The Standard Model (part I) Particles, quantum fields, symmetries. Electroweak interactions and their phenomenology



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[https://www.ugr.es/~jillana/SM1.php]





Outline

- 1. Particles, fields and symmetries
 - ▷ Basics: Poincaré symmetry
 - ▷ Particle physics with quantum fields
 - ▷ Global and gauge symmetries
 - Internal symmetries and the gauge principle
 - Quantization of gauge theories
 - Spontaneous Symmetry Breaking

2. The Standard Model

Neutrinos

- > Electroweak interactions: one generation of quarks *or* leptons [strong int. in part II]
- ▷ Electroweak SSB: Higgs sector, gauge boson and fermion masses
- ▷ Additional generations: fermion mixings (quarks *vs* leptons)
- 3. Electroweak phenomenology
 - > Input parameters, experiments, observables, precise predictions, global fits
 - [quark flavour physics in part II]

4. Tools

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1. Particles, fields and symmetries

Why Quantum Field Theory to describe Particle Physics?

- QFT is the (only) way to reconcile Quantum Mechanics and Special Relativity
 - [-] Wave equations (relativistic or not) cannot account for changing # of particles. And the relativistic versions suffer pathologies:
 - * negative probability densities
 - negative-energy solutions
 - violation of causality
 - [+] Quantum fields:
 - provide a natural framework (Fock space of multiparticle states)
 make sense of negative-energy solutions (antiparticles)
 solve causality problem (Feynman propagator)
 explains spin-statistics connection (theorem)
 arguably, solve the wave-particle duality puzzle (no particles, only fields)

Basics: Poincaré symmetry

Guided by symmetry

• **Relativistic fields** are *irreps* of Poincaré group (rotations, boosts, translations)

scalar $\phi(x)$, vector $V_{\mu}(x)$, tensor $h_{\mu\nu}(x)$, ... Weyl $\psi_L(x)$, $\psi_R(x)$; Dirac $\psi(x)$, ...

• Lagrangian densities: local $\mathcal{L}(x) = \mathcal{L}(\phi, \partial_{\mu}\phi)$ (maybe several " ϕ_i ", ψ , V_{μ} , ...) invariant under Poincaré transformations

– e.g. for a free Dirac field $\psi(x)$:

$$\mathcal{L}_0 = \overline{\psi}(\mathrm{i}\partial \!\!\!/ - m)\psi \quad \partial \!\!\!/ \equiv \gamma^\mu \partial_\mu , \quad \overline{\psi} \equiv \psi^\dagger \gamma^0$$

* Field dynamics

* Noether's theorem: (continuous) symmetry implies conservation laws (energy, momentum, angular momentum)

LagrangiansDynamics(classical)

• Principle of least action: $\delta S = 0$ where $S = \int d^4x \mathcal{L}(x)$ \Rightarrow Field EoM (E-L equations)

$$\delta S = \int d^4 x \sum_i \left(\frac{\partial \mathcal{L}}{\partial \phi_i} \delta \phi_i + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_i)} \delta (\partial_\mu \phi_i) \right)$$

= $\int d^4 x \sum_i \left(\frac{\partial \mathcal{L}}{\partial \phi_i} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_i)} \right) \delta \phi_i = 0 \quad , \quad \forall \phi_i$

(integrating by parts and assuming fields vanish at boundary)

– e.g. EoM of a free Dirac field is the Dirac equation

$$(\mathbf{i}\partial - m)\psi(x) = 0$$

$$\psi(x) = \int \frac{\mathrm{d}^3 p}{(2\pi)^3 \sqrt{2E_p}} \sum_{s=1,2} \left(a_{p,s} u^{(s)}(p) \mathrm{e}^{-\mathrm{i}px} + b_{p,s}^* v^{(s)}(p) \mathrm{e}^{\mathrm{i}px} \right)$$
with $p^2 = E_p^2 - p^2 = m^2$, $(\not p - m) u(p) = 0$, $(\not p + m) v(p) = 0$.

LagrangiansQuantization(«particle» concept emerges)

• Impose canonical quantization rules:

commutation/anticommutation of fields with conjugate momenta $\Pi_i(x) = \frac{\partial \mathcal{L}}{\partial(\partial_0 \phi_i)}$

$$[\phi(t,\boldsymbol{x}),\Pi_{\phi}(t,\boldsymbol{y})] = \mathrm{i}\delta^{3}(\boldsymbol{x}-\boldsymbol{y}), \quad \{\psi(t,\boldsymbol{x}),\Pi_{\psi}(t,\boldsymbol{y})\} = \mathrm{i}\delta^{3}(\boldsymbol{x}-\boldsymbol{y})$$

so that the Hamiltonian is bounded from below.

– e.g for a free fermion field, *anticommutation* is enforced! implying

$$\{a_{\boldsymbol{p},r}, a_{\boldsymbol{k},s}^{\dagger}\} = \{b_{\boldsymbol{p},r}, b_{\boldsymbol{k},s}^{\dagger}\} = (2\pi)^{3}\delta^{3}(\boldsymbol{p}-\boldsymbol{k})\delta_{rs}, \quad \{a_{\boldsymbol{p},r}, a_{\boldsymbol{k},s}\} = \cdots = 0$$

- After normal ordering :: (all creation to left of annihilation opts) to subtract zero-point energy,

$$H = \int d^3x : \mathcal{H}(x) := \int \frac{dp^3}{(2\pi)^3} E_p \sum_{s=1,2} (a_{p,s}^{\dagger} a_{p,s} + b_{p,s}^{\dagger} b_{p,s})$$

 \Rightarrow Fields become operators that annihilate/create particles/antiparticles

 $|0\rangle$ (vacuum), $a_{p,s}^{\dagger}|0\rangle$ (1 particle), $b_{p,s}^{\dagger}|0\rangle$ (1 antiparticle), ...

⇒ Multiparticle states symmetric/antisymmetric under exchange (spin-statistics!)

One-particle representations

• One-particle states are unitary irreps of the Poincaré group, so that

 $\langle \psi_1 | \psi_2 \rangle = \langle \psi_1 | \mathcal{P}^{\dagger} \mathcal{P} | \psi_2 \rangle$ (invariant matrix elements)

 \mathcal{P} are represented by unitary operators in this space, and the generators J^i (rotations), K^i (boosts), P^{μ} (translations) by Hermitian operators. $J_{\mu\nu} = -J_{\nu\mu}$ $(J^i = \frac{1}{2}\epsilon^{ijk}J^{jk}, K^i = J^{0i})$

- Rotations form a compact subgroup (its finite dimensional irreps are unitary). But Lorentz group and Poincaré group are non-compact. Therefore: The *unitary* representations of the Poincaré group are *infinite-dimensional*.
- Poincaré group has two Casimir operators (commute with all generators)

$$m^2 = P_{\mu}P^{\mu}, \quad W_{\mu}W^{\mu}$$
 $W_{\mu} = \frac{1}{2}\epsilon^{\mu\nu\rho\sigma}J_{\nu\rho}P_{\sigma}$ (Pauli-Lubanski vector)

whose eigenvalues label the irreps. Lorentz invariant (choose convenient frame).

One-particle representations

- Two cases, characterized by mass *m* and spin *j*
 - $m \neq 0$: choose $P^{\mu} = (m, 0, 0, 0) \Rightarrow W_{\mu}W^{\mu} = -m^2 j(j+1)$ \Rightarrow massive particles of spin j have 2j + 1 dof $(j_3 = -j, -j + 1, ..., j)$
 - because SU(2) is the *little group* (transformations leaving P^{μ} invariant)
 - m = 0: choose $P^{\mu} = (\omega, 0, 0, \omega) \Rightarrow W_{\mu}W^{\mu} = -\omega^2[(J^1 + K^2)^2 + (J^2 K^1)^2]$ \Rightarrow massless particles of spin *j* have 2 dof (helicity $h = \pm j$) because now SO(2) is the *little group* (rotations in plane \perp to P^{μ})
- <u>Note</u>: To construct a **unitary field theory with** V_{μ} (contains both spin 0 and 1) one has to choose carefully the Lagrangian so that the *physical theory* **never excites**:
 - the spin-0 component (if massive)
 - neither the longitudinal spin-1 component (if massless) \Leftrightarrow gauge invariance

Particle physics

Particle physicsS-matrix elements

 Observables (cross sections, decays widths) expressed in terms of S-matrix elements (*m* → *n* processes)

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out \langle \boldsymbol{p}_1 \boldsymbol{p}_2 \cdots \boldsymbol{p}_n | \boldsymbol{k}_1 \boldsymbol{k}_2 \cdots \boldsymbol{k}_m \rangle_{\text{in}}
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(scalar fields/particles to simplify)

Only free fields are related to particles/antiparticles (a⁺_p, b⁺_p).
 We expect

$$\phi(x) \xrightarrow[t \to -\infty]{} Z_{\phi}^{1/2} \phi_{\text{in}}(x) , \quad \phi(x) \xrightarrow[t \to +\infty]{} Z_{\phi}^{1/2} \phi_{\text{out}}(x) ,$$

 $\phi(x)$: interacting fields $\phi_{in}(x), \phi_{out}(x)$: free fields (before, after interaction) Z_{ϕ} : *wave function* renormalization Particle physicsS-matrix elementsLSZ

 $\mathbf{Z} \quad \text{(particles} \leftrightarrow \text{fields)}$

• LSZ reduction formula relates S-matrix elements with the (Fourier transform of) vacuum expectation values of *time-ordered* field products (correlators):

$$\begin{pmatrix} \prod_{i=1}^{m} \frac{i\sqrt{Z_{\phi}}}{k_{i}^{2} - m^{2}} \end{pmatrix} \begin{pmatrix} \prod_{j=1}^{n} \frac{i\sqrt{Z_{\phi}}}{p_{j}^{2} - m^{2}} \end{pmatrix}_{\text{out}} \langle \boldsymbol{p}_{1}\boldsymbol{p}_{2}\cdots\boldsymbol{p}_{n} | \boldsymbol{k}_{1}\boldsymbol{k}_{2}\cdots\boldsymbol{k}_{m} \rangle_{\text{in}}$$

$$= \int \left(\prod_{i=1}^{m} d^{4}x_{i} \ e^{-ik_{i}x_{i}} \right) \int \left(\prod_{j=1}^{n} d^{4}y_{j} \ e^{+ip_{j}y_{j}} \right) \langle 0 | T \{ \underbrace{\phi(x_{1})\cdots\phi(x_{m})\phi(y_{1})\cdots\phi(y_{n})}_{\text{interacting fields}} \} | 0 \rangle$$

The correlator = the Green's function of m + n points $G(p_1 \cdots p_n; k_1 \cdots k_m)$

▷ Physical particles (asymptotic states) are on-shell $(p^2 - m^2 = 0)$.

For on-shell incoming and outgoing particles, the rhs of LSZ formula (correlator) will have **poles** that cancel those in the prefactor of the lhs, yielding a regular S-matrix element [*residues* of the correlator].

Particle physicsPerturbation theoryFeynman diagrams

• The correlators can be expressed in terms of free fields " ϕ_0 " :

$$\langle 0 | T\{\phi(x_1)\cdots\phi(x_n)\} | 0 \rangle = \frac{\langle 0 | T\{\phi_0(x_1)\cdots\phi_0(x_n)\exp\left[i\int d^4x \mathcal{L}_{int}[\phi_0(x)]\right]\} | 0 \rangle}{\langle 0 | T\{\exp\left[i\int d^4x \mathcal{L}_{int}[\phi_0(x)]\right]\} | 0 \rangle}$$

• In perturbation theory one expands the exponential and computes every correlator using *Wick's theorem* (all possible "contractions")

contraction
$$\equiv \phi(x)\phi(y) = D_F(x-y) = \langle 0 | T\{\phi(x)\phi(y)\} | 0 \rangle = Feynman propagator$$

- Feynman diagrams/rules provide a systematic procedure to organize/compute the perturbative series in terms of *propagators* (and vertices)
- Note: functional quantization (path integral) provides an *alternative* method

$$\langle 0 | T\{\hat{\phi}(x_1)\cdots\hat{\phi}(x_n)\} | 0 \rangle = \frac{\int \mathcal{D}\phi \ \phi(x_1)\cdots\phi(x_n) \ e^{iS}}{\int \mathcal{D}\phi \ e^{iS}} \qquad \text{perturbatively} \text{ or not! (lattice)}$$

Particle physicsPerturbation theoryPropagators

- **Causality** requires $[\phi(x), \phi^{\dagger}(y)] = 0$ if $(x y)^2 < 0$ (*spacelike* interval)
- ▷ Recall that a (free) field is a combination of positive and negative energy waves:

$$\phi(x) = \int \frac{\mathrm{d}p^3}{(2\pi)^3 \sqrt{E_p}} \left(a_p \mathrm{e}^{-\mathrm{i}px} + b_p^{\dagger} \mathrm{e}^{\mathrm{i}px} \right)$$

▷ From the commutation relations of creation and annihilation operators:

$$\begin{aligned} [\phi(x),\phi^{\dagger}(y)] &= \int \frac{\mathrm{d}p^3}{(2\pi)^3\sqrt{E_p}} \int \frac{\mathrm{d}q^3}{(2\pi)^3\sqrt{E_q}} \left(\mathrm{e}^{-\mathrm{i}(px-qy)}[a_p,a_q^{\dagger}] + \mathrm{e}^{\mathrm{i}(px-qy)}[b_{p}^{\dagger},b_q] \right) \\ &= \Delta(x-y) - \Delta(y-x) \end{aligned}$$

where the first (second) contribution comes from particles (antiparticles) and

$$\Delta(x-y) = \int \frac{\mathrm{d}p^3}{(2\pi)^3 E_p} \mathrm{e}^{-\mathrm{i}p \cdot (x-y)}$$

Particle physicsPerturbation theoryPropagators

▷ If $(x - y)^2 < 0$ choose frame where $x - y \equiv (0, r)$. Then

$$\Delta(x-y) = \Delta(y-x) \propto \frac{m}{r} e^{-mr} \neq 0$$
, for $mr \gg 1$

Therefore:

If only particles: $[\phi(x), \phi^{\dagger}(y)] = \Delta(x - y) \neq 0$ (!!) If *both* particles and antiparticles: $[\phi(x), \phi^{\dagger}(y)] = \Delta(x - y) - \Delta(y - x) = 0$ (\checkmark)

• In fact the **Feynman propagator** contains *both* contributions:

 $D_F(x-y) = \langle 0 | T\{\phi(x)\phi^{\dagger}(y)\} | 0 \rangle = \theta(x^0 - y^0)\Delta(x-y) + \theta(y^0 - x^0)\Delta(y-x)$

- Probability amplitude that particle created in *y* propagates to *x*, if $x^0 > y^0$
- Probability amplitude that antiparticle created in *x* propagates to *y*, if $y^0 > x^0$

$$D_F(x-y) = \int \frac{\mathrm{d}^4 p}{(2\pi)^4} \frac{\mathrm{i}}{p^2 - m^2 + \mathrm{i}\varepsilon} \mathrm{e}^{-\mathrm{i}p \cdot (x-y)} \quad \text{where} \quad \varepsilon \to 0^+ (\text{usually omitted})$$

Particle physics

• Corrections to external legs (external propagators) can be resummed:

$$= \underbrace{i}_{p^2 - m_0^2} + \underbrace{i}_{p^2 - m_0^2} [-iM^2(p^2)] \underbrace{i}_{p^2 - m_0^2} + \dots$$

$$= \frac{i}{p^2 - m_0^2 - M^2(p^2)} \qquad (m_0 = \text{mass in } \mathcal{L})$$

Perturbation theory LSZ (rewritten)

and Taylor expanding about $p^2 = m^2$ (*physical* mass):

$$p^{2} - m_{0}^{2} - M^{2}(p^{2}) = (p^{2} - m^{2}) \left(1 - \frac{dM^{2}}{dp^{2}} \Big|_{p^{2} = m^{2}} \right)$$

$$\Rightarrow \qquad \longrightarrow \qquad = \frac{iZ_{\phi}}{p^{2} - m^{2}} + \text{regular near } p^{2} = m^{2}$$

with $m^{2} = m_{0}^{2} + M^{2}(m^{2})$, $Z_{\phi} = \left(1 - \frac{dM^{2}}{dp^{2}} \Big|_{p^{2} = m^{2}} \right)^{-1}$

Particle physicsPerturbation theoryLSZ(rewritten)

▷ Then we may factor out external legs from *amputated* diagrams:



and express the LSZ formula in a simpler form:



Perturbation theoryRenormalization

• Feynman rules require integration over loop momenta resulting *sometimes* in divergent expressions.

$$\mathcal{M} = \mathcal{M}^{(0)} + \underbrace{\mathcal{M}^{(1)}}_{\text{divergent}?} + \dots$$

(the loop expansion is also an expansion in powers of *ħ*: *quantum* corrections)

- **Regularization** and **renormalization** needed to make sense of these divergences.
- One assumes that fields and parameters in the Lagrangian (*bare*) must be redefined order by order in terms of new ones (*renormalized*) so that physical predictions are finite

$$\mathcal{M} = \mathcal{M}^{(0)} + \underbrace{\widehat{\mathcal{M}}^{(1)}}_{\text{finite}} + \dots$$

Particle physics

Particle physicsPerturbation theoryRenormalization

▷ As a consequence, renormalized *coupling constants run* (depend on a scale)

e.g.



 q^2 = *renormalization scale* (at which *e* is "measured")

Note: *e* is not an observable

Global and gauge symmetries

Internal symmetries | free Lagrangian

• In addition to **spacetime** (Poincaré) symmetries, the free Lagrangian

(Dirac)
$$\mathcal{L}_0 = \overline{\psi}(i\partial \!\!\!/ - m)\psi \quad \partial \!\!\!/ \equiv \gamma^\mu \partial_\mu , \quad \overline{\psi} \equiv \psi^\dagger \gamma^0$$

 \Rightarrow Invariant under **internal** global U(1) phase transformations:

$$\psi(x) \mapsto \psi'(x) = e^{-iQ\theta}\psi(x)$$
, Q, θ (constants) $\in \mathbb{R}$

 \Rightarrow By Noether's theorem, divergentless current:

$${\cal J}^\mu = Q \; \overline{\psi} \gamma^\mu \psi \;, \;\;\; \partial_\mu {\cal J}^\mu = 0$$

and a conserved «charge»

$$\mathcal{Q} = \int \mathrm{d}^3 x \ \mathcal{J}^0, \quad \partial_t \mathcal{Q} = 0$$

Internal symmetries | free Lagrangian

- For a free fermion **quantum** field:
 - \Rightarrow The Noether charge is an operator:*

$$\mathcal{Q} = Q \int d^3 x : \overline{\psi} \gamma^0 \psi := Q \int \frac{d^3 p}{(2\pi)^3} \sum_{s=1,2} \left(a^{\dagger}_{\boldsymbol{p},s} a_{\boldsymbol{p},s} - b^{\dagger}_{\boldsymbol{p},s} b_{\boldsymbol{p},s} \right)$$
$$\mathcal{Q} a^{\dagger}_{\boldsymbol{k},s} \left| 0 \right\rangle = +Q a^{\dagger}_{\boldsymbol{k},s} \left| 0 \right\rangle \text{ (particle) , } \mathcal{Q} b^{\dagger}_{\boldsymbol{k},s} \left| 0 \right\rangle = -Q b^{\dagger}_{\boldsymbol{k},s} \left| 0 \right\rangle \text{ (antiparticle)}$$

* normal ordering prescription for fermionic operators

$$:a_{p,r}a_{q,s}^{\dagger}:\equiv -a_{q,s}^{\dagger}a_{p,r}$$
, $:b_{p,r}b_{q,s}^{\dagger}:\equiv -b_{q,s}^{\dagger}b_{p,r}$

The gauge principlegauge symmetry dictates interactions

• To make \mathcal{L}_0 invariant under local \equiv gauge transformations of U(1):

$$\psi(x) \mapsto \psi'(x) = \mathrm{e}^{-\mathrm{i}Q\theta(x)}\psi(x) \ , \quad \theta = \theta(x) \in \mathbb{R}$$

perform the minimal substitution:

 $\partial_{\mu} \rightarrow D_{\mu} = \partial_{\mu} + ieQA_{\mu}$ (covariant derivative)

where a gauge field $A_{\mu}(x)$ is introduced transforming as:

$$A_{\mu}(x) \mapsto A'_{\mu}(x) = A_{\mu}(x) + \frac{1}{e} \partial_{\mu} \theta(x) \quad \Leftarrow \quad \boxed{D_{\mu} \psi \mapsto e^{-iQ\theta(x)} D_{\mu} \psi} \quad \overline{\psi} D \psi \text{ inv.}$$
 (1)

 \Rightarrow The new Lagrangian contains interactions between ψ and A_{μ} :

$$\mathcal{L}_{\text{int}} = -e \ Q \ \overline{\psi} \gamma^{\mu} \psi A_{\mu} \qquad \propto \begin{cases} \text{coupling } e \\ \text{charge } Q \end{cases}$$

 $(=-e \ \mathcal{J}^{\mu}A_{\mu})$

The gauge principle

gauge invariance dictates interactions

• Dynamics for the gauge field ⇒ add gauge invariant kinetic term:

(Maxwell)
$$\mathcal{L}_1 = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \qquad \Leftarrow \quad F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \mapsto F_{\mu\nu}$$

– The full U(1) gauge invariant Lagrangian for a fermion field $\psi(x)$ reads:

$$\mathcal{L}_{\text{sym}} = \overline{\psi}(i\mathcal{D} - m)\psi - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} \qquad (=\mathcal{L}_0 + \mathcal{L}_{\text{int}} + \mathcal{L}_1) \quad (\text{QED})$$

– The same applies to a complex scalar field $\phi(x)$:

$$\mathcal{L}_{\text{sym}} = (D_{\mu}\phi)^{\dagger}D^{\mu}\phi - m^{2}\phi^{\dagger}\phi - \lambda(\phi^{\dagger}\phi)^{2} - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} \quad (\text{sQED})$$

The gauge principlenon-Abelian gauge theories

• A general gauge symmetry group *G* is an *compact N*-dimensional Lie group

$$\mathbf{g}\in G$$
 , $\mathbf{g}(\boldsymbol{ heta})=\mathrm{e}^{-\mathrm{i}T_a\theta^a}$, $a=1,\ldots,N$

 $\theta^{a} = \theta^{a}(x) \in \mathbb{R}$, T_{a} = Hermitian generators, $[T_{a}, T_{b}] = if_{abc}T_{c}$ (Lie algebra) structure constants: $f_{abc} = 0$ Abelian $f_{abc} \neq 0$ non-Abelian

⇒ Unitary finite-dimensional irreducible representations:

 $g(\theta) \text{ represented by } U(\theta)$ $d \times d \text{ matrices} : \quad U(\theta) \text{ [given by } \{T_a\} \text{ algebra representation]}$ $d\text{-multiplet} : \quad \Psi(x) \mapsto \Psi'(x) = U(\theta)\Psi(x) \text{ , } \quad \Psi = \begin{pmatrix} \psi_1 \\ \vdots \\ \psi_1 \end{pmatrix}$ The gauge principlenon-Abelian gauge theories

- Examples: G N Abelian U(1) 1 Yes SU(n) $n^2 - 1$ No $(n \times n \text{ unitary matrices with det} = 1)$
 - U(1): 1 generator (*q*), one-dimensional irreps only
 - SU(2): 3 generators

 $f_{abc} = \epsilon_{abc}$ (Levi-Civita symbol)

- * Fundamental irrep (d = 2): $T_a = \frac{1}{2}\sigma_a$ (3 Pauli matrices)
- * Adjoint irrep (d = N = 3): $(T_a^{adj})_{bc} = -if_{abc}$
- SU(3): 8 generators

$$f^{123} = 1, f^{458} = f^{678} = \frac{\sqrt{3}}{2}, f^{147} = f^{156} = f^{246} = f^{247} = f^{345} = -f^{367} = \frac{1}{2}$$

- * Fundamental irrep (d = 3): $T_a = \frac{1}{2}\lambda_a$ (8 Gell-Mann matrices)
- * Adjoint irrep (d = N = 8): $(T_a^{adj})_{bc} = -if_{abc}$

(for SU(n): f_{abc} totally antisymmetric)

• To make \mathcal{L}_0 invariant under local \equiv gauge transformations of *G*:

$$\mathcal{L}_0 = \overline{\Psi}(\mathrm{i}\partial - m)\Psi$$
, $\Psi(x) \mapsto \Psi'(x) = U(\theta)\Psi(x)$, $\theta = \theta(x) \in \mathbb{R}$

substitute the covariant derivative:

$$\partial_{\mu} \rightarrow D_{\mu} = \partial_{\mu} - ig\widetilde{W}_{\mu}$$
, $\widetilde{W}_{\mu} \equiv T_a W^a_{\mu}$

where a gauge field $W_{\mu}^{a}(x)$ per generator is introduced, transforming as:

$$\widetilde{W}_{\mu}(x) \mapsto \widetilde{W}'_{\mu}(x) = \underbrace{U\widetilde{W}_{\mu}(x)U^{\dagger}}_{\text{adjoint irrep}} - \frac{i}{g}(\partial_{\mu}U)U^{\dagger} \quad \Leftarrow \quad \boxed{D_{\mu}\Psi \mapsto UD_{\mu}\Psi} \quad \overline{\Psi}D\Psi \text{ inv.}$$

 \Rightarrow The new Lagrangian contains interactions between Ψ and W_{u}^{a} :

$$\begin{aligned} \mathcal{L}_{\text{int}} &= g \, \overline{\Psi} \gamma^{\mu} T_a \Psi W^a_{\mu} \\ & (= g \, \mathcal{J}^{\mu}_a W^a_{\mu}) \end{aligned} \propto \begin{cases} \text{coupling } g \\ \text{charge } T_a \end{cases}$$

The gauge principle no

• Dynamics for the gauge fields \Rightarrow add gauge invariant kinetic terms:

(Yang-Mills)
$$\mathcal{L}_{\rm YM} = -\frac{1}{2} \operatorname{Tr} \left\{ \widetilde{W}_{\mu\nu} \widetilde{W}^{\mu\nu} \right\} = -\frac{1}{4} W^a_{\mu\nu} W^{a,\mu\nu}$$

$$\widetilde{W}_{\mu\nu} \equiv T_a W^a_{\mu\nu} \equiv D_\mu \widetilde{W}_\nu - D_\nu \widetilde{W}_\mu = \partial_\mu \widetilde{W}_\nu - \partial_\nu \widetilde{W}_\mu - \mathbf{i}g[\widetilde{W}_\mu, \widetilde{W}_\nu] \iff \widetilde{W}_{\mu\nu} \mapsto U\widetilde{W}_{\mu\nu} U^+$$

$$\Rightarrow \quad W^a_{\mu\nu} = \partial_\mu W^a_\nu - \partial_\nu W^a_\mu + gf_{abc} W^b_\mu W^c_\nu \qquad (2)$$

 $\Rightarrow \mathcal{L}_{YM}$ contains cubic and quartic self-interactions of the gauge fields W^a_{μ} :

$$\mathcal{L}_{kin} = -\frac{1}{4} (\partial_{\mu} W^{a}_{\nu} - \partial_{\nu} W^{a}_{\mu}) (\partial^{\mu} W^{a,\nu} - \partial^{\nu} W^{a,\mu}$$
$$\mathcal{L}_{cubic} = -\frac{1}{2} g f_{abc} (\partial_{\mu} W^{a}_{\nu} - \partial_{\nu} W^{a}_{\mu}) W^{b,\mu} W^{c,\nu}$$
$$\mathcal{L}_{quartic} = -\frac{1}{4} g^{2} f_{abe} f_{cde} W^{a}_{\mu} W^{b}_{\nu} W^{c,\mu} W^{d,\nu}$$

Quantization propagators

• The (Feynman) propagator of a scalar field:

$$D_F(x-y) = \langle 0 | T\{\phi(x)\phi^{\dagger}(y)\} | 0 \rangle = \int \frac{d^4p}{(2\pi)^4} \frac{i}{p^2 - m^2 + i\varepsilon} e^{-ip \cdot (x-y)}$$

(Feynman *prescription* $\varepsilon \rightarrow 0^+$)

is a Green's function of the Klein-Gordon operator:

$$(\Box_x + m^2)D_F(x - y) = -i\delta^4(x - y) \quad \Leftrightarrow \quad \widetilde{D}_F(p) = \frac{1}{p^2 - m^2 + i\varepsilon}$$

• The propagator of a fermion field:

$$S_F(x-y) = \langle 0 | T\{\psi(x)\overline{\psi}(y)\} | 0 \rangle = (\mathbf{i}\partial_x + m) \int \frac{\mathrm{d}^4 p}{(2\pi)^4} \frac{\mathbf{i}}{p^2 - m^2 + \mathbf{i}\varepsilon} \mathrm{e}^{-\mathbf{i}p \cdot (x-y)}$$

is a Green's function of the Dirac operator:

$$(i\partial_x - m)S_F(x - y) = i\delta^4(x - y) \quad \Leftrightarrow \quad \widetilde{S}_F(p) = \frac{1}{\not p - m + i\varepsilon}$$

Quantization of gauge theories propagators

• HOWEVER a gauge field propagator cannot be defined unless \mathcal{L} is modified:

(e.g. modified Maxwell)
$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2\xi}(\partial^{\mu}A_{\mu})^{2}$$

Euler-Lagrange:
$$\frac{\partial \mathcal{L}}{\partial A_{\nu}} - \partial_{\mu} \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} A_{\nu})} = 0 \quad \Rightarrow \quad \left[g^{\mu\nu} \Box - \left(1 - \frac{1}{\xi}\right) \partial^{\mu} \partial^{\nu}\right] A_{\mu} = 0$$

– In momentum space the propagator is the inverse of:

$$-k^2 g^{\mu\nu} + \left(1 - \frac{1}{\xi}\right) k^{\mu} k^{\nu} \quad \Rightarrow \quad \widetilde{D}_{\mu\nu}(k) = \frac{\mathrm{i}}{k^2 + \mathrm{i}\varepsilon} \left[-g_{\mu\nu} + (1 - \xi)\frac{k_{\mu}k_{\nu}}{k^2}\right]$$

 \Rightarrow Note that $(-k^2g^{\mu\nu} + k^{\mu}k^{\nu})$ is singular!

