

Minor changes in biomarker assemblages in the aftermath of the Cretaceous-Paleogene mass extinction event at the Agost distal section (Spain)

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ABSTRACT

The Cretaceous/Paleogene Boundary (KPgB) represents one of the five major mass extinctions in Earth's history, and determining the nature of associated environmental change and biotic recovery is critical for understanding the history of life on our planet. To explore that, we examined the distributions of selected biomarkers (*n*-alkanes, acyclic isoprenoids, steranes and hopanes), organic and carbonate $\delta^{13}\text{C}$ values, total organic carbon contents and major and trace elements in a distal section spanning the KPgB (Agost, SE Spain). The studied Agost section is an ~32 cm-thick bathyal sequence of marlstones, clays and marly limestones, and high sedimentation rates allow high (cm) resolution analysis. These analyses exhibit sample-to-sample variability, but there are few significant differences between pre- and post-KPgB biomarker assemblages, suggesting a rapid recovery of the non-fossilizing phytoplankton community after the KPgB. Despite the persistence of life, the organic matter assemblage is rather variable through the first 10 kyr after the impact event. This interval is associated with changing terrigenous and petrogenic inputs as well as varying redox conditions as reflected by the enrichment factor of uranium (U_{EF}) vs that of molybdenum (Mo_{EF}) as well as biomarker indices (gammacerane and homohopane indices). Moreover, sterane distributions do differ between pre- and post-KPgB sediments. Thus, the KPgB impact did affect environmental conditions and non-fossilizing algal and bacterial communities even in distal sites, but these organisms appear to have rapidly recovered, within 10 kyr after the KPgB.

1. Introduction

The mass extinction marking the Cretaceous/Paleogene boundary (KPgB), ~ 66.021 ± 0.024 million years ago (Clyde et al., 2016; Renne et al., 2013; Sprain et al., 2019) was one of the most devastating events in the history of life, associated with the disappearance of about 70% of marine and continental species (D'Hondt, 2005), and it is the most recent and well studied of the 'big five' Phanerozoic mass extinctions. Currently, the relationship between the Cretaceous/Paleogene mass extinction and a meteorite impact is broadly accepted (Hull et al., 2020; Gulick et al., 2019; Pälike, 2013; Schulte et al., 2010), with direct

evidence coming from the Chicxulub impact crater, discovered at the beginning of the 1990's on the Yucatán Peninsula in Mexico (Hildebrand et al., 1991). However, alternative or complementary mechanisms, especially volcanism associated with the Deccan Traps, are still considered (Schoene et al., 2019, 2015; Sprain et al., 2019). After decades of research, one of the most critical lines of inquiry is understanding the short- to long-term effects of the Chicxulub impact on the global environmental system, including the time frame required for the re-establishment of pre-impact environmental conditions and marine biological productivity (Bralower et al., 2020; Gulick et al., 2019; Lowery et al., 2020, 2018; Lyons et al., 2020; Rodríguez-Tovar et al.,

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2020; Schaefer et al., 2020; Sepúlveda et al., 2019).

There are numerous sites worldwide where the KPgB has been recorded (e.g. Schulte et al., 2010), but the characteristics of these sections differ depending on the distance from the Chicxulub crater. Four types of ejecta deposits have been distinguished: very proximal (e.g. Guayal, Bochil, Moncada, IODP Site 364), proximal (e.g. Beloc, El Mimbral, Brazos River), intermediate (e.g. Sugarite, Starkville South, ODP Leg 171 Sites 1049, 1050, 1052) and distal (e.g. Gubbio, El Kef, Agost and Caravaca). In all of them, major geochemical anomalies have been recognized, but the extraterrestrial contribution is particularly evident in the ejecta layer of distal marine deposits. These sites are also particularly important for assessing the global (as opposed to proximal) consequences of the KPgB and the nature of environmental and biotic recovery.

In particular, the Agost marine distal section in southeast Spain is one of the most continuous, well-preserved, and well-studied KPgB sections in the world (e.g. Molina et al., 2005; Rodríguez-Tovar et al., 2006). Our previous studies of sections in Spain (Caravaca and Agost) (Sosa-Montes de Oca et al., 2013, 2016, Sosa-Montes de Oca et al., 2018a,b, Sosa-Montes de Oca et al., 2020) and Italy (Bottaccione and Contessa) (Sosa-Montes de Oca et al., 2017) have revealed the paleo-environmental conditions before, during, and after the KPgB impact. Inorganic geochemistry and ichnological analyses on the Agost section revealed variations in bottom water oxygenation, documenting that anoxic conditions were exclusively linked to the ejecta layer produced at the event and that oxic conditions were re-established almost instantaneously, allowing the macrobenthic community to quickly recolonize the substrate (Laska et al., 2017; Rodríguez-Tovar, 2005; Rodríguez-Tovar et al., 2006, 2004; Rodríguez-Tovar and Uchman, 2006, 2004; Sosa-Montes de Oca, 2018).

Many different microorganisms became extinct at the KPgB; among them, marine Cretaceous calcareous nannoplankton were particularly impacted (Alvarez et al., 2019; Lowery et al., 2018; Vellekoop et al., 2019). However, it remains unclear how the non-fossilizing algal and microbial communities were affected. In fact, there have been relatively few reconstructions of changes in organic matter sources through the KPgB at any site (Lowery et al., 2020; Lyons et al., 2020; Schaefer et al., 2020; Sepúlveda et al., 2009) and those performed are often at a very low resolution (Taylor et al., 2018). Biomarkers preserved in sedimentary sequences and structurally linked to specific biological sources (e.g. Eglinton and Eglinton, 2008; Peters et al., 2005), are powerful tools for exploring biotic changes, especially in algal and microbial communities (Grice et al., 2005; Xie et al., 2010). Moreover, biomarker degradation and alteration is governed by environmental conditions, such that they can be also used to reconstruct a wide range of paleoenvironmental conditions (e.g. Eglinton and Eglinton, 2008; Peters et al., 2005).

Previous studies on the KPgB have used different biomarkers to reconstruct organic matter sources, terrestrial plant distributions or seawater temperatures in a small number of very proximal [i.e. IODP Site 364 (Chicxulub); Hull et al., 2020; Lyons et al., 2020] proximal [i.e. Brazos River (Texas-USA), and Beloc (Haiti); Mizukami et al., 2014, Vellekoop et al., 2016, 2014] and distal sections [i.e. Caravaca (Spain), El Kef (Tunisia), Stevns Klint, (Denmark), mid-Waipara River (New Zealand) and Seymour Island (Antartica); Mizukami et al., 2013; Scasso et al., 2020, 2005; Sepúlveda et al., 2009; Taylor et al., 2018; Vellekoop et al., 2015]. Applications of specific biomarkers to explore post-KPgB changes in biotic assemblages are even more limited. Recently, Schaefer et al. (2020) used lipid biomarker analyses at the Chicxulub crater core material (IODP Site 364) to reveal a rapid recovery of microbial life. Similarly, Bralower et al. (2020a) analyzed the micrite-rich layer (boundary clay) at 31 sites across the world, from very proximal to distal locations and through a range of depths in all ocean basins, revealing abundant bacterial and eukaryotic algal biomarkers; they conclude that most micrite in the carbonate layer at distal sites was produced by ocean-wide blooms of cyanobacteria and algae in the millennia after the KPgB impact event. Similarly, Sepúlveda et al. (2009) used sterane

distributions to argue for blooms of algae and other non-fossilizing phytoplankton just above the boundary in the Fish Clay of Denmark (Stevns Klint distal section), and then a rapid recovery to pre-event distributions. However, biotic and environmental responses are likely to be highly variable spatially, and obtaining further records is critical to understanding post-impact biotic recovery.

Here, we present a continuous cm-scale resolution and multiproxy analysis across ~32 cm-thick KPgB interval at the Agost distal section (SE Spain) that spans the ~24 kyr following the KPgB event (Groot et al., 1989). Combined with C-isotope and major and trace element profiles (Ca/Al, K/Al, Mg/Al, Cr/Al, Ni/Al, Cu/Al, Zn/Al, Rb/Al, Zr/Al, U/Al and Mo/Al ratios), the distribution of specific phytoplankton (e.g. pristane, phytane, steranes) and bacterial (hopanes) biomarkers at the Agost section allows us to: i) determine changes in the organic matter (OM) inputs before and in the immediate aftermath of the extinction; ii) examine the thermal maturity of the source rocks (changes that likely reflect reworked sedimentary OM inputs in this section); iii) evaluate photosynthetic productivity variations and changes in algal/microbial assemblages across the KPgB event; and most importantly iv) reconstruct how long it took for these interlinked processes to recover after the impact event.

2. Geological setting

The distal KPgB section of Agost (38°27'3.31"N; 0°38'9.71"E) is located at km 9 of road CV-827 (West side), North of the town of Agost, Alicante (southeast Spain) (Fig. 1). At the time of the KPgB impact, the Agost site was located in an upper-bathyal setting (~600 to 1000 m depth) (Schulte et al., 2010) at around 27–30° N palaeolatitude (Van Hinsbergen et al., 2015).

Lithologically, the Agost section comprises a succession of four units (Fig. 1) (see Sosa-Montes de Oca et al., 2018a,b, Sosa-Montes de Oca et al., 2020, for details): the Cretaceous sediments that consist of gray calcareous marlstones and marlstones from the late Maastrichtian (unit a); the overlying Paleogene sediments with a 2–3 mm thick red clay (the ejecta layer; unit b) overlain by a blackish-gray clay layer (the boundary clay; unit c) with gradually increasing carbonate contents related to the recovery of biological productivity; and finally the typical light marly limestones of the early Danian (unit d). The ejecta layer, at the distal sections, marks the sharp contrast between the Maastrichtian and Danian, and contains impact evidence such as spherules (Glass and Burns, 1987; Smit, 1999; Smit and Klaver, 1981), iridium and other platinum-group element anomalies (Alvarez et al., 1980, 1990), elevated concentrations of Cr, Fe, Ni, and Ti which may have come from the meteorite, a sulphide-rich carbonaceous chondrite (Kyte, 1998; Martínez-Ruiz et al., 1997; Shukolyukov and Lugmair, 2004), and a significant decrease in Ca due to the mass extinction among pelagic calcifiers (Goderis et al., 2013; Schulte et al., 2010).

According to the magnetostratigraphy-derived sedimentation rates (Groot et al., 1989) – estimated to be 1.98 cm kyr⁻¹ for the uppermost Maastrichtian sediments and calculated to be 0.83 cm kyr⁻¹ for the lowermost Danian sediments – the ~32 cm-thick KPgB interval at Agost is thought to span a time interval of ~29 kyr, ranging from 5 kyr before to 24 kyr after the KPgB. Furthermore, it is thought that the ejecta layer deposition itself was a snapshot on a geological time scale, potentially, on the order of days to months (Artemieva and Morgan, 2009).

