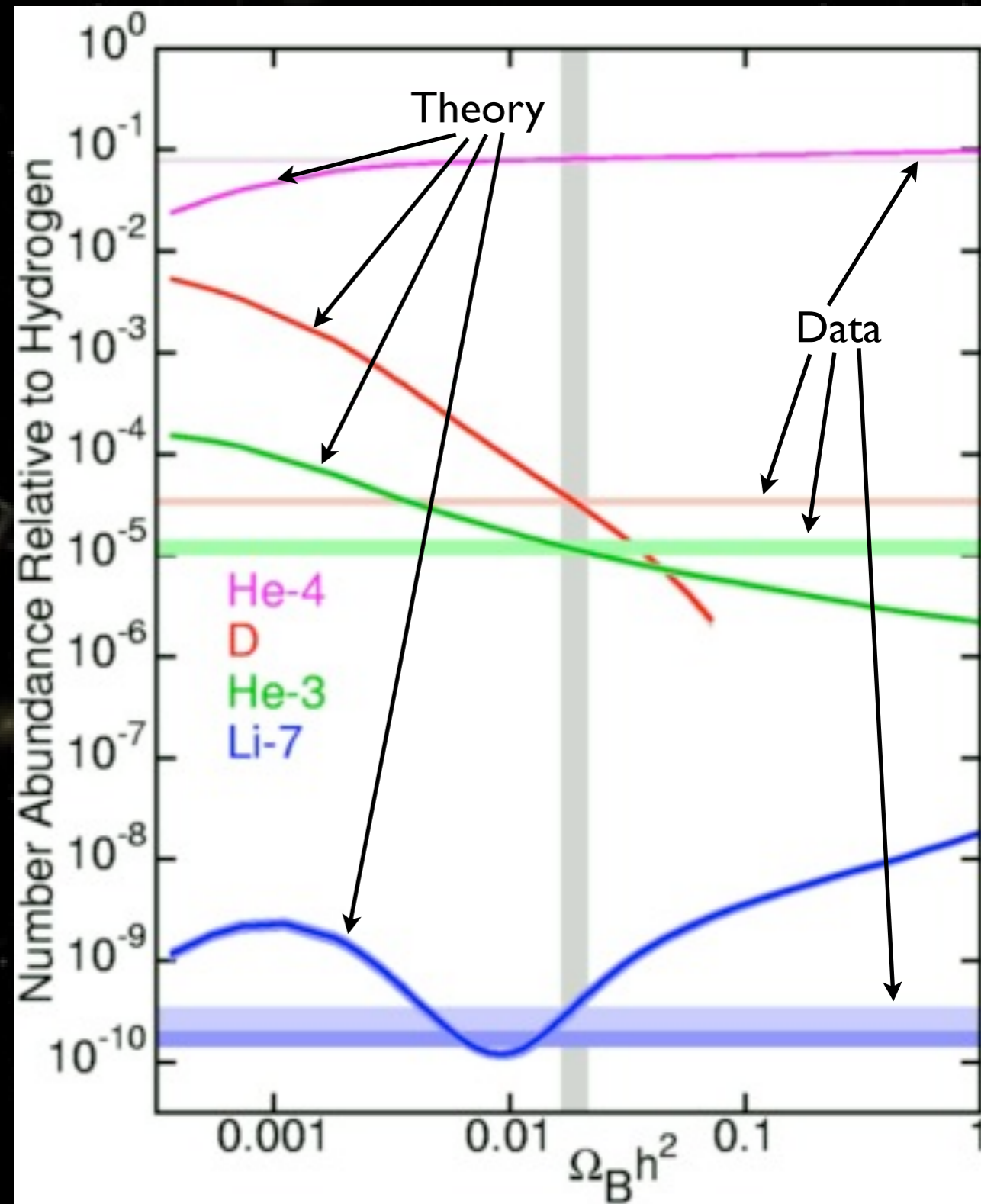
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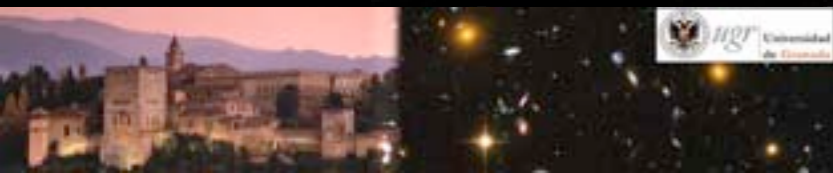