 $\Rightarrow \text{ One may argue that } \mathcal{L} \text{ above will not lead to Maxwell equations } \dots$ unless we fix a (Lorenz) gauge where: (remove redundancy)

$$\partial^{\mu}A_{\mu} = 0 \quad \Leftarrow \quad A_{\mu} \mapsto A'_{\mu} = A_{\mu} + \partial_{\mu}\Lambda \text{ with } \partial^{\mu}\partial_{\mu}\Lambda \equiv -\partial^{\mu}A_{\mu}$$

Quantization of gauge theories | gauge fixing (Abelian case)

The extra term is called Gauge Fixing:

$${\cal L}_{
m GF} = - {1 \over 2 {ar \xi}} (\partial^\mu A_\mu)^2$$

 \Rightarrow modified \mathcal{L} equivalent to Maxwell Lagrangian just in the gauge $\partial^{\mu}A_{\mu} = 0$

- \Rightarrow the ξ -dependence always cancels out in physical amplitudes
- Several choices for the gauge fixing term (simplify calculations): $R_{\tilde{c}}$ gauges

$$\zeta' \text{t Hooft-Feynman gauge}) \quad \xi = 1: \quad \widetilde{D}_{\mu\nu}(k) = -\frac{\mathrm{i}g_{\mu\nu}}{k^2 + \mathrm{i}\varepsilon}$$

$$(\text{Landau gauge}) \quad \xi = 0: \quad \widetilde{D}_{\mu\nu}(k) = \frac{\mathrm{i}}{k^2 + \mathrm{i}\varepsilon} \left[-g_{\mu\nu} + \frac{k_{\mu}k_{\nu}}{k^2} \right]$$

Quantization of gauge theories | gauge fixing (non-Abelian case)

- For a non-Abelian gauge theory, the gauge fixing terms:

$$\mathcal{L}_{\mathrm{GF}} = -\sum_{a} rac{1}{2\xi_{a}} (\partial^{\mu} W^{a}_{\mu})^{2}$$

allow to define the propagators:

$$\widetilde{D}^{ab}_{\mu\nu}(k) = \frac{\mathrm{i}\delta_{ab}}{k^2 + \mathrm{i}\varepsilon} \left[-g_{\mu\nu} + (1 - \xi_a) \frac{k_{\mu}k_{\nu}}{k^2} \right]$$

HOWEVER, unlike the Abelian case, this is not the end of the story ...

Quantization of gauge theories

• Add Faddeev-Popov *ghost* fields $c_a(x)$, a = 1, ..., N: ('t Hooft-Feynman gauge)

$$\mathcal{L}_{\rm FP} = (\partial^{\mu} \bar{c}_{a}) (D^{\rm adj}_{\mu})_{ab} c_{b} = (\partial^{\mu} \bar{c}_{a}) (\partial_{\mu} c_{a} - g f_{abc} c_{b} W^{c}_{\mu}) \qquad \Leftrightarrow \qquad D^{\rm adj}_{\mu} = \partial_{\mu} - \mathrm{i} g T^{\rm adj}_{c} W^{c}_{\mu}$$

Computational trick: *anticommuting* scalar fields, just in loops as virtual particles \Rightarrow Faddeev-Popov ghosts needed to preserve gauge symmetry:



Quantization of gauge theories

• Then the full quantum Lagrangian is

$$\mathcal{L}_{sym} + \mathcal{L}_{GF} + \mathcal{L}_{FP}$$

Full quantum Lagrangian

 \Rightarrow Note that in the case of a massive vector field

(Proca)
$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \frac{1}{2}M^2A_{\mu}A^{\mu}$$

it is not gauge invariant!!!



What about the gauge principle???

– The propagator is:

$$\widetilde{D}_{\mu\nu}(k) = \frac{\mathrm{i}}{k^2 - M^2 + \mathrm{i}\varepsilon} \left(-g_{\mu\nu} + \frac{k^{\mu}k^{\nu}}{M^2} \right)$$
Spontaneous Symmetry Breaking discrete symmetry

Consider a real scalar field $\phi(x)$ with Lagrangian:

$$\mathcal{L} = \frac{1}{2} (\partial_{\mu} \phi) (\partial^{\mu} \phi) - \frac{1}{2} \mu^{2} \phi^{2} - \frac{\lambda}{4} \phi^{4} \quad \text{invariant under} \quad \phi \mapsto -\phi$$

$$\Rightarrow \mathcal{H} = \frac{1}{2}(\dot{\phi}^2 + (\nabla\phi)^2) + V(\phi)$$
(a)
$$V = \frac{1}{2}\mu^2\phi^2 + \frac{1}{4}\lambda\phi^4$$
(b)

 μ^2 , $\lambda \in \mathbb{R}$ (Real/Hermitian Hamiltonian) and $\lambda > 0$ (existence of a ground state) (a) $\mu^2 > 0$: min of $V(\phi)$ at $\phi = 0$ (b) $\mu^2 < 0$: min of $V(\phi)$ at $\phi = v \equiv \pm \sqrt{\frac{-\mu^2}{\lambda}}$, in QFT $\langle 0 | \phi | 0 \rangle = v \neq 0$ (VEV) – A quantum field must have v = 0

$$a |0\rangle = 0$$
 \Rightarrow $\phi(x) \equiv v + \eta(x)$, $\langle 0|\eta|0\rangle = 0$

Spontaneous Symmetry Breaking discrete symmetry

• At the quantum level, the same system is described by $\eta(x)$ with Lagrangian:

$$\mathcal{L} = \frac{1}{2} (\partial_{\mu} \eta) (\partial^{\mu} \eta) - \lambda v^{2} \eta^{2} - \lambda v \eta^{3} - \frac{\lambda}{4} \eta^{4} + \frac{1}{4} \lambda v^{4} \text{ not invariant under } \eta \mapsto -\eta$$
$$(m_{\eta} = \sqrt{2\lambda} v)$$

 \Rightarrow Lesson:

 $\mathcal{L}(\phi)$ has the symmetry but the parameters can be such that the ground state of the Hamiltonian is not symmetric (Spontaneous Symmetry Breaking)

\Rightarrow Note:

One may argue that $\mathcal{L}(\eta)$ exhibits an explicit breaking of the symmetry. However this is not the case since the coefficients of terms η^2 , η^3 and η^4 are determined by just two parameters, λ and v (remnant of the original symmetry)

Spontaneous Symmetry Breaking

• Consider a complex scalar field $\phi(x)$ with Lagrangian:

 $\mathcal{L} = (\partial_{\mu}\phi^{\dagger})(\partial^{\mu}\phi) - \mu^{2}\phi^{\dagger}\phi - \lambda(\phi^{\dagger}\phi)^{2} \text{ invariant under U(1): } \phi \mapsto e^{-iq\theta}\phi$

continuous symmetry

$$\lambda > 0, \ \mu^2 < 0: \quad \langle 0 | \phi | 0 \rangle \equiv \frac{v}{\sqrt{2}}, \quad |v| = \sqrt{\frac{-\mu^2}{\lambda}}$$

Take $v \in \mathbb{R}^+$. In terms of quantum fields:

$$\phi(x) \equiv \frac{1}{\sqrt{2}} [v + \eta(x) + i\chi(x)], \quad \langle 0 | \eta | 0 \rangle = \langle 0 | \chi | 0 \rangle = 0$$

$$\sqrt{2} = \frac{1}{2} (\partial_{\mu} \eta) (\partial^{\mu} \eta) + \frac{1}{2} (\partial_{\mu} \chi) (\partial^{\mu} \chi) - \lambda v^2 \eta^2 - \lambda v \eta (\eta^2 + \chi^2) - \frac{\lambda}{4} (\eta^2 + \chi^2)^2 + \frac{1}{4} \lambda v^4$$

Note: if $ve^{i\alpha}$ (complex) replace η by $(\eta \cos \alpha - \chi \sin \alpha)$ and χ by $(\eta \sin \alpha + \chi \cos \alpha)$

⇒ The actual quantum Lagrangian $\mathcal{L}(\eta, \chi)$ is not invariant under U(1) U(1) broken ⇒ one scalar field remains massless: $m_{\chi} = 0$, $m_{\eta} = \sqrt{2\lambda} v$

 \mathcal{L}



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Spontaneous Symmetry Breaking con

- continuous symmetry
- Another example: consider a real scalar SU(2) triplet $\Phi(x)$

$$\mathcal{L} = \frac{1}{2} (\partial_{\mu} \Phi^{\mathsf{T}}) (\partial^{\mu} \Phi) - \frac{1}{2} \mu^{2} \Phi^{\mathsf{T}} \Phi - \frac{\lambda}{4} (\Phi^{\mathsf{T}} \Phi)^{2} \quad \text{inv. under SU(2):} \quad \Phi \mapsto e^{-iT_{a}\theta^{a}} \Phi$$

that for $\lambda > 0$, $\mu^2 < 0$ acquires a VEV $\langle 0 | \Phi^T \Phi | 0 \rangle = v^2$ $(\mu^2 = -\lambda v^2)$ Assume $\Phi(x) = \begin{pmatrix} \varphi_1(x) \\ \varphi_2(x) \\ v + \varphi_3(x) \end{pmatrix}$ and define $\varphi \equiv \frac{1}{\sqrt{2}}(\varphi_1 + i\varphi_2)$

 $\mathcal{L} = (\partial_{\mu}\varphi^{\dagger})(\partial^{\mu}\varphi) + \frac{1}{2}(\partial_{\mu}\varphi_{3})(\partial^{\mu}\varphi_{3}) - \lambda v^{2}\varphi_{3}^{2} - \lambda v(2\varphi^{\dagger}\varphi + \varphi_{3}^{2})\varphi_{3} - \frac{\lambda}{4}(2\varphi^{\dagger}\varphi + \varphi_{3}^{2})^{2} + \frac{1}{4}\lambda v^{4}$

 \Rightarrow Not symmetric under SU(2) but invariant under U(1):

$$\varphi \mapsto e^{-iQ\theta} \varphi \quad (Q = arbitrary) \qquad \qquad \varphi_3 \mapsto \varphi_3 \quad (Q = 0)$$

SU(2) broken to U(1) \Rightarrow 3 – 1 = 2 broken generators

 \Rightarrow 2 (real) scalar fields (= 1 complex) remain massless: $m_{\varphi} = 0$, $m_{\varphi_3} = \sqrt{2\lambda} v$

Spontaneous Symmetry Breaking continuous symmetry

\Rightarrow Goldstone's theorem:

[Nambu '60; Goldstone '61]

The number of massless particles (Nambu-Goldstone bosons) is equal to the number of spontaneously broken generators of the symmetry

Hamiltonian symmetric under group $G \Rightarrow [T_a, H] = 0$, a = 1, ..., NBy definition: $H |0\rangle = 0 \implies H(T_a |0\rangle) = T_a H |0\rangle = 0$

- If $|0\rangle$ is such that $T_a |0\rangle = 0$ for all generators \Rightarrow non-degenerate minimum: *the* vacuum

- If $|0\rangle$ is such that $T_{a'}|0\rangle \neq 0$ for some (broken) generators a'

 \Rightarrow degenerate minimum: chose one (*true* vacuum) and $e^{-iT_{a'}\theta^{a'}} |0\rangle \neq |0\rangle$

 \Rightarrow excitations (particles) from $|0\rangle$ to $e^{-iT_{a'}\theta^{a'}}|0\rangle$ cost no energy: massless!

Spontaneous Symmetry Breaking gauge symmetry

• Consider a U(1) gauge invariant Lagrangian for a complex scalar field $\phi(x)$:

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + (D_{\mu}\phi)^{\dagger}(D^{\mu}\phi) - \mu^{2}\phi^{\dagger}\phi - \lambda(\phi^{\dagger}\phi)^{2}, \quad D_{\mu} = \partial_{\mu} + \mathrm{i}eQA_{\mu}$$

inv. under $\phi(x) \mapsto \phi'(x) = e^{-iQ\theta(x)}\phi(x)$, $A_{\mu}(x) \mapsto A'_{\mu}(x) = A_{\mu}(x) + \frac{1}{e}\partial_{\mu}\theta(x)$ If $\lambda > 0$, $\mu^2 < 0$, the \mathcal{L} in terms of quantum fields η and χ with null VEVs:

$$\phi(x) \equiv \frac{1}{\sqrt{2}} [v + \eta(x) + i\chi(x)], \quad \mu^2 = -\lambda v^2 \qquad \text{Comments:}$$

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{1}{2} (\partial_\mu \eta) (\partial^\mu \eta) + \frac{1}{2} (\partial_\mu \chi) (\partial^\mu \chi) \qquad (i) \quad m_\eta = \sqrt{2\lambda} v \\ m_\chi = 0 \qquad (i) \quad M_A = |eQv| (!) \\ + eQvA_\mu \partial^\mu \chi + eQA_\mu (\eta \partial^\mu \chi - \chi \partial^\mu \eta) \qquad (ii) \quad \text{Term } A_\mu \partial^\mu \chi (?) \\ + \frac{1}{2} (eQv)^2 A_\mu A^\mu + \frac{1}{2} (eQ)^2 A_\mu A^\mu (\eta^2 + 2v\eta + \chi^2) \qquad (iv) \quad \text{Add} \ \mathcal{L}_{\text{GF}}$$

Spontaneous Symmetry Breaking

- gauge symmetry
- Removing the cross term and the (new) gauge fixing Lagrangian:

$$\mathcal{L}_{\mathrm{GF}} = -rac{1}{2\xi} (\partial_{\mu}A^{\mu} - \xi M_A \chi)^2$$

$$\Rightarrow \quad \mathcal{L} + \mathcal{L}_{\mathrm{GF}} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{1}{2} M_A^2 A_\mu A^\mu - \frac{1}{2\xi} (\partial_\mu A^\mu)^2 + \underbrace{\mathcal{M}_A \partial_\mu (A^\mu \chi)}_{+\frac{1}{2}} + \frac{1}{2} (\partial_\mu \chi) (\partial^\mu \chi) - \frac{1}{2} \xi M_A^2 \chi^2 + \dots$$

and the propagators of A_{μ} and χ are:

$$\begin{split} \widetilde{D}_{\mu\nu}(k) &= \frac{\mathrm{i}}{k^2 - M_A^2 + \mathrm{i}\varepsilon} \left[-g_{\mu\nu} + (1 - \xi) \frac{k_\mu k_\nu}{k^2 - \xi M_A^2} \right] \\ \widetilde{D}(k) &= \frac{\mathrm{i}}{k^2 - \xi M_A^2 + \mathrm{i}\varepsilon} \end{split}$$

 $\Rightarrow \chi$ has a gauge-dependent mass: actually it is not a physical field!

 $M_A[\partial_u A^\mu \chi + A_u \partial^\mu \chi]$

Spontaneous Symmetry Breaking gauge symmetry

• A more transparent parameterization of the quantum field ϕ is

$$\phi(x) \equiv e^{iQ\zeta(x)/v} \frac{1}{\sqrt{2}} [v + \eta(x)], \quad \langle 0|\eta |0\rangle = \langle 0|\zeta |0\rangle = 0$$

$$\phi(x) \mapsto e^{-iQ\zeta(x)/v}\phi(x) = \frac{1}{\sqrt{2}}[v + \eta(x)] \Rightarrow \zeta$$
 gauged away!

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \frac{1}{2}(\partial_{\mu}\eta)(\partial^{\mu}\eta)$$

$$(i) \quad m_{\eta} = \sqrt{2\lambda} v$$

$$(i) \quad M_{A} = |eQv|$$

$$+\frac{1}{2}(eQv)^{2}A_{\mu}A^{\mu} + \frac{1}{2}(eQ)^{2}A_{\mu}A^{\mu}(2v\eta + \eta^{2})$$

$$(ii) \quad No \text{ need for } \mathcal{L}_{GF}$$

 \Rightarrow This is the unitary gauge ($\xi \rightarrow \infty$): just physical fields

$$\widetilde{D}_{\mu\nu}(k) \to \frac{\mathrm{i}}{k^2 - M_A^2 + \mathrm{i}\varepsilon} \left[-g_{\mu\nu} + \frac{k_\mu k_\nu}{M_A^2} \right] \quad \text{and} \quad \widetilde{D}(k) \to 0$$

\Rightarrow Brout-Englert-Higgs mechanism:

[Anderson '62] [Higgs '64; Englert, Brout '64; Guralnik, Hagen, Kibble '64]

The gauge bosons associated with the spontaneously broken generators become massive, the corresponding would-be Goldstone bosons are unphysical and can be absorbed, the remaining massive scalars (Higgs bosons) are physical (the smoking gun!)

- The would-be Goldstone bosons are 'eaten up' by the gauge bosons ('get fat') and disappear (gauge away) in the unitary gauge ($\xi \to \infty$)

 \Rightarrow Degrees of freedom are preserved

Before SSB: 2 (massless gauge boson) + 1 (Goldstone boson)

After SSB: 3 (massive gauge boson) + 0 (absorbed would-be Goldstone)

- For loops calculations, 't Hooft-Feynman gauge ($\xi = 1$) is more convenient: \Rightarrow Gauge boson propagators are simpler, but
 - \Rightarrow Goldstone bosons must be included in internal lines

- **Spontaneous Symmetry Breaking** gauge symmetry
- Comments:
 - After SSB the FP ghost fields (unphysical) acquire a gauge-dependent mass, due to interactions with the scalar field(s):

$$\widetilde{D}_{ab}(k) = rac{\mathrm{i}\delta_{ab}}{k^2 - \xi_a M_{W^a}^2 + \mathrm{i}\varepsilon}$$

- Gauge theories with SSB are renormalizable

- ['t Hooft, Veltman '72]
- UV divergences appearing at loop level can be removed by renormalization of parameters and fields of the classical Lagrangian \Rightarrow predictive!

2. The Standard Model

Gauge group and field representations

[Glashow '61; Weinberg '67; Salam '68] [D. Gross, F. Wilczek; D. Politzer '73]

• The Standard Model is a gauge theory based on the local symmetry group:



with the electroweak symmetry spontaneously broken to the electromagnetic $U(1)_Q$ symmetry by the Brout-Englert-Higgs mechanism

• The particle (field) content: (ingredients: 12 *flavors* + 12 gauge bosons + H)

Fermions			I	II	III	Q	Bosons		
spin $\frac{1}{2}$	Quarks	f	uuu	CCC	ttt	$\frac{2}{3}$	spin 1	8 gluons	strong interaction
		f'	ddd	SSS	bbb	$-\frac{1}{3}$		W^{\pm} , Z	weak interaction
	Leptons	f	Ve	ν_{μ}	ν_{τ}	0		γ	em interaction
		f'	e	μ	τ	-1	spin 0	Higgs	origin of mass
	Q_f	$=Q_f$	r + 1			·			

Gauge group and field representations

• The fields lay in the following representations (color, weak isospin, hypercharge):

Multiplets	$\mathrm{SU}(3)_c \otimes \mathrm{SU}(2)_L \otimes \mathrm{U}(1)_Y$	Ι	II	III	$Q = T_3 + Y$		
Quarks	$(3, 2, \frac{1}{6})$	$\begin{pmatrix} u_L \\ d_L \end{pmatrix}$	$\begin{pmatrix} c_L \\ s_L \end{pmatrix}$	$\begin{pmatrix} t_L \\ b_L \end{pmatrix}$	$\frac{\frac{2}{3}}{\frac{1}{2}} = \frac{\frac{1}{2} + \frac{1}{6}}{-\frac{1}{3}} = -\frac{1}{2} + \frac{1}{6}$		
	$(3, 1, \frac{2}{3})$	u_R	C _R	t_R	$\frac{2}{3} = 0 + \frac{2}{3}$		
	$(3, 1, -\frac{1}{3})$	d_R	s _R	b_R	$-\frac{1}{3} = 0 - \frac{1}{3}$		
Leptons	$(1, 2, -\frac{1}{2})$	$\begin{pmatrix} \nu_{e_L} \\ e_L \end{pmatrix}$	$\begin{pmatrix} \nu_{\mu_L} \\ \mu_L \end{pmatrix}$	$\begin{pmatrix} \nu_{\tau_L} \\ \tau_L \end{pmatrix}$	$0 = \frac{1}{2} - \frac{1}{2} \\ -1 = -\frac{1}{2} - \frac{1}{2}$		
	(1 , 1 , −1)	e_R	μ_R	$ au_R$	-1 = 0 - 1		
	(1, 1, 0)	ν_{e_R}	ν_{μ_R}	$v_{ au_R}$	0 = 0 + 0		
Higgs	$(1, 2, \frac{1}{2})$	2, $\frac{1}{2}$) (3 families of quarks & leptons)					

 \Rightarrow Electroweak (QFD): SU(2)_L \otimes U(1)_Y Strong (QCD): SU(3)_c

Electroweak interactions

The EWSM with one family (of quarks or leptons)

• Consider two massless fermion fields f(x) and f'(x) with electric charges $Q_f = Q_{f'} + 1$ in three irreps of SU(2)_L \otimes U(1)_Y:

$$\mathcal{L}_{F}^{0} = i\overline{f}\partial f + i\overline{f}'\partial f' \qquad f_{R,L} = \frac{1}{2}(1\pm\gamma_{5})f, \quad f_{R,L}' = \frac{1}{2}(1\pm\gamma_{5})f'$$
$$= i\overline{\Psi}_{1}\partial \Psi_{1} + i\overline{\psi}_{2}\partial \psi_{2} + i\overline{\psi}_{3}\partial \psi_{3} \quad ; \quad \Psi_{1} = \underbrace{\begin{pmatrix}f_{L}\\f_{L}'\end{pmatrix}}_{(\mathbf{2},y_{1})}, \quad \psi_{2} = \underbrace{f_{R}}_{(\mathbf{1},y_{2})}, \quad \psi_{3} = \underbrace{f_{R}'}_{(\mathbf{1},y_{3})}$$

• To get a Langrangian invariant under gauge transformations:

$$\begin{split} \Psi_1(x) &\mapsto U_L(x) e^{-iy_1\beta(x)} \Psi_1(x), \quad U_L(x) = e^{-iT_i\alpha^i(x)}, \quad T_i = \frac{\sigma_i}{2} \quad \text{(weak isospin gen.)} \\ \psi_2(x) &\mapsto e^{-iy_2\beta(x)} \psi_2(x) \\ \psi_3(x) &\mapsto e^{-iy_3\beta(x)} \psi_3(x) \end{split}$$

The EWSM with one family gauge invariance

 \Rightarrow Introduce gauge fields $W^i_{\mu}(x)$ (*i* = 1,2,3) and $B_{\mu}(x)$ through covariant derivatives:

$$\begin{aligned} D_{\mu}\Psi_{1} &= (\partial_{\mu} - ig\widetilde{W}_{\mu} + ig'y_{1}B_{\mu})\Psi_{1}, \quad \widetilde{W}_{\mu} \equiv \frac{\sigma_{i}}{2}W_{\mu}^{i} \\ D_{\mu}\psi_{2} &= (\partial_{\mu} + ig'y_{2}B_{\mu})\psi_{2} \\ D_{\mu}\psi_{3} &= (\partial_{\mu} + ig'y_{3}B_{\mu})\psi_{3} \end{aligned} \qquad \Rightarrow \qquad \mathcal{L}_{F} \quad (\mathcal{P}, \mathcal{C})$$

where two couplings g and g' have been introduced and

$$\widetilde{W}_{\mu}(x) \mapsto U_{L}(x)\widetilde{W}_{\mu}(x)U_{L}^{\dagger}(x) - \frac{\mathrm{i}}{g}(\partial_{\mu}U_{L}(x))U_{L}^{\dagger}(x)$$
$$B_{\mu}(x) \mapsto B_{\mu}(x) + \frac{1}{g'}\partial_{\mu}\beta(x)$$

 \Rightarrow Add Yang-Mills: gauge invariant kinetic terms for the gauge fields

$$\mathcal{L}_{\rm YM} = -\frac{1}{4} W^i_{\mu\nu} W^{i,\mu\nu} - \frac{1}{4} B_{\mu\nu} B^{\mu\nu} , \quad W^i_{\mu\nu} = \partial_\mu W^i_\nu - \partial_\nu W^i_\mu + g \epsilon_{ijk} W^j_\mu W^k_\nu$$

(include self-interactions of the SU(2) gauge fields) and $B_{\mu\nu} = \partial_{\mu}B_{\nu} - \partial_{\nu}B_{\mu}$

The EWSM with one family mass terms forbidden

⇒ Note that mass terms are not invariant under $SU(2)_L \otimes U(1)_Y$, since LH and RH components do not transform the same:

$$m\overline{f}f = m(\overline{f_L}f_R + \overline{f_R}f_L)$$

- \Rightarrow Mass terms for the gauge bosons are not allowed either
- ⇒ Next the different types of interactions are analyzed, and later the EWSB will be discussed

The EWSM with one family | charged current interactions

•
$$\mathcal{L}_F \supset g \overline{\Psi}_1 \gamma^{\mu} \widetilde{W}_{\mu} \Psi_1 , \quad \widetilde{W}_{\mu} = \frac{1}{2} \begin{pmatrix} W_{\mu}^3 & \sqrt{2} W_{\mu}^{\dagger} \\ \sqrt{2} W_{\mu} & -W_{\mu}^3 \end{pmatrix}$$

 \Rightarrow charged current interactions of LH fermions with complex vector boson field W_{μ} :



The EWSM with one family neutral current of the end of

- neutral current interactions
- The diagonal part of $g\overline{\Psi}_1\gamma^{\mu}\widetilde{W}_{\mu}\Psi_1$ and the remaining terms

$$\mathcal{L}_F \supset \frac{1}{2} g \overline{\Psi}_1 \gamma^{\mu} \sigma_3 W^3_{\mu} \Psi_1 - g' B_{\mu} (y_1 \overline{\Psi}_1 \gamma^{\mu} \Psi_1 + y_2 \overline{\psi}_2 \gamma^{\mu} \psi_2 + y_3 \overline{\psi}_3 \gamma^{\mu} \psi_3)$$

 \Rightarrow neutral current interactions with neutral vector boson fields W_{μ}^3 and B_{μ} We would like to identify B_{μ} with the photon field A_{μ} but that requires:

$$y_1 = y_2 = y_3$$
 and $g'y_j = eQ_j \Rightarrow$ impossible!