3. Materials and methods

3.1. Core description and sampling

A ~ 32 cm-thick bore hole was drilled in the Agost section on May of 2017 to obtain a complete and continuous record of the KPgB interval with the four-above-mentioned units (a, b, c and d, see Fig. 1), ranging from the uppermost Maastrichtian to the lowermost Danian. A fresh/unexposed and intact KPgB section was recovered using a Rolatec RL 48

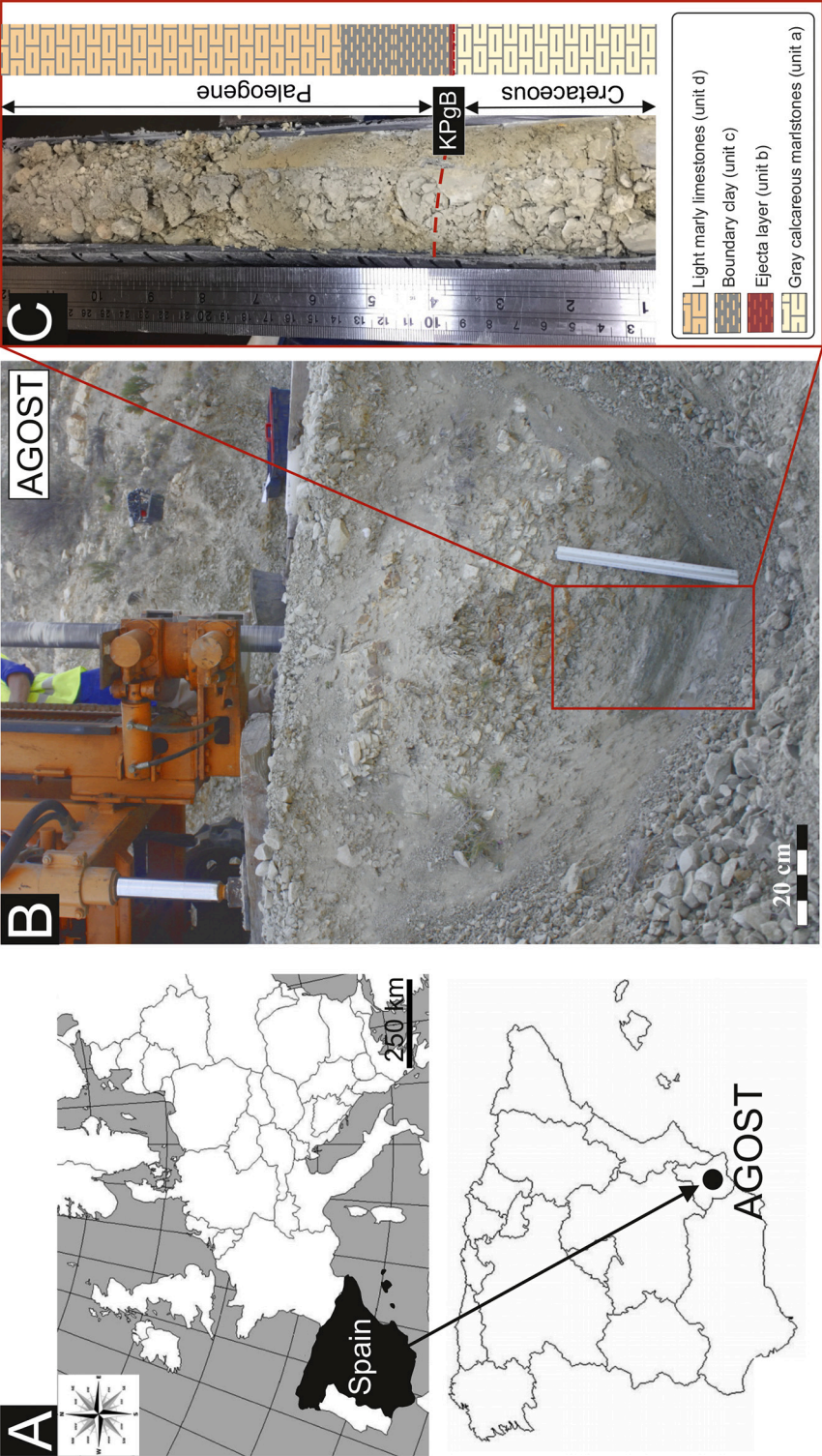


Fig. 1. The Agost outcrop and KPgB section. A-Location of the Cretaceous-Paleogene (KPg) section at the Agost site (Alicante, Southeast Spain). B- Photograph of the section taken during drilling. C-Detail of the KPgB recovered core, comprising: (unit a) gray calcareous marlstones and marlstones from the uppermost Maastrichtian, (unit b) 2 mm-thick ejecta layer, (unit c) blackish-gray boundary clay layer and (unit d) the light marly limestones, the last three belonging to the lowermost Danian. Modified from [Sosa-Montes de Oca et al. \(2018a\)](#).

L drilling machine with rubber tracks from the Center for Scientific Instrumentation (CIC) at the University of Granada (Spain). Details about the extraction methodology were previously presented in Sosa-Montes de Oca (2018) and Sosa-Montes de Oca et al. (2018a). Once the core was obtained, it was sealed and stored at 4 °C in a cold room.

The ~32 cm-thick of the KPgB core was sampled continuously each 1 to 2 cm depending of the unit interval (~20 g were taken for the Maastrichtian and Danian samples and ~0.001 g for both KPg boundary samples), yielding a total of 22 samples for inorganic geochemical analyses, 16 samples for biomarker analyses and 20 samples for total organic carbon and organic and carbonate $\delta^{13}\text{C}$ analyses. Specifically, 7 samples were taken in the 10 cm from the gray calcareous marlstones of late Maastrichtian age (unit a), and 15 samples were taken from the lower 22 cm of the early Danian, the latter comprising: 2 samples from the 2 mm-thick ejecta layer (unit b), with only enough material to determine major and trace elements analyses; 5 samples from the 5 cm-thick boundary clay layer (unit c); and 8 samples from the 15 cm-thick light marly limestones (unit d).

3.2. Inorganic geochemistry

3.2.1. ICP-OES and ICP-MS analyses

Inorganic geochemical analyses were conducted on 22 samples crushed in an agate mortar and digested with nitric acid (HNO_3) and hydrofluoric acid (HF) in order to obtain the sample solution. Trace element concentrations were determined using ICP-Mass Spectrometry (NEXION 300D). Results were calibrated using blanks and international standards, with analytical precision better than $\pm 2\%$ for 50 ppm elemental concentrations and $\pm 5\%$ for 5 ppm elemental concentrations. The concentrations of major elements (Al, Ca, Fe, K and Mg) were measured in the same sample solutions using an ICP-Optical Emission Spectroscope (Perkin-Elmer Optima 8300) with a Rh anode X-ray tube. Blanks and international standards were used for quality control and the analytical precision was better than $\pm 2.8\%$ and 1.9% for 50 ppm elemental concentrations of Al and Ca, respectively; better than $\pm 0.5\%$ for 20 ppm elemental concentrations of Fe; better than $\pm 0.4\%$ for 5 ppm elemental concentrations of K; and better than $\pm 1.5\%$ for 2.5 ppm elemental concentrations of Mg. These procedures were performed in the CIC at the University of Granada.

Through the KPgB interval, large shifts in carbonate content occur, impacting most element distributions. To study elemental variability without this carbonate dilution effect, Al-normalized concentrations or recalculation on a carbonate-free basis is usually applied (Calvert and Pedersen, 1993; De Lange et al., 1987; Morford and Emerson, 1999; Tribouillard et al., 2006; Van der Weijden, 2002). We have selected the following ratios Ca/Al, K/Al, Mg/Al, Cr/Al, Ni/Al, Cu/Al, Zn/Al, Rb/Al, Zr/Al, U/Al, and Mo/Al.

3.2.2. Stable carbon isotope analyses

Stable carbon isotope analyses were performed on 20 samples distributed along the ~32 cm-thick section. Samples were powdered using a Pulverisette 5 agate mill. The $\delta^{13}\text{C}$ values of carbonates ($\delta^{13}\text{C}_{\text{carb}}$) were determined by treatment with phosphoric acid (H_3PO_4 at 103%) using a VG Isocarb system thermostated at 90 °C. The liberated CO_2 was analyzed with an Isotope Ratio Mass Spectrometer (IRMS) VG Prism II at the University Complutense of Madrid (Spain).

For determination of organic matter $\delta^{13}\text{C}$ values ($\delta^{13}\text{C}_{\text{org}}$), samples were treated with hydrogen chloride (HCl at 3%) for 24 h to remove carbonates and then analyzed with an elemental analyzer Carlo Erba 1108 coupled to a IRMS VG Isochrom in continuous flow mode in the Stable Isotope Laboratory (SIDI) at the Universidad Autónoma de Madrid (Spain). The results were calibrated to the Vienna-Pee Dee Belemnite (V-PDB) standard, with a precision better than 0.1‰. Results are expressed in the common δ -notation in per mil (‰) relative to the V-PDB standard. The reference material used for $\delta^{13}\text{C}$ analysis was IA-R001 [$\delta^{13}\text{C}_{\text{V-PDB}}$ -26.43‰]. Replicate analysis of samples indicated an average standard

deviation of 0.08‰. The international carbonate standard NBS-19 (National Bureau of Standards; $\delta^{13}\text{C} = 1.95\%$ and $\delta^{18}\text{O} = -2.20\%$) was used to calibrate to V-PDB, with an average precision of 0.1‰ for $\delta^{13}\text{C}_{\text{carb}}$ and 0.15‰ for $\delta^{13}\text{C}_{\text{org}}$ analyses.

3.3. Organic geochemistry

3.3.1. Total organic carbon (TOC) analysis

20 samples were selected for determination of total organic carbon (TOC) contents. The analyses were performed in the Analytical Service of the Instituto de Recursos Naturales (IRNAS-CSIC) in Sevilla (Spain) using a SHIMADZU TOC-VCSH. TOC concentration was determined by subtracting the inorganic carbon (IC) from the total carbon (TC) content in each sample.

3.3.2. Biomarker analysis

Organic geochemical analyses were performed in the Department of Geology and the Centro de Instrumentación Científica-Técnica (CICT) at the University of Jaén (Spain). The 16 samples were regularly distributed along the ~32 cm-thick core (equivalent to a stratigraphic length resolution of 2 cm). First, all samples were washed with dichloromethane (DCM) to remove contamination from handling and then crushed using a Pulverisette 5 agate mill. To minimize laboratory contamination, all glassware, aluminum foil, silica, and glass wool were combusted at 550 °C for eight hours and additionally, glassware and metal tools used to process samples were rinsed first with dichloromethane (DCM). Blanks confirmed that there was no laboratory contamination. We recognize that previous work has highlighted the potential contamination risks associated with sample collection (Sosa-Montes de Oca, 2018; Sosa-Montes de Oca et al., 2018a); although we cannot exclude that, we consider it unlikely given the not use of drilling fluids associated with obtaining this short core from outcrop.

About 12 g of dry sediment were extracted using an Accelerated Solvent Extractor (ASE) DIONEX 350 using a solvent mixture of dichloromethane:methanol (DCM:MeOH) (v/v, 4:1) at 100 °C and 7.6×10^6 Pa. We included one blank for every 5 samples. After extraction, the solvent volume was reduced under pressure using a Syncore Polyvap Buchi with a Vacuum Controller V-855 and a Vacuum Pump-V700 at 60 mbar, 30 °C and 70 rpm during 40 min approx.

Activated copper turnings were added to the total lipid extract (TLE) and stirred overnight to remove elemental sulphur from the samples. The Cu turnings were removed, and the TLE was separated using flash column chromatography, using activated alumina as stationary phase, via sequential elution with *n*-hexane (*n*-Hex; 14 ml, yielding saturated hydrocarbons), *n*-Hex:DCM (v/v, 9:1, 14 ml, yielding aromatic compounds), and MeOH (14 ml, yielding polar compounds).

An internal standard, 2.5 $\mu\text{g}/\text{ml}$ of androstane, was added to the saturated hydrocarbon fractions prior to analysis by gas chromatography-mass spectrometry (GC-MS). Analyses were carried out on a Thermo Trace Ultra gas chromatograph connected to a Thermo DSQ II mass spectrometer (GC-MS). The GC was fitted with a fused silica capillary column (Supelco Equity-5; 30 m \times 0.25 mm \times 0.25 μm) and was operated with helium as carrier gas. The samples (in 50 μl Hex) were injected at 70 °C and the oven was subsequently programmed to 130 °C at 20 °C per min and then at 4 °C per min to 300 °C where it was held for 25 min. Biomarkers were identified by comparison of mass spectra and retention time with those reported in the literature (i.e. Peters et al., 2005) and with data from other studies performed in the same equipment (i.e. Quijano et al., 2012); for the identification of gammacerane an authentic standard purchased by Chiron was co-injected.