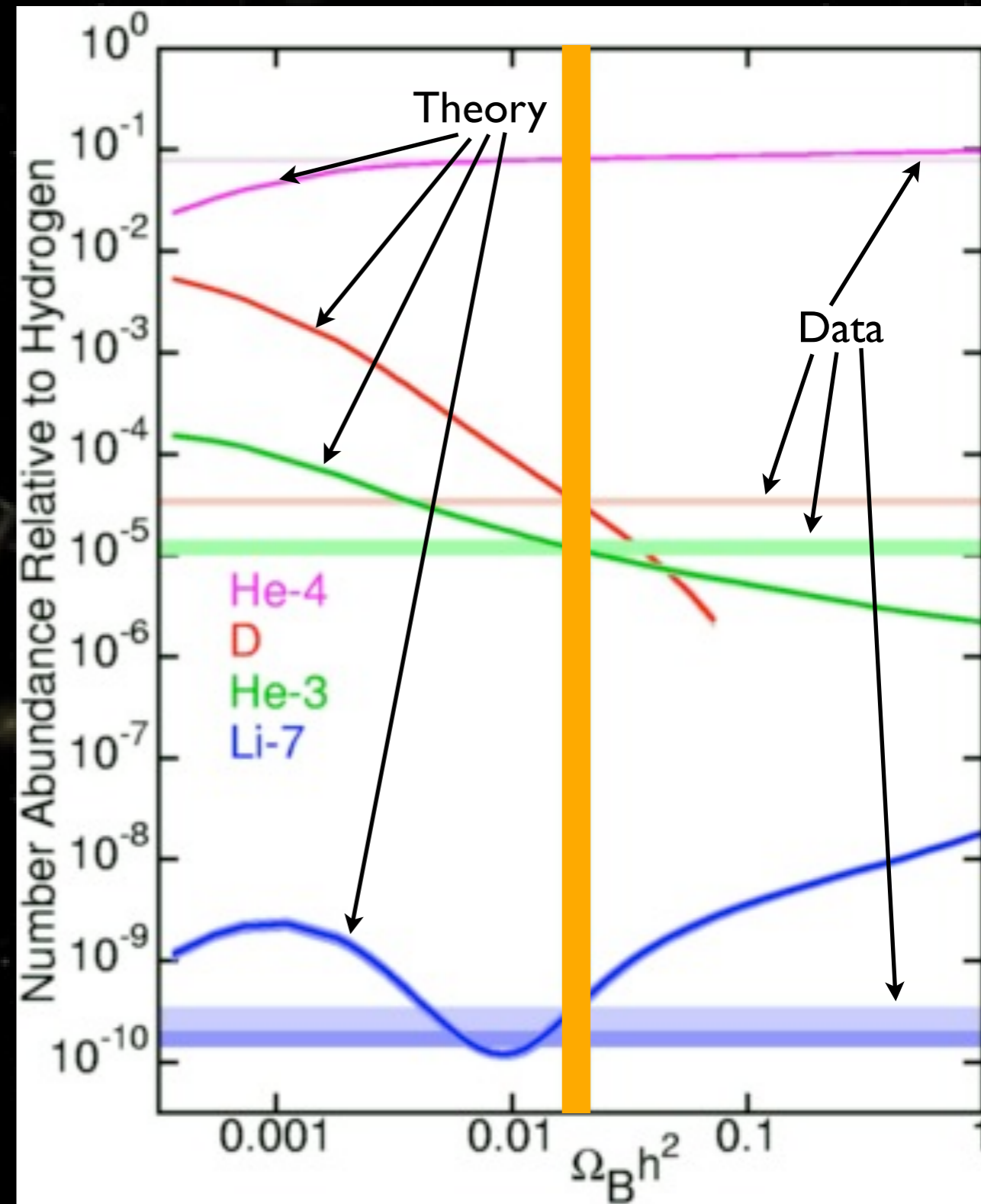
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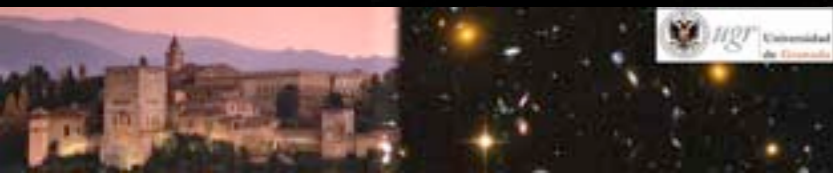