 \Rightarrow Since they are both neutral, try a combination:

$$\begin{pmatrix} W_{\mu}^{3} \\ B_{\mu} \end{pmatrix} \equiv \begin{pmatrix} c_{W} & -s_{W} \\ s_{W} & c_{W} \end{pmatrix} \begin{pmatrix} Z_{\mu} \\ A_{\mu} \end{pmatrix} \qquad s_{W} \equiv \sin \theta_{W} , \quad c_{W} \equiv \cos \theta_{W} \\ \theta_{W} = \text{weak mixing angle} \\ \mathcal{L}_{\text{NC}} = \sum_{j=1}^{3} \overline{\psi}_{j} \gamma^{\mu} \left\{ - \left[gT_{3}s_{W} + g'y_{j}c_{W} \right] A_{\mu} + \left[gT_{3}c_{W} - g'y_{j}s_{W} \right] Z_{\mu} \right\} \psi_{j} \\ T_{W} = \begin{pmatrix} \sigma_{3} \\ \sigma_{3} \end{pmatrix} (0) \text{ the third coefficiency is expressed of the characteristic structure}$$

with $T_3 = \frac{\sigma_3}{2}$ (0) the third weak isospin component of the doublet (singlet)

The EWSM with one family

- neutral current interactions
- To make A_{μ} the photon field:

(1)
$$e = gs_W = g'c_W$$
 (2) $Q = T_3 + Y$

where the electric charge operator is: $Q_1 = \begin{pmatrix} Q_f & 0 \\ 0 & Q_{f'} \end{pmatrix}$, $Q_2 = Q_f$, $Q_3 = Q_{f'}$

 $\Rightarrow (1) \text{ Electroweak unification: } g \text{ of SU(2) and } g' \text{ of U(1) related to } e = \frac{gg'}{\sqrt{g^2 + g'^2}}$

 \Rightarrow (2) The hyperchages are fixed in terms of electric charges and weak isospin:

$$y_1 = Q_f - \frac{1}{2} = Q_{f'} + \frac{1}{2}$$
, $y_2 = Q_f$, $y_3 = Q_{f'}$

$$\mathcal{L}_{\text{QED}} = -e \ Q_f \overline{f} \gamma^{\mu} f \ A_{\mu} \quad + (f \to f')$$

 \Rightarrow RH neutrinos are sterile: $y_2 = Q_f = 0$

The EWSM with one family neutral current interactions

• The Z_{μ} is the neutral weak boson field:

$$\mathcal{L}_{\rm NC}^Z = e \, \overline{f} \gamma^\mu (v_f - a_f \gamma_5) f \, Z_\mu \quad + (f \to f')$$

with

$$v_f = rac{T_3^{f_L} - 2Q_f s_W^2}{2s_W c_W}$$
 , $a_f = rac{T_3^{f_L}}{2s_W c_W}$

• The complete neutral current Lagrangian reads:

$$\mathcal{L}_{\mathrm{NC}} = \mathcal{L}_{\mathrm{QED}} + \mathcal{L}_{\mathrm{NC}}^{Z}$$

The EWSM with one family gauge boson self-interactions

• Cubic:

$$\mathcal{L}_{\rm YM} \supset \mathcal{L}_3 = -\frac{\mathrm{i}ec_W}{s_W} \left\{ W^{\mu\nu} W^{\dagger}_{\mu} Z_{\nu} - W^{\dagger}_{\mu\nu} W^{\mu} Z^{\nu} - W^{\dagger}_{\mu} W_{\nu} Z^{\mu\nu} \right\}$$
$$+ \mathrm{i}e \left\{ W^{\mu\nu} W^{\dagger}_{\mu} A_{\nu} - W^{\dagger}_{\mu\nu} W^{\mu} A^{\nu} - W^{\dagger}_{\mu} W_{\nu} F^{\mu\nu} \right\}$$

with

$$F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu} \qquad Z_{\mu\nu} = \partial_{\mu}Z_{\nu} - \partial_{\nu}Z_{\mu} \qquad W_{\mu\nu} = \partial_{\mu}W_{\nu} - \partial_{\nu}W_{\mu}$$

The EWSM with one family

gauge boson self-interactions

• Quartic:

$$\begin{split} \mathcal{L}_{\rm YM} \supset \mathcal{L}_4 &= -\frac{e^2}{2s_W^2} \left\{ \left(W^{\dagger}_{\mu} W^{\mu} \right)^2 - W^{\dagger}_{\mu} W^{\mu \dagger} W_{\nu} W^{\nu} \right\} \\ &- \frac{e^2 c_W^2}{s_W^2} \left\{ W^{\dagger}_{\mu} W^{\mu} Z_{\nu} Z^{\nu} - W^{\dagger}_{\mu} Z^{\mu} W_{\nu} Z^{\nu} \right\} \\ &+ \frac{e^2 c_W}{s_W} \left\{ 2 W^{\dagger}_{\mu} W^{\mu} Z_{\nu} A^{\nu} - W^{\dagger}_{\mu} Z^{\mu} W_{\nu} A^{\nu} - W^{\dagger}_{\mu} A^{\mu} W_{\nu} Z^{\nu} \right\} \\ &- e^2 \left\{ W^{\dagger}_{\mu} W^{\mu} A_{\nu} A^{\nu} - W^{\dagger}_{\mu} A^{\mu} W_{\nu} A^{\nu} \right\} \end{split}$$



Note: even number of *W* and no vertex with just γ or *Z*

Electroweak symmetry breaking setup

- Out of the 4 gauge bosons of $SU(2)_L \otimes U(1)_Y$ with generators T_1 , T_2 , T_3 , Y we need all to be broken except the combination $Q = T_3 + Y$ so that A_μ remains massless and the other three gauge bosons get massive after SSB
 - \Rightarrow Introduce a complex SU(2) Higgs doublet

$$\Phi = egin{pmatrix} \phi^+ \ \phi^0 \end{pmatrix} \;, \;\;\; \langle 0 | \, \Phi \, | 0
angle = rac{1}{\sqrt{2}} egin{pmatrix} 0 \ v \end{pmatrix}$$

with gauge invariant Lagrangian ($\mu^2 = -\lambda v^2$):

$$\mathcal{L}_{\Phi} = (D_{\mu}\Phi)^{\dagger}D^{\mu}\Phi - \mu^{2}\Phi^{\dagger}\Phi - \lambda(\Phi^{\dagger}\Phi)^{2}, \qquad D_{\mu}\Phi = (\partial_{\mu} - ig\widetilde{W}_{\mu} + ig'y_{\Phi}B_{\mu})\Phi$$

take
$$y_{\Phi} = \frac{1}{2} \Rightarrow (T_3 + Y) |0\rangle = 0 \Leftrightarrow \begin{pmatrix} 1 & 0 \\ 0 & 0 \end{pmatrix} \begin{pmatrix} 0 \\ v \end{pmatrix} = 0 \checkmark$$

 $\{T_1, T_2, T_3 - Y\} |0\rangle \neq 0$

Electroweak symmetry breaking

• Quantum fields in the unitary gauge:

$$\Phi(x) \equiv \exp\left\{i\frac{\sigma_i}{2v}\theta^i(x)\right\} \frac{1}{\sqrt{2}} \begin{pmatrix} 0\\ v+H(x) \end{pmatrix}$$

$$\Phi(x) \mapsto \exp\left\{-i\frac{\sigma_i}{2v}\theta^i(x)\right\}\Phi(x) = \frac{1}{\sqrt{2}}\begin{pmatrix}0\\v+H(x)\end{pmatrix} \Rightarrow$$

 physical Higgs field H(x)
 would-be Goldstones θⁱ(x) gauged away

– The 3 dof apparently lost become the longitudinal polarizations of W^{\pm} and Z that get massive after SSB:

$$\mathcal{L}_{\Phi} \supset \mathcal{L}_{M} = \underbrace{\frac{g^{2}v^{2}}{4}}_{M_{W}^{2}} W^{\dagger}_{\mu} W^{\mu} + \underbrace{\frac{g^{2}v^{2}}{8c_{W}^{2}}}_{\frac{1}{2}M_{Z}^{2}} Z_{\mu} Z^{\mu} \quad \Rightarrow \quad \underbrace{M_{W} = M_{Z}c_{W}}_{\text{custodial}} = \frac{1}{2}gv$$

Electroweak symmetry breaking | Higgs sector

 \Rightarrow In the unitary gauge (just physical fields): $\mathcal{L}_{\Phi} = \mathcal{L}_{H} + \mathcal{L}_{M} + \mathcal{L}_{HV^{2}} + \frac{1}{4}\lambda v^{4}$



$$\mathcal{L}_{M} + \mathcal{L}_{HV^{2}} = M_{W}^{2} W_{\mu}^{\dagger} W^{\mu} \left\{ 1 + \frac{2}{v} H + \frac{H^{2}}{v^{2}} \right\} + \frac{1}{2} M_{Z}^{2} Z_{\mu} Z^{\mu} \left\{ 1 + \frac{2}{v} H + \frac{H^{2}}{v^{2}} \right\}$$

$$H - \cdots \int_{W}^{W} H = \cdots \int_{W}^{W} H + \cdots \int_{Z}^{W} H + \cdots \int_{$$

Electroweak symmetry breaking

• Quantum fields in the R_{ξ} gauges:

$$\Phi(x) = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix} \equiv \begin{pmatrix} \phi^+(x) \\ \frac{1}{\sqrt{2}} [v + H(x) + i\chi(x)] \end{pmatrix} , \quad \phi^-(x) = [\phi^+(x)]^*$$

$$\begin{split} \mathcal{L}_{\Phi} &= \mathcal{L}_{H} + \mathcal{L}_{M} + \mathcal{L}_{HV^{2}} + \frac{1}{4}\lambda v^{4} \\ &+ (\partial_{\mu}\phi^{+})(\partial^{\mu}\phi^{-}) + \frac{1}{2}(\partial_{\mu}\chi)(\partial^{\mu}\chi) \\ &+ iM_{W} \left(W_{\mu}\partial^{\mu}\phi^{+} - W_{\mu}^{\dagger}\partial^{\mu}\phi^{-}\right) + M_{Z} Z_{\mu}\partial^{\mu}\chi \end{split}$$

+ trilinear interactions [SSS, SSV, SVV]

+ quadrilinear interactions [SSSS, SSVV]

Electroweak symmetry breaking gauge fixing

• To remove the cross terms $W_{\mu}\partial^{\mu}\phi^{+}$, $W_{\mu}^{\dagger}\partial^{\mu}\phi^{-}$, $Z_{\mu}\partial^{\mu}\chi$ and define propagators add:

$$\mathcal{L}_{GF} = -\frac{1}{2\xi_{\gamma}} (\partial_{\mu}A^{\mu})^2 - \frac{1}{2\xi_{Z}} (\partial_{\mu}Z^{\mu} - \xi_{Z}M_{Z}\chi)^2 - \frac{1}{\xi_{W}} |\partial_{\mu}W^{\mu} + i\xi_{W}M_{W}\phi^{-}|^2$$

 \Rightarrow Massive propagators for gauge and (unphysical) would-be Goldstone fields:

$$\begin{split} \widetilde{D}_{\mu\nu}^{\gamma}(k) &= \frac{i}{k^2 + i\varepsilon} \left[-g_{\mu\nu} + (1 - \xi_{\gamma}) \frac{k_{\mu}k_{\nu}}{k^2} \right] \\ \widetilde{D}_{\mu\nu}^{Z}(k) &= \frac{i}{k^2 - M_Z^2 + i\varepsilon} \left[-g_{\mu\nu} + (1 - \xi_Z) \frac{k_{\mu}k_{\nu}}{k^2 - \xi_Z M_Z^2} \right] \quad ; \quad \widetilde{D}^{\chi}(k) \ = \frac{i}{k^2 - \xi_Z M_Z^2 + i\varepsilon} \\ \widetilde{D}_{\mu\nu}^{W}(k) &= \frac{i}{k^2 - M_W^2 + i\varepsilon} \left[-g_{\mu\nu} + (1 - \xi_W) \frac{k_{\mu}k_{\nu}}{k^2 - \xi_W M_W^2} \right] \quad ; \quad \widetilde{D}^{\phi}(k) \ = \frac{i}{k^2 - \xi_W M_W^2 + i\varepsilon} \end{split}$$

('t Hooft-Feynman gauge: $\xi_{\gamma} = \xi_Z = \xi_W = 1$)

Electroweak symmetry breaking | Faddeev-Popov ghosts

• The SM is a non-Abelian theory \Rightarrow add Faddeev-Popov ghosts $c_i(x)$ (i = 1, 2, 3)

$$c_{1} \equiv \frac{1}{\sqrt{2}}(u_{+} + u_{-}), \quad c_{2} \equiv \frac{i}{\sqrt{2}}(u_{+} - u_{-}), \quad c_{3} \equiv c_{W} u_{Z} - s_{W} u_{\gamma}$$
$$\underbrace{\mathcal{L}_{FP}}_{U \text{ FP}} = \underbrace{(\partial^{\mu} \overline{c}_{i})(\partial_{\mu} c_{i} - g \epsilon_{ijk} c_{j} W_{\mu}^{k})}_{U \text{ kinetic } + [UUV]} + \underbrace{\text{interactions with } \Phi}_{U \text{ masses } + [SUU]}$$

 \Rightarrow Massive propagators for (unphysical) FP ghost fields:

$$\widetilde{D}^{u_{\gamma}}(k) = \frac{\mathrm{i}}{k^2 + \mathrm{i}\varepsilon} , \quad \widetilde{D}^{u_{Z}}(k) = \frac{\mathrm{i}}{k^2 - \xi_Z M_Z^2 + \mathrm{i}\varepsilon} , \quad \widetilde{D}^{u_{\pm}}(k) = \frac{\mathrm{i}}{k^2 - \xi_W M_W^2 + \mathrm{i}\varepsilon}$$

('t Hooft-Feynman gauge: $\xi_Z = \xi_W = 1$)

$$\begin{aligned} \mathcal{L}_{\mathrm{FP}} &= (\partial_{\mu}\overline{u}_{\gamma})(\partial^{\mu}u_{\gamma}) + (\partial_{\mu}\overline{u}_{Z})(\partial^{\mu}u_{Z}) + (\partial_{\mu}\overline{u}_{+})(\partial^{\mu}u_{+}) + (\partial_{\mu}\overline{u}_{-})(\partial^{\mu}u_{-}) \\ &+ \mathrm{i}e[(\partial^{\mu}\overline{u}_{+})u_{+} - (\partial^{\mu}\overline{u}_{-})u_{-}]A_{\mu} - \frac{\mathrm{i}ec_{W}}{s_{W}}[(\partial^{\mu}\overline{u}_{+})u_{+} - (\partial^{\mu}\overline{u}_{-})u_{-}]Z_{\mu} \\ &- \mathrm{i}e[(\partial^{\mu}\overline{u}_{+})u_{\gamma} - (\partial^{\mu}\overline{u}_{\gamma})u_{-}]W_{\mu}^{\dagger} + \frac{\mathrm{i}ec_{W}}{s_{W}}[(\partial^{\mu}\overline{u}_{+})u_{Z} - (\partial^{\mu}\overline{u}_{Z})u_{-}]W_{\mu}^{\dagger} \\ &+ \mathrm{i}e[(\partial^{\mu}\overline{u}_{-})u_{\gamma} - (\partial^{\mu}\overline{u}_{\gamma})u_{+}]W_{\mu} - \frac{\mathrm{i}ec_{W}}{s_{W}}[(\partial^{\mu}\overline{u}_{-})u_{Z} - (\partial^{\mu}\overline{u}_{Z})u_{+}]W_{\mu} \\ &- \xi_{Z}M_{Z}^{2} \ \overline{u}_{Z}u_{Z} - \xi_{W}M_{W}^{2} \ \overline{u}_{+}u_{+} - \xi_{W}M_{W}^{2} \ \overline{u}_{-}u_{-} \\ &\left\{ -e\xi_{Z}M_{Z} \ \overline{u}_{Z} \left[\frac{1}{2s_{W}c_{W}}Hu_{Z} - \frac{1}{2s_{W}}(\phi^{+}u_{-} + \phi^{-}u_{+}) \right] \\ &- e\xi_{W}M_{W} \ \overline{u}_{+} \left[\frac{1}{2s_{W}}(H + \mathrm{i}\chi)u_{+} - \phi^{+} \left(u_{\gamma} - \frac{c_{W}^{2} - s_{W}^{2}}{2s_{W}c_{W}}u_{Z} \right) \right] \\ &- e\xi_{W}M_{W} \ \overline{u}_{-} \left[\frac{1}{2s_{W}}(H - \mathrm{i}\chi)u_{-} - \phi^{-} \left(u_{\gamma} - \frac{c_{W}^{2} - s_{W}^{2}}{2s_{W}c_{W}}u_{Z} \right) \right] \end{aligned}$$

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Electroweak symmetry breaking fermion masses

- We need masses for quarks and leptons without breaking gauge symmetry
 - \Rightarrow Introduce Yukawa interactions:

$$\begin{aligned} \mathcal{L}_{\mathrm{Y}} &= -\lambda_{d} \begin{pmatrix} \overline{u}_{L} & \overline{d}_{L} \end{pmatrix} \begin{pmatrix} \phi^{+} \\ \phi^{0} \end{pmatrix} d_{R} - \lambda_{u} \begin{pmatrix} \overline{u}_{L} & \overline{d}_{L} \end{pmatrix} \begin{pmatrix} \phi^{0*} \\ -\phi^{-} \end{pmatrix} u_{R} \\ &-\lambda_{\ell} \begin{pmatrix} \overline{\nu}_{L} & \overline{\ell}_{L} \end{pmatrix} \begin{pmatrix} \phi^{+} \\ \phi^{0} \end{pmatrix} \ell_{R} - \lambda_{\nu} \begin{pmatrix} \overline{\nu}_{L} & \overline{\ell}_{L} \end{pmatrix} \begin{pmatrix} \phi^{0*} \\ -\phi^{-} \end{pmatrix} \nu_{R} &+ \mathrm{h.c.} \end{aligned}$$

where
$$\widetilde{\Phi} \equiv i\sigma_2 \Phi^* = \begin{pmatrix} \phi^{0*} \\ -\phi^- \end{pmatrix}$$
 transforms under SU(2) like $\Phi = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix}$

 \Rightarrow After EW SSB, fermions acquire masses ($\overline{f}f = \overline{f_L}f_R + \overline{f_R}f_L$):

$$\mathcal{L}_{\mathbf{Y}} \supset -\frac{1}{\sqrt{2}}(v+H) \left\{ \lambda_d \ \overline{d}d + \lambda_u \ \overline{u}u + \lambda_\ell \ \overline{\ell}\ell + \lambda_\nu \ \overline{\nu}\nu \right\} \quad \Rightarrow \quad m_f = \lambda_f \frac{v}{\sqrt{2}}$$

Additional generations Yukawa matrices

- There are 3 generations of quarks and leptons in Nature. They are identical copies with the same properties under $SU(2)_L \otimes U(1)_Y$ differing only in their masses
 - ⇒ Take a general case of *n* generations and let u_i^I , d_i^I , v_i^I , ℓ_i^I be the members of family *i* (*i* = 1,...,*n*). Superindex *I* (interaction basis) was omitted so far
 - \Rightarrow General gauge invariant Yukawa Lagrangian:

$$\begin{split} \mathbf{\mathcal{L}}_{\mathbf{Y}} &= -\sum_{ij} \left\{ \begin{pmatrix} \overline{u}_{iL}^{I} & \overline{d}_{iL}^{I} \end{pmatrix} \begin{bmatrix} \begin{pmatrix} \phi^{+} \\ \phi^{0} \end{pmatrix} \lambda_{ij}^{(d)} d_{jR}^{I} + \begin{pmatrix} \phi^{0*} \\ -\phi^{-} \end{pmatrix} \lambda_{ij}^{(u)} u_{jR}^{I} \end{bmatrix} \\ &+ \begin{pmatrix} \overline{\nu}_{iL}^{I} & \overline{\ell}_{iL}^{I} \end{pmatrix} \begin{bmatrix} \begin{pmatrix} \phi^{+} \\ \phi^{0} \end{pmatrix} \lambda_{ij}^{(\ell)} \ell_{jR}^{I} + \begin{pmatrix} \phi^{0*} \\ -\phi^{-} \end{pmatrix} \lambda_{ij}^{(\nu)} \nu_{jR}^{I} \end{bmatrix} \right\} + \text{h.c.} \end{split}$$

where $\lambda_{ij}^{(d)}$, $\lambda_{ij}^{(u)}$, $\lambda_{ij}^{(\ell)}$, $\lambda_{ij}^{(\nu)}$ are arbitrary Yukawa matrices

Additional generations | mass matrices

• After EW SSB, in *n*-dimensional matrix form:

$$\mathcal{L}_{\mathrm{Y}} \supset -\left(1+\frac{H}{v}\right) \left\{ \,\overline{\mathbf{d}}_{L}^{I} \,\mathbf{M}_{d} \,\mathbf{d}_{R}^{I} \,+\, \overline{\mathbf{u}}_{L}^{I} \,\mathbf{M}_{u} \,\mathbf{u}_{R}^{I} \,+\, \overline{\ell}_{L}^{I} \,\mathbf{M}_{\ell} \,\boldsymbol{\ell}_{R}^{I} \,+\, \overline{\boldsymbol{\nu}}_{L}^{I} \,\mathbf{M}_{\nu} \,\boldsymbol{\nu}_{R}^{I} \,+\, \mathrm{h.c.} \right\}$$

with mass matrices

$$(\mathbf{M}_d)_{ij} \equiv \lambda_{ij}^{(d)} \frac{\upsilon}{\sqrt{2}} \quad (\mathbf{M}_u)_{ij} \equiv \lambda_{ij}^{(u)} \frac{\upsilon}{\sqrt{2}} \quad (\mathbf{M}_\ell)_{ij} \equiv \lambda_{ij}^{(\ell)} \frac{\upsilon}{\sqrt{2}} \quad (\mathbf{M}_\nu)_{ij} \equiv \lambda_{ij}^{(\nu)} \frac{\upsilon}{\sqrt{2}}$$

- ⇒ Diagonalization determines mass eigenstates d_j , u_j , ℓ_j , ν_j in terms of interaction states d_j^I , u_j^I , ℓ_j^I , ν_j^I , respectively
- \Rightarrow Each **M**_{*f*} can be written as

$$\mathbf{M}_f = \mathbf{H}_f \,\mathcal{U}_f = \mathbf{V}_f^{\dagger} \,\mathcal{M}_f \,\mathbf{V}_f \,\mathcal{U}_f \quad \Longleftrightarrow \quad \mathbf{M}_f \mathbf{M}_f^{\dagger} = \mathbf{H}_f^2 = \mathbf{V}_f^{\dagger} \,\mathcal{M}_f^2 \,\mathbf{V}_f$$

with $\mathbf{H}_f \equiv \sqrt{\mathbf{M}_f \mathbf{M}_f^{\dagger}}$ a Hermitian positive definite matrix and \mathcal{U}_f unitary

- Every \mathbf{H}_{f} can be diagonalized by a unitary matrix \mathbf{V}_{f}
- The resulting \mathcal{M}_f is diagonal and positive definite

Additional generations | fermion masses and mixings

• In terms of diagonal mass matrices (mass eigenstate basis):

$$\mathcal{M}_{d} = \operatorname{diag}(m_{d}, m_{s}, m_{b}, \ldots) , \quad \mathcal{M}_{u} = \operatorname{diag}(m_{u}, m_{c}, m_{t}, \ldots)$$
$$\mathcal{M}_{\ell} = \operatorname{diag}(m_{e}, m_{\mu}, m_{\tau}, \ldots) , \quad \mathcal{M}_{\nu} = \operatorname{diag}(m_{\nu_{e}}, m_{\nu_{\mu}}, m_{\nu_{\tau}}, \ldots)$$

$$\mathcal{L}_{\mathrm{Y}} \supset -\left(1 + \frac{H}{v}\right) \left\{ \,\overline{\mathbf{d}} \,\mathcal{M}_{d} \,\mathbf{d} \,+\, \overline{\mathbf{u}} \,\mathcal{M}_{u} \,\mathbf{u} \,+\, \overline{\ell} \,\mathcal{M}_{\ell} \,\ell + \overline{\nu} \,\mathcal{M}_{v} \,\nu \,\right\}$$

where fermion couplings to Higgs are proportional to masses and

$$\mathbf{d}_{L} \equiv \mathbf{V}_{d} \ \mathbf{d}_{L}^{I} \qquad \mathbf{u}_{L} \equiv \mathbf{V}_{u} \ \mathbf{u}_{L}^{I} \qquad \boldsymbol{\ell}_{L} \equiv \mathbf{V}_{\ell} \ \boldsymbol{\ell}_{L}^{I} \qquad \boldsymbol{\nu}_{L} \equiv \mathbf{V}_{\nu} \ \boldsymbol{\nu}_{L}^{I} \\ \mathbf{d}_{R} \equiv \mathbf{V}_{d} \mathcal{U}_{d} \ \mathbf{d}_{R}^{I} \qquad \mathbf{u}_{R} \equiv \mathbf{V}_{u} \mathcal{U}_{u} \ \mathbf{u}_{R}^{I} \qquad \boldsymbol{\ell}_{R} \equiv \mathbf{V}_{\ell} \mathcal{U}_{\ell} \ \boldsymbol{\ell}_{R}^{I} \qquad \boldsymbol{\nu}_{R} \equiv \mathbf{V}_{\nu} \mathcal{U}_{\nu} \ \boldsymbol{\nu}_{R}^{I}$$

 $\Rightarrow \text{ Neutral Currents preserve chirality}} \\ \overline{\mathbf{f}}_{L}^{I} \gamma^{\mu} \mathbf{f}_{L}^{I} = \overline{\mathbf{f}}_{L} \gamma^{\mu} \mathbf{f}_{L} \text{ and } \overline{\mathbf{f}}_{R}^{I} \gamma^{\mu} \mathbf{f}_{R}^{I} = \overline{\mathbf{f}}_{R} \gamma^{\mu} \mathbf{f}_{R} \end{array} \right\} \Rightarrow \mathcal{L}_{\text{NC}} \text{ does not change family}$

 \Rightarrow GIM mechanism

[Glashow, Iliopoulos, Maiani '70]

Additional generations quark sector

• However, in Charged Currents (also chirality preserving and only LH):

$$\overline{\mathbf{u}}_{L}^{I}\gamma^{\mu}\,\mathbf{d}_{L}^{I}=\overline{\mathbf{u}}_{L}\gamma^{\mu}\,\mathbf{V}_{u}\,\mathbf{V}_{d}^{\dagger}\mathbf{d}_{L}=\overline{\mathbf{u}}_{L}\gamma^{\mu}\,\mathbf{V}\mathbf{d}_{L}$$

with $\mathbf{V} \equiv \mathbf{V}_u \mathbf{V}_d^{\dagger}$ the (unitary) CKM mixing matrix [Cabibbo '63; Kobayashi, Maskawa '73]

$$\Rightarrow \quad \mathcal{L}_{\rm CC} = \frac{g}{2\sqrt{2}} \sum_{ij} \overline{u}_i \gamma^{\mu} (1 - \gamma_5) \, \mathbf{V}_{ij} \, d_j \, W^{\dagger}_{\mu} + \text{h.c.}$$



⇒ If u_i or d_j had degenerate masses one could choose $\mathbf{V}_u = \mathbf{V}_d$ (field redefinition) and quark families would not mix. But they are *not degenerate*, so they mix! ⇒ \mathbf{V}_u and \mathbf{V}_d are not observable. Just masses and CKM mixings are observable

Additional generations quark sector

- How many physical parameters in this sector?
 - Quark masses and CKM mixings determined by mass (or Yukawa) matrices
 - A general $n \times n$ unitary matrix, like the CKM, is given by

 n^2 real parameters = n(n-1)/2 moduli + n(n+1)/2 phases

Some phases are unphysical since they can be absorbed by field redefinitions:

$$u_i \to e^{i\phi_i} u_i$$
, $d_j \to e^{i\theta_j} d_j \Rightarrow \mathbf{V}_{ij} \to \mathbf{V}_{ij} e^{i(\theta_j - \phi_i)}$

Therefore 2n - 1 unphysical phases and the physical parameters are:

$$(n-1)^2 = n(n-1)/2 \text{ moduli } + (n-1)(n-2)/2 \text{ phases}$$
Additional generations quark sector

 \Rightarrow Case of n = 2 generations: 1 parameter, the Cabibbo angle θ_C :

$$\mathbf{V} = \begin{pmatrix} \cos \theta_C & \sin \theta_C \\ -\sin \theta_C & \cos \theta_C \end{pmatrix}$$

 \Rightarrow Case of n = 3 generations: 3 angles + 1 phase. In the standard parameterization:

$$\mathbf{V} = \begin{pmatrix} \mathbf{V}_{ud} & \mathbf{V}_{us} & \mathbf{V}_{ub} \\ \mathbf{V}_{cd} & \mathbf{V}_{cs} & \mathbf{V}_{cb} \\ \mathbf{V}_{td} & \mathbf{V}_{ts} & \mathbf{V}_{tb} \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$
$$= \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \Rightarrow \begin{array}{l} \delta \text{ only source} \\ \Rightarrow \text{ of CP violation} \\ \text{ in the SM !} \\ \text{with } c_{ij} \equiv \cos \theta_{ij} \ge 0, \quad s_{ij} \equiv \sin \theta_{ij} \ge 0 \quad (i < j = 1, 2, 3) \quad \text{and } 0 \le \delta \le 2\pi \end{cases}$$

- If neutrinos were massless we could redefine the (LH) fields ⇒ no lepton mixing.
 However they *are* massive (though *very light* masses) ← neutrino oscillations!
 - vSM (introduce v_R and get masses from *tiny* Yukawa couplings like quarks) Alternatively ...
- Neutrinos are special: [see next section] they *may* be their own antiparticle (Majorana) since they are neutral fermions
 - \Rightarrow New mechanisms for generation of masses and mixings
 - * Mass terms are different to Dirac case
 - * Neutrinos and antineutrinos *may* mix
 - * Intergenerational mixings are richer (more CP phases)
 - If they are Majorana physics Beyond ν SM

(seesaw mechanism?)