4. Results

4.1. Major and trace elemental composition

Geochemical data analyzed by ICP-OES and ICP-MS (ICP-OES/MS)

Table 1
Elemental content (major and trace) and elemental ratios, measured by ICP-OES and ICP-MS across the ~32 cm-thick interval in the KPgB at Agost section.

| Samples | | Distance KPgB | Unit | Major elements | | | | Geochemical proxies | | | | | | | | | | | | |
|----------------|-----------------|---------------|-------|------------------|------------------|-----------------|------------------|--------------------------------|-------|-------|------|-------|-------|-------|-------|-------|-------|-------|------|------|
| | | | | (%) | | | | (x10 ⁻⁴) | | | | | | | | | | | | |
| Reference Lab. | (g) | (cm) | | Mg ¹² | Al ¹³ | K ¹⁹ | Ca ²⁰ | CaCO ₃ [*] | Mg/Al | Ca/Al | K/Al | Rb/Al | Zr/Al | Cr/Al | Ni/Al | Cu/Al | Zn/Al | Mo/Al | U/Al | |
| 22 | Ag3 MI 30–32 | 3.80 | 20.0 | Unit d | 0.6 | 2.5 | 0.8 | 26.0 | 64.9 | 0.3 | 10.4 | 0.3 | 4.8 | 8.9 | 19.2 | 13.5 | 4.8 | 15.6 | 0.3 | 0.59 |
| 21 | Ag3 MI 28–30 | 3.90 | 18.0 | | 0.5 | 2.0 | 0.7 | 23.3 | 58.3 | 0.3 | 11.7 | 0.4 | 5.4 | 10.6 | 22.1 | 15.6 | 5.4 | 17.9 | 0.3 | 0.7 |
| 20 | Ag3 MI 26–28 | 3.85 | 16.0 | | 0.6 | 2.2 | 0.8 | 27.2 | 68.0 | 0.3 | 12.4 | 0.4 | 4.9 | 9.1 | 19.6 | 14.3 | 4.9 | 16.2 | 0.2 | 0.7 |
| 19 | Ag3 MI 24–26 | 3.95 | 14.0 | | 0.5 | 1.9 | 0.9 | 28.4 | 71.1 | 0.3 | 14.9 | 0.5 | 4.7 | 8.9 | 19.3 | 14.4 | 4.9 | 16.5 | 0.3 | 0.7 |
| 18 | Ag3 MI 22–24 | 3.68 | 12.0 | | 0.5 | 1.8 | 0.8 | 27.7 | 69.4 | 0.3 | 15.6 | 0.5 | 5.4 | 10.1 | 19.0 | 15.9 | 5.1 | 16.5 | 0.3 | 0.8 |
| 17 | Ag3 MI 20–22 | 3.90 | 10.0 | | 0.6 | 1.8 | 0.9 | 28.8 | 72.0 | 0.3 | 16.4 | 0.5 | 5.1 | 9.5 | 17.2 | 14.6 | 4.4 | 15.0 | 0.2 | 0.8 |
| 16 | Ag3 MI 18–20 | 3.80 | 8.0 | | 0.5 | 1.7 | 0.8 | 28.8 | 71.9 | 0.3 | 17.3 | 0.5 | 5.2 | 10.7 | 17.0 | 14.4 | 4.3 | 13.9 | 0.1 | 0.9 |
| 15 | Ag3 MI 16–18 | 3.88 | 6.0 | | 0.5 | 1.6 | 0.8 | 28.7 | 71.8 | 0.3 | 17.8 | 0.5 | 5.3 | 9.6 | 16.7 | 14.9 | 4.9 | 15.8 | 0.1 | 1.0 |
| 14 | Ag3 MI 15–16 | 3.80 | 4.5 | Unit c | 0.5 | 1.7 | 0.8 | 28.1 | 70.3 | 0.3 | 16.5 | 0.5 | 5.1 | 9.7 | 16.5 | 14.5 | 5.0 | 15.3 | 0.2 | 1.0 |
| 13 | Ag3 MI 14–15 | 3.79 | 3.5 | | 0.5 | 1.9 | 0.8 | 27.9 | 69.7 | 0.3 | 15.0 | 0.4 | 5.3 | 10.5 | 16.8 | 13.7 | 5.4 | 14.3 | 0.2 | 1.0 |
| 12 | Ag3 MI 13–14 | 3.27 | 2.5 | | 0.5 | 2.1 | 0.9 | 27.4 | 68.5 | 0.3 | 13.1 | 0.4 | 5.4 | 10.3 | 16.5 | 12.7 | 5.2 | 15.4 | 0.1 | 1.0 |
| 11 | Ag3 MI 12–13 | 3.14 | 1.5 | | 0.6 | 2.1 | 1.0 | 26.7 | 66.7 | 0.3 | 12.6 | 0.5 | 5.0 | 9.0 | 15.0 | 11.2 | 4.4 | 13.2 | 0.2 | 0.8 |
| 10 | Ag3 MI 10–12 Pg | 2.23 | 0.5 | | 0.5 | 2.0 | 0.9 | 28.5 | 71.2 | 0.3 | 14.6 | 0.5 | 5.8 | 9.7 | 17.2 | 11.6 | 5.5 | 13.8 | 0.1 | 0.9 |
| 9 | K/Pg (10–12) | 0.41 | 0.0 | Unit b | 0.5 | 1.8 | 0.8 | 28.2 | 70.5 | 0.3 | 15.3 | 0.5 | 6.0 | 10.0 | 18.5 | 12.8 | 6.0 | 14.7 | 0.1 | 1.0 |
| 8 | K/Pg (8–9) | 0.07 | 0.0 | | 0.4 | 1.5 | 1.1 | 27.6 | 69.1 | 0.3 | 18.5 | 0.7 | 10.8 | 13.6 | 25.0 | 17.6 | 7.2 | 19.0 | 0.2 | 1.4 |
| 7 | Ag3 MI 10–12 K | 2.87 | –0.5 | Unit a | 0.6 | 1.9 | 0.9 | 28.7 | 71.7 | 0.3 | 14.9 | 0.5 | 5.5 | 7.4 | 16.2 | 8.9 | 4.9 | 12.9 | 0.1 | 0.9 |
| 6 | Ag3 MI 9–10 | 3.50 | –1.5 | | 0.5 | 1.5 | 0.8 | 30.0 | 75.1 | 0.3 | 19.4 | 0.6 | 5.3 | 9.9 | 16.5 | 13.2 | 4.9 | 12.4 | 0.1 | 1.0 |
| 5 | Ag3 MI 8–9 | 3.41 | –2.5 | | 0.5 | 1.7 | 0.7 | 30.3 | 75.7 | 0.3 | 18.3 | 0.4 | 5.3 | 9.4 | 16.6 | 13.0 | 5.2 | 14.8 | 0.2 | 1.0 |
| 4 | Ag3 MI 6–8 | 3.97 | –4.0 | | 0.5 | 1.6 | 0.9 | 30.9 | 77.2 | 0.3 | 19.3 | 0.6 | 5.0 | 9.0 | 15.3 | 13.5 | 5.2 | 15.2 | 0.1 | 0.9 |
| 3 | Ag3 MI 4–6 | 3.82 | –6.0 | | 0.5 | 1.5 | 0.9 | 30.8 | 77.0 | 0.3 | 19.9 | 0.6 | 5.7 | 9.9 | 16.5 | 14.6 | 5.8 | 15.4 | 0.1 | 1.1 |
| 2 | Ag3 MI 2–4 | 3.78 | –8.0 | | 0.5 | 1.3 | 0.8 | 29.2 | 73.1 | 0.4 | 22.8 | 0.6 | 5.5 | 8.8 | 16.2 | 13.7 | 6.3 | 15.5 | 0.2 | 1.1 |
| 1 | Ag3 MI 0–2 | 3.90 | –10.0 | | 0.5 | 1.7 | 0.9 | 31.4 | 78.4 | 0.3 | 18.6 | 0.5 | 7.9 | 13.6 | 22.0 | 18.2 | 6.9 | 21.8 | 0.2 | 1.3 |

^a Calculated from Ca²⁰.

are presented as elemental ratios (Ca/Al, K/Al, Mg/Al, Cr/Al, Ni/Al, Cu/Al, Zn/Al, Rb/Al, Zr/Al, U/Al and Mo/Al) across the KPgB and shown in Table 1 and Fig. 2. Selected Al-normalized ratios provide information about environmental variations at continuous, cm-resolution along the ~32 cm-thick studied interval. Elemental ratios related to the meteorite input, such as Cr/Al, Ni/Al, Cu/Al and Zn/Al, all clearly identify the ejecta layer (Fig. 2). Some geochemical ratios used as redox indicators, such as Mo/Al and U/Al (Algeo and Tribouillard, 2009; Tribouillard et al., 2012; Zhou et al., 2012) also peak in the ejecta layer, but exhibit complex and contrasting profiles through the rest of the section. In contrast, the Ca/Al ratio and CaCO₃ content decrease within the ejecta layer, with values of 15.3 for the Ca/Al ratio and 69.1% for the CaCO₃ content (Fig. 2). Typical detrital input ratios, K/Al, Mg/Al, Rb/Al, and Zr/Al, have more complex profiles, with several peaks in the Maastrichtian but also at the ejecta layer, and some variations occurring in the lower Danian, with relatively stable and low values above the boundary clay interval, at 6 cm above the KPgB.

4.2. Stable carbon isotopes

The $\delta^{13}\text{C}_{\text{carb}}$ profile shows some variability along the section, with values ranging between 1.2 and 1.7‰ (Table 2 and Fig. 3). A $\delta^{13}\text{C}_{\text{carb}}$ excursion is recorded at the KPgB, with a negative shift of 0.5‰, and a later recovery at ca. 5 cm above the KPgB. The $\delta^{13}\text{C}_{\text{org}}$ values range between –24.3 and –26.2‰ and in the upper part of the section have a similar trend to the $\delta^{13}\text{C}_{\text{carb}}$ curve (Table 2 and Fig. 3); however, they exhibit a positive rather than negative carbon isotope excursion in the clay boundary layer above KPgB.

4.3. TOC content

The TOC content varies from 0.001 to 1.3% along the section (Table 2, Figs. 3, 4). However, most of the values are low and the variability largely arises from three relatively high TOC peaks that occur: at 8 cm below the KPgB in the unit a, gray calcareous marlstone from the uppermost Maastrichtian (0.6%), and at 6 cm and 20 cm above the KPgB, both in the unit d, light marly limestone from the lowermost Danian (0.25 and 1.3%, respectively).

4.4. Biomarkers

The saturated hydrocarbon fractions are dominated by *n*-alkanes (with varying distributions), the acyclic isoprenoids pristane and phytane, and low abundances of steranes and hopanes, of likely eukaryotic and bacterial origin, respectively (Peters et al., 2005).

4.4.1. *n*-Alkanes

The *n*-alkanes vary in abundance between 3 and 120 µg/g dry weight (hereafter DW) (Fig. 5). The carbon number ranges from 16 to 40, but the distribution and dominant components vary significantly (Table 3a). Long-chain *n*-alkanes (up to >C₃₄) occur in all samples (except in one at 1.5 cm below the ejecta layer), but the dominant homologue varies, such that the average chain length (ACL) (Poynter and Eglinton, 1990) varies from 30 to 32 (Table 3a and Fig. 7). This is driven primarily by changes in the proportion of C₂₉ vs C₃₁ *n*-alkanes, such that the $n\text{-C}_{31}/[n\text{-C}_{29} + n\text{-C}_{31}]$ ratio (Schefuß et al., 2003) exhibits similar variations as the ACL through the section (Table 3a and Fig. 7).

