# Dynamics of interacting vortices on trapped Bose－Einstein condensates 

Pedro J．Torres<br>University of Granada

## Joint work with:

- P.G. Kevrekidis (University of Massachusetts, USA)
- Ricardo Carretero-González (San Diego State University, USA)
- D.J. Frantzeskakis (University of Athens, Greece)
- S. Middelkamp and P. Schmelcher (Universität Hamburg, Germany)
- David S. Hall and D.V. Freilich (University of Massachusetts, USA)


## Publications:

- P.J. Torres, R. Carretero-González, S. Middelkamp, P. Schmelcher, D.J. Frantzeskakis and P.G. Kevrekidis, Vortex Interaction Dynamics in Trapped Bose-Einstein Condensates, Communications on Pure and Applied Analysis 10 (2011) 1589-1615.
- S. Middelkamp, P. J. Torres, P. G. Kevrekidis, D.J. Frantzeskakis, R. Carretero-González, P. Schmelcher, D.V. Freilich, and D.S. Hall, Guiding-center dynamics of vortex dipoles in Bose-Einstein condensates., Phys. Rev. A, Vol. 84, 011605(R) (2011).
- P. J. Torres, P. G. Kevrekidis, D.J. Frantzeskakis, R. Carretero-González, P. Schmelcher, and D.S. Hall, Dynamics of Vortex Dipoles in Confined Bose-Einstein Condensates., Phys. Lett. A, Volume 375, Issue 33 (2011), pp. 3044-3050


## Physical background

- A Bose-Einstein condensate (BEC) is a state of matter of a dilute gas of weakly interacting bosons confined in an external potential and cooled to temperatures very near absolute zero ( 0 K or $-273.15-\mathrm{C}$ ). Under such conditions, a large fraction of the bosons occupy the lowest quantum state of the external potential, at which point quantum effects become apparent on a macroscopic scale.


## Physical background

- A Bose-Einstein condensate (BEC) is a state of matter of a dilute gas of weakly interacting bosons confined in an external potential and cooled to temperatures very near absolute zero ( 0 K or $-273.15-\mathrm{C}$ ). Under such conditions, a large fraction of the bosons occupy the lowest quantum state of the external potential, at which point quantum effects become apparent on a macroscopic scale.
- Theoretically predicted by Satyendra Nath Bose and Albert Einstein in 1924-25. First experiment performed in 1995 by Eric Cornell and Carl Wieman (Nobel Prize in Physics 2001).



## Physical background

Vortices can be created on a BEC.


# The dynamical evolution of such vortices is a natural question 

## The model for n vortices

$$
\begin{aligned}
\dot{x}_{k} & =-c\left(r_{k}, t\right) S_{k} y_{k}-b \sum_{j \neq k} S_{j} \frac{y_{k}-y_{j}}{r_{j k}^{2}}, \\
\dot{y}_{k} & =+c\left(r_{k}, t\right) S_{k} x_{k}+b \sum_{j \neq k} S_{j} \frac{x_{k}-x_{j}}{r_{j k}^{2}}, \quad k=1 \ldots n
\end{aligned}
$$

where
$\left(x_{k}, y_{k}\right) \equiv$ coordinate of vortex $k$
$S_{k}= \pm 1$ charge of vortex $k$
$r_{j k}=\sqrt{\left(x_{j}-x_{k}\right)^{2}+\left(y_{j}-y_{k}\right)^{2}} \equiv$ separation between vortex $j$ and vortex $k$
$r_{k}=\sqrt{x_{k}^{2}+y_{k}^{2}} \equiv$ distance of vortex $k$ to the center.
$c\left(r_{k}, t\right) \equiv$ trap coefficient, positive and $T$ - periodic in time.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two co-rotating vortices.