- What we know about neutrinos:
 - From Z lineshape: n = 3 generations of *active* v_L [v_i (i = 1, ..., n)] (but we do not know (*yet*) if neutrinos are Dirac or Majorana fermions)
 - From oscillations: active neutrinos are very light, non degenerate and mix
 PMNS matrix U [Pontecorvo '57; Maki, Nakagawa, Sakata '62; Pontecorvo '68]

$$|\nu_{\alpha}\rangle = \sum_{i} \mathbf{U}_{\alpha i} |\nu_{i}\rangle \quad \Longleftrightarrow \quad |\nu_{i}\rangle = \sum_{\alpha} \mathbf{U}_{\alpha i}^{*} |\nu_{\alpha}\rangle$$

mass eigenstates v_i (i = 1, 2, 3) / interaction states v_{α} ($\alpha = e, \mu, \tau$)

- $\Rightarrow If neutrinos were Majorana U seems unitary (for negligible light-heavy mixings) and analogous to <math>V_u$, V_d , V_ℓ defined for quarks and charged leptons except for:
 - ν fields have both chiralities: $\nu_i = \nu_{iL} + \eta_i \nu_{iL}^c$
 - If ν 's are Majorana, **U** may contain two additional physical (Majorana) phases that *cannot be absorbed* since then field phases are fixed by $\nu_i = \eta_i \nu_i^c$

 \Rightarrow Standard parameterization of the PMNS matrix:

$$\mathbf{U} = \begin{pmatrix} \mathbf{U}_{e1} & \mathbf{U}_{e2} & \mathbf{U}_{e3} \\ \mathbf{U}_{\mu 1} & \mathbf{U}_{\mu 2} & \mathbf{U}_{\mu 3} \\ \mathbf{U}_{\tau 1} & \mathbf{U}_{\tau 2} & \mathbf{U}_{\tau 3} \end{pmatrix}$$

$$= \begin{pmatrix} c_{12} c_{13} & s_{12} c_{13} & s_{13} e^{-i\delta} \\ -s_{12} c_{23} - c_{12} s_{23} s_{13} e^{i\delta} & c_{12} c_{23} - s_{12} s_{23} s_{13} e^{i\delta} & s_{23} c_{13} \\ s_{12} s_{23} - c_{12} c_{23} s_{13} e^{i\delta} & -c_{12} s_{23} - s_{12} c_{23} s_{13} e^{i\delta} & c_{23} c_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha_{21}/2} & 0 \\ 0 & 0 & e^{i\alpha_{31}/2} \end{pmatrix}$$

(Majorana phases)

 $[\theta_{13} \equiv \theta_{\odot}, \quad \theta_{23} \equiv \theta_{atm}, \quad \theta_{13} \quad and \quad \delta \quad measured in oscillations]$

• U introduces family mixings in \mathcal{L}_{CC} (like CKM), but in this case:



 $v_{\alpha} \text{ are coherent superpositions of mass eigenstates } \nu_i$ $(produced/detected in association with <math>\ell_{\alpha}$) $\ell_{\alpha} (e, \mu, \tau)$ are mass eigenstates (do *not* oscillate) because $\Delta m_{ij}^2 \ll \Delta m_{\mu e}^2$ [0706.1216]



[see section on Neutrino phenomenology]

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Neutrinos are special

Dirac vs Majorana fermions

• A Dirac fermion field is a spinor with 4 independent components: 2 LH+2 RH (left/right-handed particles and antiparticles)

 $\psi_L = P_L \psi$, $\psi_R = P_R \psi$, $\psi_L^c \equiv (\psi_L)^c = P_R \psi^c$, $\psi_R^c \equiv (\psi_R)^c = P_L \psi^c$ where $\psi^c \equiv C \overline{\psi}^{\mathsf{T}} = -i\gamma^2 \psi^*$ (charge conjugate), $C = -i\gamma^2 \gamma^0$, $P_{R,L} = \frac{1}{2}(1 \pm \gamma_5)$

• A Majorana fermion field has just 2 independent components since $\psi^c \equiv \eta^* \psi$:

$$\psi_L = \eta \psi^c_R$$
 , $\; \psi_R = \eta \psi^c_L$

where $\eta = -i\eta_{CP}$ (CP parity) with $|\eta|^2 = 1$. Only possible if neutral

[Useful relations:
$$C^{\dagger} = C^T = C^{-1} = -C$$
, $C\gamma_{\mu}C^{-1} = -\gamma_{\mu}^T$, $\overline{\psi^c} = \psi^T C$]

• Lorentz invariant terms:

$$\overline{\psi_R}\psi_L = \overline{\psi_L^c}\psi_R^c \quad \stackrel{\text{hc}}{\longleftrightarrow} \quad \overline{\psi_L}\psi_R = \overline{\psi_R^c}\psi_L^c \quad (\Delta F = 0)$$

$$\frac{\overline{\psi_L^c}\psi_L}{\overline{\psi_R}\psi_R^c} = \quad \overline{\psi_L^c}\psi_L^c \quad \{|\Delta F| = 2\}$$

$$\Rightarrow \quad -\mathcal{L}_m = \underbrace{m_D \ \overline{\psi_R}\psi_L}_{\text{Dirac term}} + \underbrace{\frac{1}{2}m_L \ \overline{\psi_L^c}\psi_L + \frac{1}{2}m_R \ \overline{\psi_R}\psi_R^c}_{\text{Majorana terms}} + \text{h.c.}$$

- A Dirac fermion can only have a Dirac mass term (fermion number preserving)
- Majorana fermions may have Majorana mass terms
- ⇒ In the SM: * m_D from Yukawa coupling after EW SSB $(m_D = \lambda_v v / \sqrt{2})$ * m_L forbidden by gauge symmetry
 - * m_R compatible with gauge symmetry! (ν_R are sterile)

General mass terms (a more transparent parameterization)

• Rewrite previous mass terms introducing an array of two Majorana fermions:

$$\chi_{L}^{0} = \begin{pmatrix} \psi_{L} \\ \psi_{R}^{c} \end{pmatrix}, \quad \chi^{0} = \chi^{0c} = \chi_{L}^{0} + \chi_{L}^{0c} \equiv \begin{pmatrix} \chi_{1}^{0} \\ \chi_{2}^{0} \end{pmatrix}, \quad \chi_{1}^{0} = \chi_{1L}^{0c} = \chi_{1L}^{0} + \chi_{1L}^{0c} \equiv \psi_{L} + \psi_{L}^{c} \\ \chi_{2}^{0} \end{pmatrix}, \quad \chi_{2}^{0} = \chi_{2L}^{0c} = \chi_{2L}^{0} + \chi_{2L}^{0c} \equiv \psi_{R}^{c} + \psi_{R}^{c}$$

$$\Rightarrow -\mathcal{L}_m = \frac{1}{2} \overline{\chi_L^{0c}} \mathbf{M} \chi_L^0 + \text{h.c. with } \mathbf{M} = \begin{pmatrix} m_L & m_D \\ m_D & m_R \end{pmatrix}$$

M is a square symmetric matrix \Rightarrow diagonalizable by a unitary matrix $\widetilde{\mathcal{U}}$:

$$\widetilde{\mathcal{U}}^{\mathsf{T}}\mathbf{M}\ \widetilde{\mathcal{U}} = \mathcal{M} = \operatorname{diag}(m_1', m_2'), \quad \chi_L^0 = \widetilde{\mathcal{U}}\chi_L \quad (\chi_L^{0c} = \widetilde{\mathcal{U}}^*\chi_L^c)$$

To get positive eigenvalues $m_i = \eta_i m'_i$ (physical masses) replace $\chi_{iL} = \sqrt{\eta_i} \xi_{iL}$

$$\chi_{L}^{0} = \mathcal{U}\xi_{L}, \quad \mathcal{U} = \widetilde{\mathcal{U}}\text{diag}(\sqrt{\eta_{1}}, \sqrt{\eta_{2}}), \qquad \begin{array}{l} \xi_{1} = \xi_{1L} + \xi_{1L}^{c} \\ \xi_{2} = \xi_{2L} + \xi_{2L}^{c} \end{array} \quad \text{(physical fields)}$$

• Case of only Dirac term $(m_L = m_R = 0)$

$$\chi_L^0 = (\nu_L, \nu_R^c)$$

$$\mathbf{M} = \begin{pmatrix} 0 & m_D \\ m_D & 0 \end{pmatrix} \quad \Rightarrow \quad \widetilde{\mathcal{U}} = \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & 1 \\ -1 & 1 \end{pmatrix} , \quad m'_1 = -m_D , \quad m'_2 = m_D$$

Eigenstates
$$\Rightarrow$$
 Physical states
 $\chi_{1L} = \frac{1}{\sqrt{2}}(\nu_L - \nu_R^c)$
 $\zeta_{1L} = -i\chi_{1L} \quad [\eta_1 = -1]$
 $\zeta_{2L} = \chi_{2L} \quad [\eta_2 = +1]$
with masses $m_1 = m_2 = m_D$
 $\Rightarrow -\mathcal{L}_m = m_D(\overline{\nu_R}\nu_L + \overline{\nu_L}\nu_R) = \frac{1}{2}m_D(\overline{\zeta_{1L}^c}\xi_{1L} + \overline{\zeta_{2L}^c}\xi_{2L}) + \text{h.c.}$

One Dirac fermion = two Majorana of equal mass and opposite CP parities

• Case of seesaw (type I) $(m_D \ll m_R)$

$$\chi^0_L = (\nu_L, N^c_R)$$

[Yanagida '79; Gell-Mann, Ramond, Slansky '79; Mohapatra, Senjanovic '80]

$$\mathbf{M} = \begin{pmatrix} 0 & m_D \\ m_D & m_R \end{pmatrix} \quad \Rightarrow \quad \widetilde{\mathcal{U}} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix}$$

$$m_1 \equiv m_{\nu} \simeq \frac{m_D^2}{m_R} \ll m_2 \equiv m_N \simeq m_R$$
$$\theta \simeq \frac{m_D}{m_R} \simeq \sqrt{\frac{m_{\nu}}{m_N}} \ll 1$$



$$\chi_{1L} \approx \nu_L - \frac{m_D}{m_R} N_R^c \approx \nu_L \quad \Rightarrow \quad \xi_{1L} \approx -i\nu_L$$

$$\chi_{2L} \approx \frac{m_D}{m_R} \nu_L + N_R^c \approx N_R^c \quad \Rightarrow \quad \xi_{2L} \approx N_R^c \quad \Rightarrow \quad -\mathcal{L}_m = \underbrace{\frac{1}{2} m_\nu}_{\text{gauge invariant}??} \underbrace{\frac{1}{2} m_N}_{\text{gauge invariant}??} \overline{N_R} + \text{h.c.}$$

Case of seesaw (type I)

$$\chi^0_L = (
u_L, N^c_R)$$

 $\frac{1}{2}m_{\nu} \overline{\nu_L^c} \nu_L$ comes after EW SSB from a dim-5 effective interaction, that is gauge-invariant but lepton-number violating (Weinberg operator):

$$\mathcal{L}_{\text{Weinberg}} = -\frac{1}{2} \frac{\lambda_{\nu}^{2}}{m_{R}} (\overline{L}\widetilde{\Phi}) (\widetilde{\Phi}^{T} L^{c}) + \text{h.c.} \qquad \langle \Phi \rangle \qquad \qquad \langle \Phi \rangle$$
with $L = \begin{pmatrix} \nu_{L} \\ e_{L} \end{pmatrix}$

$$\underbrace{\nu_{L} \quad m_{D} \quad N_{R} \quad m_{R} \quad N_{R} \quad m_{D} \quad \nu_{L}}$$

Perhaps the observed neutrino v_L is the LH component of a light Majorana v (then $\overline{v} = RH$) and light because of a very heavy Majorana neutrino N

e.g.
$$m_D = \lambda_v \frac{v}{\sqrt{2}} \sim 100 \text{ GeV}$$
, $m_R \sim m_N \sim 10^{14} \text{ GeV} \Rightarrow m_v \sim 0.1 \text{ eV}$ \checkmark

Case of seesaw (type I): $\chi_{L}^{0} = (\nu_{\alpha L}, N_{Rj}^{c})$ several generations $\alpha = e, \mu, \tau \text{ (active)} \quad j = 1, \dots, n_{R} \ge 2 \text{ (sterile)}$ $\mathbf{M} = \begin{pmatrix} 0 & M_{D} \\ M_{D}^{T} & M_{R} \end{pmatrix} \text{ with blocks } \begin{cases} 0: 3 \times 3 & M_{D}: 3 \times n_{R} \\ M_{D}^{T}: n_{R} \times 3 & M_{R}: n_{R} \times n_{R} \end{cases}$

For $M_D \ll M_R$, and taking M_R diagonal to simplify:

$$\mathcal{U}^T \mathbf{M} \mathcal{U} \approx \begin{pmatrix} \mathbf{U}^T M_D M_R^{-1} M_D^T \mathbf{U} & 0\\ 0 & M_R \end{pmatrix} \equiv \begin{pmatrix} M_{\nu}^{\text{diag}} & 0\\ 0 & M_N^{\text{diag}} \end{pmatrix}$$

The 3 × 3 block **U** is *approximately unitary* because it is contained in U:

$$\mathcal{U} \approx \begin{pmatrix} \mathbf{U} & \mathcal{O}(m_D/m_R) \\ \mathcal{O}(m_D/m_R) & \mathbb{1} \end{pmatrix} \text{ and } \nu_{\alpha} = \nu_{\alpha L} + \nu_{\alpha L}^c \text{ with } \nu_{\alpha L} = \mathbf{U}_{\alpha i} \nu_{iL} \\ (\chi_L^0 = \mathcal{U}\xi_L) \end{pmatrix}$$

Anomalies?

• www.ugr.es/local/jillana

About anomalies

- Anomaly: a symmetry of the classical Lagrangian broken by quantum corrections.
- Anomalies appear when *both* axial $(\psi \gamma^{\mu} \gamma_5 \psi)$ and vector $(\psi \gamma^{\mu} \psi)$ currents involved.
- Anomalies of global symmetries are welcome. For example: $\pi^0 \rightarrow \gamma \gamma$ thanks to coupling of an axial current $j^{\mu}_A = (\overline{u}\gamma^{\mu}\gamma_5 u - \overline{d}\gamma^{\mu}\gamma_5 d)$ to two electromagnetic (vector) currents, breaking the conservation of the axial current $(\partial^{\mu}j^{\mu}_A \neq 0)$ at 1 loop, even in the limit of massless quarks.



 However, gauge anomalies are a disaster: they break Ward-Takahashi identities spoiling renormalizability.

Gauge anomalies

• The gauge anomalies are generated by triangle diagrams connecting three gauge bosons V^a , V^b , V^c , each coupled to fermions by $(\overline{\Psi}_L \gamma^{\mu} T_L^a \Psi_L + \overline{\Psi}_R \gamma^{\mu} T_R^a \Psi_R) V_{\mu}^a$ with T_L^a (T_R^a) the associated generators:



$$\mathcal{A}^{abc} = \operatorname{Tr}\left(\{T_L^a, T_L^b\}T_L^c\right) - \operatorname{Tr}\left(\{T_R^a, T_R^b\}T_R^c\right)$$

[traces include summation over *all* fermions]

Gauge symmetry is preserved at quantum level if *every* $A^{abc} = 0$.

• In $SU(3)_c \times SU(2)_L \times U(1)_Y$ we have $T^a \in \{\frac{1}{2}\lambda^i, \frac{1}{2}\sigma^i, Y\}$ with

$$Tr(\lambda^{i}\lambda^{j}) = 2\delta^{ij}$$
$$\{\sigma^{i}, \sigma^{j}\} = 2\delta^{ij} \mathbb{1}$$
$$Tr(\lambda^{i}) = Tr(\sigma^{i}) = 0$$

Gauge anomalies cancel!

- Since $SU(3)_c$ is non-chiral (not anomalous), the only non trivial combinations are $SU(3)^2U(1): \operatorname{Tr}(\{\lambda^i,\lambda^j\}Y\}) \Rightarrow \mathcal{A}^{abc} \propto \sum (Y_L - Y_R) = 0 \checkmark$ quarks $SU(2)^2U(1): \operatorname{Tr}(\{\sigma^i,\sigma^j\}Y) \Rightarrow \mathcal{A}^{abc} \propto \sum Y_L + N_c \sum Y_L = 0 \checkmark$ leptons quarks $U(1)^3: \quad \mathrm{Tr}(Y^3) \qquad \qquad \Rightarrow \quad \mathcal{A}^{abc} \propto \sum (Y_L^3 - Y_R^3) + N_c \sum (Y_L^3 - Y_R^3) = 0 \quad \checkmark$ leptons quarks v_e e u d where $\begin{vmatrix} Y_L & -\frac{1}{2} & -\frac{1}{2} & \frac{1}{6} & \frac{1}{6} \\ Y_R & 0 & -1 & \frac{2}{3} & -\frac{1}{3} \end{vmatrix}$ and anomalies cancel if $N_c = 3$. In particular the second constraint is equivalent to
 - $Q_{\nu} + Q_{e} + N_{c}(Q_{u} + Q_{d}) = -1 + \frac{1}{3}N_{c} = 0 \quad \Rightarrow \quad N_{c} = 3 \quad (!!)$
 - \Rightarrow The electroweak SM needs leptons + quarks in every generation !!
 - \Rightarrow The electroweak SM needs the QCD part !!

3. Electroweak Pheno

$$\mathcal{L} = \mathcal{L}_F + \mathcal{L}_{YM} + \mathcal{L}_{\Phi} + \mathcal{L}_Y + \mathcal{L}_{GF} + \mathcal{L}_{FP}$$

 $\mathcal{L}_F \supset \mathcal{L}_{\mathrm{CC}} + \mathcal{L}_{\mathrm{NC}}$ $\mathcal{L}_{YM} \supset \mathcal{L}_{VVV} + \mathcal{L}_{VVVV}$ $\mathcal{L}_{\Phi} \supset$ gauge boson masses $\mathcal{L}_{\gamma} \supset$ fermion masses and mixings

- Fields: [F] fermions [S] scalars (Higgs and unphysical Goldstones) [V] vector bosons [U] unphysical ghosts
- Interactions: [FFV] [FFS] [SSV] [SVV] [SSVV] [VVV] [VVVV] [SSS] [SSSS] [SUU] [UUVV]

Full SM Lagrangiantypes of interactions

• Lorentz structure of generic interactions (normalized to *e*):

$$\begin{aligned} \mathcal{L}_{\text{FFV}} &= e \,\overline{\psi}_i \gamma^{\mu} (g_V - g_A \gamma_5) \psi_j \, V_{\mu} = e \,\overline{\psi}_i \gamma^{\mu} (g_L P_L + g_R P_R) \psi_j \, V_{\mu} \\ \mathcal{L}_{\text{FFS}} &= e \,\overline{\psi}_i (g_S - g_P \gamma_5) \psi_j \phi = e \,\overline{\psi}_i (c_L P_L + c_R P_R) \psi_j \phi \\ \mathcal{L}_{\text{VVV}} &= -ie \, c_{VVV} \left(W^{\mu\nu} W^{\dagger}_{\mu} V_{\nu} - W^{\dagger}_{\mu\nu} W^{\mu} V^{\nu} - W^{\dagger}_{\mu} W_{\nu} V^{\mu\nu} \right) \\ \mathcal{L}_{\text{VVVV}} &= e^2 \, c_{VVVV} \left(2W^{\dagger}_{\mu} W^{\mu} V_{\nu} V'^{\nu} - W^{\dagger}_{\mu} V^{\mu} W_{\nu} V'^{\nu} - W^{\dagger}_{\mu} V'^{\mu} W_{\nu} V^{\nu} \right) \\ \mathcal{L}_{\text{SSV}} &= -ie \, c_{SSV} \, \phi \, \overline{\partial_{\mu}} \phi' \, V^{\mu} \\ \mathcal{L}_{\text{SVV}} &= e \, c_{SVV} \, \phi \, V^{\mu} V'_{\mu} \\ \mathcal{L}_{\text{SSVV}} &= e^2 \, c_{SSVV} \, \phi \phi' V^{\mu} V'_{\mu} \\ \mathcal{L}_{\text{SSS}} &= e \, c_{SSS} \, \phi \phi' \phi'' V''_{\mu} \\ \mathcal{L}_{\text{SSSS}} &= e \, c_{SSS} \, \phi \phi' \phi'' \phi''', \\ \varphi \, \overline{\partial_{\mu}} \phi' &\equiv \phi_i \partial_{\mu} \phi' - (\partial_{\mu} \phi_i) \phi' \quad \text{and} \quad V_{\mu} \in \{A_{\mu}, Z_{\mu}, W_{\mu}, W^{\dagger}_{\mu}\}. \end{aligned}$$

where

Full SM LagrangianFeynman Rules

• Feynman rules for generic vertices normalized to *e* (all momenta incoming):

$$(i\mathcal{L}) \qquad [FFV_{\mu}] = ie\gamma^{\mu}(g_{L}P_{L} + g_{R}P_{R}) \\ [FFS] = ie(c_{L}P_{L} + c_{R}P_{R}) \\ [V_{\mu}(k_{1})V_{\nu}(k_{2})V_{\rho}(k_{3})] = iec_{VVV} \left[g_{\mu\nu}(k_{2} - k_{1})_{\rho} + g_{\nu\rho}(k_{3} - k_{2})_{\mu} + g_{\mu\rho}(k_{1} - k_{3})_{\nu}\right] \\ [V_{\mu}V_{\nu}V_{\rho}V_{\sigma}] = iec_{VVVV} \left[2g_{\mu\nu}g_{\rho\sigma} - g_{\mu\rho}g_{\nu\sigma} - g_{\mu\sigma}g_{\nu\rho}\right] \\ [S(p)S(p')V_{\mu}] = iec_{SSV} (p_{\mu} - p'_{\mu}) \\ [SV_{\mu}V_{\nu}] = iec_{SSVV}g_{\mu\nu} \\ [SSV_{\mu}V_{\nu}] = iec_{SSV}g_{\mu\nu} \\ [SSS] = iec_{SSS} \\ [SSSS] = ie^{2}c_{SSSS}$$

Note: $g_{L,R} = g_V \pm g_A$ $\partial_\mu \rightarrow -ip_\mu$ Attention to symmetry factors! $c_{L,R} = g_S \pm g_P$ $e.g. \ 2 \times HZZ$

Feynman rules ('t Hooft-Feynman gauge)

FFV
$$\overline{f}_i f_j \gamma$$
 $\overline{f}_i f_j Z$ $\overline{u}_i d_j W^+$ $\overline{d}_j u_i W^ \overline{v}_i \ell_j W^+$ $\overline{\ell}_j v_i W^ g_L$ $-Q_f \delta_{ij}$ $g_+^f \delta_{ij}$ $\frac{1}{\sqrt{2}s_W} \mathbf{V}_{ij}$ $\frac{1}{\sqrt{2}s_W} \mathbf{V}_{ij}^*$ $\frac{1}{\sqrt{2}s_W} \delta_{ij}$ $\frac{1}{\sqrt{2}s_W} \delta_{ij}$ g_R $-Q_f \delta_{ij}$ $g_-^f \delta_{ij}$ 0000

$$g_{\pm}^{f} \equiv v_{f} \pm a_{f}$$
 $v_{f} = \frac{T_{3}^{f_{L}} - 2Q_{f}s_{W}^{2}}{2s_{W}c_{W}}$ $a_{f} = \frac{T_{3}^{f_{L}}}{2s_{W}c_{W}}$

Feynman rules ('t Hooft-Feynman gauge)



$$(f = u, d, \ell)$$



SVV	HZZ	HW^+W^-	$\phi^{\pm}W^{\mp}\gamma$	$\phi^{\pm}W^{\mp}Z$
C _{SVV}	$M_W/(s_W c_W^2)$	M_W/s_W	$-M_W$	$-M_W s_W / c_W$

SSV	χHZ	$\phi^\pm\phi^\mp\gamma$	$\phi^\pm\phi^\mp Z$	$\phi^{\mp}HW^{\pm}$	$\phi^{\mp}\chi W^{\pm}$
c _{SSV}	$-\frac{\mathrm{i}}{2s_W c_W}$	干1	$\pm \frac{c_W^2 - s_W^2}{2s_W c_W}$	$\mp \frac{1}{2s_W}$	$-\frac{\mathrm{i}}{2s_W}$

VVV	$W^+W^-\gamma$	W^+W^-Z
C _{VVV}	-1	c_W/s_W

VVVV	$W^+W^+W^-W^-$	W^+W^-ZZ	$W^+W^-\gamma Z$	$W^+W^-\gamma\gamma$
C _{VVVV}	$\frac{1}{s_W^2}$	$-rac{c_W^2}{s_W^2}$	$\frac{c_W}{s_W}$	-1



- Would-be Goldstone bosons in [SSVV], [SSS] and [SSSS] omitted
- Faddeev-Popov ghosts in [SUU] and [UUVV] omitted
- All Feynman rules from FeynArts (same conventions; $\chi, \phi^{\pm} \rightarrow G^0, G^{\pm}$):

http://www.ugr.es/local/jillana/SM/FeynmanRulesSM.pdf

Input parameters

• Parameters:

$$\frac{17+9 = 1 \quad 1 \quad 1 \quad 1 \quad 1 \quad 9+3 \quad 4 \quad 6}{\text{formal:} \quad g \quad g' \quad v \quad \lambda \quad \lambda_f} \quad \mathbf{V}_{\text{CKM}} \quad \mathbf{U}_{\text{PMNS}}$$
where $g = \frac{e}{s_W} \quad g' = \frac{e}{c_W}$ and
$$\underbrace{\alpha = \frac{e^2}{4\pi} \quad M_W = \frac{1}{2}gv \quad M_Z = \frac{M_W}{c_W}}_{g, g', v} \quad M_H = \sqrt{2\lambda}v \quad m_f = \frac{v}{\sqrt{2}}\lambda_f$$

 \Rightarrow Many (more) experiments

 \Rightarrow After Higgs discovery, for the first time *all* parameters measured!