Short-chain *n*-alkanes (C₁₇–C₁₉) are absent or occur in low abundances throughout the section (Fig. 5), despite the thermal maturity of the sediments (see below). The lack of short-chain and predominance of long-chain *n*-alkanes suggests that the organic matter has a significant terrigenous source. To evaluate that further, we calculated the high-molecular-weight (HMW) *n*-alkanes vs \sum total *n*-alkanes ratio ($[n\text{-C}_{25} + n\text{-C}_{26} + n\text{-C}_{27} + n\text{-C}_{28} + n\text{-C}_{29} + n\text{-C}_{30} + n\text{-C}_{31} + n\text{-C}_{32} + n\text{-C}_{33} + n\text{-C}_{34} + n\text{-C}_{35} + n\text{-C}_{36} + n\text{-C}_{37} + n\text{-C}_{38} + n\text{-C}_{40}]/[\sum \text{total } n\text{-alkanes}]$) (Fig. 5). Values are generally high, exceeding 0.7, but are particularly high in the

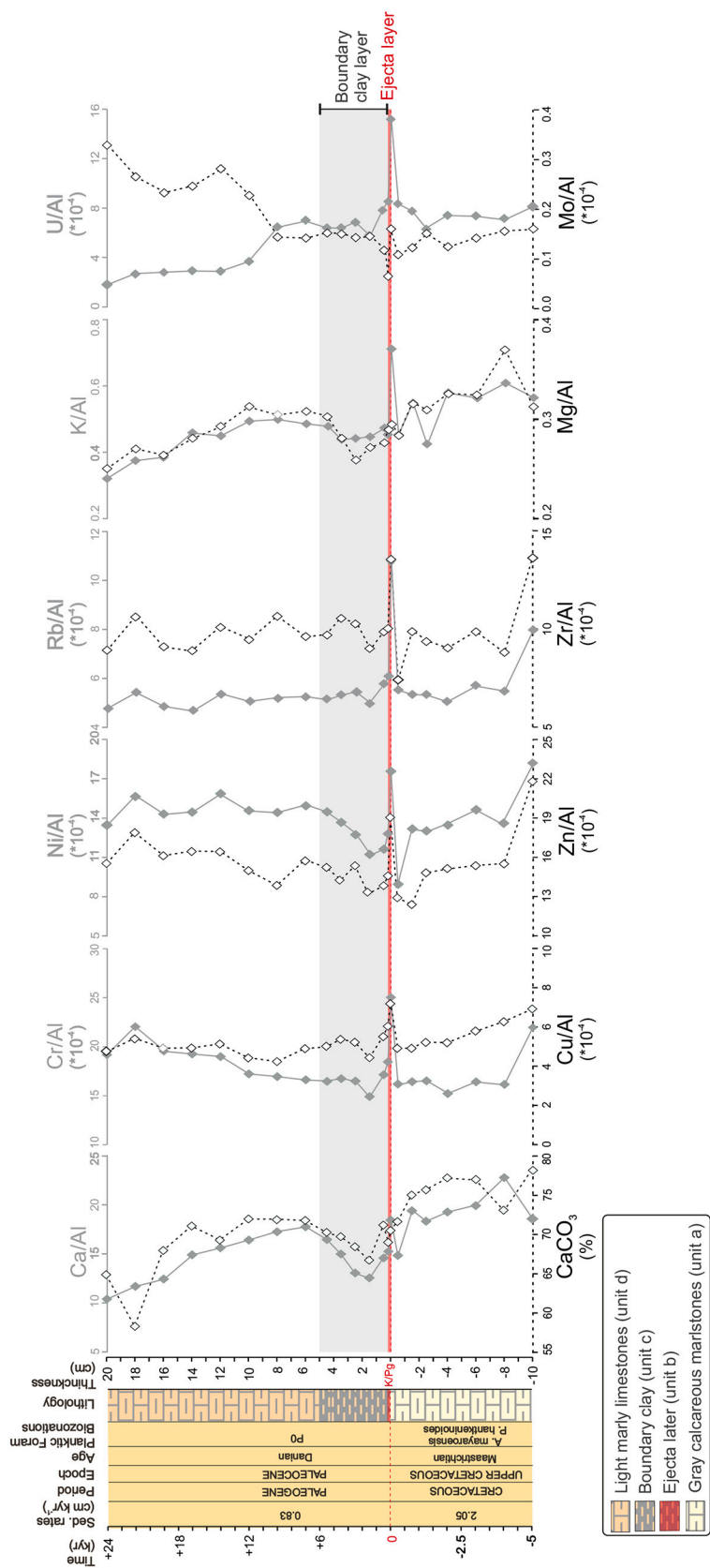


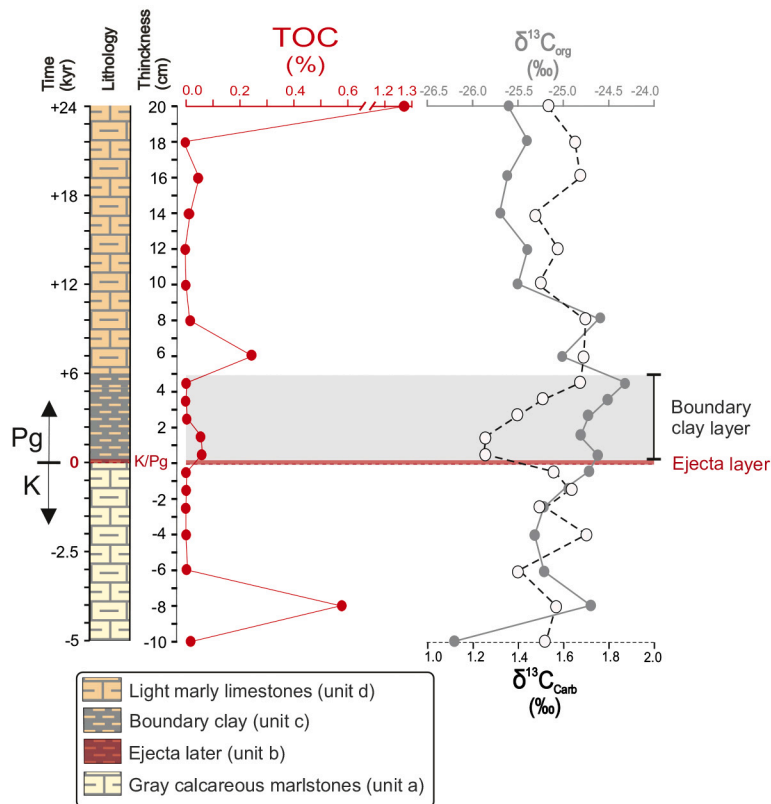
Fig. 2. Elemental/Al ($\times 10^{-4}$) profiles from the KPgB at Agost section. Cr/Al, Ni/Al, Cu/Al, Zn/Al ratios are related to the meteorite input; Mo/Al and U/Al ratios are used as redox proxies; and K/Al, Mg/Al, Rb/Al, and Zr/Al ratios are related to detrital input. Top geochemical ratios are represented by continuous lines and bottom geochemical ratios by dashed lines. Lithologically, from the bottom to the top: (unit a) gray calcareous marlstones and marlstones from the late Maastrichtian (Cretaceous) and the three overlying Paleogene units from the early Danian: (unit b) a 2–3 mm-thick red clay (the ejecta layer), (unit c) blackish-gray clay layer (the boundary clay), and (unit d) light marly limestones. Planktic foraminiferal zones are based on [Molina et al. \(2005\)](#), while magnetostratigraphy is based on [Groot et al. \(1989\)](#). (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

Table 2

Total organic carbon content (TOC, %) and stable carbon isotopic composition (‰ V-PDB) across the KPgB at Agost section.

| Samples* | | Distance KPgB (cm) | Unit | C _{inorganic} (%) | C _{organic} (TOC) (%) | C _{total} (%) | Stable carbon isotopes | |
|----------|-----------------|-----------------------|--------|-------------------------------|-----------------------------------|---------------------------|------------------------------------|-------------------------------------|
| | | | | | | | $\delta^{13}\text{C}_{\text{org}}$ | $\delta^{13}\text{C}_{\text{carb}}$ |
| 22 | Ag3 MI 30–32 | 20.0 | Unit d | 7.7 | 1.3 | 9.0 | –25.6 | 1.5 |
| 21 | Ag3 MI 28–30 | 18.0 | | 9.5 | 0.0 | 9.5 | –25.4 | 1.7 |
| 20 | Ag3 MI 26–28 | 16.0 | | 8.8 | 0.1 | 8.9 | –25.6 | 1.7 |
| 19 | Ag3 MI 24–26 | 14.0 | | 9.6 | 0.0 | 9.6 | –25.7 | 1.5 |
| 18 | Ag3 MI 22–24 | 12.0 | | 9.2 | 0.0 | 9.2 | –25.0 | 1.6 |
| 17 | Ag3 MI 20–22 | 10.0 | | 9.6 | 0.0 | 9.6 | –25.5 | 1.5 |
| 16 | Ag3 MI 18–20 | 8.0 | | 9.6 | 0.0 | 9.6 | –24.6 | 1.7 |
| 15 | Ag3 MI 16–18 | 6.0 | Unit c | 9.2 | 0.3 | 9.5 | –25.0 | 1.7 |
| 14 | Ag3 MI 15–16 | 4.5 | | 9.7 | 0.0 | 9.7 | –24.3 | 1.7 |
| 13 | Ag3 MI 14–15 | 3.5 | | 8.1 | 0.0 | 8.1 | –24.5 | 1.5 |
| 12 | Ag3 MI 13–14 | 2.5 | | 9.0 | 0.0 | 9.0 | –24.7 | 1.4 |
| 11 | Ag3 MI 12–13 | 1.5 | Unit a | 8.7 | 0.1 | 8.8 | –24.8 | 1.3 |
| 10 | Ag3 MI 10–12 Pg | 0.5 | | 9.3 | 0.1 | 9.4 | –24.6 | 1.2 |
| 7 | Ag3 MI 10–12 K | –0.5 | | 9.1 | 0.0 | 9.1 | –24.7 | 1.7 |
| 6 | Ag3 MI 9–10 | –1.5 | | 9.5 | 0.0 | 9.5 | –25.0 | 1.6 |
| 5 | Ag3 MI 8–9 | –2.5 | | 9.8 | 0.0 | 9.8 | –25.2 | 1.5 |
| 4 | Ag3 MI 6–8 | –4.0 | | 9.8 | 0.0 | 9.8 | –25.3 | 1.7 |
| 3 | Ag3 MI 4–6 | –6.0 | | 9.8 | 0.0 | 9.9 | –25.2 | 1.4 |
| 2 | Ag3 MI 2–4 | –8.0 | | 9.1 | 0.6 | 9.7 | –24.7 | 1.6 |
| 1 | Ag3 MI 0–2 | –10.0 | | 9.7 | 0.0 | 9.7 | –26.2 | 1.5 |

* The ejecta layer was not analyzed because there was not enough sample material.

**Fig. 3.** Total organic carbon content (TOC, %) and stable carbon isotopes profiles (‰ V-PDB) from the KPgB at Agost section. $\delta^{13}\text{C}_{\text{org}}$ is represented by continuous line and $\delta^{13}\text{C}_{\text{carb}}$ by dashed line. For lithology, see caption in Fig. 2.

interval after the KPgB (unit c), which supports a predominant terrigenous source of OM in this interval.

To facilitate the interpretation of *n*-alkane distributions in relation to the source, diagenetic alteration and thermal maturity, we have also calculated the Carbon Preference Index ($\text{CPI} = 1/2 \cdot \frac{\sum(X_i + X_{i+2} + \dots + X_n)}{\sum(X_{i-1} + X_{i+1} + \dots + X_{n-1})} + 0.5 \cdot \frac{\sum(X_i + X_{i+2} + \dots + X_n)}{\sum(X_{i+1} + X_{i+3} + \dots + X_{n+1})}$, with $i = 25$ and $n = 35$) (Bray and Evans, 1961; Peters et al., 2005) (Table 3a and Fig. 6). CPIs are typically high in leaf waxes

(>5; e.g. Eglinton and Hamilton, 1963; Mazurek and Simoneit, 1984), decrease in soils and sediments to values between 2 and 5 (Kennicutt et al., 1987) due to diagenetic alteration, and approach 1 in thermally mature rocks (e.g. Gearing et al., 1976; Pendoley, 1992). Here, CPIs are all slightly above 1 and vary through the section (Fig. 6). However, they are particularly variable in the gray calcareous marlstones (unit a), and in the boundary clay layer (unit c) from the lowermost Danian, with values as high as 1.6, confirming at least a partial higher plant origin. In