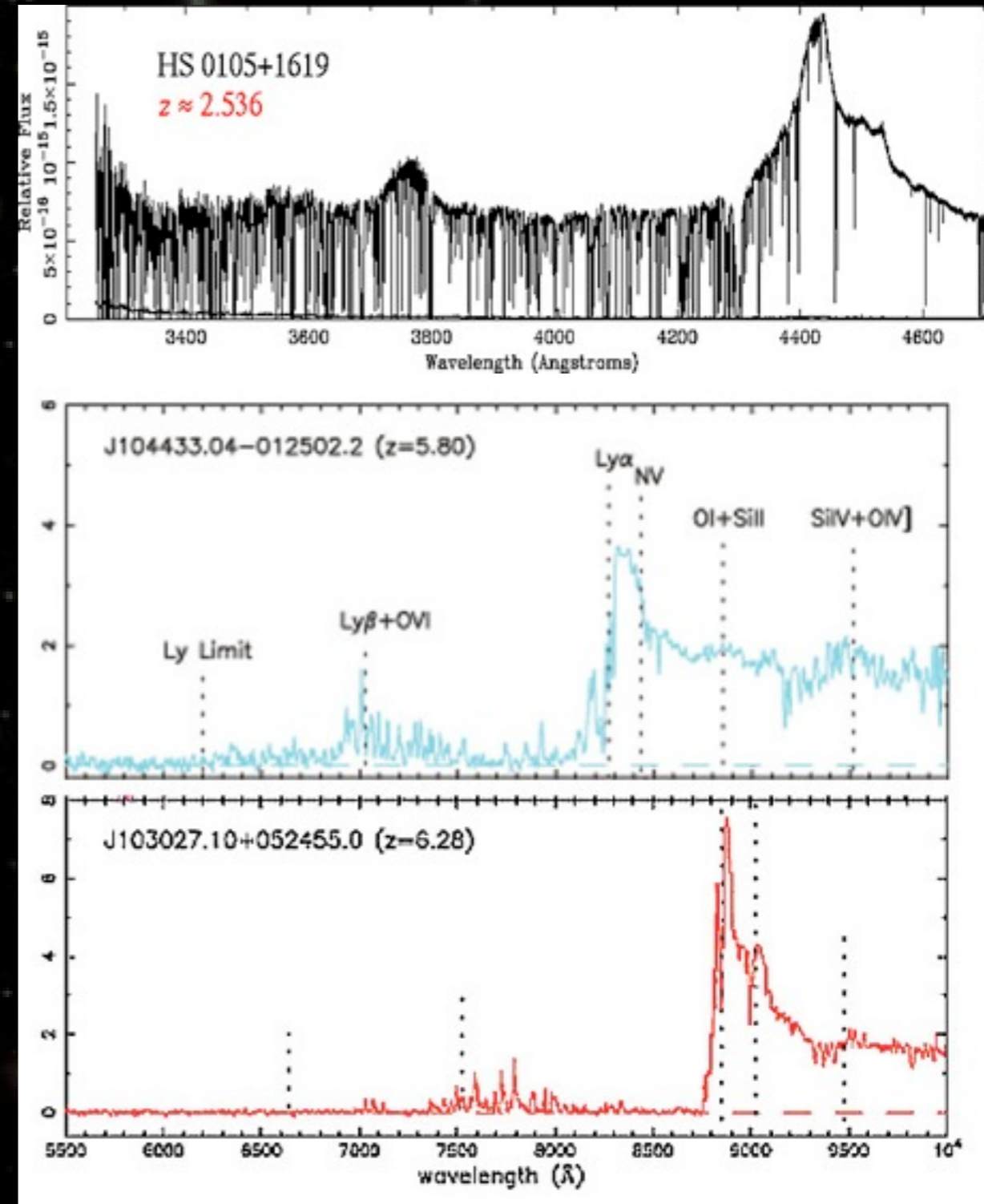
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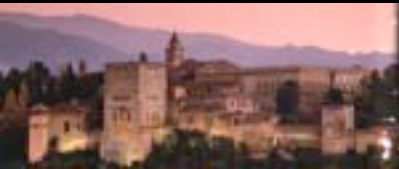