Fix $n=2, S_{1}=S_{2}=1$.

$$
\begin{array}{cc}
\dot{x}_{1}=-c(t) y_{1}-b \frac{y_{1}-y_{2}}{r^{2}}, & \dot{x}_{2}=-c(t) y_{2}-b \frac{y_{2}-y_{1}}{r^{2}} \\
\dot{y}_{1}=+c(t) x_{1}+b \frac{x_{1}-x_{2}}{r^{2}}, & \dot{y}_{2}=+c(t) x_{2}+b \frac{x_{2}-x_{1}}{r^{2}},
\end{array}
$$

## The case $c\left(r_{i}, t\right) \equiv c(t)$ ．Two co－rotating vortices．

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\end{array}
$$

## Change：

$$
\begin{array}{ll}
s_{1}=x_{1}+x_{2} & d_{1}=x_{1}-x_{2} \\
s_{2}=y_{1}+y_{2} & d_{2}=y_{1}-y_{2}
\end{array}
$$

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\end{array}
$$

## Gives:

$$
\begin{aligned}
s_{1}^{\prime} & =-c(t) s_{2} \\
s_{2}^{\prime} & =c(t) s_{1}
\end{aligned}
$$

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\end{array}
$$

## Solution:

$$
\begin{aligned}
& s_{1}=A \cos (C(t)+B) \\
& s_{2}=A \sin (C(t)+B),
\end{aligned}
$$

with $C(t)=\int_{0}^{t} c(s) d s$.

The case $c\left(r_{i}, t\right) \equiv c(t)$. Two co-rotating vortices.

$$
\begin{aligned}
& \dot{d}_{1}=-c(t) d_{2}-\frac{2 b d_{2}}{\left(d_{1}^{2}+d_{2}^{2}\right)}, \\
& \dot{d}_{2}=+c(t) d_{1}+\frac{2 b d_{1}}{\left(d_{1}^{2}+d_{2}^{2}\right)} .
\end{aligned}
$$

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& \dot{d}_{2}=+c(t) d_{1}+\frac{2 b d_{1}}{\left(d_{1}^{2}+d_{2}^{2}\right)} .
\end{aligned}
$$

Change to polar coordinates:

$$
d_{1}=r \cos \varphi, \quad d_{2}=r \sin \varphi,
$$

From $r^{2}=d_{1}^{2}+d_{2}^{2}$, we get

$$
r \dot{r}=d_{1} \dot{d}_{1}+d_{2} \dot{d}_{2}=0 .
$$

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$$
d_{1}=r \cos \varphi, \quad d_{2}=r \sin \varphi,
$$

From $r^{2}=d_{1}^{2}+d_{2}^{2}$, we get

$$
r \dot{r}=d_{1} \dot{d}_{1}+d_{2} \dot{d}_{2}=0 .
$$

Hence, $r$ is a constant of motion.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two co-rotating vortices.

The angular component $\varphi=\arctan \left(d_{2} / d_{1}\right)$ yields

$$
\dot{\varphi}=\frac{d_{1} \dot{d}_{2}-d_{2} \dot{d}_{1}}{r^{2}}=\frac{2 b}{r}+c(t) .
$$

Since $r$ is constant,

$$
\varphi(t)=\frac{2 b}{r} t+C(t)+D
$$

where $D$ is an arbitrary constant.

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Since $r$ is constant,

$$
\varphi(t)=\frac{2 b}{r} t+C(t)+D
$$

where $D$ is an arbitrary constant.
The general solution reads

$$
\begin{aligned}
d_{1} & =R \cos \left(\frac{2 b}{R} t+C(t)+D\right), \\
d_{2} & =R \sin \left(\frac{2 b}{R} t+C(t)+D\right)
\end{aligned}
$$

where $D$ and $R$ are arbitrary constants.
Hence, the case of two corotating vortices is explicitly solvable. Generically, one finds quasi-periodic solutions and there is a sequence of periodic solutions that can be obtained by fine-tuning $R$.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two co-rotating vortices.



Figure: Periodic orbit corresponding to two vortices with constant trapping coefficient $c(t)=0.1$, vortex-vortex interaction coefficient $b=1$. The vortices are initially placed at $\left(x_{1}(0), y_{1}(0)\right)=(2,0)$ (see [blue] square) and $\left(x_{2}(0), y_{2}(0)\right)=(4,0)$ (see [green] circle).

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two co-rotating vortices.



Figure: Same as before but for a quasi-periodic orbit for a constant trapping coefficient $c(t)=\pi / 20$.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two co-rotating vortices.





Figure: A quasi-periodic orbit for $c(t)=0.1(1+\epsilon \sin (\omega t))$ and $b=1$ with $\epsilon=0.25$ and $\omega=\pi / 30$.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two co-rotating vortices.