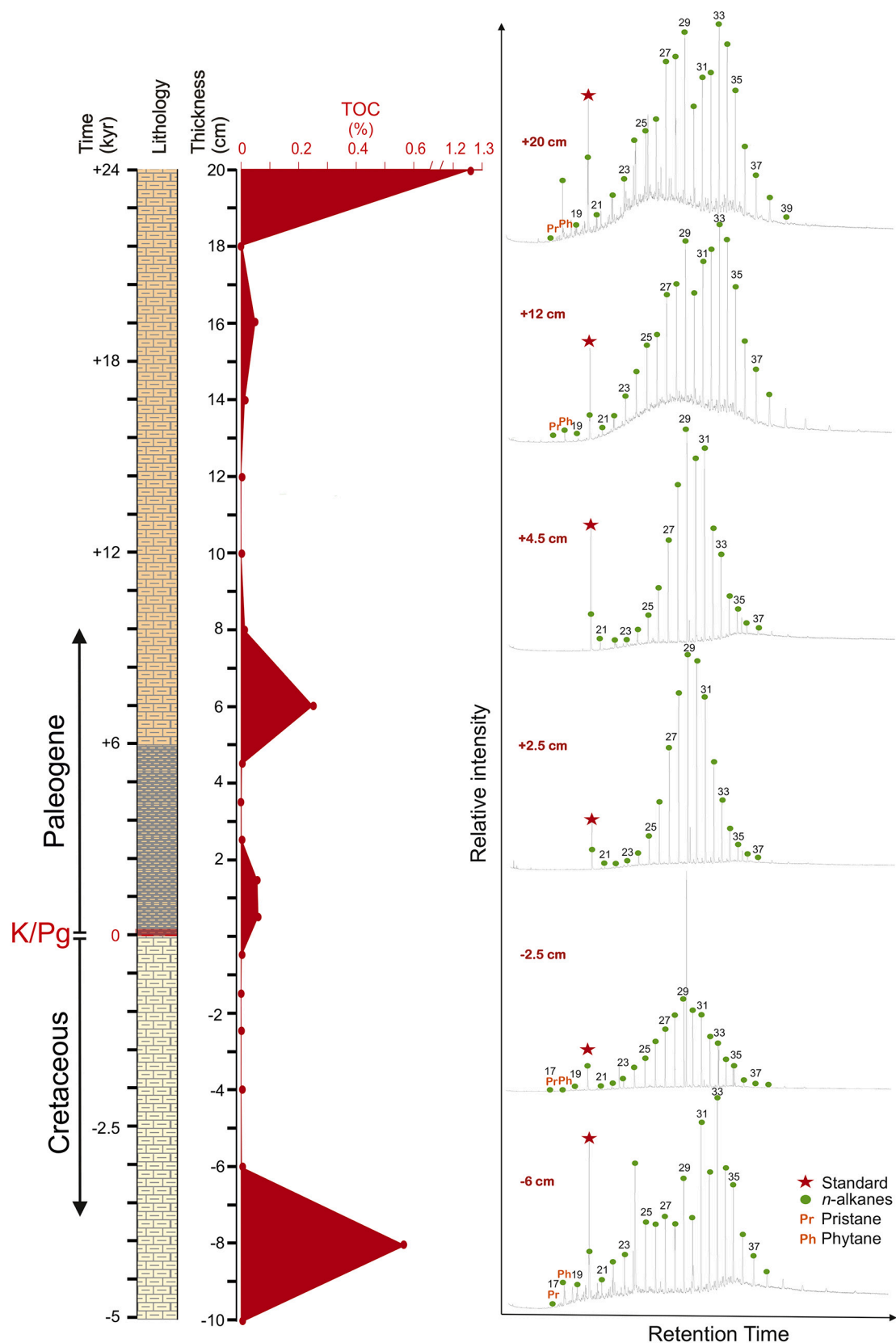


Fig. 4. Total organic carbon (TOC, %) profile from the KPgB at Agost section alongside six partial gas chromatograms derived from GC-MS total ion currents (TIC); the selected traces are representative of units a, c and d described above, with the main biomarkers identified. For lithology, see caption in Fig. 2.

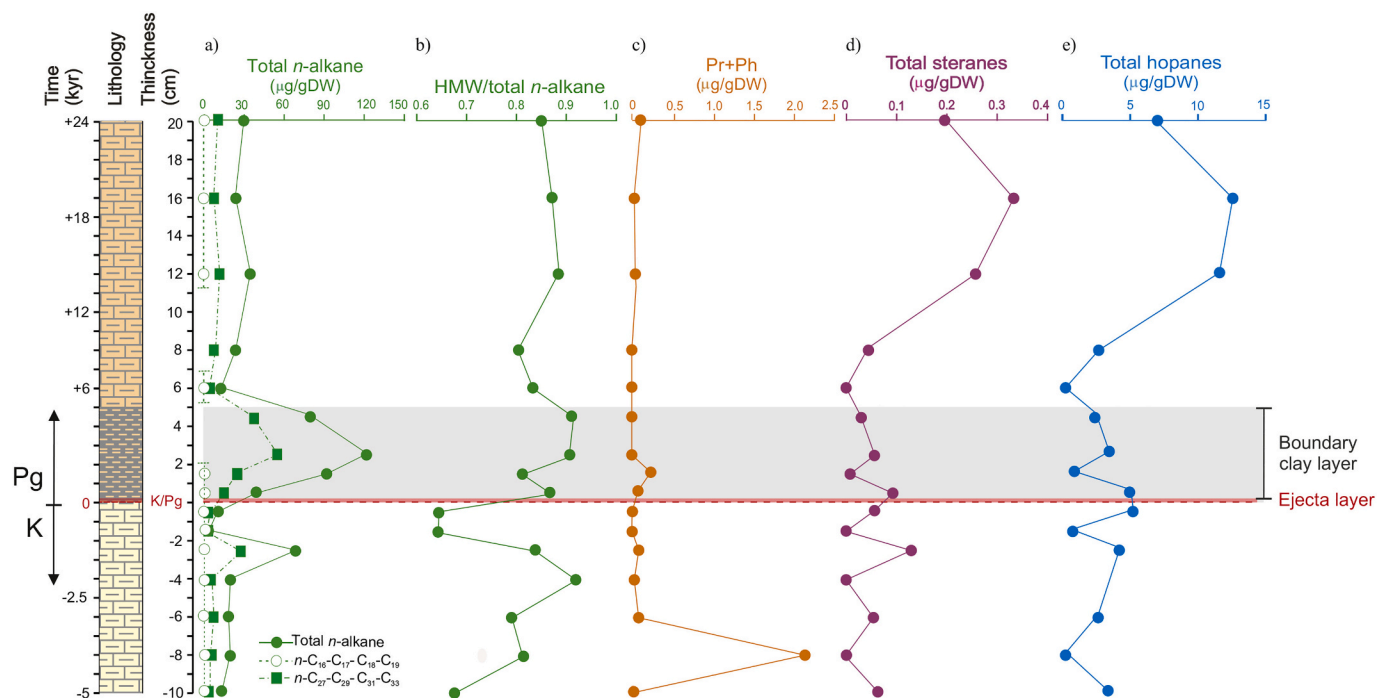


Fig. 5. Abundances of main biomarkers through the KPgB at Agost section. Concentration in µg/g dry weight (DW) of a) total *n*-alkane [$\Sigma(n\text{-}C_{16} + \dots + n\text{-}C_{40})$], b) HMW/ Σ total *n*-alkane ratio [$\Sigma(n\text{-}C_{27} + \dots + n\text{-}C_{40}) / \Sigma(n\text{-}C_{16} + \dots + n\text{-}C_{40})$] (green colour), c) acyclic isoprenoids concentration, (Pr + Ph; orange colour), d) total sterane concentration [$\Sigma(C_{27} + C_{28} + C_{29})$] (pink colour), and e) total hopane concentrations [$\Sigma(C_{27} + \dots + C_{35})$] (blue colour). For lithology details see caption in Fig. 2. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

Table 3a

Main *n*-alkanes and acyclic isoprenoids ratios across the KPgB at Agost section. Acyclic isoprenoids concentration, (Pr + Ph) is in µg/g dry weight (DW).

| Samples | Distance KPgB (cm) | Unit | <i>n</i> -alkanes | | | | | Acyclic isoprenoids | | | |
|--------------------|--------------------|--------|--------------------|-------------------------------|------------------|---------------------------|---|---------------------|-------------------------------|-------------------------------|---------|
| | | | Range ^a | C _{max} ^b | ACL ^c | 31/(29 + 31) ^d | HMW/ Σ total <i>n</i> -alkane ^e | CPI ^f | Pr/ <i>n</i> -C ₁₇ | Ph/ <i>n</i> -C ₁₈ | Pr + Ph |
| 22 Ag3 MI 30–32 | 20.0 | Unit d | 16–38 | 33 | 31 | 0.42 | 0.81 | 1.31 | 0.27 | 0.47 | 0.13 |
| 20 Ag3 MI 26–28 | 16.0 | | 16–40 | 33 | 31 | 0.49 | 0.84 | 1.28 | 0.15 | 0.35 | 0.04 |
| 18 Ag3 MI 22–24 | 12.0 | | 17–39 | 33 | 31 | 0.48 | 0.86 | 1.25 | 0.25 | 0.58 | 0.06 |
| 16 Ag3 MI 18–20 | 8.0 | | 20–38 | 31 | 30 | 0.53 | 0.76 | 1.27 | – | – | – |
| 15 Ag3 MI 16–18 | 6.0 | Unit c | 19–37 | 34 | 31 | 0.57 | 0.79 | 1.14 | – | – | – |
| 14 Ag3 MI 15–16 | 4.5 | | 20–37 | 29 | 30 | 0.47 | 0.89 | 1.15 | – | – | – |
| 12 Ag3 MI 13–14 | 2.5 | | 20–37 | 29 | 30 | 0.42 | 0.89 | 1.07 | – | – | – |
| 11 Ag3 MI 12–13 | 1.5 | | 18–40 | 33 | 31 | 0.56 | 0.76 | 1.48 | – | 0.29 | 0.23 |
| 10 Ag3 MI 10–12 Pg | 0.5 | Unit a | 17–38 | 33 | 31 | 0.51 | 0.84 | 1.18 | 0.20 | 0.57 | 0.10 |
| 7 Ag3 MI 10–12 K | –0.5 | | 17–37 | 31 | 30 | 0.53 | 0.55 | 1.56 | 0.19 | 0.41 | 0.02 |
| 6 Ag3 MI 9–10 | –1.5 | | 17–34 | 24 | 30 | 0.56 | 0.56 | 1.39 | 0.53 | 0.26 | 0.03 |
| 5 Ag3 MI 8–9 | –2.5 | | 17–38 | 29 | 30 | 0.47 | 0.80 | 1.15 | 0.34 | 0.40 | 0.12 |
| 4 Ag3 MI 6–8 | –4.0 | | 18–39 | 34 | 32 | 0.60 | 0.90 | 1.22 | – | 2.85 | 0.05 |
| 3 Ag3 MI 4–6 | –6.0 | | 17–38 | 33 | 31 | 0.59 | 0.74 | 1.55 | 0.43 | 0.57 | 0.12 |
| 2 Ag3 MI 2–4 | –8.0 | | 17–39 | 34 | 31 | 0.58 | 0.77 | 1.36 | 0.98 | 2.04 | 2.14 |
| 1 Ag3 MI 0–2 | –10.0 | | 18–39 | 33 | 30 | 0.59 | 0.60 | 1.29 | – | 0.27 | 0.02 |

^a The detected carbon number range of *n*-alkanes.

^b Carbon number of the homologue with highest abundance.

^c Average chain length $ACL = \Sigma(i \cdot X_i) / \Sigma X_i$, where *X* is abundance and *i* ranges from C₂₅ to C₃₅.

^d Ratio of the *n*-C₃₁ to the sum of the *n*-C₂₉ and *n*-C₃₁ alkane.

^e $HMW / \Sigma \text{total } n\text{-alkanes} = \Sigma(C_{27} + \dots + C_{40}) / \Sigma(C_{16} + \dots + C_{40})$. ^f Carbon Preference Index $CPI = 1/2 \cdot \Sigma(X_i + X_{i+2} + \dots + X_n) / \Sigma(X_{i-1} + X_{i+1} + \dots + X_{n-1}) + 0.5 \cdot \Sigma(X_i + X_{i+2} + \dots + X_n) / \Sigma(X_{i+1} + X_{i+3} + \dots + X_{n+1})$, with *i* = 25 and *n* = 35.

the Danian, from 8 cm above the KPgB and within the light marly limestones (unit d), CPIs have stable and low values of ~1.3.