Input parameters

- Experimental values
 - Fine structure constant:

 $\alpha^{-1} = 137.035\,999\,206\,(11)$

[Particle Data Group '20] @

 $\alpha^{-1} = 137.035\,999\,150\,(33)$ Harvard cyclotron (g_e) [1712.06060] $\alpha^{-1} = 137.035\,999\,046\,(27)$ atom interferometry (Cesium) [1

atom interferometry (Cesium) [1812.04130] atom interferometry (Rubidium) [Nature 588, 61(2020)]

– The SM predicts $M_W < M_Z$ in agreement with measurements:

 $M_Z = (91.1876 \pm 0.0021) \text{ GeV}$ LEP1/SLD

- $M_W = (80.379 \pm 0.012) \text{ GeV}$ LEP2/Tevatron/LHC
- Top quark mass:

 $m_t = (172.25 \pm 0.30) \text{ GeV}$ Tevatron/LHC

– Higgs boson mass:

 $M_H = (125.46 \pm 0.17) \text{ GeV}$ LHC

- Low energy observables $(Q^2 \ll M_Z^2)$
 - ν -nucleon (NuTeV) and νe (CERN) scattering asymmetries CC/NC and $\nu/\bar{\nu} \Rightarrow s_W^2$



Weak NC discovery (1973)

- Parity and Atomic Parity violation (SLAC, CERN, Jefferson Lab, Mainz) *LR* asymmetries $e_{R,L}N \rightarrow eX$ and Z effects on atomic transitions $\Rightarrow s$
- muon decay: $\mu \rightarrow e \,\overline{\nu}_e \nu_\mu$ (PSI) lifetime

• Low energy observables

 \Rightarrow Fermi constant provides the Higgs VEV (electroweak scale):

$$v = \left(\sqrt{2}G_F\right)^{-1/2} \approx 246\,\text{GeV}$$

and constrains the product $M_W^2 s_W^2$, which implies

$$M_Z^2 > M_W^2 = \frac{\pi \alpha}{\sqrt{2}G_F s_W^2} > \frac{\pi \alpha}{\sqrt{2}G_F} \approx (37.4 \text{ GeV})^2$$

 \Rightarrow Consistency checks: e.g. from muon lifetime:

$$G_F = 1.166\,378\,7(6) \times 10^{-5} \text{ GeV}^{-2}$$

If one compares with (tree level result)

$$\frac{G_F}{\sqrt{2}} = \frac{\pi\alpha}{2s_W^2 M_W^2} = \frac{\pi\alpha}{2(1 - M_W^2 / M_Z^2)M_W^2} \approx 1.125 \times 10^{-5}$$

a discrepancy that disappears when *quantum corrections* are included

• $e^+e^- \rightarrow \bar{f}f$ (PEP, PETRA, TRISTAN, ..., LEP1, SLD)

$$e^{+} \qquad \qquad \frac{\mathrm{d}\sigma}{\mathrm{d}\Omega} = N_{c}^{f} \frac{\alpha^{2}}{4s} \beta_{f} \Big\{ \Big[1 + \cos^{2}\theta + (1 - \beta_{f}^{2}) \sin^{2}\theta \Big] G_{1}(s) \\ + 2(\beta_{f}^{2} - 1) G_{2}(s) + 2\beta_{f} \cos\theta G_{3}(s) \Big\}$$



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9

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Observables and experiments

• Z pole observables (LEP1/SLD)

 $M_Z, \Gamma_Z, \sigma_{\text{had}}, A_{FB}, A_{LR}, R_b, R_c, R_\ell \implies M_Z, s_W^2$

from $e^+e^- \rightarrow \bar{f}f$ at the Z pole ($\gamma - Z$ interference vanishes). Neglecting m_f :



Forward-Backward and (if polarized e⁻) Left-Right asymmetries due to Z:

$$A_{FB} = \frac{\sigma_F - \sigma_B}{\sigma_F + \sigma_B} = \frac{3}{4} A_f \frac{A_e + P_e}{1 + P_e A_e} \qquad A_{LR} = \frac{\sigma_L - \sigma_R}{\sigma_L + \sigma_R} = A_e P_e \quad \text{with } A_f \equiv \frac{2v_f a_f}{v_f^2 + a_f^2}$$

• W-pair production (LEP2) $e^+e^- \rightarrow WW \rightarrow 4 f (+\gamma)$





• W production (Tevatron/LHC) $pp/p\bar{p} \rightarrow W \rightarrow \ell \nu_{\ell} (+\gamma)$



• Top-quark production (Tevatron/LHC) $pp/p\bar{p} \rightarrow t\bar{t} \rightarrow 6 f$



• Higgs (LHC)

Single and Double H production and decay to different channels $\Rightarrow M_H$



• Higgs (LHC)

 Signal strength $\mu = \frac{(\sigma \cdot BR)_{obs}}{(\sigma \cdot BR)_{SM}}$ Run 1
 Run 2

 ATLAS
 1.17 ± 0.27 1.02 ± 0.14

 CMS
 $1.18^{+0.26}_{-0.23}$ $1.18^{+0.17}_{-0.14}$

Per channel:



 $b\bar{b} > 5\sigma$ [Jul '18!]



 $\sim 3\sigma$ [Jul '20!!]

Electroweak phenomenology

• Higgs mass (LHC)



• Higgs mass (LHC)


Observables and experiments

• Higgs couplings (LHC)



proben over more that 3 orders of mangitude!

Electroweak phenomenology

- Experimental precision requires accurate predictions ⇒ quantum corrections (complication: loop calculations involve renormalization)
- Correction to G_F from muon lifetime:

$$\frac{G_F}{\sqrt{2}} \to \frac{G_F}{\sqrt{2}} = \frac{\pi \alpha}{2(1 - M_W^2 / M_Z^2) M_W^2} [1 + \Delta r(m_t, M_H)]$$

when loop corrections are included:



Since muon lifetime is measured more precisely than M_W , it is traded for G_F :

$$\Rightarrow M_W^2(\alpha, G_F, M_Z, m_t, M_H) = \frac{M_Z^2}{2} \left(1 + \sqrt{1 - \frac{4\pi\alpha}{\sqrt{2}G_F M_Z^2} [1 + \Delta r(m_t, M_H)]} \right)$$

(correlation between M_W , m_t and M_H , given α , G_F and M_Z)

Ξ



– Corrections to vector and axial couplings from Z pole observables:

$$v_f \to g_V^f = v_f + \Delta g_V^f \qquad a_f \to g_A^f = a_f + \Delta g_A^f$$
$$\Rightarrow \sin^2 \theta_{\text{eff}}^f \equiv \frac{1}{4|Q_f|} \left| 1 - \operatorname{Re}(g_V^f/g_A^f) \right| \equiv \underbrace{(1 - M_W^2/M_Z^2)}^{s_W^2} \kappa_Z^f$$

(Two) loop calculations are crucial and point to a light Higgs:



- In addition, experiments and observables testing the flavor structure of the SM: flavor conserving: dipole moments, ... flavor changing: $b \rightarrow s\gamma$, ...
 - \Rightarrow very sensitive to new physics through loop corrections

Extremely precise measurements are:

- electron magnetic moment:

exp:
$$g_e/2 = 1.001\,159\,652\,182\,032\,(720)$$

theo: QED (5 loops!) $\Rightarrow \alpha^{-1} = 137.035\,999\,150\,(33)$

– muon anomalous magnetic moment: $a_{\mu} = (g_{\mu} - 2)/2$

 $\begin{aligned} a_{\mu}^{\exp} &= 116\,592\,089\,(63) \times 10^{-11} & [\text{Brookhaven '06}] \\ a_{\mu}^{\text{QED}} &= 116\,584\,719 & \times 10^{-11} & [\text{QED: 5 loops}] \\ a_{\mu}^{\text{EW}} &= 154 \ (1) \times 10^{-11} & [\text{W, Z, H: 2 loops}] \\ a_{\mu}^{\text{had}} &= 6\,937\,(43) \times 10^{-11} & [\text{e}^+\text{e}^- \rightarrow \text{had}] \\ a_{\mu}^{\text{SM}} &= 116\,591\,810\,(43) \times 10^{-11} & [\text{Theory Initiative '20}] \end{aligned}$

$$a_{\mu}^{\exp} - a_{\mu}^{SM} = 279 (76) \times 10^{-11}$$

3.7 σ !



[2103.11769] www.ugr.es/local/jillana **Precise determination of parameters Flavor anomalies**

• Test of lepton universality in *b* decays at LHCb



 Fit parameters from a list of observables: find the χ²_{min} varying some of them
 [n_{dof} = # of observables minus # of parameters]

http://gfitter.desy.de [1803.01853]

n _{dof}	$\chi^2_{ m min}$	<i>p</i> -value
15	18.6	0.23

(goodness of fit)

Parameter	Input value	Free in fit
M_H [GeV]	125.1 ± 0.2	Yes
M_W [GeV]	80.379 ± 0.013	_
Γ_W [GeV]	2.085 ± 0.042	_
M_Z [GeV]	91.1875 ± 0.0021	Yes
Γ_Z [GeV]	2.4952 ± 0.0023	_
$\sigma_{ m had}^0$ [nb]	41.540 ± 0.037	_
R^0_ℓ	20.767 ± 0.025	_
$A_{ m FB}^{0,\ell}$	0.0171 ± 0.0010	_
$A_\ell \ ^{(\star)}$	0.1499 ± 0.0018	_
$\sin^2 \theta_{\rm eff}^{\ell}(Q_{\rm FB})$	0.2324 ± 0.0012	_
$\sin^2 \theta_{\rm eff}^{\ell}$ (Tevt.)	0.23148 ± 0.00033	_
A_c	0.670 ± 0.027	_
A_b	0.923 ± 0.020	_
$A_{ m FB}^{0,c}$	0.0707 ± 0.0035	_
$A_{ m FB}^{0,b}$	0.0992 ± 0.0016	_
R_c^0	0.1721 ± 0.0030	_
R_b^0	0.21629 ± 0.00066	_
\overline{m}_c [GeV]	$1.27^{+0.07}_{-0.11}$	Yes
\overline{m}_b [GeV]	$4.20^{+0.17}_{-0.07}$	Yes
$m_t \; [\text{GeV}]^{(\bigtriangledown)}$	172.47 ± 0.68	Yes
$\Delta lpha_{ m had}^{(5)}(M_Z^2) \ ^{(\dagger \bigtriangleup)}$	2760 ± 9	Yes
$\alpha_s(M_Z^2)$	_	Yes
2		

Global fits (Comparisons)

• Compare direct measurements of the observables with fit values:



 \Rightarrow some tensions (none above 3σ): $A_{\ell}(SLD)$, $A_{FB}^{b}(LEP)$, R_{b} , ...

Global fits (Comparisons)

• Compare indirect determinations with fit values (error bars are direct measmts.):



[indirect determination means fit without using constraint from given direct measurement]

Global fits (Conclusions)

 \Rightarrow Fits prefer a somewhat lighter Higgs:



Global fits (Conclusions)

 \Rightarrow In general, impressive consistency of the SM, e.g.:



Neutrinos

Neutrinos Dirac vs Majorana

• General mass terms compatible with $SU_L(2) \times U_Y(1)$ gauge symmetry (1 family):

$$-\mathcal{L}_m = \frac{1}{2} \begin{pmatrix} \overline{\nu_L^c} & \overline{\nu_R} \end{pmatrix} \begin{pmatrix} 0 & m_D \\ m_D & m_R \end{pmatrix} \begin{pmatrix} \nu_L \\ \nu_R^c \end{pmatrix} + \text{h.c.}$$

- 1 Dirac = light ν $[m_R = 0]$ $-\mathcal{L}_m = m_D \overline{\nu_R} \nu_L + \text{h.c.}$

$$m_{\nu} = m_D \quad (4 \text{ dof})$$

- 2 Majorana = light ν + heavy N [$m_D \ll m_R$] (seesaw) LNV

$$-\mathcal{L}_{m} = m_{D}\overline{\nu_{R}}\nu_{L} + \frac{1}{2}m_{R}\overline{\nu_{R}}\nu_{R}^{c} + \text{h.c.}$$

$$= \frac{1}{2}m_{\nu}\overline{\nu_{L}^{c}}\nu_{L} + \frac{1}{2}m_{N}\overline{N_{R}}N_{R}^{c} + \text{h.c.}$$

$$m_{\nu} = \frac{m_{D}^{2}}{m_{R}} \quad (2 \text{ dof}) , \quad m_{N} = m_{R} \quad (2 \text{ dof})$$

Neutrinos Dirac vs Majorana

- Neutrinos at low energies (3 families of light neutrinos):
 - 3 light Dirac: $v_{\alpha} = v_{\alpha L} + v_{\alpha R}$ (like quarks)

 $u_{\alpha L} = U_{\alpha i} \nu_{iL}, \quad \nu_{\alpha R} = V_{\alpha i} \nu_{iR}, \quad V^{\dagger} M_{\nu} U = \mathcal{M}_{\text{diag}}$

- 3 light Majorana: $\nu_{\alpha} = \nu_{\alpha L} + \nu_{\alpha L}^c$ $(\nu_{\alpha R} = \nu_{\alpha L}^c)$

 $\nu_{\alpha L} = U_{\alpha i} \nu_{iL}$, $\nu_{\alpha R} = U_{\alpha i}^* \nu_{iR}$, $U^T M_{\nu} U = \mathcal{M}_{\text{diag}} \quad \Leftarrow M_{\nu} = M_D M_R^{-1} M_D^T$

▷ In both cases, mixing matrix shows up in Charged Currents:

$$\mathcal{L}_{CC} = \frac{g}{\sqrt{2}} \sum_{\alpha i} \overline{\ell_{L\alpha}} \gamma^{\mu} U_{\alpha i} \nu_{iL} W_{\mu} + \text{h.c.}$$

(basis where charged leptons are diagonal)

▷ In Majorana case there are also Neutral Current interactions with the *Z*:

$$\mathcal{L}_{\mathrm{NC}}^{Z} \supset \frac{g}{4c_{W}} \sum_{ij} \overline{\nu}_{j} \gamma^{\mu} (P_{L}C_{ij} - P_{R}C_{ij}^{*}) \nu_{j} Z_{\mu}, \quad C_{ij} = \sum_{\alpha=1}^{3} U_{\alpha i}^{*} U_{\alpha j}$$

Neutrinos Dirac vs Majorana

- About (light) neutrino flavour mixing (3 families):
 - Dirac: All phases absorbed by field redefinitions but one phase remains.
 In the standard parameterization

$$U_{\text{Dirac}} = \begin{pmatrix} c_{12} c_{13} & s_{12} c_{13} & s_{13} e^{-i\delta} \\ -s_{12} c_{23} - c_{12} s_{23} s_{13} e^{i\delta} & c_{12} c_{23} - s_{12} s_{23} s_{13} e^{i\delta} & s_{23} c_{13} \\ s_{12} s_{23} - c_{12} c_{23} s_{13} e^{i\delta} & -c_{12} s_{23} - s_{12} c_{23} s_{13} e^{i\delta} & c_{23} c_{13} \end{pmatrix}$$

– Majorana: Since $\nu_{\alpha R} = \nu_{\alpha L}^c$ two extra phases remain

$$U_{\text{Majorana}} = U_{\text{Dirac}} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha_{21}} & 0 \\ 0 & 0 & e^{i\alpha_{31}} \end{pmatrix}$$

Neutrinos Solar and atmospheric neutrino problems

- The solar neutrino problem
 - The Sun produces v_e 's whose flux can be calculated using solar models
 - The flux of ν_e measured on the Earth in all expts reduced by a factor 0.3–0.5
 - \Rightarrow Explained by oscillations $\nu_e \rightarrow \nu_{\mu,\tau}$
- The atmospheric neutrino problem
 - Cosmic rays produce π 's in the atmosphere that should give a flux of ν_{μ} 's and ν_{e} 's in (2:1)

$$\pi \to \bar{\nu}_{\mu} \mu \to \bar{\nu}_{\mu} \nu_{\mu} \bar{\nu}_{e} e$$

- The observed flux of ν_{μ} is largely reduced
- \Rightarrow Explained by oscillations $\nu_{\mu} \rightarrow \nu_{\tau}$

• Case of two family mixing:

Define $|\nu_e\rangle$ and $|\nu_{\mu}\rangle$ the flavour eigenstates producing *e* and μ respectively in a CC[®] Assume:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

where $|\nu_1\rangle$ and $|\nu_2\rangle$ are the energy eigenstates. Their time evolution is given by

$$|\nu_{e}, t\rangle = \cos\theta e^{-iE_{1}t} |\nu_{1}\rangle + \sin\theta e^{-iE_{2}t} |\nu_{2}\rangle$$
$$|\nu_{\mu}, t\rangle = -\sin\theta e^{-iE_{1}t} |\nu_{1}\rangle + \cos\theta e^{-iE_{2}t} |\nu_{2}\rangle$$

Then the probability of oscillation from ν_e to ν_{μ} after a time *t* is

$$P(\nu_e \to \nu_\mu; t) = |\langle \nu_\mu | \nu_e, t \rangle|^2 = \sin^2(2\theta) \sin^2\left(\frac{\Delta E}{2}t\right) , \quad \Delta E = E_2 - E_1$$

• Case of two family mixing:

For ultrarelativistic neutrinos with definite momentum ($p \gg m_i$)

$$E_i = \sqrt{m_i^2 + p^2} pprox p + rac{m_i^2}{2p}$$
 , $L pprox t$, $p pprox E$

$$P(\nu_e \to \nu_\mu; L) = \sin^2(2\theta) \sin^2\left(\pi \frac{L}{L_{\rm osc}}\right)$$

where the oscillation length is

$$L_{\rm osc} = \pi \frac{4E}{\Delta m^2} = \frac{\pi}{1.27} \frac{E/\text{GeV}}{\Delta m^2/\text{eV}^2} \text{ km} , \quad \Delta m^2 = |m_2^2 - m_1^2|$$

- Oscillations are only sensitive to (squared) mass differences
- If $L \gg L_{\text{osc}}$ (too fast oscillations) then average: $\langle P(\nu_e \rightarrow \nu_\mu; t) \rangle = \frac{1}{2} \sin^2(2\theta)$
- For subtleties on the theory of neutrino oscillations see E. Akhmedov: http://arxiv.org/pdf/0706.1216.pdf, http://arxiv.org/pdf/0905.1903.pdf

• Case of three family mixing: $|\nu_{\alpha}\rangle = U_{\alpha i} |\nu_i\rangle$ (see next slide):

$$P(\nu_{\alpha} \rightarrow \nu_{\beta}; L) = |\langle \nu_{\beta} | \nu_{\alpha}, L \rangle|^{2} = \left| \sum_{ij} \langle \nu_{j} | U_{\beta j}^{*} e^{-iE_{i}L} U_{\alpha i} | \nu_{i} \rangle \right|^{2} = \left| \sum_{i} U_{\beta i}^{*} U_{\alpha i} e^{-im_{i}^{2}L/(2E)} \right|$$
$$= \sum_{ij} U_{\beta j} U_{\alpha j}^{*} U_{\beta i} U_{\alpha i}^{*} e^{-i\Delta m_{ij}^{2}L/(2E)}$$
$$= \delta_{\alpha\beta} - 4 \sum_{i>j} \operatorname{Re}(U_{\beta j} U_{\alpha j}^{*} U_{\beta i}^{*} U_{\alpha i}) \sin^{2} \frac{\Delta m_{ij}^{2}L}{4E}$$
$$+ 2 \sum_{i>j} \operatorname{Im}(U_{\beta j} U_{\alpha j}^{*} U_{\beta i}^{*} U_{\alpha i}) \sin \frac{\Delta m_{ij}^{2}L}{2E} , \qquad \Delta m_{ij}^{2} = m_{i}^{2} - m_{j}^{2}$$

– Majorana phases are irrelevant

-
$$P(\nu_{\beta} \to \nu_{\alpha}; U) = P(\nu_{\alpha} \to \nu_{\beta}; U^{*})$$
 and $CPT \Rightarrow P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta}) = P(\nu_{\beta} \to \nu_{\alpha})$
Therefore $P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta}; U) = P(\nu_{\alpha} \to \nu_{\beta}; U^{*})$
 \Rightarrow if CP conserved ($U = U^{*}$) then $P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta}) = P(\nu_{\alpha} \to \nu_{\beta})$

 \triangleright Use:

$$\langle v_j | v_i \rangle = \delta_{ij}, \quad \Delta m_{ij}^2 \equiv m_i^2 - m_j^2, \quad \sum_{ij} = \sum_{i=j} + \sum_{i>j} + \sum_{i< j} + \sum$$

 \triangleright Use the unitarity of *U* in:

$$\begin{split} \sum_{i=j} (U_{\beta j} U_{\alpha j}^* U_{\beta i}^* U_{\alpha i}) &= \sum_j (U_{\beta j} U_{\alpha j}^*) \sum_i (U_{\beta i}^* U_{\alpha i}) - \sum_{i \neq j} (U_{\beta j} U_{\alpha j}^* U_{\beta i}^* U_{\alpha i}) \\ &= \delta_{\alpha \beta} - 2 \sum_{i>j} \operatorname{Re}(U_{\beta j} U_{\alpha j}^* U_{\beta i}^* U_{\alpha i}) \;. \end{split}$$

▷ Combine:

$$\left(\sum_{i>j} + \sum_{ij} \left[2\operatorname{Re}(U_{\beta j}U_{\alpha j}^*U_{\beta i}^*U_{\alpha i})\cos\frac{\Delta m_{ij}^2 L}{2E} + 2\operatorname{Im}(U_{\beta j}U_{\alpha j}^*U_{\beta i}^*U_{\alpha i})\sin\frac{\Delta m_{ij}^2 L}{2E}\right]$$

 \triangleright And substitute: $1 - \cos \frac{\Delta m_{ij}^2 L}{2E} = 2 \sin^2 \frac{\Delta m_{ij}^2 L}{4E}$

- Comments:
 - Nature seems to have chosen 3 flavours with two very different squared mass differences: $\Delta m_{\odot}^2 \equiv \Delta m_{21}^2 \ll \Delta m_{atm}^2 \equiv \Delta m_{31}^2 \simeq \Delta m_{32}^2$. Then

$$P(\nu_{\alpha} \to \nu_{\beta} \neq \nu_{\alpha}) \approx P_{\alpha\beta}^{\text{short}} + P_{\alpha\beta}^{\text{long}}$$

with

$$P_{\alpha\beta}^{\text{short}} = 4U_{\beta3}^2 U_{\alpha3}^2 \sin^2 [1.27\Delta m_{\text{atm}}^2 L/E]$$
$$P_{\alpha\beta}^{\text{long}} = -4U_{\beta2} U_{\alpha2} U_{\beta1} U_{\alpha1} \sin^2 [1.27\Delta m_{\odot}^2 L/E]$$

where we have neglected CPV effects and used that unitarity of *U*:

$$U_{\beta 1}U_{\alpha 1}+U_{\beta 2}U_{\alpha 2}=-U_{\beta 3}U_{\alpha 3}$$

 Selecting the appropriate *L/E* the oscillation is sensitive to the short or the long component only (decoupled) and can be effectively treated as a two-flavour mixing



- Values of Δm^2 that can be *explored* in different experiments. SBL (LBL) stands for Short (Long) Baseline

Experiment	<i>L</i> [m]	E [MeV]	$\Delta m^2 [{ m eV}^2]$
Reactors SBL	10 ²	1	10^{-2}
Reactors LBL	10^{3}	1	10^{-3}
	10^{5}	1	10^{-5}
Accelerators SBL	10^{3}	10^{3}	1
Accelerators LBL	10 ⁶	10^{3}	10^{-3}
Atmospheric	$10^4 - 10^7$	10 ³	$10^{-1} - 10^{-4}$
Solar *	10 ¹¹	1	10 ⁻¹¹

* Since $\Delta m_{\odot}^2 \sim 10^{-5} \gg 10^{-11} \text{ eV}^2 \Rightarrow L \gg L_{\text{osc}} \sim 100 \text{ km and } P \sim \frac{1}{2} \sin^2(2\theta)$

Neutrinos Oscillations in matter

Mikheyev-Smirnov-Wolfenstein (MSW) effect

- Electron neutrinos travelling through dense matter (Sun, Earth, supernovas) suffer Charged Current interactions enhancing their oscillation probability
- This effect is important for high energy solar neutrinos, above 2 MeV, and it transforms most of the v_e into v_μ when leaving the Sun

We skip their description here

Neutrinos Oscillation phenomenology

• Solar neutrino fluxes [at $L = 1 \text{ AU} \sim 10^{11} \text{ km}$] $E \sim \text{MeV}$



 Δm^2_{\odot}



Neutrinos Oscillation phenomenology Δm_{\odot}^2



NeutrinosOscillation phenomenology Δm_{\odot}^2

• Experiments

Experiment	Reaction	Energy threshold [MeV]
Homestake	$\nu_e \ ^{37}\mathrm{Cl} ightarrow e \ ^{37}\mathrm{Ar}$	0.814
SAGE, Gallex/GNO	$\nu_e \ ^{71}{ m Ga} ightarrow e \ ^{71}{ m Ge}$	0.233
SuperKamiokande	$\nu_{e,x} \ e \to \nu_{e,x} \ e$	5.5
SNO	ES: $\nu_{e,x} \ e \to \nu_{e,x} \ e$	
	CC: $\nu_e D \rightarrow ppe$	5.5
	NC: $\nu_x D \rightarrow \nu_x pn$	



Electroweak phenomenology: neutrinos

Neutrinos Oscillation phenomenology

• SuperKamiokande + SNO



 Δm_{\odot}^2

Neutrinos Oscillation phenomenology

• Test with reactor neutrinos ($E \sim \text{MeV}, \phi \sim 10^{20} \text{ s}^{-1} \text{ GW}^{-1}$): KamLAND A very LBL reactor $L \sim 180$ km measuring $P(\bar{\nu}_e \rightarrow \bar{\nu}_e)$ as a function of the energy

 Δm_{\odot}^2



Neutrinos Oscillation phenomenology Δm_{\odot}^2

• Global fit including KamLAND

10 global $\Delta m_{21}^2 \ [10^{-5} eV^2]$ 8 .′ ∧ KamLAND 6 solar 90, 99% C.L. 2 <u></u> 0.2 0.3 0.4 0.5 $\sin^2 \theta_{12}$

Neutrinos Oscillation phenomenology

• Atmospheric fluxes of ν_{μ} and ν_{e} in (2:1) from $\pi \rightarrow \bar{\nu}_{\mu}\mu \rightarrow \bar{\nu}_{\mu}\nu_{\mu}\bar{\nu}_{e} e$ $L \sim 10 \text{ km [downgoing]} - 10^{4} \text{ km [upgoing]}$ $E \sim 10^{2} - 10^{4} \text{ MeV}$ $\phi \sim 100 \text{ m}^{-2} \text{ s}^{-1}$

(much less abundant but much more energetic than solar)

Experiment: SuperKamiokande (atmospheric neutrinos)
 Reaction: ν_iN → ℓ_iN' detecting ℓ_i by Cherenkov ⇒ direction, flavour (not charge)
 Result: ν_e flux unchanged and ν_μ → ν_x

 $\Delta m_{\rm atm}^2$

• Tests with (LBL) accelerator neutrinos (ν_{μ} disappearance):

Opera	L = 732 km	$E \sim 10 \text{ GeV}$	(CERN \rightarrow Gran Sasso)
K2K	L = 250 km	$E \sim 1 \text{ GeV}$	$(\text{KEK} \rightarrow \text{Kamioka})$
T2K	L = 295 km	$E \sim 1 \text{ GeV}$	$(JPARC \rightarrow Kamioka)$
MINOS	L = 735 km	$E \sim 3 \text{ GeV}$	(Fermilab \rightarrow Soudan (MN))
ΝΟνΑ	L = 810 km	$E \sim 2 \text{ GeV}$	(Fermilab \rightarrow Ash River (MN) [new])

Neutrinos Oscillation phenomenology $\Delta m_{\rm atm}^2$

• Global fit [atm (SK), K2K] and comparison with T2K and MINOS (IceCube/DeepCore and ANTARES less constraining)



Two solutions:

- Normal ordering (NO): $\Delta m_{31}^2 > 0$
- Inverted ordering (IO): $\Delta m_{31}^2 < 0$ (not shown here)



Neutrinos Oscillation phenomenology θ_{13} and δ



- Reactor experiments (Daya Bay, RENO and Double Chooz)
- Accelerator experiments (MINOS and T2K)



 \Rightarrow From global fits some indirect information on δ

[deSalas '19]
Neutrinos Oscillation phenomenology

Status

Parameters

www.nu-fit.org

[1811.05487]