Figure: Same as in Fig. 3 but for a larger perturbation strength of $\epsilon=2.5$. The apparent complex motion displayed in the top-left panel (in original coordinates) is nothing but a quasi-periodic orbit that is better elucidated in transformed coordinates in the bottom panel.

The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.

Fix $n=2, S_{1}=1, S_{2}=-1$.

$$
\begin{array}{cc}
\dot{x}_{1}=-c(t) y_{1}+b \frac{y_{1}-y_{2}}{r^{2}}, & \dot{x}_{2}=c(t) y_{2}-b \frac{y_{2}-y_{1}}{r^{2}} \\
\dot{y}_{1}=+c(t) x_{1}-b \frac{x_{1}-x_{2}}{r^{2}}, & \dot{y}_{2}=-c(t) x_{2}+b \frac{x_{2}-x_{1}}{r^{2}},
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$$
\begin{gathered}
\dot{x}_{1}=-c(t) y_{1}+b \frac{y_{1}-y_{2}}{r^{2}}, \quad \dot{x}_{2}=c(t) y_{2}-b \frac{y_{2}-y_{1}}{r^{2}} \\
\dot{y}_{1}=+c(t) x_{1}-b \frac{x_{1}-x_{2}}{r^{2}}, \quad \dot{y}_{2}=-c(t) x_{2}+b \frac{x_{2}-x_{1}}{r^{2}} \\
\dot{s}_{1}=+2 b \frac{d_{2}}{r^{2}}-c(t) d_{2} \\
\dot{s}_{2}=-2 b \frac{d_{1}}{r^{2}}+c(t) d_{1} \\
\dot{d}_{1}=-c(t) s_{2} \\
\dot{d}_{2}=+c(t) s_{1}
\end{gathered}
$$

and

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.

Changing time $\tau=C(t) \equiv \int_{0}^{t} c(s) d s$ yields the equivalent system

$$
\begin{align*}
& \dot{s}_{1}=+f(\tau) \frac{d_{2}}{r^{2}}-d_{2}, \\
& \dot{s}_{2}=-f(\tau) \frac{d_{1}}{r^{2}}+d_{1},  \tag{1}\\
& \dot{d}_{1}=-s_{2}, \\
& \dot{d}_{2}=+s_{1},
\end{align*}
$$

where $f(\tau) \equiv 2 b / c(t(\tau))$.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.

$$
\begin{align*}
& \ddot{d}_{1}+d_{1}=f(\tau) \frac{d_{1}}{r^{2}},  \tag{2}\\
& \ddot{d}_{2}+d_{2}=f(\tau) \frac{d_{2}}{r^{2}} .
\end{align*}
$$

Recalling that $r=\sqrt{d_{1}^{2}+d_{2}^{2}}$, one notes that this Newtonian system has a radial symmetry. This type of systems plays a central role in Celestial Mechanics.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.

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Recalling that $r=\sqrt{d_{1}^{2}+d_{2}^{2}}$, one notes that this Newtonian system has a radial symmetry. This type of systems plays a central role in Celestial Mechanics.
Change to polar coordinates: $d_{1}=r \cos (\varphi), \quad d_{2}=r \sin (\varphi)$, leads to

$$
\begin{align*}
\ddot{r}+r & =\frac{\ell^{2}}{r^{3}}+\frac{f(\tau)}{r},  \tag{3}\\
\dot{\varphi} & =\frac{\ell}{r^{2}}, \tag{4}
\end{align*}
$$

where $\ell \equiv r^{2} \dot{\varphi}=d_{1} \dot{d}_{2}-d_{2} \dot{d}_{1}=r_{1}^{2}-r_{2}^{2}=$ const. is the angular momentum of the solution $\left(d_{1}, d_{2}\right)$.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.

Equation (3) is decoupled and can be studied separately. If $r(t+T)=r(t)$ :

$$
\varphi(\tau)=\int_{0}^{\tau} \frac{\ell}{r^{2}} d s
$$

and

$$
r(t+T)=r(t) \quad \varphi(t+T)=\varphi(t)+\theta,
$$

where $\theta \equiv \varphi(T)=\int_{0}^{T} \ell / r^{2} d s$ is the rotation number. Coming back to the Cartesian coordinates $\boldsymbol{d}=\left(d_{1}, d_{2}\right)$ and using the more convenient complex notation,

$$
\boldsymbol{d}(t+T)=e^{i \theta} \boldsymbol{d}(t)
$$

Therefore, from a $T$-periodic solution of (3) we get a quasi-periodic solution of the original system. If $\theta=0$ (stationary case) the solution $d$ is $T$-periodic, whereas if $\theta=2 \pi / k$ then $d$ is $k T$-periodic (subharmonic of order $k$ ).