4.4.2. Acyclic isoprenoids

The acyclic isoprenoids, pristane (Pr) and phytane (Ph), derive from multiple sources but in this setting, most likely from aquatic phytoplankton via the degradation of the phytol side-chain of chlorophyll (Peters et al., 2005). Their abundances vary dramatically but non-systematically through the section, with some intervals where they are

absent or below detection limits (Table 3a). Maximum abundances (µg/g DW) of summed Pr + Ph occur at 8 and 2.5 cm below and at 1.5 cm above the KPgB, respectively (Fig. 5); lowest abundances generally occur above the KPgB, with neither compound found between 1.5 and 12 cm. The Pr/*n*-C₁₇ ratios range between 0 and 1.5, and the Ph/*n*-C₁₈ ratios range between 0.3 and 2.9, with highest values generally in unit a (Fig. 7). These variations largely reflect the variations in *n*-alkane abundances (see above). The overall low abundances of these acyclic isoprenoids could be consistent with a predominantly terrigenous source

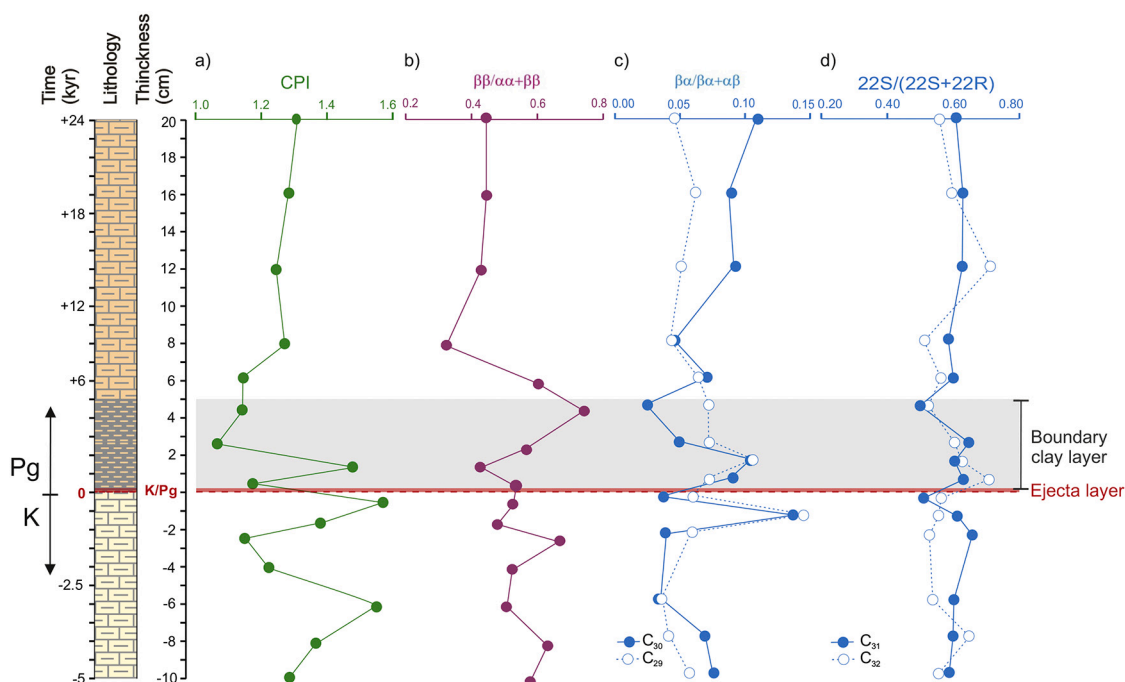


Fig. 6. Biomarker ratios indicative of thermal maturity or other chemical alteration. a) CPI from n -alkane $[1/2 * \Sigma(X_i + X_{i+2} + \dots + X_n) / \Sigma(X_{i+1} + X_{i+3} + \dots + X_{n-1}) + 0.5 * \Sigma(X_i + X_{i+2} + \dots + X_n) / \Sigma(X_{i+1} + X_{i+3} + \dots + X_{n-1})]$, with $i = 25$ and $n = 35$ (green colour), b) $C_{29} \beta / (\alpha + \beta)$ sterane ratio (pink colour), c) C_{31} and $C_{32} 22S / (22S + 22R)$, and d) $\beta / (\beta + \alpha)$ hopane ratios (blue colour). For lithology, see caption in Fig. 2. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

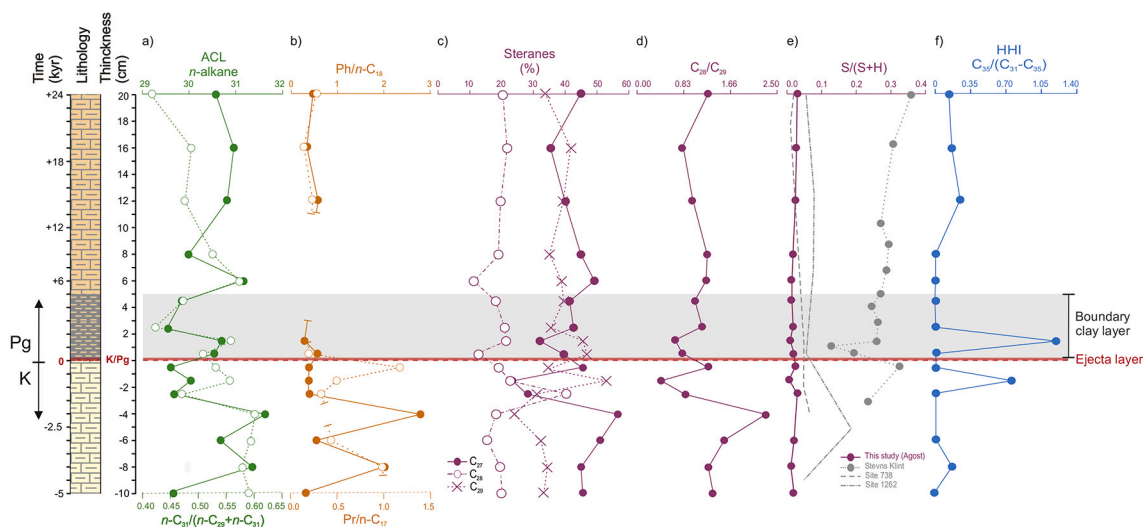


Fig. 7. Biomarker ratios indicative of source changes in the KPGB at Agost section.

a) ACL (Average Chain Length) from n -alkane $[\Sigma(i * X_i) / \Sigma X_i]$ where X is abundance and i ranges from C_{25} to C_{35} (continuous green line) and $n-C_{31} / [n-C_{29} + n-C_{31}]$ ratio (dashed green line), b) $Pr/n-C_{17}$ (dashed orange line) and $Ph/n-C_{18}$ (continuous orange line) ratios, c) % of each regular sterane $[C_{27} / \Sigma(C_{27} + C_{28} + C_{29}); C_{28} / \Sigma(C_{27} + C_{28} + C_{29}); C_{29} / \Sigma(C_{27} + C_{28} + C_{29})]$, d) C_{28} / C_{29} sterane ratio, e) sterane/hopane ratio $(S/S + H) [\Sigma(C_{27} + C_{28} + C_{29}) / (C_{27} + C_{28} + C_{29}) + (\Sigma C_{27} + \dots + C_{35})]$, pink colour for the Agost section and gray colours for Stevn Klint section (data from Sepúlveda et al., 2009) and sites 738 and 1262 (data from Bralower et al., 2020), and f) Homohopane indices (HHI) $[C_{35} / (C_{31} - C_{35})]$ (blue colour). For lithology, see caption in Fig. 2. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

of the organic matter throughout the section and preclude the use of classical redox proxies based on the Pr/Ph ratio (Didyk et al., 1978; Ten Haven et al., 1987).

4.4.3. Steranes

Steranes (C_{27} – C_{29}) are tetracyclic triterpanes arising from diagenetic alteration of sterols and consequently derive almost exclusively from

eukaryotic organisms (e.g. Peters et al., 2005; Schwark and Empt, 2006 and references therein), particularly marine phytoplankton (Volkman et al., 1998). The predominant three regular steranes, C_{27} -cholestane, C_{28} -ergostane and C_{29} -stigmastane, are present in all samples but in highly variable concentrations (Table 3b and Fig. 5). They are represented by the four epimers ($5\alpha(H), 14\alpha(H), 17\alpha(H), 20R$; $5\alpha(H), 14\alpha(H), 17\alpha(H), 20S$; $5\alpha(H), 14\beta(H), 17\beta(H), 20S$ and $5\alpha(H), 14\beta(H), 17\beta(H), 20R$).

Table 3b

Sterane ratios across the KPgB interval at Agost section.

| Samples | Distance KPgB (cm) | Unit | Steranes | | Presence of Diasterane | C_{28}/C_{29} | $S/(S + H)^c$ | n -alkane/ S^d | C_{27}/C_{29} $\beta\beta/\alpha\alpha + \beta\beta$ | C_{27}/C_{29} $20S/20S + 20R$ |
|--------------------|-----------------------|--------|--------------------|-------------|------------------------|-----------------|---------------|--------------------|---|------------------------------------|
| | | | Range ^a | C_{max}^b | | | | | | |
| 22 Ag3 MI 30–32 | 20.0 | Unit d | 27–29 | 27 | Yes | 1.33 | 0.03 | 153.74 | 1.3 | 1.5 |
| 20 Ag3 MI 26–28 | 16.0 | | 27–29 | 29 | Yes | 0.85 | 0.03 | 74.61 | 1.4 | 1.5 |
| 18 Ag3 MI 22–24 | 12.0 | | 27–29 | 27 | Yes | 1.02 | 0.02 | 135.94 | 1.5 | 1.6 |
| 16 Ag3 MI 18–20 | 8.0 | | 27–29 | 27 | Yes | 1.28 | 0.02 | 532.73 | 1.1 | 1.3 |
| 15 Ag3 MI 16–18 | 6.0 | Unit c | 27–29 | 27 | Yes | 1.26 | 0.01 | 2989.63 | 1.5 | 1.0 |
| 14 Ag3 MI 15–16 | 4.5 | | 27–29 | 27 | Yes | 1.05 | 0.01 | 2588.11 | 1.4 | 1.3 |
| 12 Ag3 MI 13–14 | 2.5 | | 27–29 | 27 | Yes | 1.20 | 0.02 | 2113.63 | 1.2 | 1.3 |
| 11 Ag3 MI 12–13 | 1.5 | | 27–29 | 29 | Yes | 0.71 | 0.01 | 9190.13 | 1.3 | 1.5 |
| 10 Ag3 MI 10–12 Pg | 0.5 | Unit a | 27–29 | 29 | Yes | 0.85 | 0.02 | 427.31 | 1.3 | 1.3 |
| 7 Ag3 MI 10–12 K | –0.5 | | 27–29 | 27 | Yes | 1.30 | 0.02 | 92.05 | 1.3 | 1.5 |
| 6 Ag3 MI 9–10 | –1.5 | | 27–29 | 29 | Yes | 0.45 | 0.01 | 451.18 | 1.3 | 1.5 |
| 5 Ag3 MI 8–9 | –2.5 | | 27–29 | 28 | Yes | 0.90 | 0.03 | 535.86 | 1.4 | 0.8 |
| 4 Ag3 MI 6–8 | –4.0 | | 27–29 | 27 | Yes | 2.33 | – | 12,217.35 | 1.6 | 1.2 |
| 3 Ag3 MI 4–6 | –6.0 | | 27–29 | 27 | Yes | 1.57 | 0.02 | 331.18 | 1.3 | 1.4 |
| 2 Ag3 MI 2–4 | –8.0 | | 27–29 | 27 | Yes | 1.31 | 0.01 | 6000.75 | 1.5 | 1.0 |
| 1 Ag3 MI 0–2 | –10.0 | | 27–29 | 27 | Yes | 1.36 | 0.02 | 212.07 | 1.4 | 1.3 |

^a The detected carbon number range of steranes.^b Carbon number with the highest abundance.^c The sterane/(sterane + hopane) ratio $[(\Sigma C_{27} + C_{28} + C_{29})/(\Sigma C_{27} + C_{28} + C_{29}) + (\Sigma C_{27} + \dots + C_{35})]$.^d The n -alkane ratio/sterane $[(\Sigma n\text{-alkane})/(\Sigma C_{27} + C_{28} + C_{29})]$.

(H),20R). The C_{30} regular sterane has also been identified in most of the samples. Also present are the $13\beta,17\alpha$ steranes or diasteranes (Table 3b). Collectively, the sterane and diasterane distributions are consistent with a relatively high thermal maturity, near the onset of the oil window (Table 3c).