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					NuFIT 4.1 (2019)	
		Normal Ore	dering (best fit)	Inverted Orde	ering $(\Delta \chi^2 = 6.2)$	
without SK atmospheric data		bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range	
	$\sin^2 \theta_{12}$	$0.310\substack{+0.013\\-0.012}$	$0.275 \rightarrow 0.350$	$0.310\substack{+0.013\\-0.012}$	$0.275 \rightarrow 0.350$	
	$ heta_{12}/^{\circ}$	$33.82_{-0.76}^{+0.78}$	$31.61 \rightarrow 36.27$	$33.82_{-0.76}^{+0.78}$	$31.61 \rightarrow 36.27$	
	$\sin^2 \theta_{23}$	$0.558\substack{+0.020\\-0.033}$	$0.427 \rightarrow 0.609$	$0.563\substack{+0.019\\-0.026}$	$0.430 \rightarrow 0.612$	
	$ heta_{23}/^{\circ}$	$48.3^{+1.1}_{-1.9}$	$40.8 \rightarrow 51.3$	$48.6^{+1.1}_{-1.5}$	$41.0 \rightarrow 51.5$	
	$\sin^2 \theta_{13}$	$0.02241\substack{+0.00066\\-0.00065}$	$0.02046 \to 0.02440$	$0.02261\substack{+0.00067\\-0.00064}$	$0.02066 \to 0.02461$	
	$ heta_{13}/^{\circ}$	$8.61^{+0.13}_{-0.13}$	$8.22 \rightarrow 8.99$	$8.65_{-0.12}^{+0.13}$	$8.26 \rightarrow 9.02$	
	$\delta_{ m CP}/^{\circ}$	222^{+38}_{-28}	$141 \rightarrow 370$	285^{+24}_{-26}	$205 \rightarrow 354$	
	$\frac{\Delta m_{21}^2}{10^{-5} \ {\rm eV}^2}$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	
	$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.523^{+0.032}_{-0.030}$	$+2.432 \rightarrow +2.618$	$-2.509^{+0.032}_{-0.030}$	$-2.603 \rightarrow -2.416$	
with SK atmospheric data		Normal Ore	dering (best fit)	Inverted Ordering ($\Delta \chi^2 = 10.4$)		
		bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range	
	$\sin^2 \theta_{12}$	$0.310\substack{+0.013\\-0.012}$	$0.275 \rightarrow 0.350$	$0.310\substack{+0.013\\-0.012}$	$0.275 \rightarrow 0.350$	
	$ heta_{12}/^\circ$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	$33.82^{+0.78}_{-0.75}$	$31.61 \rightarrow 36.27$	
	$\sin^2 \theta_{23}$	$0.563\substack{+0.018 \\ -0.024}$	$0.433 \rightarrow 0.609$	$0.565\substack{+0.017\\-0.022}$	$0.436 \rightarrow 0.610$	
	$ heta_{23}/^{\circ}$	$48.6^{+1.0}_{-1.4}$	$41.1 \rightarrow 51.3$	$48.8^{+1.0}_{-1.2}$	$41.4 \rightarrow 51.3$	
	$\sin^2 heta_{13}$	$0.02237^{+0.00066}_{-0.00065}$	$0.02044 \to 0.02435$	$0.02259\substack{+0.00065\\-0.00065}$	$0.02064 \to 0.02457$	
	$ heta_{13}/^\circ$	$8.60^{+0.13}_{-0.13}$	$8.22 \rightarrow 8.98$	$8.64_{-0.13}^{+0.12}$	$8.26 \rightarrow 9.02$	
	$\delta_{ m CP}/^{\circ}$	221^{+39}_{-28}	$144 \rightarrow 357$	282^{+23}_{-25}	$205 \rightarrow 348$	
	$\frac{\Delta m_{21}^2}{10^{-5} \ {\rm eV}^2}$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	
	$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.528^{+0.029}_{-0.031}$	$+2.436 \rightarrow +2.618$	$-2.510^{+0.030}_{-0.031}$	$-2.601 \rightarrow -2.419$	

J.I. Illana (ugr)

Electroweak phenomenology: neutrinos

Neutrinos	Oscillation phenomenology			Status	/local/jillana
Mixing matrix		WW	w.nu-fit.org		[1811.05487] [000]
	$ U _{3\sigma}^{\mathrm{w/o \ SK-atm}} = \left($	$egin{pmatrix} 0.797 & o 0.842 \\ 0.244 & o 0.496 \\ 0.287 & o 0.525 \end{cases}$	$0.518 \rightarrow 0.585$ $0.467 \rightarrow 0.678$ $0.488 \rightarrow 0.693$	$\begin{array}{c} \text{NuFIT 4.1 (2019)} \\ 0.143 \rightarrow 0.156 \\ 0.646 \rightarrow 0.772 \\ 0.618 \rightarrow 0.749 \end{array}$	
	$ U _{3\sigma}^{\text{with SK-atm}} = \left($	$egin{pmatrix} 0.797 & o 0.842 \\ 0.243 & o 0.490 \\ 0.295 & o 0.525 \end{cases}$	$0.518 \rightarrow 0.585$ $0.473 \rightarrow 0.674$ $0.493 \rightarrow 0.688$	$\begin{array}{c} 0.143 \to 0.156 \\ 0.651 \to 0.772 \\ 0.618 \to 0.744 \end{array} \right)$	

Comparison CKM vs PMNS:





J.I. Illana (ugr)

Electroweak phenomenology: neutrinos

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≥ ⊙

Neutrinos Oscillation phenomenology

• Nova (from 2016)



- Measurement of sign(Δm_{31}^2) (strong dependence on θ_{23})
- Measurement of δ from direct CP asymmetry:

$$\mathcal{A}_{\rm CP} = \frac{P(\nu_{\alpha} \to \nu_{\beta}) - P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta})}{P(\nu_{\alpha} \to \nu_{\beta}) - P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta})}$$

Recent results

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Neutrinos Absolute mass scale

• Cosmology: if $m_{\nu} \neq 0$ neutrinos contribute to the mass density of the universe

From CMB and LSS (hypothesis dependent)



 $\sum m_{\nu_i} < 0.23 - 0.59 \text{ eV} \text{ [PLANCK]}$

Neutrinos Absolute mass scale

• Cosmology: if $m_{\nu} \neq 0$ neutrinos contribute to the mass density of the universe



Neutrinos Absolute mass scale

• Beta decay of tritium: ${}^{3}\text{H} \rightarrow {}^{3}\text{He} + e^{-} + \bar{\nu}_{e}$ (end point of electron spectrum)

$$\frac{\mathrm{d}N}{\mathrm{d}E} = \sum_{i} |U_{ei}|^2 \Gamma(m_{\nu_i}^2, E) \approx \Gamma(\langle m_{\nu_e} \rangle, E)$$
$$m_\beta^2 \equiv \langle m_{\nu_e} \rangle^2 = \sum_{i} |U_{ei}|^2 m_{\nu_i}^2 = c_{13}^2 (m_1^2 c_{12}^2 + m_2^2 s_{12}^2) + m_3^2 s_{13}^2$$



Neutrinos Dirac or Majorana?

- Neutrinoless double-beta decay $(0\nu\beta\beta)$
 - $2\nu\beta\beta$ observed with $T_{2\nu\beta\beta} \sim 10^{20}$ years
 - $0\nu\beta\beta$ requires LNV (Majorana ν 's)



Suppressed by $m_{\beta\beta}^2$ but enhanced by phase space

$$m_{\beta\beta} = \left| \sum_{i} U_{ei}^2 m_i \right| = \left| c_{13}^2 (m_1 c_{12}^2 + m_2 s_{12}^2 e^{2i\alpha}) + m_3 s_{13}^2 e^{2i(\beta - \delta)} \right|$$

Neutrinos Dirac or Majorana?

• Neutrinoless double-beta decay $(0\nu\beta\beta)$



Neutrinos	Summa	ary of par	ameters					
ww.ugr.es/								
$N_{ u}$ (active an	d light)	3						
$\Delta m^2_{21} \sim 7.5 \times 10^{-5} \ { m eV}^2$		$ heta_{12}\sim 35^\circ$	Solar, KamLAND					
$\Delta m^2_{31} \sim \pm 2.5 imes$	10^{-3} eV^2	$ heta_{23}\sim 45^\circ$	Atmospheric, K2K, MINOS					
		$ heta_{13}\sim9^\circ$	T2K, MINOS, Double Chooz, Daya Bay, RENO					
$m_{etaeta}\equiv \sum_i u_i $	$J_{ei}^2 m_{\nu_i}$	$\lesssim 0.2 \text{ eV}$	KamLAND-Zen, EXO, HM, IGEX,					
$m_{eta} \equiv \sqrt{\sum_i l }$	$J_{ei} ^2 m_{\nu_i}^2$	$\lesssim 2.3 \text{ eV}$	Mainz, Troitsk					
$\sum_i m_{\nu_i}$		$\lesssim 1 \ \mathrm{eV}$	Cosmology					
$\operatorname{sign}(\Delta m_{31}^2)$?	Noνa, NF, BB, SB,					
СР, δ		?	Nova, NF, BB, SB,					
Dirac or Majorana?		?	HM?, $0\nu\beta\beta$					

Summary and conclusions

Summary and conclusions

- The SM is a gauge theory with spontaneous symmetry breaking (renormalizable)
- Confirmed by many low and high energy experiments with remarkable accuracy, at the level of quantum corrections, with (almost) no significant deviations
- In spite of its tremendous success, it leaves fundamental questions unanswered: why 3 generations? why the observed pattern of fermion masses and mixings?
- And there are several hints for physics beyond:
 - phenomenological:
 - * $(g_{\mu} 2)$
 - * neutrino masses
 - * flavor anomalies
 - * baryon asymmetry
 - * dark matter
 - * dark energy

- conceptual:
 - * gravity is not included
 - hierarchy problem
 - cosmological constant

The SM is an Effective Theory valid up to electroweak scale?

4. Tools

Kinematics

$2 \rightarrow 2$ Kinematics



 $p_1 + p_2 = p_3 + p_4$

Mandelstam variables

(Lorentz invariant)

$$s = (p_1 + p_2)^2 = (p_3 + p_4)^2 = m_1^2 + m_2^2 + 2(p_1 \cdot p_2)$$

$$t = (p_1 - p_3)^2 = (p_2 - p_4)^2 = m_1^2 + m_3^2 - 2(p_1 \cdot p_3)$$

$$u = (p_1 - p_4)^2 = (p_2 - p_3)^2 = m_1^2 + m_4^2 - 2(p_1 \cdot p_4)$$

$$s + t + u = \sum_{i=1}^4 m_i^2.$$

 \vec{p}_f

 $-\vec{p_i}$

 $\vec{p_i}$

 $2 \rightarrow 2$ Kinematics Center of mass frame (CM)

Consider particular case: $m_1 = m_2 \equiv m_i$, $m_3 = m_4 \equiv m_f$

Scalar products:

$$m_i^2 + (p_1 \cdot p_2) = m_f^2 + (p_3 \cdot p_4) = 2E^2 = \frac{s}{2}$$

$$(p_1 \cdot p_3) = (p_2 \cdot p_4) = E^2(1 - \beta_i\beta_f\cos\theta) = \frac{m_i^2 + m_f^2 - t}{\frac{m_i^2 + m_f^2 - u}{2}}$$

$$(p_1 \cdot p_4) = (p_2 \cdot p_3) = E^2(1 + \beta_i\beta_f\cos\theta) = \frac{m_i^2 + m_f^2 - u}{2}$$

Cross-section



$$d\sigma(i \to f) = \frac{1}{4\left\{(p_1 p_2)^2 - m_1^2 m_2^2\right\}^{1/2}} |\mathcal{M}|^2 (2\pi)^4 \delta^4(p_i - p_f) \prod_{j=3}^{n+2} \frac{d^3 p_j}{(2\pi)^3 2E_j}$$

- Sum over initial polarizations and/or average over final polarizations if the initial state is unpolarized and/or the final state polarization is not measured
- ▷ Divide the total cross-section by a symmetry factor $S = \prod_{i} n_i!$ if there are n_i identical particles of species *i* in the final state

Cross-section case $2 \rightarrow 2$ in CM frame

$$q = p_1 = -p_2$$

 p_1, m_1
 p_3, m_3
 p_2, m_2
 p_4, m_4
 $p = p_3 = -p_4$

$$\Rightarrow \int d\Phi_2 \equiv (2\pi)^4 \int \delta^4 (p_1 + p_2 - p_3 - p_4) \frac{d^3 p_3}{(2\pi)^3 2E_3} \frac{d^3 p_4}{(2\pi)^3 2E_4} = \int \frac{|\mathbf{p}| d\Omega}{16\pi^2 E_{\rm CM}}$$

and if $m_1 = m_2 \implies 4 \{ (p_1 p_2)^2 - m_1^2 m_2^2 \}^{1/2} = 4 E_{\text{CM}} |\mathbf{q}|$

$$\frac{\mathrm{d}\sigma}{\mathrm{d}\Omega}(1,2\to3,4) = \frac{1}{64\pi^2 E_{\mathrm{CM}}^2} \frac{|\boldsymbol{p}|}{|\boldsymbol{q}|} |\mathcal{M}|^2$$

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Decay width



 \triangleright Note that masses *M*, *m*₁ and *m*₂ fix final energies and momenta:

$$E_{1} = \frac{M^{2} - m_{2}^{2} + m_{1}^{2}}{2M} \qquad E_{2} = \frac{M^{2} - m_{1}^{2} + m_{2}^{2}}{2M}$$
$$|\mathbf{p}| = |\mathbf{p}_{1}| = |\mathbf{p}_{2}| = \frac{\left\{ [M^{2} - (m_{1} + m_{2})^{2}] [M^{2} - (m_{1} - m_{2})^{2}] \right\}^{1/2}}{2M}$$

General Feynman Rules

Para el cálculo de funciones de Green o de amplitudes invariantes de scattering \mathcal{M}_{fi} .

- Dibujar todos los diagramas conectados y topológicamente distintos en el orden deseado de teoría de perturbaciones.
 En cada diagrama:
- 2. Asociar momentos externos a todas las líneas externas y *L* momentos internos a los *L* loops. Determinar los momentos de las líneas internas de modo que el cuadri-momento se conserve en cada vértice.
- 3. Asignar un propagador a cada línea interna:

•••••• [bosón escalar]
$$\frac{i}{p^2 - m^2}$$

••••• [fermión spin 1/2] $\frac{i(\not p + m)}{p^2 - m^2}$
 $\mu \bullet \cdots \bullet \nu$ [bosón vectorial] $\frac{i}{p^2 - m^2} \left[-g_{\mu\nu} + (1 - \xi) \frac{p_{\mu}p_{\nu}}{(p^2 - \xi m^2)} \right]$ (R_{ξ})
($\xi = 1$: gauge 't Hooft-Feynman; $\xi = 0$: gauge de Landau; $\xi = \infty$: gauge unitario

- 4. A cada vértice asignar un peso compuesto por los siguientes factores:
 - (a) La constante de acoplamiento que aparezca en $i\mathcal{L}_{int}$.
 - (b) Por cada derivada de un campo ϕ cualquiera $\partial_{\mu}\phi$ asociar $(-ip_{\mu})$ donde *p* es el correspondiente momento entrante.
 - (c) Un factor proviniente de la degeneración de partículas idénticas en cada vértice. (Por ejemplo, $\times 2$ para *ZZH*, $\times 4$ para *ZZHH*.)
- 5. Por cada momento interno *q* no fijado por la conservación de momento en cada vértice (loops), introducir un factor

$$\int \frac{\mathrm{d}^4 q}{(2\pi)^4}$$

e integrar, (si es necesario) después de regularizar.

- 6. Multiplicar la contribución de cada diagrama por:
- (a) Un factor (-1) entre diagramas que difieren entre sí solo por el intercambio de dos fermiones externos idénticos. (Por ejemplo, los dos diagramas del scattering de Møller, $e^-e^- \rightarrow e^-e^-$, o los dos del scattering de Bhabha, $e^+e^- \rightarrow e^+e^-$, a nivel árbol.)
- (b) Un factor de simetría 1/S donde S el número de permutaciones de líneas internas y vértices que deja invariante el diagrama si las líneas externas permanecen fijadas.
- (c) Un factor (-1) por cada loop fermiónico, pues:

7. Para obtener i \mathcal{M}_{fi} , poner las líneas externas sobre su capa de masas, es decir $p_i^2 = m_i^2$. Poner por cada línea fermiónica externa un espinor: u(p) [o v(p)] para fermiones [o antifermiones] entrantes con momento p; $\bar{u}(p)$ [o $\bar{v}(p)$] para fermiones [o antifermiones] salientes con momento p. Poner vectores de polarización $\varepsilon_{\mu}(p, \lambda)$ [o $\varepsilon_{\mu}^*(p, \lambda)$] para bosones vectoriales entrantes [o salientes] con momento p.



Calculation of a simple process

 Consideremos la aniquilación de un electrón y un positrón para dar un muón y un antimuón. En QED este proceso viene descrito a orden más bajo de TP (nivel árbol) por el diagrama de la figura.



- ▷ El muón tiene la misma carga del electrón, $Q_{\mu} = Q_e = -1$, y una masa *M* unas 200 veces mayor que la masa *m* del electrón.
- ▷ Vamos a hallar paso a paso y en detalle la sección eficaz de este proceso.

• En primer lugar, asignamos momentos a todas las partículas del diagrama y usamos la conservación del cuadrimomento en cada vértice, lo que fija el cuadrimomento del fotón virtual que se propaga entre los dos vértices de interacción,

$$q = k_1 + k_2 = p_1 + p_2$$
.

- ▷ Las patas externas son fermiones, cuyos espines etiquetamos mediante índices r_1, r_2, s_1, s_2 que toman dos valores posibles {1,2}.
- Aplicando las reglas de Feynman, recorriendo cada línea fermiónica en sentido contrario al flujo fermiónico, el elemento de matriz invariante viene dado por

$$\mathbf{i}\mathcal{M} = \overline{u}^{(s_2)}(\mathbf{p}_2)(\mathbf{i}e\gamma^{\beta})v^{(s_1)}(\mathbf{p}_1)\frac{(-\mathbf{i})}{q^2} \left[g_{\alpha\beta} - (1-\xi)\frac{q_{\alpha}q_{\beta}}{q^2}\right]\overline{v}^{(r_1)}(\mathbf{k}_1)(\mathbf{i}e\gamma^{\alpha})u^{(r_2)}(\mathbf{k}_2) \ .$$

Nótese que como los fermiones externos están sobre su capa de masas satisfacen las respectivas ecuaciones de Dirac,

$$k_1 v^{(r_1)}(k_1) = -mv^{(r_1)}(k_1)$$
, $k_2 u^{(r_2)}(k_2) = mu^{(r_2)}(k_2)$,

así que la amplitud no depende del parámetro ξ , como deber ser, ya que

$$q_{\alpha}\overline{v}^{(r_1)}(k_1)\gamma^{\alpha}u^{(r_2)}(k_2)=\overline{v}^{(r_1)}(k_1)(k_1+k_2)u^{(r_2)}(k_2)=0.$$

▷ Podríamos haber trabajado desde el principio en el gauge de 't Hooft-Feynman ($\xi = 1$). Por tanto,

$$\mathcal{M} = \frac{e^2}{q^2} \overline{u}^{(s_2)}(\boldsymbol{p}_2) \gamma^{\alpha} v^{(s_1)}(\boldsymbol{p}_1) \ \overline{v}^{(r_1)}(\boldsymbol{k}_1) \gamma_{\alpha} u^{(r_2)}(\boldsymbol{k}_2) \ .$$

• Para hallar $|\mathcal{M}|^2$, nótese que

$$(\overline{u}\gamma^{\alpha}v)^{*} = v^{\dagger}\gamma^{\alpha\dagger}\gamma^{0\dagger}u = v^{\dagger}\gamma^{0}\gamma^{0}\gamma_{\alpha}\gamma^{0}u = \overline{v}\gamma^{\alpha}u ,$$

donde se ha usado

$$\overline{u} = u^{\dagger} \gamma^{0}$$
, $\gamma^{lpha \dagger} = \gamma_{lpha}$, $\gamma^{0} \gamma_{lpha} \gamma^{0} = \gamma^{lpha}$.

Se trata además de un número complejo que podemos multiplicar en cualquier orden. Lo mismo ocurre con la otra línea fermiónica.

▷ Conviene escribir,

$$|\mathcal{M}|^{2} = \frac{e^{4}}{q^{4}} \overline{u}^{(s_{2})}(\boldsymbol{p}_{2}) \gamma^{\alpha} v^{(s_{1})}(\boldsymbol{p}_{1}) \ \overline{v}^{(s_{1})}(\boldsymbol{p}_{1}) \gamma^{\beta} u^{(s_{2})}(\boldsymbol{p}_{2}) \times \overline{v}^{(r_{1})}(\boldsymbol{k}_{1}) \gamma_{\alpha} u^{(r_{2})}(\boldsymbol{k}_{2}) \ \overline{u}^{(r_{2})}(\boldsymbol{k}_{2}) \gamma_{\beta} v^{(r_{1})}(\boldsymbol{k}_{1})$$
(1)

Podemos ahora hacer uso de las propiedades de espinores y matrices de Dirac, que conducen a multitud de identidades (Diracología).

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 En particular, puede verse que los dos estados de espín a lo largo del eje z satisfacen

$$u^{(1)}(\mathbf{p})\overline{u}^{(1)}(\mathbf{p}) = (\not \!\!\!/ + m)\frac{1 + \gamma_5 \not \!\!\!/}{2},$$
$$u^{(2)}(\mathbf{p})\overline{u}^{(2)}(\mathbf{p}) = (\not \!\!\!/ + m)\frac{1 - \gamma_5 \not \!\!/}{2},$$
$$v^{(1)}(\mathbf{p})\overline{v}^{(1)}(\mathbf{p}) = (\not \!\!\!/ - m)\frac{1 + \gamma_5 \not \!\!/}{2},$$
$$v^{(2)}(\mathbf{p})\overline{v}^{(2)}(\mathbf{p}) = (\not \!\!\!/ - m)\frac{1 - \gamma_5 \not \!\!/}{2},$$

donde $n^{\mu} = (0, 0, 0, 1)$ en el sistema de referencia en el que $p^{\mu} = (m, 0, 0, 0)$. En general,

$$u(\boldsymbol{p},n)\overline{u}(\boldsymbol{p},n) = (\not\!\!p+m)\frac{1+\gamma_5\not\!\!n}{2}, \quad v(\boldsymbol{p},n)\overline{v}(\boldsymbol{p},n) = (\not\!\!p-m)\frac{1+\gamma_5\not\!\!n}{2}$$

proyectan sobre polarizaciones bien definidas a lo largo de una dirección n^{μ} , que cumple $n^2 = -1$ y $p_{\mu}n^{\mu} = 0$.

- ▷ Si elegimos, por simplicidad, el eje *z* como dirección del movimiento, $p^{\mu} = (E, 0, 0, |\mathbf{p}|)$, los operadores anteriores proyectan sobre los dos estados de helicidad de partícula y antipartícula, respectivamente, si tomamos $n^{\mu} = \pm (|\mathbf{p}|/m, 0, 0, E/m)$.
- ▷ En particular, en el límite ultrarrelativista ($E \gg m$) los proyectores sobre quiralidades *right* y *left* de partícula y antipartícula son:

$$\begin{split} u^{(1)}(\boldsymbol{p})\overline{u}^{(1)}(\boldsymbol{p}) &= (\not\!\!p+m)\frac{1+\gamma_5\not\!\!/}{2} \to u_R(p)\overline{u}_R(\boldsymbol{p}) = (\not\!\!p+m)\frac{1+\gamma_5}{2}, \\ u^{(2)}(\boldsymbol{p})\overline{u}^{(2)}(\boldsymbol{p}) &= (\not\!\!p+m)\frac{1-\gamma_5\not\!\!/}{2} \to u_L(p)\overline{u}_L(\boldsymbol{p}) = (\not\!\!p+m)\frac{1-\gamma_5}{2}, \\ v^{(1)}(\boldsymbol{p})\overline{v}^{(1)}(\boldsymbol{p}) &= (\not\!\!p-m)\frac{1+\gamma_5\not\!\!/}{2} \to v_L(p)\overline{v}_L(\boldsymbol{p}) = (\not\!\!p-m)\frac{1-\gamma_5}{2}, \\ v^{(2)}(\boldsymbol{p})\overline{v}^{(2)}(\boldsymbol{p}) &= (\not\!\!p-m)\frac{1-\gamma_5\not\!\!/}{2} \to v_R(\boldsymbol{p})\overline{v}_R(\boldsymbol{p}) = (\not\!\!p-m)\frac{1+\gamma_5}{2}. \end{split}$$

▷ Otra propiedad que se demuestra fácilmente de lo anterior es

$$\overline{u}(\boldsymbol{p},n)\Gamma u(\boldsymbol{p},n) = \operatorname{Tr}\left[\Gamma(\not\!\!p+m)\frac{1+\gamma_5\not\!\!n}{2}\right], \quad \overline{v}(\boldsymbol{p},n)\Gamma v(\boldsymbol{p},n) = \operatorname{Tr}\left[\Gamma(\not\!\!p-m)\frac{1+\gamma_5\not\!\!n}{2}\right]$$

donde Γ es una matriz 4 × 4 arbitraria.

 Por otro lado, si los fermiones no están polarizados el cálculo se simplifica notablemente pues podemos aplicar directamente las relaciones de completitud,

$$\sum_{s} u^{(s)}(\boldsymbol{p}) \overline{u}^{(s)}(\boldsymbol{p}) = \not \!\!\! \boldsymbol{p} + m , \quad \sum_{s} v^{(s)}(\boldsymbol{p}) \overline{v}^{(s)}(\boldsymbol{p}) = \not \!\!\! \boldsymbol{p} - m ,$$

que conducen a

$$\sum_{s} \overline{u}^{(s)}(\boldsymbol{p}) \Gamma u^{(s)}(\boldsymbol{p}) = \operatorname{Tr} \left[\Gamma(\not \!\!\!p + m) \right] , \quad \sum_{s} \overline{v}^{(s)}(\boldsymbol{p}) \Gamma v^{(s)}(\boldsymbol{p}) = \operatorname{Tr} \left[\Gamma(\not \!\!\!p - m) \right]$$

• Volvamos a nuestro cálculo y supongamos por simplicidad que tanto los fermiones iniciales como los finales no están polarizados. Tenemos entonces que promediar sobre espines iniciales y sumar sobre espines finales:

$$\begin{split} \widetilde{\sum_{r_i}} \sum_{s_i} |\mathcal{M}|^2 &= \frac{1}{4} \sum_{r_i} \sum_{s_i} |\mathcal{M}|^2 \\ &= \frac{e^4}{4q^4} \operatorname{Tr}[\gamma^{\alpha}(\not \!\!\! p_1 - M)\gamma^{\beta}(\not \!\!\! p_2 + M)] \operatorname{Tr}[\gamma_{\alpha}(\not \!\!\! k_2 + m)\gamma_{\beta}(\not \!\!\! k_1 - m)] , \end{split}$$

que aparece como el producto de las trazas de las dos cadenas fermiónicas.