# Reionization

- Dark ages & reionization
  - ★ Neutral medium after CMB release
    - UV light is absorbed by neutral H
    - Universe opaque to starlight
  - ★ First stars & QSOs produce UV
    - Universe starts to reionize
    - UV absorption less efficient
    - Reionization complete at  $z \sim 6$
    - Universe transparent to UV
- Test: Gunn-Peterson effect
  - ★ Light from 1st quasars should be completely absorbed up to H-alpha line
  - ★ In more recent quasars one should see partial absorption only (Ly- $\alpha$  forest)
  - ★ Predicted long before observation
- Reionization epoch dating matches CMB inferred one

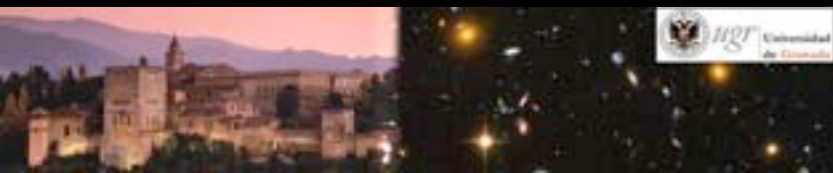
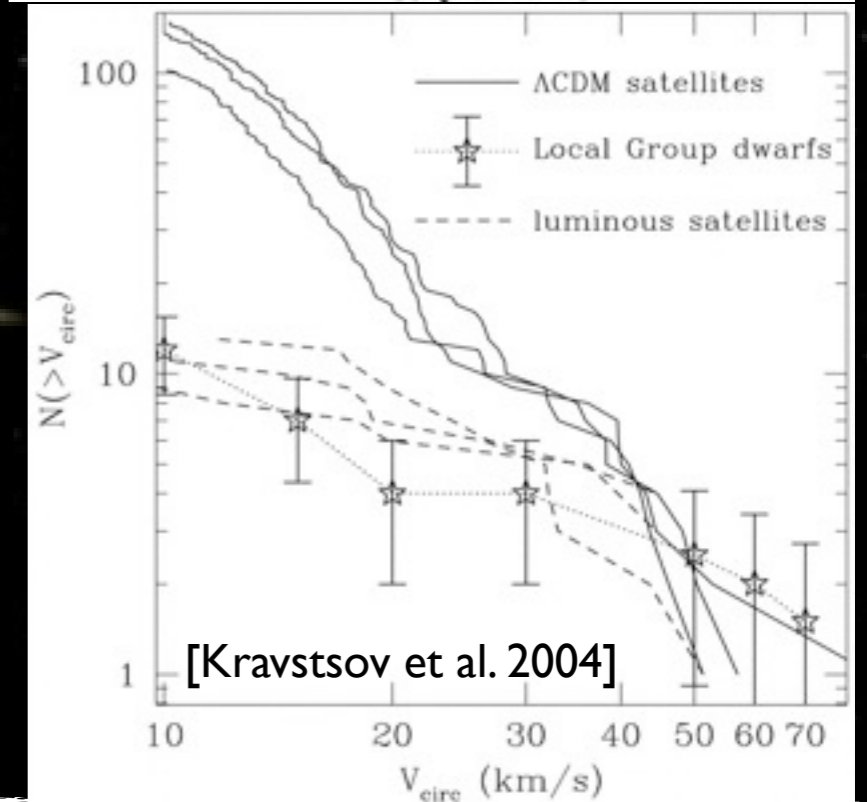
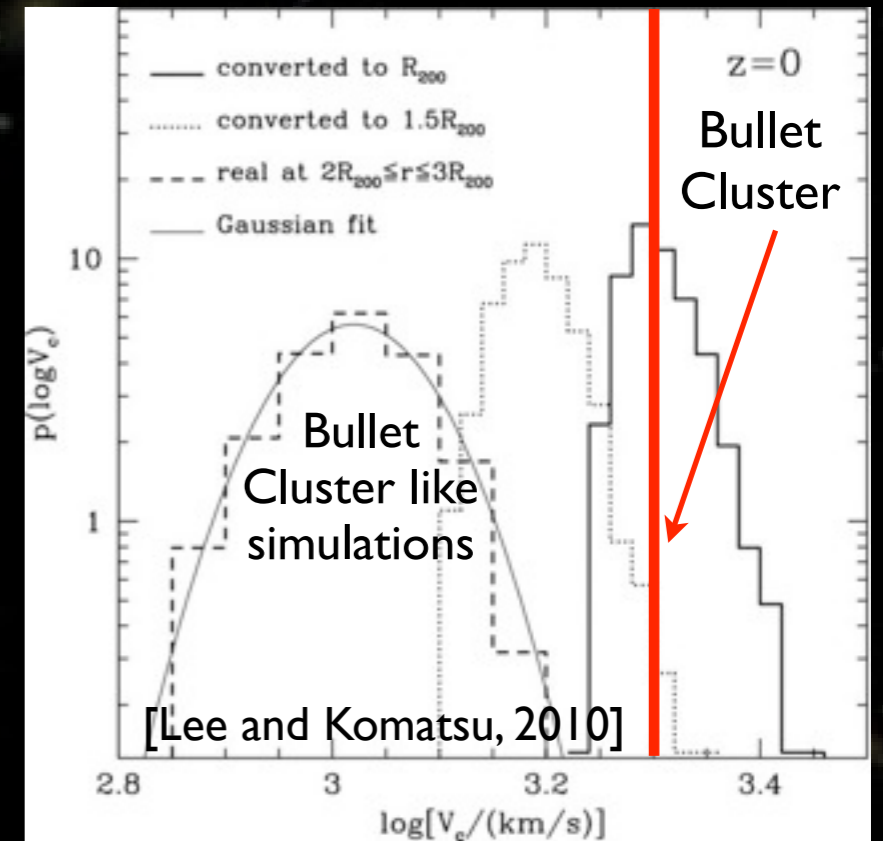