Regular sterane (C_{27} , C_{28} and C_{29}) distributions vary significantly through the section. The most abundant sterane in most Maastrichtian sediments is the C_{27} component; across the KPgB, at 1.5 cm below the

KPgB and between 0.5 and 1.5 cm above it, the C_{29} sterane dominates; in the uppermost part of the section, the C_{27} and C_{29} components occur in similar abundances (Fig. 7). Relative abundances of the C_{28} sterane are more stable, reflected both by its overall proportion as well as invariant C_{28}/C_{29} sterane ratios (barring a value of 2.5 below the KPgB; Fig. 7). Sterane abundances also vary relative to those of bacterially-derived hopanes, such that the sterane/(sterane+hopane) ratio $[S/(S + H)]$, i.e. $\Sigma C_{27} + C_{28} + C_{29}$ steranes $/(\Sigma C_{27} + C_{28} + C_{29}$ steranes) $+ (\Sigma C_{27} + \dots +$

Table 3c

Hopanes ratios across the KPgB interval at Agost section.

| Samples | Distance KPgB (cm) | Unit | Hopanes | | ACL ^b | Ts/ Tm | C_{29}/C_{30} | C_{29} $\beta\alpha/(\alpha\beta + \beta\alpha)$ | C_{30} $\beta\alpha/(\alpha\beta + \beta\alpha)$ | C_{31} $22S/(22S + 22R)$ | C_{32} $22S/(22S + 22R)$ | HHI ^c $C_{35}/(C_{31} - C_{35})$ |
|--------------------|-----------------------|--------|--------------------|-------------------------|------------------|-----------|-----------------|---|---|-------------------------------|-------------------------------|---|
| | | | Range ^a | Presence of Gammacerane | | | | | | | | |
| 22 Ag3 MI 30–32 | 20.0 | Unit d | 27–35 | – | 31 | 0.55 | 1.27 | 0.05 | 0.11 | 0.55 | 0.48 | 0.17 |
| 20 Ag3 MI 26–28 | 16.0 | | 27–35 | – | 30 | 0.98 | 1.32 | 0.06 | 0.09 | 0.58 | 0.54 | 0.21 |
| 18 Ag3 MI 22–24 | 12.0 | | 27–35 | – | 30 | 0.87 | 1.32 | 0.05 | 0.09 | 0.58 | 0.69 | 0.30 |
| 16 Ag3 MI 18–20 | 8.0 | | 27–33 | – | 30 | 1.11 | 1.11 | 0.04 | 0.05 | 0.52 | 0.42 | 0.00 |
| 15 Ag3 MI 16–18 | 6.0 | Unit c | 27–33 | – | 30 | 1.25 | 1.39 | 0.06 | 0.07 | 0.54 | 0.49 | 0.00 |
| 14 Ag3 MI 15–16 | 4.5 | | 27–33 | – | 30 | 1.04 | 1.27 | 0.07 | 0.02 | 0.41 | 0.44 | 0.00 |
| 12 Ag3 MI 13–14 | 2.5 | | 27–33 | – | 30 | 0.63 | 1.29 | 0.07 | 0.05 | 0.61 | 0.55 | 0.00 |
| 11 Ag3 MI 12–13 | 1.5 | | 27–35 | – | 31 | 1.08 | 0.77 | 0.11 | 0.10 | 0.55 | 0.58 | 1.52 |
| 10 Ag3 MI 10–12 Pg | 0.5 | Unit a | 27–34 | Yes | 30 | 1.05 | 1.18 | 0.07 | 0.09 | 0.58 | 0.69 | 0.00 |
| 7 Ag3 MI 10–12 K | –0.5 | | 27–32 | Yes | 30 | 1.05 | 0.98 | 0.06 | 0.04 | 0.42 | 0.49 | 0.00 |
| 6 Ag3 MI 9–10 | –1.5 | | 27–35 | – | 31 | 1.28 | 0.66 | 0.15 | 0.14 | 0.56 | 0.48 | 0.76 |
| 5 Ag3 MI 8–9 | –2.5 | | 27–33 | – | 30 | 0.54 | 0.99 | 0.06 | 0.04 | 0.62 | 0.44 | 0.00 |
| 4 Ag3 MI 6–8 | –4.0 | | – | – | – | – | – | – | – | – | – | – |
| 3 Ag3 MI 4–6 | –6.0 | | 27–33 | – | 30 | 1.59 | 1.44 | 0.04 | 0.04 | 0.54 | 0.46 | 0.00 |
| 2 Ag3 MI 2–4 | –8.0 | | 27–35 | – | 31 | 0.81 | 1.11 | 0.04 | 0.07 | 0.54 | 0.61 | 0.19 |
| 1 Ag3 MI 0–2 | –10.0 | | 27–33 | – | 30 | 1.02 | 1.13 | 0.06 | 0.08 | 0.52 | 0.48 | 0.00 |

^a The detected carbon number range of hopanes.^b $ACL = \Sigma(i \cdot X_i) / \Sigma X_i$, where X is abundance and i ranges from C_{27} to C_{35} .^c The homohopane indices (HHI) $[C_{35}/(C_{31} - C_{35})]$.

C₃₅ hopanes)], calculated from the *m/z* 217 and 191 traces, respectively, is low in Maastrichtian sediments, variable across the KPgB and relatively high in the uppermost part of the section (Fig. 7).

Intriguingly, sterane-derived thermal maturity parameters also differ between the lower and upper parts of the section. The C₂₉ $\beta\beta/(\alpha\alpha + \beta\beta)$ ratio, which increases with thermal maturity (Mackenzie et al., 1980), is generally higher in the lower part of unit a (Maastrichtian), variable across the boundary layer, high again at the top of unit c (clay boundary layer), and then decreases at the base of unit d to relatively low and stable values of 0.41 upwards (Danian) (Table 3b and Fig. 6). Given the thickness of this section, it is unlikely that these values reflect thermal maturity variations of the rocks themselves; below in the discussion, changes in OM source input are considered as an alternative mechanism.

4.4.4. Hopanes

Hopanes are derived from functionalized bacteriohopanoids that are common in a wide range of predominantly but not exclusively aerobic bacteria (Belin et al., 2018; Rohmer et al., 1984). The hopanes and homohopanes (>C₃₀ hopanes) range in carbon number from C₂₇ to C₃₅ and occur in almost all of the samples across the KPgB (Table 3c and Fig. 5). Hopanes were not detected at 4.0 cm below the KPgB, and C₃₄-C₃₅ homohopanes were absent/below detection in about half of the samples (Table 3c). Therefore, homohopane indices (HHI) [C₃₅/(C₃₁-C₃₅)], which reflect the increased preservation of extended hopanes (>C₃₃) under anoxic conditions (Bishop and Farrimond, 1995; Peters et al., 2005), are generally very low. However, minor peaks in HHIs do occur just above and below the KPgB (Fig. 7). Intriguingly, the distribution of hopanes varies through the section, such that the <C₃₁ hopanes are dominant before the event and >C₃₀ hopanes are dominant afterwards (Table 3c). The 17 α (H)-22,29,30-trisnorhopanes (Tm) and 18 α (H)-22,29,30-trisnorhopanes (Ts) are present in all samples, with a Ts/Tm ratio ranging between 0.54 and 1.3, with a mean value of 1.1 from the upper Maastrichtian sediments and 0.95 from the lower Danian sediments (Table 3c). Gammacerane (C₃₀ triterpane) was identified only in two samples around the ejecta layer, at 0.5 cm below and above the KPgB (Table 3c), but only at low concentrations close to the detection limit.

Several hopane ratios predominantly (but not exclusively) reflect thermal maturity (Mackenzie et al., 1980; Seifert and Moldowan, 1980). The C₂₉ and C₃₀ 17 β ,21 α (H)/(17 β ,21 α (H) + 17 α ,21 β (H)) hopane ratios—hereafter $\beta\alpha/(\beta\alpha + \alpha\beta)$ ratios (or moretane/hopane ratio)—largely covary between 0.02 and 0.11 (Table 3c and Fig. 6), and the C₃₁ and C₃₂ 22S/(22S + 22R) 17 α ,21 β (H)-homohopane ratios range between 0.41 and 0.69 (Table 3c and Fig. 6). Both the low $\beta\alpha/(\beta\alpha + \alpha\beta)$ hopane ratios and near-equilibrium 22S/(22S + 22R) ratios indicate a relatively high thermal maturity close to oil generation (equivalent vitrinite reflectance at or near 0.6, Peters et al., 2005). This is also consistent with the absence of the $\beta\beta$ hopane configuration. However, there are large variations in all ratios, and especially the $\beta\alpha/(\beta\alpha + \alpha\beta)$ hopane ratios; as with steranes, this variability is not thought to reflect changes in the thermal history of these rocks.

5. Discussion

The KPgB mass extinction strongly impacted phytoplankton and calcareous nannoplankton (Alvarez et al., 2019; Lowery et al., 2020), but the response of non-fossilizing plankton remains less clear (Lowery et al., 2020). Until now, there are few studies based on biomarkers (Bralower et al., 2020; Schaefer et al., 2020; Sepúlveda et al., 2009; Taylor et al., 2018). Here, we first integrate new biomarker data with some elemental ratios to evaluate changes in environmental conditions. We then use specific biomarkers to characterize changes in the OM assemblages along the continuous 32 cm sequence of the KPgB in the Agost distal section (SE Spain), and compare them to those environmental changes. We further evaluate the relative contribution of terrigenous, algal and bacterial OM as well as reworked OM (the latter inferred from

changes in the thermal maturity of the biomarkers). Then we examine hopane and sterane distributions in greater detail, using those to explore the recovery of primary producers, mainly bacterial and algal communities, after the KPgB extinction.

5.1. Changes in environmental conditions

The elemental analysis performed at the Agost section confirms the preservation of the ejecta layer in our study section, defined by a peak in Cr/Al, Ni/Al, Cu/Al, and Zn/Al ratios related to extraterrestrial input. However, the ejecta layer also exhibits peaks in others ratios such as Mo/Al and U/Al, widely considered as proxies for paleoredox conditions (Algeo and Tribovillard, 2009; Tribovillard et al., 2012). These ratios indicate that changes in the oxygen water conditions occurred at the KPg impact (detected only in the ejecta layer), although minor variations in both profiles also occur elsewhere in the section (Fig. 2). In the nearby Caravaca section, Sosa-Montes de Oca et al. (2013) showed that depositional conditions for the ejecta layer were anoxic, inferred to be a consequence of the enhanced contribution of metals to the basins as well as an increased supply of terrestrial and marine organic material. However, shortly after the impact, oxygen levels rapidly recovered, favoring the earliest macrobenthic opportunist colonization (Rodríguez-Tovar, 2005).

At the Agost site, the enrichment factor (EF) of selected trace elements, in particular uranium and molybdenum, which are enriched in anoxic conditions, U_{EF} vs Mo_{EF} (Fig. 8), indicates a change in redox conditions immediately following the KPgB but without anoxic conditions ever being reached (Fig. 8). This is likely due to the Agost section (upper-bathyal) being shallower than the Caravaca section (bathyal) (Schulte et al., 2010). Furthermore, Mo increases from 8 cm onwards above the KPgB, indicating further minor changes in redox conditions. We do not have biomarker data from the thin ejecta layer, but biomarkers in sub- and super-adjacent layers are consistent with this being the interval with the most reducing conditions. Gammacerane, thought to derive from bacteriovorous ciliates and an indicator of anoxia (Peters et al., 2005; Sinninghe-Damste et al., 1995), is present only in rocks 0.5 cm below and above the KPgB (Table 3c). Similarly, the redox sensitive homohopane indices (HHI = [C₃₅/(C₃₁-C₃₅)]; Bishop and Farrimond, 1995; Peters et al., 2005) are generally low, but two maxima occur at 2.5 below and 1.5 cm above the KPgB (Fig. 7), also suggesting episodic anoxia or dysoxia in the bottom waters in the latest Cretaceous and early Danian after deposition of the ejecta layer.