• Para hallar las trazas volvemos a recurrir a la Diracología. Necesitamos en particular,

$$Tr[\# \text{ impar } \gamma' s] = 0$$

$$Tr[\gamma^{\mu}\gamma^{\nu}] = 4g^{\mu\nu}$$

$$Tr[\gamma^{\mu}\gamma^{\nu}\gamma^{\rho}\gamma^{\sigma}] = 4(g^{\mu\nu}g^{\rho\sigma} - g^{\mu\rho}g^{\nu\sigma} + g^{\mu\sigma}g^{\nu\rho})$$

de donde

$$\operatorname{Tr}[\gamma^{\alpha}(\not p_{1} - M)\gamma^{\beta}(\not p_{2} + M)] = \operatorname{Tr}[\gamma^{\alpha}\not p_{1}\gamma^{\beta}\not p_{2}] - M^{2}\operatorname{Tr}[\gamma^{\alpha}\gamma^{\beta}]$$

$$= 4(p_{1}^{\alpha}p_{2}^{\beta} - (p_{1}p_{2})g^{\alpha\beta} + p_{1}^{\beta}p_{2}^{\alpha}) - 4M^{2}g^{\alpha\beta}$$

$$\operatorname{Tr}[\gamma_{\alpha}(\not k_{2} + m)\gamma_{\beta}(\not k_{1} - m)] = \operatorname{Tr}[\gamma_{\alpha}\not k_{1}\gamma_{\beta}\not k_{2}] - m^{2}\operatorname{Tr}[\gamma_{\alpha}\gamma_{\beta}]$$

$$= 4(k_{1\alpha}k_{2\beta} - (k_{1}k_{2})g_{\alpha\beta} + k_{1\beta}k_{2\alpha}) - 4m^{2}g_{\alpha\beta}$$

y por tanto,

$$\begin{split} \widetilde{\sum_{r_i}} \sum_{s_i} |\mathcal{M}|^2 &= \frac{16e^4}{4q^4} [(p_1k_1)(p_2k_2) - (p_1p_2)(k_1k_2) + (p_1k_2)(p_2k_1) - m^2(p_1p_2) \\ &- (p_1p_2)(k_1k_2) + 4(p_1p_2)(k_1k_2) - (p_1p_2)(k_1k_2) + 4m^2(p_1p_2) \\ &+ (p_1k_2)(p_2k_1) - (p_1p_2)(k_1k_2) + (p_1k_1)(p_2k_2) - m^2(p_1p_2) \\ &- M^2(k_1k_2) + 4M^2(k_1k_2) - M^2(k_1k_2) + 4M^2m^2] \\ &= \frac{8e^4}{q^4} [(p_1k_1)(p_2k_2) + (p_1k_2)(p_2k_1) + m^2(p_1p_2) + M^2(k_1k_2) + 2M^2m^2] \;. \end{split}$$

• El siguiente paso es elegir un sistema de referencia. Supongamos el sistema centro de masas y sea θ el ángulo que forma el μ^+ saliente con el e^+ incidente,

$$\begin{split} k_1^{\mu} &= E(1,0,0,\beta_i) ,\\ k_2^{\mu} &= E(1,0,0,-\beta_i) , \quad \beta_i = \sqrt{1-m^2/E^2} ,\\ p_1^{\mu} &= E(1,\beta_f \sin \theta,0,\beta_f \cos \theta) ,\\ p_2^{\mu} &= E(1,-\beta_f \sin \theta,0,-\beta_f \cos \theta) , \quad \beta_f = \sqrt{1-M^2/E^2} . \end{split}$$

Entonces,

$$q^{2} = (k_{1} + k_{2})^{2} = (p_{1} + p_{2})^{2} = E_{CM}^{2} = 4E^{2} ,$$

$$(p_{1}k_{1}) = (p_{2}k_{2}) = E^{2}(1 - \beta_{i}\beta_{f}\cos\theta) ,$$

$$(p_{1}k_{2}) = (p_{2}k_{1}) = E^{2}(1 + \beta_{i}\beta_{f}\cos\theta) ,$$

$$(p_{1}p_{2}) = E^{2}(1 + \beta_{f}^{2}) = E^{2}(2 - M^{2}/E^{2}) ,$$

$$(k_{1}k_{2}) = E^{2}(1 + \beta_{i}^{2}) = E^{2}(2 - m^{2}/E^{2})$$
Un proceso sencillo: $e^+e^- \rightarrow \mu^+\mu^-$

y la expresión anterior queda

$$\widetilde{\sum_{r_i}} \sum_{s_i} |\mathcal{M}|^2 = \frac{e^4}{2E^4} [2E^4(1+\beta_i^2\beta_f^2\cos^2\theta) + 2E^2(m^2+M^2)]$$
$$= e^4 \left[1 + 4\frac{m^2+M^2}{E_{\rm CM}^2} + \left(1 - \frac{4m^2}{E_{\rm CM}^2}\right) \left(1 - \frac{4M^2}{E_{\rm CM}^2}\right)\cos^2\theta \right]$$

• La sección eficaz diferencial del proceso se obtiene a partir de

$$\frac{d\sigma}{d\Omega} = \frac{1}{64\pi^2 E_{\rm CM}^2} \frac{|\mathbf{p}|}{|\mathbf{k}|} |\mathcal{M}|^2$$
$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2}{4E_{\rm CM}^2} \sqrt{\frac{E_{\rm CM}^2 - 4M^2}{E_{\rm CM}^2 - 4m^2}} \left[1 + 4\frac{m^2 + M^2}{E_{\rm CM}^2} + \left(1 - \frac{4m^2}{E_{\rm CM}^2}\right) \left(1 - \frac{4M^2}{E_{\rm CM}^2}\right) \cos^2\theta \right]$$

donde se ha sustituido la constante de estructura fina $\alpha = e^2/(4\pi)$.

Un proceso sencillo: $e^+e^- \rightarrow \mu^+\mu^-$

▷ Nótese que $E_{CM} > 2M > 2m$, la energía umbral del proceso. La sección eficaz total es

$$\sigma = \int \mathrm{d}\Omega \frac{\mathrm{d}\sigma}{\mathrm{d}\Omega} = 2\pi \int \mathrm{d}\cos\theta \; \frac{\mathrm{d}\sigma}{\mathrm{d}\Omega} \; .$$

▷ En el límite ultrarrelativista ($E_{CM} \gg M > m$),

$$rac{\mathrm{d}\sigma}{\mathrm{d}\Omega}
ightarrow rac{lpha^2}{4E_{\mathrm{CM}}^2} (1 + \cos^2 \theta)$$
 $\sigma
ightarrow rac{4\pi lpha^2}{3E_{\mathrm{CM}}^2} \, .$

Comentarios Sobre los signos relativos entre diagramas

- En QED se trabaja con campos espinoriales y ya hemos visto que hay que tener cuidado porque las contracciones de Wick de estos campos pueden dar lugar a signos relativos entre los distintos diagramas que contribuyen a la amplitud de un proceso. Recordemos que hay que mirar si la reordenación de los espinores corresponde a una permutación par o impar. Veamos unos cuantos ejemplos.
- Scattering de Bhabha: $e^+e^- \rightarrow e^+e^-$



Comentarios Sobre los signos relativos entre diagramas

– Scattering de Møller: $e^-e^- \rightarrow e^-e^-$



– Scattering de Compton: $e\gamma \rightarrow e\gamma$ (¡no hay cambio de signo!)



Comentarios | Sobre partículas idénticas

Recordemos también que si hay dos partículas idénticas en el estado final (por ejemplo, γγ, e⁺e⁺, e⁻e⁻) la sección eficaz total es

$$\sigma = \frac{1}{2} \int \mathrm{d}\Omega \; \frac{\mathrm{d}\sigma}{\mathrm{d}\Omega}$$

- Caso del fotón.

Tiene dos estados de polarización (transversos). Supongamos el sistema de referencia en el que $k^{\mu} = (\omega, 0, 0, \omega)$ (nuestras conclusiones serán independientes de esta elección gracias a la covariancia Lorentz). Entonces, pueden ser

lineales: $\epsilon^{\mu}(k,1) = (0,1,0,0)$, $\epsilon^{\mu}(k,2) = (0,0,1,0)$ elípticas: $\epsilon^{\mu}(k,L) = (0,\cos\theta, i\sin\theta, 0)$, $\epsilon^{\mu}(k,R) = (0,\cos\theta, -i\sin\theta, 0)$.

En cualquier caso, si sumamos sobre los dos estados de polarización,

Veamos que, debido a la invariancia gauge, en la práctica podemos ignorar el término $Q_{\mu\nu}$.

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- Caso del fotón.

En efecto, la amplitud de un proceso arbitrario de QED que involucre un fotón externo con momento k (tomamos un fotón saliente) puede escribirse con toda generalidad como

$$\mathcal{M}(m{k},\lambda)=m{\epsilon}^*_\mu(m{k},\lambda)\mathcal{M}^\mu(m{k})$$

y cualquier observable, en este sistema de referencia, será proporcional a

$$\sum_{\lambda} |\mathcal{M}(\boldsymbol{k},\lambda)|^{2} = \sum_{\lambda=1,2} \epsilon_{\mu}^{*}(\boldsymbol{k},\lambda) \epsilon_{\nu}(\boldsymbol{k},\lambda) \mathcal{M}^{\mu}(\boldsymbol{k}) \mathcal{M}^{\nu*}(\boldsymbol{k})$$
$$= |\mathcal{M}^{1}(\boldsymbol{k})|^{2} + |\mathcal{M}^{2}(\boldsymbol{k})|^{2}.$$
(3)

Caso del fotón.

Ahora bien, sabemos que el campo del fotón se acopla a una corriente conservada⁶ mediante una interacción $\int d^4x \ j^{\mu}(x) A_{\mu}(x)$, con $\partial_{\mu} j^{\mu}(x) = 0$, así que

$$\mathcal{M}^{\mu}(\boldsymbol{k}) = \int \mathrm{d}^{4}x \, \mathrm{e}^{\mathrm{i}kx} \left\langle f \left| j^{\mu}(x) \left| i \right\rangle \right.$$

donde los estados inicial y final incluyen todas las partículas externas excepto el fotón en cuestión.

Como la simetría gauge se debe preservar también a nivel cuántico, de la conservación de la corriente y la expresión anterior deducimos^a

$$egin{aligned} &k_\mu \mathcal{M}^\mu(m{k}) = \mathrm{i} \int \mathrm{d}^4 x \; \mathrm{e}^{\mathrm{i} k x} \left< f \right| \partial_\mu j^\mu(x) \left| i \right> = 0 \ &\Rightarrow k_\mu \mathcal{M}^\mu(m{k}) = \omega \mathcal{M}^0(m{k}) - \omega \mathcal{M}^3(m{k}) = 0 \Rightarrow \mathcal{M}^0(m{k}) = \mathcal{M}^3(m{k}) \;. \end{aligned}$$

$${}^{a}0 = \int \mathrm{d}^{4}x \,\partial_{\mu} \left[\mathrm{e}^{\mathrm{i}kx} \langle f | j^{\mu}(x) | i \rangle \right] = \mathrm{i}k_{\mu} \int \mathrm{d}^{4}x \, \mathrm{e}^{\mathrm{i}kx} \langle f | j^{\mu}(x) | i \rangle + \int \mathrm{d}^{4}x \, \mathrm{e}^{\mathrm{i}kx} \langle f | \partial_{\mu}j^{\mu}(x) | i \rangle.$$

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- Caso del fotón.

Así que podemos reescribir $\sum_{\lambda} |\mathcal{M}(\boldsymbol{k},\lambda)|^2$ como

$$\sum_{\lambda=1,2} \epsilon_{\mu}^{*}(\boldsymbol{k},\lambda) \epsilon_{\nu}(\boldsymbol{k},\lambda) \mathcal{M}^{\mu}(\boldsymbol{k}) \mathcal{M}^{\nu*}(\boldsymbol{k})$$
$$= |\mathcal{M}^{1}(\boldsymbol{k})|^{2} + |\mathcal{M}^{2}(\boldsymbol{k})|^{2} + |\mathcal{M}^{3}(\boldsymbol{k})|^{2} - |\mathcal{M}^{0}(\boldsymbol{k})|^{2}$$

que equivale a reemplazar

$$\sum_{\lambda} \epsilon^*_{\mu}(\boldsymbol{k},\lambda) \epsilon_{\nu}(\boldsymbol{k},\lambda)
ightarrow -g_{\mu
u} \; .$$

- Caso de un bosón vectorial masivo.

Tiene tres estados de polarización (uno longitudinal y dos transversos). En este caso podemos elegir el sistema de referencia en reposo, $k^{\mu} = (M, 0, 0, 0)$ y los estados de polarización

 $\epsilon^{\mu}(\mathbf{k},1) = (0,1,0,0)$, $\epsilon^{\mu}(\mathbf{k},2) = (0,0,1,0)$, $\epsilon^{\mu}(\mathbf{k},3) = (0,0,0,1)$.

Si sumamos sobre polarizaciones,

en el sistema de referencia en reposo.

- Caso de un bosón vectorial masivo.

Podemos obtener la expresión válida para $k^{\mu} = (k^0, \mathbf{k}) \operatorname{con} M^2 = (k^0)^2 - \mathbf{k}^2$ haciendo un boost con $\gamma = k^0 / M$, $\gamma \beta = \mathbf{k} / M$,

$$\Lambda^{\mu}_{\ \mu'} = \begin{pmatrix} \gamma & \gamma \beta_1 & \gamma \beta_2 & \gamma \beta_3 \\ \gamma \beta_1 & & \\ \gamma \beta_2 & \delta_{ij} + (\gamma - 1) \frac{\beta_i \beta_j}{\beta^2} \\ \gamma \beta_3 & & \end{pmatrix}$$

que conduce a

$$\sum_{\lambda} \epsilon_{\mu}^{*}(\boldsymbol{k},\lambda) \epsilon_{\nu}(\boldsymbol{k},\lambda) = -g_{\mu\nu} + \Lambda^{0}_{\ \mu} \Lambda^{0}_{\ \nu} = -g_{\mu\nu} + \frac{k_{\mu}k_{\nu}}{M^{2}}$$

Loop calculations

Structure of one-loop amplitudes

• Consider the following generic one-loop diagram with *N* external legs:



$$k_1 = p_1, \quad k_2 = p_1 + p_2, \quad \dots \quad k_{N-1} = \sum_{i=1}^{N-1} p_i$$

• It contains general integrals of the kind:

$$\frac{i}{16\pi^2} T^N_{\mu_1\dots\mu_P} \equiv \mu^{4-D} \int \frac{d^D q}{(2\pi)^D} \frac{q_{\mu_1}\cdots q_{\mu_P}}{[q^2 - m_0^2][(q+k_1)^2 - m_1^2]\cdots[(q+k_{N-1})^2 - m_{N-1}^2]}$$

Structure of one-loop amplitudes

- ▷ *D* dimensional integration in dimensional regularization
- ▷ Integrals are symmetric under permutations of Lorentz indices
- \triangleright Scale μ introduced to keep the proper mass dimensions
- ▷ *P* is the number of *q*'s in the numerator and determines the tensor structure of the integral (scalar if *P* = 0, vector if *P* = 1, etc.). Note that $P \le N$
- ▷ Notation: *A* for T^1 , *B* for T^2 , etc. For example, the scalar integrals A_0 , B_0 , etc.
- ▷ The tensor integrals can be decomposed as a linear combination of the Lorentz covariant tensors that can be built with $g_{\mu\nu}$ and a set of linearly independent momenta [Passarino, Veltman '79]
- ▷ The choice of basis is not unique

Here we use the basis formed by $g_{\mu\nu}$ and the momenta k_i , where the tensor coefficients are totally symmetric in their indices [Denner '93]

This is the basis used by the computer package LoopTools [www.feynarts.de/looptools]

Structure of one-loop amplitudes

• We focus here on:

$$B_{\mu} = k_{1\mu}B_{1}$$

$$B_{\mu\nu} = g_{\mu\nu}B_{00} + k_{1\mu}k_{1\nu}B_{11}$$

$$C_{\mu} = k_{1\mu}C_{1} + k_{2\mu}C_{2}$$

$$C_{\mu\nu} = g_{\mu\nu}C_{00} + \sum_{i,j=1}^{2} k_{i\mu}k_{j\nu}C_{ij}$$

$$C_{\mu\nu\rho} = \dots$$

- We will see that the scalar integrals A_0 and B_0 and the tensor integral coefficients B_1 , B_{00} , B_{11} and C_{00} are divergent in D = 4 dimensions (ultraviolet divergence, equivalent to take cutoff $\Lambda \rightarrow \infty$ in q)
- It is possible to express every tensor coefficient in terms of scalar integrals (scalar reduction) [Denner '93]

- Basic ingredients:
- Euler Gamma function:

$$\Gamma(x+1) = x\Gamma(x)$$

Taylor expansion around poles at x = 0, -1, -2, ...:

$$x = 0: \quad \Gamma(x) = \frac{1}{x} - \gamma + \mathcal{O}(x)$$
$$x = -1: \quad \Gamma(x) = -\frac{1}{(x+1)} + \gamma - 1 + \dots + \mathcal{O}(x+1)$$

where $\gamma \approx 0.5772\ldots$ is Euler-Mascheroni constant

– Feynman parameters:

$$\frac{1}{a_1 a_2 \cdots a_n} = \int_0^1 \mathrm{d} x_1 \cdots \mathrm{d} x_n \ \delta\left(\sum_{i=1}^n x_i - 1\right) \frac{(n-1)!}{[x_1 a_1 + x_2 a_2 + \cdots + x_n a_n]^n}$$

– The following integrals (with $\varepsilon \to 0^+$) will be needed:

$$\int \frac{d^{D}q}{(2\pi)^{D}} \frac{1}{(q^{2} - \Delta + i\varepsilon)^{n}} = \frac{(-1)^{n}i}{(4\pi)^{D/2}} \frac{\Gamma(n - D/2)}{\Gamma(n)} \left(\frac{1}{\Delta}\right)^{n - D/2}$$
$$\Rightarrow \int \frac{d^{D}q}{(2\pi)^{D}} \frac{q^{2}}{(q^{2} - \Delta + i\varepsilon)^{n}} = \frac{(-1)^{n - 1}i}{(4\pi)^{D/2}} \frac{D}{2} \frac{\Gamma(n - D/2 - 1)}{\Gamma(n)} \left(\frac{1}{\Delta}\right)^{n - D/2 - 1}$$

▷ Let's solve the first integral in Euclidean space: $q^0 = iq_E^0$, $q = q_E$, $q^2 = -q_E^2$,

$$\int \frac{d^{D}q}{(2\pi)^{D}} \frac{1}{(q^{2} - \Delta + i\varepsilon)^{n}} = i(-1)^{n} \int \frac{d^{D}q_{E}}{(2\pi)^{D}} \frac{1}{(q_{E}^{2} + \Delta)^{n}}$$

(equivalent to a Wick rotation of 90°). The second integral follows from this one



In *D*-dimensional spherical coordinates:

$$\int \frac{\mathrm{d}^{D} q_{E}}{(2\pi)^{D}} \frac{1}{(q_{E}^{2} + \Delta)^{n}} = \int \mathrm{d}\Omega_{D} \int_{0}^{\infty} \mathrm{d}q_{E} q_{E}^{D-1} \frac{1}{(q_{E}^{2} + \Delta)^{n}} \equiv \mathcal{I}_{A} \times \mathcal{I}_{B}$$
where
$$\mathcal{I}_{A} = \int \mathrm{d}\Omega_{D} = \frac{2\pi^{D/2}}{\Gamma(D/2)}$$

since
$$(\sqrt{\pi})^D = \left(\int_{-\infty}^{\infty} dx \ e^{-x^2}\right)^D = \int d^D x \ e^{-\sum_{i=1}^D x_i^2} = \int d\Omega_D \int_0^{\infty} dx \ x^{D-1} e^{-x^2}$$

$$= \left(\int d\Omega_D\right) \frac{1}{2} \int_0^{\infty} dt \ t^{D/2-1} e^{-t} = \left(\int d\Omega_D\right) \frac{1}{2} \Gamma(D/2)$$

and, changing variables: $t = q_E^2$, $z = \Delta/(t + \Delta)$, we have

$$\mathcal{I}_{B} = \frac{1}{2} \left(\frac{1}{\Delta}\right)^{n-D/2} \int_{0}^{1} dz \ z^{n-D/2-1} (1-z)^{D/2-1} = \frac{1}{2} \left(\frac{1}{\Delta}\right)^{n-D/2} \frac{\Gamma(n-D/2)\Gamma(D/2)}{\Gamma(n)}$$

where Euler Beta function was used: $B(\alpha, \beta) = \int_0^1 dz \ z^{\alpha-1} (1-z)^{\beta-1} = \frac{\Gamma(\alpha)\Gamma(\beta)}{\Gamma(\alpha+\beta)}$

Explicit calculationTwo-point functionsp $q + k_1$ p m_1 p m_1 q

$$\frac{\mathrm{i}}{16\pi^2} \{ B_0, \ B^{\mu}, \ B^{\mu\nu} \} (\mathrm{args}) = \mu^{4-D} \int \frac{\mathrm{d}^D q}{(2\pi)^D} \frac{\{ 1, \ q^{\mu}, \ q^{\mu} q^{\nu} \}}{\left(q^2 - m_0^2 \right) \left[(q+p)^2 - m_1^2 \right]}$$

$$\triangleright \ k_1 = p$$

 \triangleright The integrals depend on the masses m_0 , m_1 and the invariant p^2 :

$$(args) = (p^2; m_0^2, m_1^2)$$

• Using Feynman parameters,

$$\frac{1}{a_1 a_2} = \int_0^1 \mathrm{d}x \frac{1}{\left[a_1 x + a_2 (1 - x)\right]^2}$$

$$\Rightarrow \frac{\mathrm{i}}{16\pi^2} \{ B_0, B^{\mu}, B^{\mu\nu} \} = \mu^{4-D} \int_0^1 \mathrm{d}x \int \frac{\mathrm{d}^D q}{(2\pi)^D} \frac{\{ 1, -A^{\mu}, q^{\mu}q^{\nu} + A^{\mu}A^{\nu} \}}{(q^2 - \Delta_2)^2}$$

with

$$\Delta_2 = x^2 p^2 + x(m_1^2 - m_0^2 - p^2) + m_0^2$$

$$a_1 = (q+p)^2 - m_1^2$$

 $a_2 = q^2 - m_0^2$

and a loop momentum shift to obtain a perfect square in the denominator:

$$q^{\mu} \rightarrow q^{\mu} - A^{\mu}$$
, $A^{\mu} = x p^{\mu}$

• Then, the scalar function is:

$$\frac{\mathrm{i}}{16\pi^2} B_0 = \mu^{4-D} \int_0^1 \mathrm{d}x \int \frac{\mathrm{d}^D q}{(2\pi)^D} \frac{1}{(q^2 - \Delta_2)^2}$$

$$\Rightarrow \quad B_0 = \Delta_{\epsilon} - \int_0^1 \mathrm{d}x \, \ln \frac{\Delta_2}{\mu^2} + \mathcal{O}(\epsilon) \qquad [D = 4 - \epsilon]$$

where $\Delta_{\epsilon} \equiv \frac{2}{\epsilon} - \gamma + \ln 4\pi$ and the Euler Gamma function was expanded around x = 0 for $D = 4 - \epsilon$, using $x^{\epsilon} = \exp{\{\epsilon \ln x\}} = 1 + \epsilon \ln x + O(\epsilon^2)$:

$$\mu^{4-D} \frac{\mathrm{i}\Gamma(2-D/2)}{(4\pi)^{D/2}} \left(\frac{1}{\Delta_2}\right)^{2-D/2} = \frac{\mathrm{i}}{16\pi^2} \left(\Delta_{\epsilon} - \ln\frac{\Delta_2}{\mu^2}\right) + \mathcal{O}(\epsilon)$$

• Comparing with the definitions of the tensor coefficientes we have:

$$\frac{\mathrm{i}}{16\pi^2} B^{\mu} = -\mu^{4-D} \int_0^1 \mathrm{d}x \int \frac{\mathrm{d}^D q}{(2\pi)^D} \frac{A^{\mu}}{(q^2 - \Delta_2)^2}$$

$$\Rightarrow \quad B_1 = -\frac{1}{2} \Delta_{\epsilon} + \int_0^1 \mathrm{d}x \, x \ln \frac{\Delta_2}{\mu^2} + \mathcal{O}(\epsilon) \qquad [D = 4 - \epsilon]$$

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Two-point functions

and

$$\begin{aligned} \frac{i}{16\pi^2} B^{\mu\nu} &= \mu^{4-D} \int_0^1 dx \int \frac{d^D q}{(2\pi)^D} \frac{(q^2/D)g^{\mu\nu} + A^{\mu}A^{\nu}}{(q^2 - \Delta_2)^2} \\ \Rightarrow B_{00} &= -\frac{1}{12} (p^2 - 3m_0^2 - 3m_1^2) \Delta_{\epsilon} + \mathcal{O}(\epsilon) \qquad [D = 4 - \epsilon] \\ B_{11} &= \frac{1}{3} \Delta_{\epsilon} - \int_0^1 dx \ x^2 \ln \frac{\Delta_2}{\mu^2} + \mathcal{O}(\epsilon) \qquad [D = 4 - \epsilon] \end{aligned}$$

where $q^{\mu}q^{\nu}$ have been replaced by $(q^2/D)g^{\mu\nu}$ in the integrand and the Euler Gamma function was expanded around x = -1 for $D = 4 - \epsilon$:

$$-\mu^{4-D}\frac{\mathrm{i}\Gamma(1-D/2)}{(4\pi)^{D/2}2\Gamma(2)}\left(\frac{1}{\Delta_2}\right)^{1-D/2} = \frac{\mathrm{i}}{16\pi^2}\frac{1}{2}\left(\Delta_{\epsilon} - \ln\frac{\Delta_2}{\mu^2} + 1\right)\Delta_2 + \mathcal{O}(\epsilon)$$

Explicit calculation | Three-point functions p_1 $q+k_1$ m_0 $p_2 - p_1$ m_2 $q + k_2$ $-p_{2}$ $\frac{\mathrm{i}}{16\pi^2} \{ C_0, \ C^{\mu}, \ C^{\mu\nu} \} (\mathrm{args}) = \mu^{4-D} \int \frac{\mathrm{d}^D q}{(2\pi)^D} \frac{\{1, \ q^{\mu}, \ q^{\mu}q^{\nu}\}}{(q^2 - m_0^2) \left[(q + p_1)^2 - m_1^2\right] \left[(q + p_2)^2 - m_2^2\right]}$

It is convenient to choose the external momenta so that: \triangleright

$$k_1 = p_1, \quad k_2 = p_2.$$

The integrals depend on the masses m_0 , m_1 , m_2 and the invariants: \triangleright

$$(args) = (p_1^2, Q^2, p_2^2; m_0^2, m_1^2, m_2^2), \quad Q^2 \equiv (p_2 - p_1)^2.$$

• Using Feynman parameters,

$$\frac{1}{a_1 a_2 a_3} = 2 \int_0^1 \mathrm{d}x \int_0^{1-x} \mathrm{d}y \frac{1}{\left[a_1 x + a_2 y + a_3 (1-x-y)\right]^3}$$

$$\Rightarrow \frac{\mathrm{i}}{16\pi^2} \{ C_0, \ C^{\mu}, \ C^{\mu\nu} \} = 2\mu^{4-D} \int_0^1 \mathrm{d}x \int_0^{1-x} \mathrm{d}y \int \frac{\mathrm{d}^D q}{(2\pi)^D} \frac{\{1, \ -A^{\mu}, \ q^{\mu}q^{\nu} + A^{\mu}A^{\nu}\}}{(q^2 - \Delta_3)^3}$$

with

 $\Delta_3 = x^2 p_1^2 + y^2 p_2^2 + xy(p_1^2 + p_2^2 - Q^2) + x(m_1^2 - m_0^2 - p_1^2) + y(m_2^2 - m_0^2 - p_2^2) + m_0^2$

$$a_{1} = (q + p_{1})^{2} - m_{1}^{2}$$
$$a_{2} = (q + p_{2})^{2} - m_{2}^{2}$$
$$a_{3} = q^{2} - m_{0}^{2}$$

and a loop momentum shift to obtain a perfect square in the denominator:

$$q^\mu
ightarrow q^\mu - A^\mu$$
, $A^\mu = x p_1^\mu + y p_2^\mu$

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Tools: loop calculations

• Then the scalar function is:

$$\frac{i}{16\pi^2}C_0 = 2\mu^{4-D} \int_0^1 dx \int_0^{1-x} dy \int \frac{d^D q}{(2\pi)^D} \frac{1}{(q^2 - \Delta_3)^3}$$

$$\Rightarrow C_0 = -\int_0^1 dx \int_0^{1-x} dy \frac{1}{\Delta_3} \qquad [D=4]$$

• Comparing with the definitions of the tensor coefficientes we have:

$$\frac{i}{16\pi^2} C^{\mu} = -2\mu^{4-D} \int_0^1 dx \int_0^{1-x} dy \int \frac{d^D q}{(2\pi)^D} \frac{A^{\mu}}{(q^2 - \Delta_3)^3}$$

$$\Rightarrow C_1 = \int_0^1 dx \int_0^{1-x} dy \frac{x}{\Delta_3} \qquad [D=4]$$

$$C_2 = \int_0^1 dx \int_0^{1-x} dy \frac{y}{\Delta_3} \qquad [D=4]$$

Explicit calculation | Three-point functions

$$\begin{aligned} \frac{\mathrm{i}}{16\pi^2} C^{\mu\nu} &= 2\mu^{4-D} \int_0^1 \mathrm{d}x \int_0^{1-x} \mathrm{d}y \int \frac{\mathrm{d}^D q}{(2\pi)^D} \frac{(q^2/D)g^{\mu\nu} + A^\mu A^\nu}{(q^2 - \Delta_3)^3} \\ \Rightarrow & C_{11} &= -\int_0^1 \mathrm{d}x \int_0^{1-x} \mathrm{d}y \frac{x^2}{\Delta_3} \qquad [D=4] \\ & C_{22} &= -\int_0^1 \mathrm{d}x \int_0^{1-x} \mathrm{d}y \frac{y^2}{\Delta_3} \qquad [D=4] \\ & C_{12} &= -\int_0^1 \mathrm{d}x \int_0^{1-x} \mathrm{d}y \frac{xy}{\Delta_3} \qquad [D=4] \\ & C_{00} &= \frac{1}{4} \Delta_{\epsilon} - \frac{1}{2} \int_0^1 \mathrm{d}x \int_0^{1-x} \mathrm{d}y \ln \frac{\Delta_3}{\mu^2} + \mathcal{O}(\epsilon) \qquad [D=4-\epsilon] \end{aligned}$$

where $\Delta_{\epsilon} \equiv \frac{2}{\epsilon} - \gamma + \ln 4\pi$ and $q^{\mu}q^{\nu}$ was replaced by $(q^2/D)g^{\mu\nu}$ in the integrand In C_{00} the Euler Gamma function was expanded around x = 0 for $D = 4 - \epsilon$:

$$\mu^{4-D} \frac{\mathrm{i}\Gamma(2-D/2)}{(4\pi)^{D/2}\Gamma(3)} \left(\frac{1}{\Delta_3}\right)^{2-D/2} = \frac{\mathrm{i}}{16\pi^2} \frac{1}{2} \left(\Delta_{\epsilon} - \ln\frac{\Delta_3}{\mu^2}\right) + \mathcal{O}(\epsilon)$$