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.

$$
\ddot{r}+r=\frac{\ell^{2}}{r^{3}}+\frac{f(\tau)}{r}
$$

For simplicity, take $c(t)=1+\epsilon \sin (\omega t)$.

## Theorem

Assume that $\omega>2$. Then, for any $\ell \geq 0$, (3) has a $T$-periodic solution such that

$$
r(t)>r_{*}=\sqrt{\frac{2 b}{1+\epsilon}} \cos \left(\frac{\pi}{\omega}\right)
$$

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## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.

## Sketch of the proof.

We summarize the technique employed in [Torres, JDE, 2003].

- $\omega>2$ implies that the linear operator $L r:=\ddot{r}+r$ has a positive Green's function $G(t, s)$ such that

$$
m:=\cot \left(\frac{\pi}{\omega}\right) \leq G(t, s) \leq M:=1 / \sin \left(\frac{\pi}{\omega}\right)
$$

- A $T$-periodic solution of Eq. (3) is a fixed point of the operator

$$
\mathcal{A} r=\int_{0}^{T} G(\tau, s)\left[\frac{\ell^{2}}{r^{3}}+\frac{f(\tau)}{r}\right] d s
$$

- Apply Krasnoselskii's fixed point Theorem for cone compression-expansion with the cone

$$
P=\left\{u \in X: \min _{x} u \geq \frac{m}{M}\|u\|\right\},
$$

where $X$ the Banach space of the continuous and $T$-periodic functions with the norm of the supremum.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.

In consequence, there is a continuous branch of $T$-periodic solutions $\left(\ell, r_{\ell}\right)$ of Eq. (3). Now we analyze the stability of such solutions.

## Theorem

Take $\ell \geq 0$. Then, there exists an explicitly computable $\omega_{\ell}$ such that if $\omega>\omega_{\ell}$ then $r_{\ell}$ is linearly stable as a T-periodic solution of Eq. (3).

Proof. The linearized equation around $r_{\ell}$ is $\ddot{y}+a(\tau) y=0$ with

$$
a(\tau)=1+\frac{3 \ell^{2}}{r_{\ell}^{4}}+\frac{f(\tau)}{r_{\ell}^{2}}
$$

Now, use $r(t)>r_{*}$ and apply Lyapunov's stability criterion.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.

It is possible to prove analytically the presence of KAM dynamics in Eq. (3), by using a Ortega's third order approximation method as in [Torres,Adv.Nonlinear Stud.,2002].

## Theorem

Let us assume $\ell=0, \omega>\sqrt{2}, \omega \neq\{3 \sqrt{2}, 4 \sqrt{2}\}$. Then the $T$-periodic solution $r_{0}$ of Eq. (3) is of twist type except possibly for a finite number of values of $\epsilon$.

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.



Figure: Opposite charge vortices $S_{1}=1$ (see solid [blue] line) and $S_{2}=-1$ (see dashed [green] line) with constant trapping coefficient $c(t)=0.1$, vortex-vortex interaction coefficient $b=1$. The vortices are initially placed at $\left(x_{1}(0), y_{1}(0)\right)=(2,0)$ (see [blue] square) and $\left(x_{2}(0), y_{2}(0)\right)=(4,0)$ (see [green] circle) .

## The case $c\left(r_{i}, t\right) \equiv c(t)$. Two counter-rotating vortices.





Figure: An irregular orbit corresponding to $c(t)=0.1(1+\epsilon \sin (\omega t))$ and $b=1$ with $\epsilon=1$ and $\omega=0.12$.

## The Fetter case.

In a more realistic situation, the precession frequency depends also on $r$. Fetter [2001] proposed

$$
c(r, t) \equiv \Omega(r)=\frac{2 \hbar \omega_{r}^{2}}{8 \mu\left(1-r^{2} / R_{\perp}^{2}\right)}\left(3+\frac{\omega_{r}^{2}}{5 \omega_{z}^{2}}\right) \ln \left(\frac{2 \mu}{\hbar \omega_{r}}\right),
$$

where $\omega_{r}\left(\omega_{z}\right)$ is the radial (axial) trap frequency, $R_{\perp}$ is the Thomas-Fermi radial extent of the condensate (can be normalized to 1 ), and $\mu$ is the chemical potential.