# $\Lambda$ CDM Problems



# $\Lambda$ CDM Issues

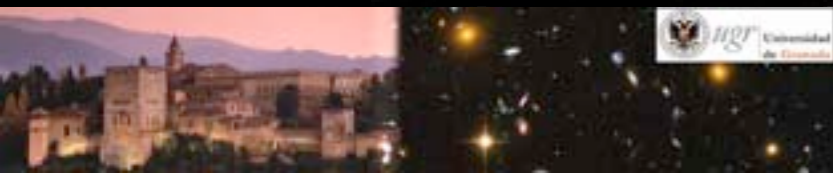
- Dark Matter
  - Dark Energy
  - BBN Lithium problem
    - ★ + star SDSS J102915+172927 ?
  - Galactic-scale dynamics:
    - Even with D.M., the «bullet cluster» is weird:
      - Gaz shock seems to involve huge velocities that are not expected in  $\Lambda$ CDM
    - not enough satellites w.r.t. simulations
    - Tully Fisher relation is empirical in  $\Lambda$ CDM but explained in MOND
    - cusps at center of galaxies unobserved
- ➔ Most of this could be explained by lack of realism or resolution of simulations...
- CMB low multipoles/axis of evil
  - ➔ not serious (just my two cents)





# Conclusions (as an observer)

- $\Lambda$ CDM has firm theoretical and observational basis
  - ★ Confirmed by independent probes
  - ★ Successfully predictive
    - CMB fluctuations/polarization, odd/even peaks, scalar index slightly below 1
    - BAO
    - Gunn-Peterson effect
- Requires large amounts of unobserved/unexplained stuff
  - ★ Dark Matter
  - ★ Dark Energy
  - ➔ Both very convincing from the observational point of view
- Also requires complex analysis of complex data
  - ★ In most cases analyses are well tested for robustness w.r.t. assumptions
  - ★ Complexity doesn't make it wrong
- Cosmology is obviously not finished, maybe just starting...
  - ★ «C'est à la fin du bal qu'on paie les musiciens ...»



# COSMOLOGY MARCHES ON