5.2. Variability in organic matter input across the KPgB event

The *n*-alkanes primarily derive from leaf waxes of vascular plants, and this is especially the case for the long-chain *n*-alkanes (>C₂₅) with an odd carbon number (e.g. Eglinton and Hamilton, 1967; Eglinton and Eglinton, 2008). However, *n*-alkyl moieties also occur in various macromolecules, including algaenan, cutin, cutan and suberin (e.g. (Gelin et al., 1999, 1998; Tegelaar et al., 1989). Therefore, HMW components in geological materials, especially when the odd-over-predominance has been largely lost, can derive from a range of sources (Eglinton and Eglinton, 2008). Conversely, short-chain *n*-alkanes (C₁₅-C₁₉) are thought to mainly derive from aquatic sources, including algae and photosynthetic bacteria (Cranwell et al., 1987; Gelpi et al., 1970); however, like their long-chain homologues, in thermally mature sediments they also derive from the diagenetic and thermal degradation of a wide range of biological sources. Consequently, because of its high thermal maturity, interpretation of *n*-alkane distributions in Agost sediments is challenging. Nonetheless, the dominance of >C₂₃ components (and high HMW/ Σ total *n*-alkane ratios >0.7) throughout the sequence, as well as very low abundances of Pr and Ph (Fig. 5), suggest that terrigenous inputs were relatively dominant. This could be particularly true in the lower part of the section, Maastrichtian (unit a) and lowermost boundary clay layer (unit c), where an elevated CPI above 1.4 persists in some

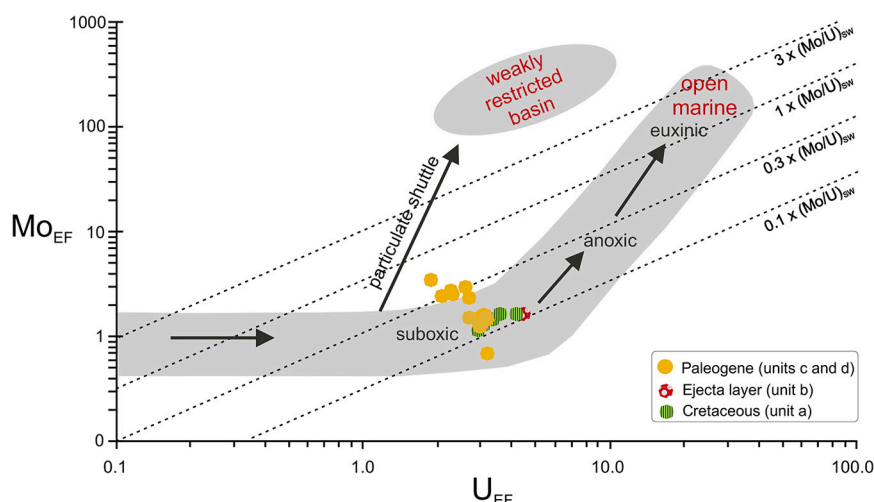


Fig. 8. Mo_{EF} vs U_{EF} covariation for the KPgB at Agost section. The uppermost Maastrichtian sediments (unit a) are indicated by green circles with vertical stripes, the ejecta layer (unit b) is noted by red circles with dots and the lowermost Danian sediments (units c and d) by yellow circles. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

horizons (Fig. 6). The HMW/Σ total n -alkane ratio, however, does not show similar trends, with higher values mostly associated with the upper part of the section, from 2.5 to 20 cm above the KPgB (Fig. 5). Consequently, it appears that higher plant inputs were variable and complex, and likely overprinted by inputs from more thermally mature sources.

Further evidence for erosive change comes from the unexpected variations in biomarker thermal maturity parameters, especially in a ~ 32 cm-thick section with no differences in burial history. The lower values are likely representative of the history of the basin, whereas higher values could be evidence of reworked allochthonous organic matter (e.g. Carmichael et al., 2018; Castro et al., 2019; Hackley et al., 2020; van Breugel et al., 2007). Thermal maturity parameters are also impacted by lithology (Peters et al., 2005; Peters and Moldowan, 1991), but we discount that explanation given the minor lithological variations in this section and their lack of correspondence with biomarker parameters. In fact, the biomarker maturity parameters exhibit rather inconsistent trends. The sterane $\beta\beta/(\alpha\alpha + \beta\beta)$ ratio is highest in the lower part of the section, but the hopane $22S/(22S + 22R)$ is lowest in isolated horizons between 2 cm below and 8 cm above the KPgB (Fig. 6). In contrast, the $\beta\alpha/(\beta\alpha + \alpha\beta)$ hopane ratio, which decreases with thermal maturity, is highest at 1.5 cm below and above the KPgB (Fig. 6). We note, however, that high moretane/hopane ratios are also associated with terrestrial inputs (Xie et al., 2007). Thus, it appears that the interval immediately preceding the KPgB through to the boundary clay layer is characterized by highly variable terrestrial inputs, from both higher plants and reworked organic matter in older sedimentary rocks.

Consequently, the proportion of putative aquatic organic matter inputs also appears to vary through the section. Both $Pr/n-C_{17}$ and $Ph/n-C_{18}$ ratios are higher and more variable prior to the KPgB than after it, perhaps suggesting decreased marine productivity (Fig. 7). The sterane/hopane ratio $[S/(S + H)]$ exhibits a more complex variation, varying in the pre-KPgB horizons, decreasing in the boundary layer and stabilizing at maximum values in the uppermost part of the section (unit d) (Fig. 7).

These changes in the OM assemblage likely impacted the bulk organic matter $\delta^{13}C$ record. Negative carbon isotope excursions (-CIEs) in both $\delta^{13}C_{carb}$ and $\delta^{13}C_{org}$ are typical in immediate post-KPgB intervals (Birch et al., 2016; Sepúlveda et al., 2019). Generally, at those sites, the -CIEs in $\delta^{13}C_{org}$ are smaller in amplitude and shorter in duration than those in $\delta^{13}C_{carb}$. However, in the Agost section an unexpected +CIE occurs in the $\delta^{13}C_{org}$ record (but not in the $\delta^{13}C_{carb}$; Fig. 3), and this was also observed by Sepúlveda et al. (2019). Like Sepúlveda et al. (2019) we attribute this to complex post-KPgB inputs of ^{13}C -enriched organic matter, possibly from older reworked organic material being eroded and

deposited on the basin. Similar processes could also explain why $\delta^{13}C_{org}$ records at other sites exhibit smaller -CIEs than corresponding $\delta^{13}C_{carb}$.

5.3. Changes in algal and bacterial communities across KPgB

Eukaryote and bacterial biomarkers are present in the Agost section and persist through the event and afterwards. In fact, steranes, hopanes, and n -alkanes occur in every sample. There is some evidence for decreased algal production from the decrease in $CaCO_3$ contents, decrease in sterane/hopane ratios and the absence of detectable pristane and phytane in post-KPgB sediments; however, pristane and phytane are also absent from some pre-KPgB sediments and are abundant in samples in unit d. Collectively, these data are consistent with previous biomarker studies wherein disruption of primary production was relatively transient (Bralower et al., 2020; Schaefer et al., 2020; Sepúlveda et al., 2009). This can be further explored via distributions of those compounds.

Changes in the relative contribution of terrigenous OM have implications for the interpretation of sterane biomarker ratios as strictly algal biomarkers, but this does not seem to be the case in the Agost section. For example, we expect elevated higher plant inputs to be associated with elevated proportions of the C_{29} sterane. However, the opposite is observed, i.e. the highest HMW/Σ total n -alkane ratio is associated with high C_{27} sterane proportions (Fig. 5). Nor do sterane proportions match putative changes in thermal maturity. Therefore, variations in the distribution of steranes and hopanes of different carbon number appear to reflect changes in phytoplankton and bacterial sources, respectively.

Despite the persistence of life across the KPgB, variability in OM inputs does occur across the KPgB at the Agost section. Sterane distributions switch from an assemblage dominated by C_{27} prior to the event to an assemblage dominated by both C_{27} and C_{29} post-KPgB (Fig. 7). There is little variation in C_{28}/C_{29} sterane ratios, however, similar to observations from the Fish Clay in Denmark (Sepúlveda et al., 2009). Although hopane distributions are different before and after the ejecta layer deposition, the dominant hopane is $>C_{31}$ prior the KPgB and $<C_{30}$ afterwards (Table 3c). Therefore, diverse biomarkers persist through the section and their abundances rapidly recover by unit d, but the organic matter assemblage in Danian sediments overlying the ejecta layer is different than that of the late Maastrichtian.

The sterane/hopane ratios $[S/(S + H)]$ vary minimally, by about a factor of 3, with highest ratios in the uppermost parts of unit a and unit d, i.e. the latest Maastrichtian and the Danian post-boundary layer interval. We have compared our $[S/(S + H)]$ ratio with that from the Fish

Clay in Denmark (Sepúlveda et al., 2009), although it is located in a different setting with higher sedimentation rates, with both sections showing a minimum value immediately above the KPgB and maximum values in the uppermost part of the section (Fig. 7). In contrast, value variations of the $[S/(S + H)]$ ratio along the KPgB published by Bralower et al. (2020) from Sites 738 (Kerguelen Plateau) and 1262 (Walvis Ridge), are in agreement with values from Agost, since all are distal and bathyal sections, showing a rapid recovery of bacterial and algal production after the impact (Fig. 7). This likely reflects changes in algal inputs, with a transient decrease post-KPgB followed by a rapid recovery. However, recent work in the Chicxulub core (Schaefer et al., 2020) reveals large changes in bacterial assemblages in the immediate KPgB aftermath, such that the observed changes could be related to the distance to the impact site or could reflect more dynamic variation in both algal and bacterial populations.

In summary, and consistent with previous investigations of non-fossilized communities (i.e. Bralower et al., 2020; Sepúlveda et al., 2009), the Agost section exhibits only modest evidence for a decrease in aquatic productivity in the aftermath of the KPgB. This is reflected by a decline in sterane to hopane ratios and isoprenoid to *n*-alkane ratios, and it is consistent with the decrease in %CaCO₃. Our work, therefore, expands the findings of Sepúlveda et al. (2009) from neritic to bathyal settings. In apparent contrast to the neritic Fish Clay, Agost does record a change in sterane distributions that could document a long-term change in algal assemblages. It is difficult to define exactly what these changes were, but it is possible that either eukaryotic assemblages changed and/or became simpler, hence less variable and more stable. Nonetheless, our results reaffirm the conclusion that the recovery of algal and microbial assemblages occurred relatively rapidly after the KPgB extinction event.

6. Conclusions

We have determined the distribution of selected biomarkers (*n*-alkanes, acyclic isoprenoids, steranes and hopanes), C-isotope profiles, TOC content and major and trace elements across the KPgB in the Agost section (SE Spain), throughout a continuous ~32 cm-thick section extending from 5 kyr before to 24 kyr after the event. U_{EF} vs Mo_{EF} covariation, the presence of gammacerane and elevated homohopane indices suggest more reducing conditions just within the KPgB. Biomarker ratios, including those indicative of higher plants (i.e. HMW/ Σ total *n*-alkane ratios and CPIs) and thermal maturity (but here interpreted as petrogenic input, i.e. $\beta\beta/(\alpha\alpha + \beta\beta)$ sterane and 22S/(22S + 22R) and $\beta\alpha/(\beta\alpha + \alpha\beta)$ hopane ratios), indicate that the interval immediately preceding the KPgB through to the boundary clay layer was characterized by highly variable terrigenous inputs. This appears to have impacted the bulk organic carbon isotopic record, as observed elsewhere. Most significantly, we observe a transient decrease in eukaryotic (algal) inputs in the boundary clay layer, reflected by both lower %CaCO₃ but also by lower sterane/hopane ratios and linear isoprenoid to *n*-alkane ratios. Therefore, we conclude that, at the Agost distal section, algal and bacterial assemblages were affected by the impact event, but these particular biotic changes were minor and transient. Note that many of these biomarker changes were much smaller than those observed for other biotic events in Earth's history. Furthermore, steranes occurred in all samples, and many biomarker ratios returned to pre-KPgB values by the early Danian, i.e. within 10 kyr. Some biomarker ratios, such as the proportions of C₂₇ relative to C₂₉ steranes, remained different post-KPgB providing some evidence for a longer-term impact on biotic communities. However, the Agost section biomarker records generally document the persistence of non-fossilized algae and bacteria and a rapid recovery from the KPgB impact. This relatively muted biotic response is likely because Agost is distant from the impact crater.

Author declaration

We wish to confirm that there are no known conflicts of interest associated with this publication and there has been no significant financial support for this work that could have influenced its outcome.

We confirm that the manuscript has been read and approved by all named authors and that there are no other persons who satisfied the criteria for authorship but are not listed. We further confirm that the order of authors listed in the manuscript has been approved by all of us.

We confirm that we have given due consideration to the protection of intellectual property associated with this work and that there are no impediments to publication, including the timing of publication, with respect to intellectual property. In so doing we confirm that we have followed the regulations of our institutions concerning intellectual property.

We understand that the Corresponding Author is the sole contact for the Editorial process (including Editorial Manager and direct communications with the office). She is responsible for communicating with the other authors about progress, submissions of revisions and final approval of proofs. We confirm that we have provided a current, correct email address which is accessible by the Corresponding Author and which has been configured to accept email from (claudia.sosamontes.deoca@bristol.ac.uk).

Declaration of Competing Interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

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