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$$

where $\omega_{r}\left(\omega_{z}\right)$ is the radial (axial) trap frequency, $R_{\perp}$ is the Thomas-Fermi radial extent of the condensate (can be normalized to 1 ), and $\mu$ is the chemical potential. From now on,

$$
\Omega(r)=\frac{\Omega_{0}}{1-r^{2}} .
$$

## The Fetter case.

The model is

$$
\begin{aligned}
\dot{x}_{k} & =-S_{k} \Omega\left(r_{k}\right) y_{k}-\frac{b}{2} \sum_{j \neq k}^{n} S_{j} \frac{y_{k}-y_{j}}{r_{j k}^{2}}, \\
\dot{y}_{k} & =S_{k} \Omega\left(r_{k}\right) x_{k}+\frac{b}{2} \sum_{j \neq k}^{n} S_{j} \frac{x_{k}-x_{j}}{r_{j k}^{2}} .
\end{aligned}
$$

or in complex variables

$$
i \dot{z}_{k}=-S_{k} \Omega\left(r_{k}\right) z_{k}+\frac{b}{2} \sum_{j \neq k}^{n} S_{j} \frac{z_{k}-z_{j}}{r_{j k}^{2}}
$$

## Conserved quantities and integrabitlity.

The system is hamiltonian. Define

$$
H\left(z_{1}, \ldots, z_{n}\right)=-\frac{\Omega_{0}}{2} \sum_{k=1}^{n} \ln \left(1-r_{k}^{2}\right)+\frac{b}{4} \sum_{k=1}^{n} \sum_{j \neq k}^{n} S_{j} S_{k} \ln \left(r_{j k}^{2}\right)
$$

where $\Omega_{0} \equiv \Omega(0)$. Then,

$$
\begin{aligned}
S_{k} \dot{x}_{k} & =-\frac{\partial H}{\partial y_{k}}, \\
S_{k} \dot{y}_{k} & =\frac{\partial H}{\partial x_{k}} \quad \text { for every } k=1, \ldots, n
\end{aligned}
$$

$H$ is a the first integral, i.e., a first conserved quantity along the orbits of the system.
A second integral of motion is

$$
V=\sum_{k=1}^{n} S_{k} r_{k}^{2},
$$

which is a form of inertial momentum.

## Conserved quantities and integrabitlity.

The presence of two conserved quantities or dynamical invariants guarantees integrability in the classical Liouville sense for the case $n=2$. This implies that the energy level sets are compact and the phase space is foliated by invariant tori. On each of these, the motion is quasi-periodic with two frequencies.

## The Fetter case with $n=2, S_{1}=1, S_{2}=-1$.

## Theorem

The two vortices never collide, i.e., they are separated by a computable minimal distance.

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Taking exponentials on the hamiltonian, we have

$$
\left(1-r_{1}^{2}\right)^{\Omega_{0}}\left(1-r_{2}^{2}\right)^{\Omega_{0}} r_{12}^{b / 2}=C^{2}>0 \quad \forall t,
$$

where $C^{2}=\left(1-r_{1}(0)^{2}\right)^{\Omega_{0}}\left(1-r_{2}(0)^{2}\right)^{\Omega_{0}} r_{12}(0)^{b / 2}$. Note that $0<C^{2}<2^{b / 2}$ because $0 \leq r_{i}(0)^{2}<1(i=1,2)$ and $0<r_{12}(0)<2$. Then,

$$
r_{12}(t)>C^{4 / b} \quad \forall t,
$$

## The Fetter case with $n=2, S_{1}=1, S_{2}=-1$.

## Theorem

There is an equilibrium, which is unique up to rotations, given by

$$
\left(x_{1}^{0}, y_{1}^{0}\right)=\left(\sqrt{\frac{b}{4 \Omega_{0}+b}}, 0\right), \quad\left(x_{2}^{0}, y_{2}^{0}\right)=\left(-\sqrt{\frac{b}{4 \Omega_{0}+b}}, 0\right)
$$

## The Fetter case with $n=2, S_{1}=1, S_{2}=-1$.

To emphasize the rotational invariance of our system, it is convenient to work on the polar coordinates $\left(r_{1}, r_{2}, \theta\right)$. The (unique) equilibrium is $\left(r_{1}^{0}, r_{2}^{0}, \theta_{0}\right)=\left(\sqrt{\frac{b}{4 \Omega_{0}+b}}, \sqrt{\frac{b}{4 \Omega_{0}+b}}, \pi\right)$.