Note about Diracology in *D* **dimensions**

• Attention should be paid to the traces of Dirac matrices when working in *D* dimensions (dimensional regularization) since

$$\gamma^{\mu}\gamma^{\nu} + \gamma^{\nu}\gamma^{\mu} = 2g^{\mu\nu}\mathbf{1}_{4\times 4}, \quad g^{\mu\nu}g_{\mu\nu} = \operatorname{Tr}\{g^{\mu\nu}\} = D$$

Thus, the following identities involving contractions of Lorentz indices can be proven:

$$\begin{split} \gamma^{\mu}\gamma_{\mu} &= D \\ \gamma^{\mu}\gamma^{\nu}\gamma_{\mu} &= -(D-2)\gamma^{\nu} \\ \gamma^{\mu}\gamma^{\nu}\gamma^{\rho}\gamma_{\mu} &= 4g^{\nu\rho} - (4-D)\gamma^{\nu}\gamma^{\rho} \\ \gamma^{\mu}\gamma^{\nu}\gamma^{\rho}\gamma^{\sigma}\gamma_{\mu} &= -2\gamma^{\sigma}\gamma^{\rho}\gamma^{\nu} + (4-D)\gamma^{\nu}\gamma^{\rho}\gamma^{\sigma} \end{split}$$

Casos sencillos Integrales para $(g - 2)_{\ell}$ en QED

- Para el cálculo del momento dipolar magnético anómalo del electrón en QED, necesitaremos las siguientes funciones de tres puntos, evaluadas en:
 - $p_1^2 = p_2^2 = m^2 \qquad (\text{electrones on-shell})$ $Q^2 = 0 \qquad (\text{fot} on \text{ on-shell})$ $m_0 = 0 \qquad (\text{masa del fot} on)$ $m_1 = m_2 = m \qquad (\text{masa del electr} on)$ $\Rightarrow \Delta_3 = m^2 (x + y)^2.$

Las integrales básicas son entonces

$$C_0 = \text{divergente en el infrarrojo (no se necesita),}$$

$$C_1 = C_2 = \frac{1}{2m^2},$$

$$C_{11} = C_{22} = 2 C_{12} = -\frac{1}{6m^2}.$$

$$C_{00} = \text{divergente en el ultravioleta (no se necesita),}$$

$$\text{donde } C \equiv C(m^2, 0, m^2; 0, m^2, m^2).$$

Casos sencillos Integrales para $(g - 2)_{\ell}$ en SM y MSSM

- Para el cálculo de las contribuciones débiles (y de supersimetría) a los momentos dipolares magnéticos se necesitan las siguientes funciones de tres puntos, evaluadas en:
 - $p_1^2 = p_2^2 = 0$ (se desprecian las masas de los fermiones externos) $Q^2 = 0$ (fotón *on-shell*)
 - (masa de la partícula virtual no acoplada al fotón externo)
 - $m_1 = m_2 = M_2$ (masa de las otras partículas virtuales)

$$\Rightarrow \Delta_3 = (M_2^2 - M_1^2)(x + y) + M_1^2.$$

 $m_0 = M_1$

Casos sencillos Integrales para $(g - 2)_{\ell}$ en SM y MSSM

▷ Las integrales básicas son

$$C_{0} = \frac{1}{M_{1}^{2}} \frac{1 - x_{21} + \ln x_{21}}{(1 - x_{21})^{2}},$$

$$C_{1} = C_{2} = \frac{1}{M_{1}^{2}} \frac{-3 + 4x_{21} - x_{21}^{2} - 2\ln x_{21}}{4(1 - x_{21})^{3}},$$

$$C_{11} = C_{22} = 2 C_{12} = \frac{1}{M_{1}^{2}} \frac{11 - 18x_{21} + 9x_{21}^{2} - 2x_{21}^{3} + 6\ln x_{21}}{18(1 - x_{21})^{4}},$$

$$C_{00} = \text{divergente en el ultravioleta (no se necesita).}$$

 \triangleright O bien

$$\overline{C}_{0} = \frac{1}{M_{1}^{2}} \frac{-1 + x_{21} - x_{21} \ln x_{21}}{(1 - x_{21})^{2}},$$

$$\overline{C}_{1} = \overline{C}_{2} = \frac{1}{M_{1}^{2}} \frac{1 - 4x_{21} + 3x_{21}^{2} - 2x_{21}^{2} \ln x_{21}}{4(1 - x_{21})^{3}},$$

$$\overline{C}_{11} = \overline{C}_{22} = 2 \overline{C}_{12} = \frac{1}{M_{1}^{2}} \frac{-2 + 9x_{21} - 18x_{21}^{2} + 11x_{21}^{3} - 6x_{21}^{3} \ln x_{21}}{18(1 - x_{21})^{4}},$$

donde $C \equiv C(0, 0, 0; M_1^2, M_2^2, M_2^2), \overline{C} \equiv C(0, 0, 0; M_2^2, M_1^2, M_1^2)$ y $x_{21} \equiv M_2^2 / M_1^2$.

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Casos sencillos | Transiciones magnéticas en SM y MSSM

- Para el cálculo de las contribuciones débiles (y de supersimetría) a las transiciones radiativas del tipo $\mu \rightarrow e\gamma$ se usan las funciones de tres puntos anteriores.
- Para comprobar que esta transición es puramente magnética, es decir no contribuyen las auto-energías de las patas externas, conviene conocer explícitamente C₀₀ evaluada en la misma configuración anterior:

$$C_{00}(0,0,0;M_1^2,M_2^2,M_2^2) = -\frac{1}{2}B_1(0;M_1^2,M_2^2)$$

y las siguientes funciones de dos puntos evaluadas en $p^2 = 0$, $m_0 = M_1$, $m_1 = M_2$:

$$B_{0}(0; M_{1}^{2}, M_{2}^{2}) = \Delta_{\epsilon} + 1 - \frac{M_{1}^{2} \ln \frac{M_{1}^{2}}{\mu^{2}} - M_{2}^{2} \ln \frac{M_{2}^{2}}{\mu^{2}}}{M_{1}^{2} - M_{2}^{2}}$$
$$B_{1}(0; M_{1}^{2}, M_{2}^{2}) = -\frac{1}{2}\Delta_{\epsilon} + \frac{4M_{1}^{4} - 3M_{2}^{4} - M_{1}^{2}M_{2}^{2} - 2\ln \frac{M_{1}^{2}}{M_{2}^{2}}}{4(M_{1}^{2} - M_{2}^{2})^{2}}$$
$$= -B_{0}(0; M_{2}^{2}, M_{1}^{2}) - B_{1}(0; M_{2}^{2}, M_{1}^{2})$$

Factores de forma dipolares a un loop



La estructura Lorentz más general del vértice vector-fermión contiene 24 términos independientes, que son combinaciones de los cuadrivectores *p* ≡ *p*₁ + *p*₂, *q* ≡ *p*₂ − *p*₁ y las 16 matrices de Dirac (indicadas abajo entre paréntesis):

• Con frecuencia el vértice se escribe de la siguiente forma:

$$i\Gamma^{\mu}(p_{1}, p_{2}) = ie \left[\gamma^{\mu}(F_{V} - F_{A}\gamma_{5}) + (iF_{M} + F_{E}\gamma_{5})\sigma^{\mu\nu}q_{\nu} + (iF_{S} + F_{P}\gamma_{5})q^{\mu} + (F_{MV} + iF_{EV}\gamma_{5})p^{\mu} + (F_{TS} + iF_{TP}\gamma_{5})\sigma^{\mu\nu}p_{\nu} + \dots\right].$$

- ▷ Los factores de forma F_i son en general funciones de todos los escalares independientes (invariantes Lorentz) que se puedan construir con los vectores p_1 y p_2 , es decir, $F_i(p_1^2, p_2^2, q^2)$. La constante e se ha introducido por conveniencia, de modo que los acoplamientos quedan normalizados a los de la electrodinámica cuántica (QED).
- Si ambos fermiones están *on-shell* (es decir, $p^2 = m^2$), la ecuación de Dirac nos permite eliminar los términos omitidos anteriormente y también F_{MV} , F_{EV} F_{TS} y F_{TP} , pues ya no son independientes y entonces

ij on-shell : $i\Gamma^{\mu}(p_1, p_2) = ie[\gamma^{\mu}(F_V - F_A\gamma_5) + (iF_M + F_E\gamma_5)\sigma^{\mu\nu}q_{\nu} + (iF_S + F_P\gamma_5)q^{\mu}].$

▷ Basta usar la siguiente relación entre matrices de Dirac:

$$\gamma_5 \gamma_
ho \epsilon^{
ho\mu
u\sigma} = rac{\mathrm{i}}{6} (\gamma^\mu \gamma^
u \gamma^\sigma + \gamma^\sigma \gamma^\mu \gamma^
u + \gamma^
u \gamma^\sigma \gamma^\mu - \gamma^
u \gamma^\mu \gamma^\sigma - \gamma^\mu \gamma^\sigma \gamma^
u - \gamma^\sigma \gamma^
u \gamma^\mu),$$

la ecuación de Dirac (ED): $p_1 u(p_1) = m_1 u(p_1)$, $p_2 u(p_2) = m_2 u(p_2)$, y las identidades de Gordon (que se deducen de la ED y $\gamma^{\mu} \gamma^{\nu} + \gamma^{\nu} \gamma^{\mu} = 2g^{\mu\nu}$):

$$\bar{u}(p_2)\sigma^{\mu\nu}(p_2\pm p_1)_{\nu}u(p_1) = \bar{u}(p_2) \left\{ -i(m_2\mp m_1)\gamma^{\mu} + i(p_2\mp p_1)^{\mu} \right\} u(p_1),$$

$$\bar{u}(p_2)\gamma_5\sigma^{\mu\nu}(p_2\pm p_1)_{\nu}u(p_1) = \bar{u}(p_2) \left\{ -i(m_2\pm m_1)\gamma^{\mu}\gamma_5 + i\gamma_5(p_2\mp p_1)^{\mu} \right\} u(p_1)$$

• Si el bosón vectorial *V* también está *on-shell*, su polarización satisface $q^{\mu}\varepsilon_{\mu} = 0$ y por tanto los factores de forma *F*_S, *F*_P son irrelevantes y el vértice se reduce a:

Vij on-shell : $i\Gamma^{\mu}(p_1, p_2) = ie \left[\gamma^{\mu}(F_V - F_A\gamma_5) + (iF_M + F_E\gamma_5)\sigma^{\mu\nu}q_{\nu}\right].$

Los factores de forma F_V , F_A y $F_{M,E}$ se denominan vectorial, axial y dipolares, respectivamente.

Una partícula masiva de spin 1 de momento p^{μ} tiene tres grados de libertad de polarización $\varepsilon^{\mu}(\lambda = 1, 2, 3)$ que verifican $p^{\mu}\varepsilon_{\mu}(\lambda) = 0$ y $\varepsilon^{\mu}(\lambda)\varepsilon_{\mu}(\lambda') = -\delta_{\lambda\lambda'}$. En reposo: $p^{\mu} = (M, 0, 0, 0)$, $\varepsilon_{1}^{\mu} = (0, 1, 0, 0)$, $\varepsilon_{2}^{\mu} = (0, 0, 1, 0)$, $\varepsilon_{3}^{\mu} = (0, 0, 0, 1)$. En movimiento: $p^{\mu} = (E, 0, 0, p)$, $\varepsilon_{x}^{\mu} = (0, 1, 0, 0)$, $\varepsilon_{y}^{\mu} = (0, 0, 1, 0)$, $\varepsilon_{L}^{\mu} = (p/M, 0, 0, E/M)$ [circular: $\varepsilon_{\pm}^{\mu} = 1/\sqrt{2}(\varepsilon_{x}^{\mu} \pm i\varepsilon_{y}^{\mu})$]. Si tiene masa nula (ej. el fotón) no existe el s.r. en reposo y sólo existen las dos polarizaciones transversales.

• Si $V = \gamma$ (fotón) la invariancia gauge U(1) impone la conservación de la corriente, $q_{\mu}\Gamma^{\mu} = 0$, y por tanto para fermiones *on-shell*:

$$[V = \gamma] \qquad (m_i - m_j)F_V + iq^2F_S = 0,$$

ij on-shell
$$-(m_i + m_j)F_A + q^2F_P = 0.$$

En consecuencia, si también el fotón está *on-shell* (q² = 0) y los fermiones son idénticos (m = m_i = m_j), necesariamente F_A = 0. El vértice electromagnético viene entonces descrito por tres constantes, relacionadas con la carga y los momentos dipolar magnético y dipolar eléctrico:

$$\gamma ii \text{ on-shell} : i\Gamma_{i=j}^{\mu} = ie \left[\gamma^{\mu}F_{V} + (iF_{M} + F_{E}\gamma_{5})\sigma^{\mu\nu}q_{\nu}\right]$$
▷ Entonces, de acuerdo con nuestra convención para la derivada covariante,

$$eQ_{f} \equiv -eF_{V}(0) = \text{ carga eléctrica del fermión } f,$$

$$\mu \equiv -\left(\frac{e}{2m}F_{V}(0) + eF_{M}(0)\right) = \text{ momento dipolar magnético (MDM)},$$

$$a \equiv 2m\frac{F_{M}(0)}{F_{V}(0)} = \text{ momento dipolar magnético anómalo (AMDM)},$$

$$d = -eF_{E}(0) = \text{ momento dipolar eléctrico (EDM)}.$$

• Así, a nivel árbol (electrodinámica clásica), un electrón tiene acoplamientos $F_V = 1$, $F_A = F_M = F_E = 0$, y por tanto carga $Q_e = -1$ y momento dipolar magnético

$$\mu = -\frac{g}{2}\frac{e}{2m}$$
, $g \approx 2 \iff$ interacción no relativista $\mu \cdot \mathbf{B} \equiv \mu \, \boldsymbol{\sigma} \cdot \mathbf{B}$

donde $\mathbf{S} = \frac{1}{2}\sigma$ es el spin y g es la razón giromagnética o factor de Landé. Nótese que el momento anómalo y la razón giromagnética se definen sólo para partículas cargadas. Sin embargo una partícula neutra puede tener momento magnético dado por $\mu = -eF_M(0)$.

Momento dipolar magnético



(5)

• Las correcciones cuánticas inducen valores no nulos de AMDM y EDM. Las condiciones de renormalización fijan $F_V(0) = -Q_f$ (a todo orden de teoría de perturbaciones), pero aparece un AMDM que viene dado por

$$a = \frac{g-2}{2} = -2m \frac{F_M(0)}{Q_f} \Rightarrow \mu = \frac{g}{2} \frac{eQ_f}{2m} = (1+a) \frac{eQ_f}{2m}.$$

Por otro lado, las ecuaciones anteriores implican F_V = F_A = 0 para fermiones distintos. Es decir, procesos tales como μ → eγ se deben sólo a transiciones dipolares,

$$\gamma i j \text{ on-shell}: \quad \mathrm{i} \Gamma^{\mu}_{i \neq j} = \mathrm{i} e (\mathrm{i} F_M + F_E \gamma_5) \sigma^{\mu \nu} q_{\nu}$$

- En general, todos los factores de forma son reales a nivel árbol para fermiones externos iguales, por la hermiticidad del lagrangiano, pero se pueden hacer complejos al introducir las correcciones cuánticas (regla de Cutkosky).
 La amplitud se hace compleja cuando sea posible cortar el diagrama en dos diagramas tales que ambos describan procesos físicos. Se trata de una aplicación del teorema óptico. Es fácil darse cuenta de que si V = γ la amplitud ha de ser siempre real porque el fotón tiene masa nula.
- Los factores de forma que acompañan a los operadores de dimensión mayor que cuatro (todos menos F_V y F_A), ej. los dipolares, son nulos a nivel árbol en cualquier teoría renormalizable. Por tanto sus correcciones a un loop son finitas. Además acoplan fermiones de quiralidades contrarias, por lo que deben ser proporcionales a alguna masa fermiónica, ya sea interna o externa.
- Los factores de forma F_V , F_A y F_M multiplican sendos bilineales pares bajo CP, mientras que F_E acompaña a uno impar. Esto significa que si CP se conserva el momento dipolar F_E se anula si i = j, aunque esto no ocurre si los fermiones externos son distintos.
- Similarmente, si P se conserva (ej. en QED) F_A y F_E son nulos.

• Los diagramas que contribuyen a un loop al vértice efectivo vector-fermión pueden agruparse en seis clases o topologías distintas:



El momento magnético anómalo

- Los momentos magnéticos anómalos de electrón y muón son observables medidos con gran precisión. Como se deben enteramente a correcciones cuánticas

 ponen especialmente a prueba la consistencia de la teoría.
- El momento magnético del electrón se ha medido en un ciclotrón en Harvard con una precisión impresionante: [Particle Data Group, Phys. Rev. D86 (2012) 010001]

$$\frac{g_e}{2} = 1.001\,159\,652\,180\,76\,(27)$$

• El momento magnético del muón se obtiene a partir de la frecuencia de precesión del spin respecto a un campo magnético homogéneo en un anillo de almacenamiento de muones:

$$\omega_{\mathrm{a}} = \mathrm{a}_{\mu} \frac{eB}{2m}, \quad \mathrm{a}_{\mu} = \frac{(\mathrm{g}_{\mu} - 2)}{2}$$

Se ha medido con gran precisión en el experimento E821 de Brookhaven:

[G.W. Bennett et al., Phys. Rev. D 73 (2006) 072003]

$$\frac{g_{\mu}}{2} = 1.001\,165\,920\,91\,(63)$$

El momento magnético anómalo en QED



 En QED sólo existe un diagrama a un loop (clase I) y el único acoplamiento no nulo es del tipo [VFF] con g_V = 1, g_A = 0. La configuración de masas y momentos es también muy simple. El AMDM del electrón es entonces:

$$F_M(0) = \frac{\alpha}{4\pi} 2m(C_1 + C_2 + C_{11} + C_{22} + 2C_{12}) = \frac{\alpha}{4\pi} \frac{1}{m'},$$

$$a = \frac{g-2}{2} = 2m F_M(0) = \frac{\alpha}{2\pi'},$$

donde $\alpha = e^2/4\pi$ es la constante de estructura fina y hemos utilizado las funciones de tres puntos con argumentos (m^2 , 0, m^2 ; 0, m^2 , m^2) evaluadas anteriormente.

El momento magnético anómalo en el SM

• Los diagramas a un loop en el gauge de 't Hooft-Feynman (incluyendo QED) son:



Para los nuevos, necesitaremos más paciencia, las reglas de Feynman del SM y las funciones de tres puntos con argumentos del tipo (0,0,0; M₁², M₂², M₂²) evaluadas anteriormente. (Se puede despreciar la masa del leptón ℓ.)

El momento magnético anómalo en el SM

• Sumando todas las contribuciones a un loop:

$$a_{\ell} = \frac{(g_{\ell} - 2)}{2} = \underbrace{\frac{\alpha}{2\pi}}_{QED} + \frac{G_F}{8\pi^2\sqrt{2}}m_{\ell}^2 \left\{ \underbrace{\frac{10}{3}}_{W} \underbrace{-\frac{1}{3}\left[5 - (1 - 4s_W^2)^2\right]}_{Z} + \underbrace{\mathcal{O}\left(\frac{m_{\ell}^2}{M_H^2}\ln\frac{M_H^2}{m_{\ell}^2}\right)}_{H} \right\}^{\circ}$$

donde

$$G_F = \frac{\pi \alpha}{\sqrt{2} s_W^2 M_W^2} (1 + \Delta r) = 1.1663787 \, (6) \times 10^{-5} \, \text{GeV}^{-2} \Leftrightarrow \text{vida media del muon}$$

$$\alpha^{-1} = 137.035\,999\,139 \, (31) \Leftrightarrow \text{comparando con} \, (g_e - 2) \text{ en QED a 5 loops}$$

$$s_W^2 = 1 - \frac{M_W^2}{M_Z^2} = 0.223, \quad m_e = 0.511 \, \text{MeV}, \quad m_\mu = 0.106 \, \text{GeV}, \quad m_\tau = 1.777 \, \text{GeV}.$$

- ▷ Nótese que $(1 4s_W^2)^2 \simeq 0.012$ por lo que la Z contribuye aproximadamente la mitad que la W y con signo opuesto. La contribución del Higgs es despreciable.
- A continuación resumimos las predicciones del SM y las medidas actuales del momento anómalo del muón a_{μ} [Particle Data Group, Phys. Rev. D98 (2018) 030001]

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El momento magnético anómalo en el SM

Resultados

CÁLCULOS TEÓRICOS		Contribución a $a_{\mu}(\times 10^{11})$	
QED	coeficiente de $\left(\frac{\alpha}{\pi}\right)^n$		
n = 1	0.5	116 140 973	
<i>n</i> = 2	0.765857425(17)		
<i>n</i> = 3	24.050 509 96 (32)		
<i>n</i> = 4	130.8796 (63)		
<i>n</i> = 5	753. (10)		
Total		116	584719
Débil	1. loop	195	
	2. loop		-41
Hadrónica	f f f f f f	LO: N(N)LO:	6 930 (34) 7 (26)
TOTAL		$a_{\mu}^{ m teo} = 116591810(43)$	
MEDIDA EXPERIMENTAL: E821 (Brookhaven '06)		$a_{\mu}^{\exp} = 116592089(63)$	
Muon $g - 2$ (Fermilab '21)		$a_{\mu}^{\exp} = 116592061(41)$	

 \Rightarrow Discrepancia de 3.7 σ (4.2 σ)!!

El proceso raro $\mu \rightarrow e\gamma$

- Se trata de un proceso con violación de sabor leptónico (LFV), que en el SM con neutrinos sin masa está prohibido.
- Sin embargo, en el SM con neutrinos masivos o en otras extensiones del SM, como supersimetría, este proceso puede darse.
- La colaboración MEG lleva acabo un experimento en el PSI (Suiza) desde el 2004 con el haz de muones más intenso del mundo. No han observado ningún suceso, lo que pone una cota actualmente de $\mathcal{B}(\mu \to e\gamma) < 4.2 \times 10^{-13}$. El objetivo es llegar en unos años hasta 10^{-14} .
- Los experimentos BaBar en el PEP-II de SLAC (EE UU) y Belle en el KEKB (Japón) ponen cotas a desintegraciones similares del τ . Las más actuales son $\mathcal{B}(\tau \to \mu \gamma) < 4.4 \times 10^{-8}$ y $\mathcal{B}(\tau \to e \gamma) < 3.3 \times 10^{-8}$.

El proceso raro $\mu \rightarrow e\gamma$

- Recordemos que se trata de transiciones dipolares, lo que puede comprobarse explícitamente hallando las contribuciones a F_V y F_A , que son nulas.
- Las anchuras y fracciones de desintegración relevantes son

$$\begin{split} \Gamma(\ell_j \to \ell_i \gamma) &= \frac{\alpha}{2} m_{\ell_j}^3 \left(|F_M|^2 + |F_E|^2 \right), \\ \Gamma(\ell_j \to \ell_i \nu_j \bar{\nu}_i) &= \frac{G_F^2 m_{\ell_j}^5}{192 \pi^3}, \quad G_F = \frac{\pi \alpha_W}{\sqrt{2} M_W^2}, \\ \frac{\mathcal{B}(\ell_j \to \ell_i \gamma)}{\mathcal{B}(\ell_j \to \ell_i \nu_j \bar{\nu}_i)} &= \frac{\Gamma(\ell_j \to \ell_i \gamma)}{\Gamma(\ell_j \to \ell_i \nu_i \bar{\nu}_j)} \\ &= \frac{12\alpha}{\pi} \frac{M_W^4}{m_{\ell_j}^2} \left(\frac{4\pi}{\alpha_W} \right)^2 \left(|F_M|^2 + |F_E|^2 \right), \end{split}$$

donde $\mathcal{B}(\ell_j \to \ell_i \nu_j \bar{\nu}_i) = 1/0.17/0.17$ para $\ell_j \ell_i = \mu e / \tau \mu / \tau e$.

$\mu \rightarrow e\gamma$ en el SM con neutrinos masivos

• Los diagramas a un loop en el gauge de 't Hooft-Feynman son:



A continuación listamos las contribuciones de cada clase de diagramas que se obtienen usando las funciones de tres puntos con argumentos (0,0,0; m²_{vi}, M²_W, M²_W).

$\mu \rightarrow e\gamma$ en el SM con neutrinos masivos

• Definiendo
$$x_i \equiv m_{\nu_i}^2 / M_W^2$$

II: $F_M = -iF_E = -\frac{\alpha_W}{16\pi} m_\mu \sum_i U_{ei} U_{\mu i}^* \left[3\overline{C}_{11} - \overline{C}_1 \right]$
IV: $F_M = -iF_E = -\frac{\alpha_W}{16\pi} m_\mu \sum_i U_{ei} U_{\mu i}^* x_i \left[\overline{C}_0 + 3\overline{C}_1 + \frac{3}{2} \overline{C}_{11} \right]$
V: $F_M = -iF_E = 0$
VI: $F_M = -iF_E = \frac{\alpha_W}{16\pi} m_\mu \sum_i U_{ei} U_{\mu i}^* \overline{C}_1$
Total: $F_M = -iF_E = \frac{\alpha_W}{16\pi} \frac{m_\mu}{M_W^2} \sum_i U_{ei} U_{\mu i}^* F_W(x_i)$
donde $F_W(x) = \frac{10 - 33x + 45x^2 - 4x^3}{12(1 - x)^3} + \frac{3x^3}{2(1 - x)^4} \ln x \to \frac{5}{6} - \frac{x}{4} + \mathcal{O}(x^2)$

• Por tanto, para neutrinos ligeros y usando la unitariedad de *U*,

$$\mathcal{B}(\mu \to e\gamma)|_{SM} = \frac{3\alpha}{2\pi} \left| \sum_{i} U_{ei} U_{\mu i}^* F_W(x_i) \right|^2 \simeq \frac{3\alpha}{32\pi} \left| \sum_{i} U_{ei} U_{\mu i}^* x_i \right|^2 \lesssim 10^{-54} ,$$

donde se han sustituido los ángulos de mezcla y Δm_{ij}^2 medidos en oscilaciones.

$\mu \rightarrow e\gamma$ en el SM con neutrinos masivos

 Nota importante: En las expresiones anteriores se ha despreciado m_e. Para recuperar la contribución de la W al momento magnético anómalo conviene reinsertar m_e:

$$F_{M} = \frac{\alpha_{W}}{16\pi} \sum_{i} \left\{ m_{\mu} U_{ei} U_{\mu i}^{*} + m_{e} U_{ei}^{*} U_{\mu i} \right\} [\dots]$$

$$-iF_{E} = \frac{\alpha_{W}}{16\pi} \sum_{i} \left\{ m_{\mu} U_{ei} U_{\mu i}^{*} - m_{e} U_{ei}^{*} U_{\mu i} \right\} [\dots]$$

Entonces obtenemos que en efecto, para $\ell = \mu = e$, la contribución de la W a a_{ℓ} es

$$a_{\ell} = 2m_{\ell}F_M(0) = \frac{G_F}{8\pi^2\sqrt{2}}4m_{\ell}^2F_W(0) = \frac{10}{3}\frac{G_F}{8\pi^2\sqrt{2}}m_{\ell}^2$$

pues, despreciando las correcciones radiativas de G_F ,

$$\frac{G_F}{8\pi^2\sqrt{2}} = \frac{\alpha_W}{16\pi} \frac{1}{M_W^2},$$

y d_{ℓ} = 0 si no hay fases complejas en *U* o si los neutrinos no tienen masa.