## Theorem

The equilibrium $\left(r_{1}^{0}, r_{2}^{0}, \theta_{0}\right)$ is stable (in the sense of Lyapunov).

## The Fetter case with $n=2, S_{1}=1, S_{2}=-1$.

The hamiltonian $H$ is still a conserved quantity which in the new variables reads
$H\left(r_{1}, r_{2}, \theta\right)=-\frac{1}{2}\left[\Omega_{0} \ln \left(1-r_{1}^{2}\right)+\Omega_{0} \ln \left(1-r_{2}^{2}\right)+\frac{b}{2} \ln \left(r_{1}^{2}+r_{2}^{2}-2 r_{1} r_{2} \cos \theta\right)\right]$.
The equilibrium $\left(r_{1}^{0}, r_{2}^{0}, \theta_{0}\right)$ is a critical point of $H$. The hessian matrix evaluated on $\left(r_{1}^{0}, r_{2}^{0}, \theta_{0}\right)$ is

$$
\frac{1}{8}\left(\begin{array}{ccc}
7 b+12 \Omega_{0}+\frac{b^{2}}{\Omega_{0}} & b+4 \Omega_{0} & 0 \\
b+4 \Omega_{0} & 7 b+12 \Omega_{0}+\frac{b^{2}}{\Omega_{0}} & 0 \\
0 & 0 & b
\end{array}\right) .
$$

One can prove easily that this matrix is positive-definite by Sylvester's criterion. Hence $H$ attains a minimum at $\left(r_{1}^{0}, r_{2}^{0}, \theta_{0}\right), H$ is a Lyapunov function and the equilibrium is, in fact, stable.

## The Fetter case with $n=2, S_{1}=1, S_{2}=-1$ ．



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## Guiding Center Equilibria and their Stability

We are looking for "guiding centers" about which the vortices oscillate. These guiding centers might themselves be precessing around at some frequency $\omega$. For the case of symmetric guiding centers this precession frequency will be zero.

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We adopt a trial solution of the form

$$
\begin{equation*}
z_{1}(t)=r_{1} \exp (i \omega t) \quad \text { and } \quad z_{2}(t)=r_{2} \exp (i \omega t+i \pi)=-r_{2} \exp (i \omega t) \tag{5}
\end{equation*}
$$

where $r_{1}$ and $r_{2}$ are now constant，as is $\omega$ ，the precession frequency of the guiding center．

## Guiding Center Equilibria and their Stability

Fix

$$
\begin{equation*}
\alpha=\frac{1}{1-r_{1}^{2}}, \quad \beta=\frac{1}{1-r_{2}^{2}}, \quad \text { and } \quad \gamma=\frac{1}{2 r_{12}^{2}}=\frac{1}{2 s^{2}}, \tag{6}
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where $s=r_{12}=r_{1}+r_{2}$ is the separation distance between the two guiding centers.

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After some algebra, one finds that the precession frequency should be

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\begin{equation*}
\omega=\frac{1}{2}\left[\Omega_{0}(\alpha-\beta)+\gamma b_{0}\left(\frac{r_{1}}{r_{2}}-\frac{r_{2}}{r_{1}}\right)\right] . \tag{7}
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Note that if $r_{2}=r_{1}$, then $\alpha=\beta$ and $\omega=0$, as we expect.

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\begin{equation*}
\Omega_{0} \beta r^{3}+\frac{b}{2 r_{2}} r^{2}-\Omega_{0}(1+\beta) r-\frac{b}{r_{2}}=0 \tag{8}
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The stability of the guiding centers can be done as in the case of the equilibrium by passing to a rotational frame.

## Open problems.

- The Fetter case with two corotating vortices ( $n=2, S_{1}=S_{2}=1$ ) can be handled in a similar way.
- Dynamics of many vortices, both in the Fetter and non-Fetter case.
- Collisions.
- Effect of a more complicated precession frequency (for instance non symmetric).

Thank you for your attention!!!

